

1 Article

2 **Moving Source Depth Estimation Using a Horizontal Line** 3 **Array in Shallow Water Waveguides**

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11 **Abstract:** In this study, a matched-mode autoregressive source depth estimation method (MMAR)
12 based on autoregressive (AR) wavenumber estimation is proposed for a moving source in shallow
13 water waveguides. The signal original frequency and the environmental parameters, namely, the
14 sound speed profile and bottom properties are known as a prior knowledge. The mode
15 wavenumbers are estimated by the AR modal wavenumber spectrum. On the basis of the mode
16 wavenumber estimation, the mode amplitudes can be estimated by the wavenumber spectrum
17 that is obtained by generalized Hankel transform. The source depth estimation is determined by
18 the peak of source depth function wherein the data mode best matches the replica mode that is
19 calculated using a propagation model. Compared with other methods of moving source depth
20 estimation, the proposed method exhibits a better performance in source depth estimation under
21 low signal-to-noise ratio or the small range span. The selection of horizontal line array depth is
22 illustrated by simulation and normal mode theory in details.

23 **Keywords:** modal wavenumber spectrum; match mode; horizontal line array; moving source

25 1. Introduction

26 Matched field processing (MFP) [1-2] or the equivalent matched mode processing (MMP) [3] is
27 a popular method for localizing the source in complex propagation environments. Conventional
28 MFP obtains the range depth function, that is, the correlation between the replica and the measured
29 fields. The source location is determined by the peak of the range depth function. MFP can localize
30 the source accurately in a range-independent or slowly varying environment without considering
31 any environmental mismatches. When the sound propagation is not exactly modeled, the range and
32 depth estimated by MFP generally present an error, which is referred to as the environmental
33 mismatch problem [4]. The need for depth estimation is more pressing than range estimation for
34 some applications. The source range and depth are estimated simultaneously by MFP, which
35 indicates that depth estimation depends on range estimation. The robustness of depth estimation
36 decreases because of the sensitivity of range estimation to sound speed profile (SSP). MFP always
37 requires a vertical line array (VLA) or a horizontal line array (HLA) with large aperture. However,
38 real application of large aperture array is difficult because of the installation platform. The
39 application of MFP is restricted by the aperture problem and the environmental mismatch problem.
40 Consequently, an active area of research in recent years has been devoted to estimate source depth
41 robustly by a single hydrophone or a short line array [5-7].

42 A class of methods that has shown some promise in passive location exploits the properties of
43 the reliable acoustic path (RAP) [8]. Such methods reduce the requirement of array aperture. The
44 difficulty associated with the propagation of RAP limits the application of the method in shallow
45 water. Depth estimation for shallow water waveguides is generally performed in two steps: mode
46 filtering first and then MMP [3]. The modal amplitudes are determined by mode filters, and then the
47 MMP is used to estimate the source location on basis of the modal amplitudes. However, the mode

48 filters always require well-sampling of the water column by the VLA; otherwise, the mode filtering
49 will become an ill-posed problem. The difficulty in source depth estimation using a short line array
50 is the lack of spatial information. Thus, some previous studies have relied on the multi-frequency
51 characteristics of wideband signals in estimating source depth [9-10]. Considering the dispersion
52 phenomenon of signals in shallow water, some scholars have estimated source depth on the basis of
53 the related information of marine environment contained in dispersion [11]. However, the above
54 mentioned studies have restricted the signal form to be broadband and few works on source depth
55 estimation for signal with low-frequency line spectrum are available. Recently, data-based
56 MFP/MMP is being proposed for a moving source by using a full-spanning VLA, which is free of the
57 environmental mismatch problem in theory [12-13]. Yang [7] proposed a method based on the idea
58 of data-based MFP/MMP to estimate moving source depth, and this method requires only a single
59 hydrophone. It is more robust than MFP for depth estimation because of not requiring source range
60 estimation. However, the method presents some inherent drawbacks. (1) The method requires
61 source travelling a sufficient range in radial direction owing to that the mode wavenumbers are
62 estimated by synthetic aperture modal beamforming (SAB). Some of the horizontal wavenumbers
63 cannot be estimated, when the moving range does not meet the requirement. Source depth
64 estimation based on synthetic aperture beaming may be incorrect. (2) The model assumes that the
65 signal possesses high signal-to-noise (SNR). Such requirement of SNR is difficult to meet in passive
66 positioning. (3) The algorithm is difficult to apply when the source is not moving in the radial
67 direction.

