

Article

Mass in de Sitter and Anti-de-Sitter universes with regard to Dark Matter

Jean-Pierre Gazeau

Université de Paris, CNRS, Astroparticule et Cosmologie, F-75013 Paris, France; gazeau@apc.in2p3.fr

Abstract: An explanation of the origin of dark matter is suggested in this work. The argument is based on symmetry considerations about the concept of mass. In the Wigner's view, the rest mass and the spin of a free elementary particle in flat space-time are the two invariants that characterize the associated unitary irreducible representation of the Poincaré group. The Poincaré group has two and only two deformations with maximal symmetry. They describe respectively the de Sitter (dS) and Anti de Sitter (AdS) kinematic symmetries. Analogously to their shared flat space-time limit, two invariants, spin and energy scale for de Sitter and rest energy for Anti de Sitter, characterize the unitary irreducible representation associated with dS and AdS elementary systems. While the dS energy scale is a simple deformation of the Poincaré rest energy and so has a purely mass nature, AdS rest energy is the sum of a purely mass component and a kind of zero-point energy derived from the curvature. An analysis based on recent estimates on the chemical freeze-out temperature marking in Early Universe the phase transition quark-gluon plasma epoch to the hadron epoch supports the guess that dark matter energy might originate from an effective AdS curvature energy.

Keywords: de Sitter; Anti de Sitter; group representation; hadronization; dark matter; zero-point energy; quark-gluon plasma; critical point

MSC: 81R05, 85A04, 81V05

Contents

| | | |
|---|---|---|
| 1 | Introduction: Some (observational) facts about dark matter | 1 |
| 2 | Mass and symmetries | 2 |
| 3 | Minkowskian content of dS and AdS elementary systems: the Garidi mass | 3 |
| 4 | Dark matter as a relic AdS curvature energy? | 5 |
| 5 | Discussion | 5 |
| | References | 6 |

1. Introduction: Some (observational) facts about dark matter

According to the Planck 2018 analysis of CMB power spectrum [1], our Universe is spatially flat, accelerating, and composed of 5% baryonic matter, 27% cold dark matter (CDM, non baryonic) and 68% dark energy (Λ). (Cold) dark matter [2] is observed by its gravitational influence on luminous, baryonic matter. The dark matter mass halo and the total stellar mass are coupled through a function that varies smoothly with mass (see [3] and references therein), with possible noticeable exception(s) like the recent

[4,5]. Up to now, all hypothetical particle models (WIMP, Axions, Neutrinos ...) failed direct or indirect detection tests. Alternative theories (e.g. MOND), which negate the existence of dark matter as a physical entity, have failed to explain clusters and the observed pattern in the CMB.

In this article, we view dark matter as a physical entity and we propose an explanation of its current existence as the remnant, after hadronization, of the zero-point energy of the quark-gluon plasma (QGP) [6] due to an effective Anti-de-Sitterian environment experienced by the QGP massive constituents existing at the so-called quark epoch, $\sim 10^{-12}\text{s} \sim 10^{-6}\text{s}$ after the Big-Bang and subsequent to the inflation (over $\sim 10^{-33}\text{s} \sim 10^{-32}\text{s}$).

Section 2 is a survey of Poincaré, de Sitter and Anti-de-Sitter kinematic symmetries and their quantum realizations in terms of the unitary irreducible representations (UIR) of the corresponding relativity groups which are labelled by spin and \sim mass invariants. In Section 3, these de Sitter and Anti de Sitter massive representations are considered from the point of view of their minkowskian contraction limits in terms of the so-called Garidi mass formulas. We then reexamine in Section 4 the dark matter enigma by advancing three hypotheses based on the Anti de Sitter Garidi mass formula. We conclude in Section 5 with a few comments.

2. Mass and symmetries

In Minkowski (M), the concept of (rest) mass originates from the ubiquitous law of conservation of energy, a direct consequence of the Poincaré symmetry. This concept was formulated by Wigner [7] in rigorous mathematical terms as associated with the concept of elementary system. The latter is a description of a set of states which forms a unitary irreducible representation (UIR) space for the proper orthochronous Poincaré group $P_0^\uparrow = \mathbb{R}^{1,3} \times \text{SO}_0(1,3)$ (or $\mathbb{R}^{1,3} \times \text{SL}(2, \mathbb{C})$), semi-direct product of the translations in Minkowski with the Lorentz group (or its universal covering). The UIR's of the Poincaré group [8] that we are concerned with here are the massive ones with positive energy. They are denoted by $U_M(m, s)$ and are completely characterized by the eigenvalues of two Casimir operators, namely the quadratic Klein-Gordon operator

$$Q_M^{(1)} = P^\mu P_\mu = P_0^2 - \mathbf{P}^2,$$

with eigenvalues $\langle Q_M^{(1)} \rangle = m^2 c^2$, and the quartic Pauli-Lubanski operator

$$Q_M^{(2)} = W^\mu W_\mu, \quad W_\mu = \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} J^{\nu\rho} P^\sigma,$$

with eigenvalues (in the non-zero mass case) $\langle Q_M^{(2)} \rangle = -m^2 c^2 s(s+1) \hbar^2$.

As it was explained in [9], with the requirements of kinematical rotation, parity, and time-reversal invariance, there exists only one way to deform the Poincaré group P_0^\uparrow , namely, in endowing space-time with a certain curvature. Hence, there are two possible deformations that are distinguished with the curvature sign of the corresponding space-times, namely the de Sitter (dS) and Anti de Sitter (AdS) space-times. More precisely, dS (resp. AdS) are the unique maximally symmetric solutions of the vacuum Einstein's equations with positive (resp. negative) cosmological constant Λ . Their symmetries are one-parameter deformations of the minkowskian symmetry with

- negative curvature $-\varkappa_{\text{dS}} = -H/c = -\sqrt{|\Lambda_{\text{dS}}|/3}$
- positive curvature $\varkappa_{\text{AdS}} = \sqrt{|\Lambda_{\text{AdS}}|/3}$

Their respective invariances (in the relativity or kinematical sense) hold with respect to the ten-parameter dS group $\text{SO}_0(1,4)$ (or its universal covering $\text{Sp}(2,2)$) and the ten-parameter AdS group $\text{SO}_0(2,3)$ (or its two-fold covering $\text{Sp}(4, \mathbb{R})$, or even its universal covering) groups.

The “massive” UIR’s of the dS and AdS groups are those which are deformations of the above $U_M(m, s)$. Those UIR’s are denoted by $U_{\text{dS}}(\zeta_{\text{dS}}, s)$ and $U_{\text{AdS}}(\zeta_{\text{AdS}}, s)$ respectively. The real dimensionless parameters ζ_{dS} and ζ_{AdS} replace the minkowskian rest mass. Together with the spin s they determine the corresponding UIR through the eigenvalues of dS and AdS invariant Casimir operators. The dS UIR $U_{\text{dS}}(\zeta_{\text{dS}}, s)$ belongs to the so-called principal series, for which we have

$$\langle Q_{\text{dS}}^{(1)} \rangle = \zeta_{\text{dS}}^2 + \frac{9}{4} - s(s+1), \quad \langle Q_{\text{dS}}^{(2)} \rangle = \left(\zeta_{\text{dS}}^2 + \frac{1}{4} \right) s(s+1),$$

with $\zeta_{\text{dS}} \in \mathbb{R}$, and $s \in \mathbb{N}/2$. Note that $U_{\text{dS}}(\zeta_{\text{dS}}, 0)$ and $U_{\text{dS}}(-\zeta_{\text{dS}}, 0)$ are equivalent. The AdS UIR $U_{\text{AdS}}(\zeta_{\text{AdS}}, s)$ belongs to the so-called discrete series, and we have,

$$\langle Q_{\text{AdS}}^{(1)} \rangle = \zeta_{\text{AdS}}(\zeta_{\text{AdS}} - 3) + s(s+1), \quad \langle Q_{\text{AdS}}^{(2)} \rangle = -(\zeta_{\text{AdS}} - 1)(\zeta_{\text{AdS}} - 2)s(s+1),$$

with $s \in \mathbb{N}/2$ and $\zeta_{\text{AdS}} = s+2, s+3, \dots$ (or $\zeta_{\text{AdS}} \geq s+2$ for the universal covering of $\text{SO}_0(2, 3)$).