68 In this study, we propose a source depth estimation approach with an HLA for a moving source
69 in a range-independent environment, without knowing source absolute range information. When
70 beam steers to the source direction, the signal is enhanced after beamforming. We apply the
71 autoregressive (AR) model to the signal enhanced by array gain to estimate the wavenumbers. The
72 AR wavenumber spectrum is unsuitable to estimate mode amplitudes. Thus, the
73 generalized/modified Hankel transform is applied to the data. The source depth is estimated by the
74 peak of the depth ambiguity function wherein the mode depth functions best match the mode
75 amplitudes. Compared with SAB, the proposed method effectively improves the depth estimation
76 performance for the data with short range span or low SNR, and exhibits better adaptability to the
77 notion of the observation platform. In addition, the proposed method can be applied to underwater
78 unmanned vehicles or the autonomous underwater vehicles because it identifies the peaks of
79 wavenumber spectrum automatically.

80 The remainder of the paper is organized as follows. The wavenumber estimation for the
81 beamforming output is developed in Section 2. In Section 2.1, the generalized Hankel transform is
82 extended to beamforming output in a range independent environment [7]. A method for
83 wavenumber estimation based on an AR model without knowing the source range is proposed in
84 Section 2.2. Source depth estimation based on wavenumber spectrum is introduced in Sec 3. The
85 factors that influence the performance of the algorithm are explored in Section 4. The performance of
86 the algorithm on the data with different SNRs and different range spans is analyzed in Section 4.1.
87 Given that the HLA depth also affects the performance of the algorithm, the effect of the laying
88 depth of HLA on the algorithm and its selection method are examined in details in Section 4.2. A
89 summary and discussion is provided in Section 5.

90 2. Mode Wavenumber Estimation

91 We assume that source travels at a fixed depth during observation in a range-independent
92 environment. The considered HLA possesses N receivers that are uniformly spaced with d . The
93 first element of HLA is the reference. The far-field source radiates the continuous wave signal, with a
94 look direction θ_i . θ_i is the true value of the look direction at the i th moment. The range between
95 source and reference array element is r_i at the i th moment. In shallow water waveguides, we
96 suppress a time dependence of the form $e^{j\omega t}$ with ω positive, thus, the pressure field of the n th
97 element at the i th moment can be represented using normal mode theory as

$$p_n(r_i, z_r) = \sqrt{2\pi} e^{-j\frac{\pi}{4}} \sum_{m=1}^M \phi_m(z_s) \phi_m(z_r) \frac{\exp\{-j(k_m - j\alpha_m)[r_i + (n-1)d \sin \theta_i]\}}{\sqrt{k_m [r_i + (n-1)d \sin \theta_i]}}, \quad (1)$$

98 where k_m , α_m , and ϕ_m are the wavenumber, the attenuation coefficient, and the mode depth
99 function, respectively, which all belong to the m th mode; z_s and z_r are the source and HLA
100 depths, respectively; and M is the number of modes considered in the propagation model.

101 The source is at the far field, such that $r_i + (n-1)d \sin \theta \approx r_i$. Equation. (1) is simplified as

$$p_n(r_i, z_r) = \sqrt{2\pi} e^{-j\frac{\pi}{4}} \sum_{m=1}^M \phi_m(z_s) \phi_m(z_r) \frac{\exp\{-j(k_m - j\alpha_m)[r_i + (n-1)d \sin \theta_i]\}}{\sqrt{k_m r_i}}, \quad (2)$$

$$= \sum_{m=1}^M A_m \exp\{-j(k_m - j\alpha_m)[r_i + (n-1)d \sin \theta_i]\}$$

102 where $A_m = \sqrt{2\pi} e^{-j\frac{\pi}{4}} \frac{\phi_m(z_s) \phi_m(z_r)}{\sqrt{k_m r_i}}$. To improve the SNR, beamforming is performed on the
103 signal received on the HLA at the i th moment.