There exists an irreconcilable difference between the dS invariant ζ_{dS} and the AdS one ζ_{AdS} . The latter is the lowest value of the discrete spectrum of the generator of “time” rotations in AdS [10]. Hence it can be given a non ambiguous meaning of a rest energy when it is expressed in energy AdS units $\hbar\kappa_{\text{AdS}}/c$:

$$E_{\text{AdS}}^{\text{rest}} := \hbar c \kappa_{\text{AdS}} \zeta_{\text{AdS}}. \quad (1)$$

Therefore, this physical concept of energy at rest survives with the deformation Poincaré \rightarrow AdS.

The situation is radically different with dS. The spectrum of the generator of “time” hyperbolic rotations in dS is not bounded below [11]. Actually it covers the whole real line. Of course, the invariant ζ_{dS} can be given an energy dimension as

$$E_{\text{dS}} := \hbar c \kappa_{\text{dS}} \zeta_{\text{dS}}. \quad (2)$$

3. Minkowskian content of dS and AdS elementary systems: the Garidi mass

If we wish to go further into the interpretative problem of a mass in a dS/AdS background, the crucial question to be addressed concerns the interpretation of the dS/AdS UIR’s (or quantum AdS and dS elementary systems) from a (asymptotically) minkowskian point of view. We mean by this the study of the contraction limit $\kappa \rightarrow 0$ or equivalently $\Lambda \rightarrow 0$ of these representations. The notion of mass in “desitterian Physics” may appear ambiguous in term of contraction of representation [12,13], exemplified by the fact that one cannot give a precise meaning to a dS rest energy, except if one follows an approach based on a causality de-sitterian semi-group [11], or based on a analyticity prerequisite [14]. Nevertheless, a consistent mass formula has been proposed by Garidi [15] in terms of the dS UIR parameters ζ_{dS} and s :

$$m_{\text{dS}}^2 = \frac{\hbar^2 \kappa_{\text{dS}}^2}{c^2} \left(\langle Q_{\text{dS}}^{(1)} \rangle - \langle Q_{\text{dS}}^{(1)} |_{s=1/2+i\zeta_{\text{dS}}} \rangle \right) = \frac{\hbar^2 \kappa_{\text{dS}}^2}{c^2} \left(\zeta_{\text{dS}}^2 + \left(s - \frac{1}{2} \right)^2 \right). \quad (3)$$

The minimal value assumed by the eigenvalues of the first Casimir in the set of UIR in the discrete series is precisely reached at $s = 1/2 + i\zeta_{\text{dS}}$, which corresponds to the “conformal” massless case, for which s loses clearly its spin meaning. Controlling the validity of such a formula from a minkowskian observer amounts to understand the contraction (mathematically non trivial en terms of sequences of Hilbert spaces [16])

$$\text{dS UIR} \longrightarrow \text{Poincaré UIR}$$

Then the contraction $dS \rightarrow$ Poincaré in terms of mass has to be understood as

$$\varkappa_{dS} \rightarrow 0 \quad \text{and} \quad \zeta_{dS} \rightarrow +\infty \quad \text{while} \quad \zeta_{dS} \hbar \varkappa_{dS} / c \rightarrow m.$$

In terms of representations,

$$D(\zeta_{dS}, s) \xrightarrow[\zeta_{dS} \hbar \varkappa_{dS} / c \rightarrow m / \hbar]{\varkappa_{dS} \rightarrow 0, \zeta_{dS} \rightarrow \infty} U_M(m, s).$$

Thus, close by the contraction limit,

$$m_{dS} \sim m \sqrt{1 + \frac{(s - \frac{1}{2})^2}{\zeta_{dS}^2}} = \begin{cases} m & \text{for } s = 1/2 \\ m + o(1/\zeta_{dS}) & \text{for } s \neq 1/2 \end{cases}$$