$$B(\hat{\theta}_i) = \frac{1}{N} \sum_{n=1}^N w_n(\hat{\theta}_i) (p_n(r_i, z_r) + n_n(r_i, z_r)), \quad (3)$$

104 $n_n(r_i, z_r)$ is modeled as the Gaussian white noise. For conventional beamforming, the output of
105 beamforming is written as

$$B(\hat{\theta}_i) = \frac{1}{N} \sum_{n=1}^N e^{jk(n-1)d \sin \hat{\theta}_i} (p_n(r_i, z_r) + n_n(r_i, z_r)), \quad (4)$$

106 After substituting Equation. (2) into Equation. (4), while ignoring noise, we obtain

$$B(\hat{\theta}_i) = \frac{1}{N} \sum_{m=1}^M A_m e^{-jk_m r_i - \alpha_m r_i} \text{sinc}(X_m), \quad (5)$$

107 where $X_m = (-k_m - j\alpha_m) \sin \theta_i + k \sin \hat{\theta}_i$, $\text{sinc}(X_m) = \sin(\frac{Nd}{2} X_m) / \sin(\frac{d}{2} X_m)$. $\hat{\theta}_i$ is the
108 estimation of the look direction, which can be obtained by many methods [14]. In order to explain
109 the algorithm conveniently, we discuss the method of estimating the sound source depth when
110 source moves away from the HLA under the assumption that $\hat{\theta}_i = \theta_i = \theta$.

111 2.1. Wavenumber estimation using generalized Hankel transform

112 In range-independent shallow water waveguides, a Hankel transform pair presents the
113 relationship between the complex pressure field $p(r; z_s, z_r)$ and the Green's function
114 $g(k_r; z_s, z_r)$ [15]

$$p(r; z_s, z_r) = \int_0^{+\infty} g(k_r; z_s, z_r) J_0(k_r r) k_r dk_r, \quad (6)$$

$$g(k_r; z_s, z_r) = \int_0^{+\infty} p(r; z_s, z_r) J_0(k_r r) r dr$$

115 where J_0 is the zeroth-order Bessel function.

116 Considering the computation of Hankel transform, the Green's function is often approximated
117 by an inverse Fourier transform (FT) as [15]

$$g(k_r; z_s, z_r) \sim \frac{e^{\frac{i\pi}{4}}}{\sqrt{2\pi k_r}} \int_{-\infty}^{+\infty} p(r; z_s, z_r) e^{ik_r r} \sqrt{r} dr, \quad k_r r \gg 1, \quad (7)$$

118 We consider the source motion model, source range $r = r_i = r_0 + iv\Delta t$, where Δt is the
 119 sampling interval, and r_0 is the unknown initial source range at $t = 0$. We assume that the source
 120 radiates a tone signal. Thus, the source speed v can be estimated by Doppler shift with knowing
 121 original signal frequency. According to Equation. (7), the Hankel transform, which is used
 122 previously to estimate the wavenumber, requires knowledge of the source range. The source range
 123 is difficult to be measured correctly for passive position system. We apply the generalized Hankel
 124 transform proposed in [7] to the beamforming output. The generalized Hankel transform for the
 125 beamforming output is described as

$$\bar{g}(k_r, z_r) = \frac{e^{\frac{i\pi}{4}}}{\sqrt{2\pi k_r}} \int_{r_0}^{r_0+R} B(r) e^{ik_r r} S(r) dr, \quad k_r r_0 \gg 1, \quad (8)$$

126 where R is the range span wherein source moves during the observation time. $S(r)$ is intended
 127 to compensate for the cylindrical spreading loss and will be obtained directly from the data.