The analogous of the Garidi mass for the AdS case exists as well [17,18]. It precisely vanishes for massless conformal AdS fields that lie at the lowest limit $\zeta_{AdS} = s + 1$ of the discrete series.

$$\begin{aligned} m_{AdS}^2 &= \frac{\hbar^2 \varkappa_{AdS}^2}{c^2} \left(\langle Q_{AdS}^{(1)} \rangle - \langle Q_{AdS}^{(1)} |_{\zeta_{AdS}=s+1} \rangle \right) \\ &= \frac{\hbar^2 \varkappa_{AdS}^2}{c^2} \left[\left(\zeta_{AdS} - \frac{3}{2} \right)^2 - \left(s - \frac{1}{2} \right)^2 \right]. \end{aligned} \quad (4)$$

As for dS, the contraction $AdS \rightarrow$ Poincaré in terms of mass reads as

$$\varkappa_{AdS} \rightarrow 0 \quad \text{and} \quad \zeta_{AdS} \rightarrow +\infty \quad \text{while} \quad \zeta_{AdS} \hbar \varkappa_{AdS} / c \rightarrow m.$$

The contraction $AdS \rightarrow$ Poincaré in terms of masses and representations reads as

$$D(\zeta_{AdS}, s) \xrightarrow[\zeta_{AdS} \varkappa_{AdS} \rightarrow mc / \hbar]{\varkappa_{AdS} \rightarrow 0, \zeta_{AdS} \rightarrow \infty} U_M(m, s).$$

We now show how the rest energy introduced in (1) for a massive AdS elementary system reveals a universal pure curvature or vibration energy component besides a matter energy content. From the Garidi mass formula (4), an AdS scalar elementary system can indeed be viewed asymptotically [19] as a combination of both a relativistic free particle with rest energy $m_{AdS} c^2 \sim mc^2$ and a 3d isotropic quantum harmonic oscillator with zero-point energy $\frac{3}{2} \hbar \varkappa_{AdS} c \equiv \frac{3}{2} \hbar \omega_{AdS}$ at the first order in the curvature:

$$\begin{aligned} E_{AdS}^{\text{rest}} &= \hbar \varkappa_{AdS} c \zeta_{AdS} = \left[m_{AdS}^2 c^4 + \hbar^2 \omega_{AdS}^2 \left(s - \frac{1}{2} \right)^2 \right]^{1/2} + \frac{3}{2} \hbar \omega_{AdS} \\ &= m_{AdS} c^2 + \frac{3}{2} \hbar \omega_{AdS} + \frac{1}{2} \frac{\hbar^2 \omega_{AdS}^2}{m_{AdS} c^2} \left(s - \frac{1}{2} \right)^2 + o(\varkappa_{AdS}^2) \\ &= m_{AdS} c^2 + \frac{3}{2} \hbar \omega_{AdS} + o(\varkappa_{AdS}). \end{aligned}$$

By contrast, the meaning of energy in dS relativity with regard to its Poincaré limit is less tractable. It is also exemplified by the absence of any term of order \varkappa_{dS} besides the dS mass energy

$$E_{dS} = \hbar \varkappa_{dS} c \zeta_{dS} = m_{dS} c^2 + o(\varkappa_{dS})$$

One notices the remarkable position occupied by the spin $s = 1/2$ in the above formulas:

$$\begin{aligned} \text{for dS: } & E_{\text{dS}} = m_{\text{dS}}c^2, \\ \text{for AdS: } & E_{\text{AdS}}^{\text{rest}} = m_{\text{AdS}}c^2 + \frac{3}{2}\hbar\kappa_{\text{AdS}}c. \end{aligned}$$

4. Dark matter as a relic AdS curvature energy?

Dark matter is observed as an energy more or less localized in halos surrounding baryonic matter in galaxies and galaxy clusters.

Assumption 1

In our approach, the nature of this dark matter energy is supposed to be related to some effective AdS curvature. Precisely, for a spin $1/2$ elementary particle X ,

$$E_{\text{AdS}}^{\text{rest}} = \underbrace{m_{\text{AdS}}(X)c^2}_{\text{visible}} + \underbrace{\frac{3}{2}\hbar\kappa_{\text{AdS}}c}_{\sim E_{\text{dm}}(X)}, \quad E_{\text{dm}}(X) = r(X)m_{\text{AdS}}(X)c^2,$$

where the ratio $r(X) := E_{\text{dm}}(X)/m_{\text{AdS}}(X)c^2$ should reflect at a certain extent the ratio dark matter/visible matter.

Assumption 2

The appearance of dark matter held over a period when the temperature(s) was (were) compatible with a phase of entities X , compatibility being understood in the sense of validity of the equipartition theorem applied to the quantum-oscillator-like energy spectrum of AdS elementary system:

$$k_B T_X \approx \hbar\kappa_{\text{AdS}}c \approx \frac{2r(X)}{3} m_{\text{AdS}}(X)c^2.$$

The most probable candidates X in agreement with the above assumptions are stable light quarks u and d when, at the “quark epoch” (10^{-12} s \sim 10^{-6} s, $T > 10^{12}$ K), the quark-gluon plasma experienced the phase transition, i.e., hadronization, which marked the beginning of the “hadron epoch” (10^{-6} s \sim 1 s, $T > 10^{10}$ K). The current estimate of the hadronization temperature for light quarks [20] is $T_{cf} = 156.5 \pm 1.5$ MeV $\approx 1.8 \times 10^{12}$ K (“chemical freeze-out temperature”). So, with $m_{\text{AdS}} \sim m$,

$$T_X \approx 1.8 \times 10^{12} \text{ K} \approx \frac{2r(X)}{3} \frac{m(X)c^2}{k_B}, \quad r(u) \approx 108, \quad r(d) \approx 49.$$

Moreover, this value of T_X yields AdS curvature κ_{AdS} and lifetime τ :

$$k_B T_X \approx \hbar\kappa_{\text{AdS}}c \equiv \frac{\hbar}{\tau} \Rightarrow \tau \approx 2.7 \times 10^{-23} \text{ s}, \quad \kappa_{\text{AdS}}^{-1} \approx 8 \text{ fm}. \quad (5)$$

This AdS length scale $\kappa_{\text{AdS}}^{-1} \approx 8$ fm is to be compared with the QGP typical distance scales, which exceed the size of the largest atomic nuclei (and the low typical momentum scale) (in the Pb case, $R_{\text{Pb}} \approx 5.3788$ fm).

Assumption 3

The pure AdS curvature energy decouples from the rest mass energy at the critical hadronization point and abides as a free component of the Universe along its posterior epochs.

5. Discussion

The conjectural interpretation we have proposed above of $E_{dm}(X) \sim \frac{3}{2}\hbar\omega_{AdS} = \frac{3}{2}\hbar\omega_{AdS}$ as the remnant of the AdS zero-point energy of the elementary system X in the QGP period of the early Universe, which immediately follows the dS inflationary phase like an “AdS bounce” (AdS phase as anti inflation!) and which precedes the hadronization, requires of course a lot deeper analysis in terms of QCD, thermal QFT, and phase transition. Note that a QCD vacuum density due to conformal anomaly yields a Lorentz-invariant negative-valued contribution to the cosmological constant (see the review [6] and references therein). Unfortunately, this effect is negligible in comparison with the estimate (5). Moreover, the assumption of a AdS space-time should compel us to work within the framework of consistent thermodynamics, QFT and QCD in AdS and not in Minkowski, a formidable program...

Here we have advanced a guess based on some of the most basic symmetry considerations, by exploiting ideas coming from previous works like [13,17–19]. Nevertheless, one may imagine that in the primordial QGP period happening just after inflation the massive constituents experience an effective geometric environment analogous to AdS during extremely short periods, and that the total zero-point energy resulting from those AdS phases subsists after hadronization. Like the cosmic microwave background (CMB) is a remnant from the recombination epoch of the universe ($\approx 379,000$ years, at $T \approx 3000$ K), when protons and electrons combined to form neutral hydrogen atoms, the dark matter would be a “relic” of the QGP epoch, totally free of any interaction but the gravitational one. It is indeed tempting to establish a parallel between dark matter and CMB, since the latter is viewed as the emergence of the photon decoupling, precisely when photons started to travel freely through space rather than constantly being scattered by electrons and protons in plasma.