$$S(r) = \left\langle |B(r)|^2 \right\rangle^{-\frac{1}{2}}, \quad (9)$$

128 where $\langle \rangle$ is the range averaging or smoothing operation. Equation. (8) reduces to the original
 129 Hankel transform by setting $S(r) = \sqrt{r}$. In the present discussion, $S(r)$ is approximately
 130 proportional to \sqrt{r} using range averaging. After substituting Equation. (5) into Equation. (8),
 131 while assuming $S(r) \sim \sqrt{r}$, we obtain

$$\begin{aligned} \bar{g}(k_r, z_r) &\sim \sum_{m=1}^M \frac{\phi_m(z_s) \phi_m(z_r)}{\sqrt{k_r k_m}} \text{sinb}(X_m) \int_{r_0}^{r_0+R} e^{j(k_r - k_m)r - \alpha_m r} dr \\ &= \sum_{m=1}^M a_m \frac{\phi_m(z_s) \phi_m(z_r)}{k_r - k_m + j\alpha_m} \end{aligned}, \quad (10)$$

132 where

$$a_m = \frac{e^{j[(k_r - k_m) - \alpha_m](r_0+R)} - e^{j[(k_r - k_m) - \alpha_m]r_0}}{j\sqrt{k_r k_m}} \text{sinb}(X_m), \quad (11)$$

133 When $k_r = k_m$, the value of the m th item is much larger than the others in Equation. (10). Thus,
 134 the wavenumber spectral peak at $k_r = k_m$ is given by
 135

$$\bar{g}(k_m, z_r) \sim b_m \phi_m(z_r), \quad (12)$$

$$b_m = \frac{2e^{-\alpha_0 r'}}{\alpha_m k_m} \sinh\left(\frac{\alpha_m R}{2}\right) \phi_m(z_s) \text{sinb}(X_m), \quad (13)$$

136 where $r' = r_0 + \frac{R}{2}$. This expression can be described by the matrix form.

$$\mathbf{g} = \Phi \cdot \mathbf{b}, \quad (14)$$

137 where

$$\mathbf{g} = [\bar{g}(k_1, z_r), \bar{g}(k_2, z_r), \dots, \bar{g}(k_M, z_r)]^T, \quad (15)$$

$$\Phi = \text{diag}([\phi_1(z_r), \phi_2(z_r), \dots, \phi_M(z_r)]), \quad (16)$$

$$\mathbf{b} = [b_1, b_2, \dots, b_M]^T. \quad (17)$$

138 The aforementioned wavenumber estimation method is based on the FT, and is convenient to
 139 be calculated by FFT. However, the FT-based approach presets some inherent disadvantages: (1) The
 140 spectral resolution of FT is limited by the range span. To resolve the modal wavenumbers between
 141 the i th and j th modes, the range needs to be larger than the interference distance $d_{ij} = \frac{2\pi}{|k_i - k_j|}$. If
 142 all the wavenumbers are estimated, then the range must be larger than all interference distances. (2)
 143 The spectrum leakage occurs seriously when the method is applied to the data with a short range
 144 span. The sidelobes of strong spectral components can contaminate the weak spectral components or
 145 generate a false spectral peak. With the increase in source frequency or waveguide depth, the lower
 146 mode wavenumbers will closely group together. As a result, the wavenumbers are difficult to be
 147 estimated in this environment. Considering the drawbacks of FT, the modern spectral estimation
 148 methods can be applied to wavenumber estimation.

149 2.2. Wavenumber estimation using AR model

150 Several methods have been developed to extract the mode wavenumbers, such as Prony's
 151 method, the signal subspace algorithms and matrix-pencil [16,17]. However, these methods regard
 152 the number of wavenumbers as a priori. The number of wavenumbers cannot be known correctly in
 153 practice. The AR spectral estimators based on an all-pole model are often used to extract the spectral
 154 peaks in frequency estimation [17]. The AR estimator does not need to know the number of
 155 wavenumbers, and is thus attractive to estimate wavenumbers. However, the source range must be
 156 known for the AR estimator. Consequently, improved AR estimator should be used to extract the
 157 mode wavenumbers.