Finally, I cannot resist quoting Bacry and Lévy-Leblond in [9]:

...it is amusing to notice that in the Newton Universe N_- resulting from $c \rightarrow \infty$ contraction of AdS, the kinetic energy of the elementary system on the quantum level, that is,

$$E_{\text{kin}}^{N_-} := \frac{I}{2} \left[P^2 + \frac{1}{\tau^2} K^2 \right], \quad I \equiv \text{Newtonian inertia}, \quad K \equiv \text{Newtonian boost generator}$$

is quantized, which is not surprising in view of the “compactness” of the corresponding universe. The oscillator levels have a separation $\delta E_{\text{kin}}^{N_-} \approx \hbar\tau^{-1}$ in agreement with the uncertainty principle, since τ may be thought of as the “lifetime” of this oscillating universe.

References

1. Planck 2018 results XIII. Cosmological parameters, *A&A* to appear (2019); arXiv:1807.06209v1 [astro-ph.CO]
2. Baudis, L. The Search for Dark Matter, *European Review* **26** 70-81 (2017)
3. Behroozi, P. S., Wechsler, R. H., and Conroy, C. The average star formation histories in dark matter halos from $z = 0 - 8$, *Astrophys. J.*, **770**:57 (2013).
4. van Dokkum et al, A galaxy lacking dark matter, *Nature Lett.* **555** (2018).
5. van Dokkum et al, The Distance of the Dark Matter Deficient Galaxy NGC 1052-DF2, *Astrophys. J. Lett.*, **864**:L18 (2018).
6. Pasechnik, R. and Šumbera, M. Phenomenological Review on Quark-Gluon Plasma: Concepts vs. Observations, *Universe* **3** (2017).
7. Newton, T. D. and Wigner E. P. Localized States for Elementary Systems, *Rev. Mod. Phys.* **21** 400-406 (1949).
8. Wigner, E. P. On Unitary Representations of the Inhomogeneous Lorentz Group, *Ann. Math.* **40** 149-204 (1939).
9. Bacry, H. and Lévy-Leblond, J.-M. Possible Kinematics, *J. Math. Phys.* **9** 1605 (1968).
10. Fronsdal, C. Elementary particles in a curved space. II, *Phys. Rev D* **10** 589-598 (1974).

11. Mizony, M. 3 semigroupes de causalité et formalisme hilbertien de la mécanique quantique, *Publ. Dep. Math. Lyon* **3B** 47-64 (1984).
12. Mickelsson, J. and Niederle, J. Contractions of representations of de Sitter groups, *Commun. Math. Phys.* **27** 167-180 (1972).
13. Garidi, T., Huguet, E., and Renaud, J. de Sitter waves and the zero curvature limit, *Phys. Rev. D* **67** (2003). arxiv gr-qc/0304031
14. Bros J. , Gazeau, J.-P. and Moschella, U. Quantum Field Theory in the de Sitter Universe, *Phys. Rev. Lett.* **73** 1746 (1994).
15. Garidi, T. What is mass in desitterian Physics? hep-th/0309104
16. Dooley, A.H. and Rice, J.W., On contractions of semisimple Lie groups, *Trans. Amer. Math. Soc.* **289** (1985) 185-202.
17. Gazeau, J.-P. and Novello, M. The question of mass in (anti-) de Sitter spacetimes, *J. Phys. A: Math. Theor.* **41** 304008 (2008).
18. Gazeau, J.-P. and Novello, M. The Nature of Λ and the Mass of the Graviton: A Critical View, *Int. J. Mod. Phys. A* **26** 3697-3720 (2011).
19. Gazeau, J.-P. and Renaud, J. Relativistic harmonic oscillator and space curvature, *Phys. Lett. A* **179** 67 (1993).
20. Andronic, A., Braun-Munzinger, P., Redlich, K. and Stachel, J. Decoding the phase structure of QCD via particle production at high energy, *Nature* (2018). DOI: 10.1038/s41586-018-0491-6