158 The method can be divided into three steps. First, data are preprocessed by

$$y[i] = B(r_i)S(r_i) \quad i = 1, 2, \dots, L, \quad (18)$$

159 Second, considering that $y[i]$ can be described as the output of a linear system, we construct the
 160 AR model as

$$y[i] = -\sum_{k=1}^p a[k]y[i-k] + u[i], \quad (19)$$

161 where p is the order of the AR model to represent the data, and is often set to $\frac{2N}{3}$ [17]. Finally,

162 we assume that $u[i]$ is a zero mean white noise sequence with σ^2 . Thus, the wavenumber spectral
 163 density P_{AR} is expressed as:

$$P_{AR}(l) = \frac{\sigma^2}{\left| 1 + \sum_{k=1}^p a[k] \exp[-ilk] \right|^2}, \quad (20)$$

164 P_{AR} is obtained by estimating the coefficients $a[1]$, $a[2]$, \dots , $a[p]$ and σ^2 . The
 165 above-mentioned coefficients can be estimated in different ways. We select the modified covariance
 166 approach because it avoids the spectral line splitting effectively [18]. The locations of peaks in P_{AR}
 167 yield the wavenumbers estimated by AR model. The peak levels estimated by AR estimator possess
 168 large variances [18]. Thus, AR estimator is unsuitable to estimate the modal amplitudes.

169 3. Matched-Mode Source Depth Estimation

170 As discussed above, the wavenumbers can be estimated by AR spectrum. However, AR
 171 spectrum is unsuitable to estimate the modal amplitudes. The performance of estimating
 172 wavenumbers degrades by using the generalized Hankel transform. The performance is effected by
 173 false spectral peaks and spectral resolution. To combine the advantages of two methods, the modal
 174 amplitudes are estimated by generalized Hankel transform with a prior knowledge of wavenumbers
 175 that are estimated by AR model. Thus, we estimate the wavenumber k_m by AR model, and then
 176 obtain $\bar{g}(k_m, z_r)$ in the wavenumber spectrum on basis of the generalized Hankel transform.

177 According Equation. (14), we use a source depth ambiguity function to estimate source depth,
 178 and this function is expressed as

$$D(z) = \varphi(z) \mathbf{b} \mathbf{b}^H(z), \quad (21)$$

179 where

$$\varphi(z) = [\phi_1(z), \phi_2(z), \dots, \phi_M(z)]. \quad (22)$$

180 \mathbf{b} can be solved by

$$\mathbf{b} = (\Phi + U)^{-1} \mathbf{g}, \quad (23)$$

$$U = \text{diag}(\Delta / \phi_1(z_r), \Delta / \phi_2(z_r), \dots, \Delta / \phi_M(z_r)), \quad (24)$$

181 where Δ is a small amount (on the order of one half of the maximum value of the mode function)
 182 for preventing the singularity of Φ . Wavenumbers k_m and $\phi_m(z)$ are calculated by KRAKEN on
 183 the basis of a given frequency and environmental information (SSP and bottom properties). Some
 184 modes cannot be resolved even if we use the AR spectrum, such that the number of $\bar{g}(k_m, z_r)$ is
 185 less than M . We assume that only M_0 order is estimated. $\phi_m(z)$ with the same order of
 186 $\bar{g}(k_m, z_r)$ can be determined by solving the problem as follows

$$\begin{cases} \min(\mathbf{k} - \mathbf{k}_0)^H (\mathbf{k} - \mathbf{k}_0) \\ \text{s.t. } \mathbf{k}_0(1) < \mathbf{k}_0(2) < \dots < \mathbf{k}_0(M_0)' \end{cases} \quad (25)$$

187 where

$$\mathbf{k} = [k_1, k_2, \dots, k_{M_0}]^T, \quad (26)$$

$$\mathbf{k}' = [k'_1, k'_2, \dots, k'_M]^T, \quad (27)$$

$$\mathbf{k}_0 \subseteq \mathbf{k}'. \quad (28)$$

188 \mathbf{k} is an $M_0 \times 1$ vector, \mathbf{k}' is an $M \times 1$ vector, and \mathbf{k}_0 is an $M_0 \times 1$ vector.

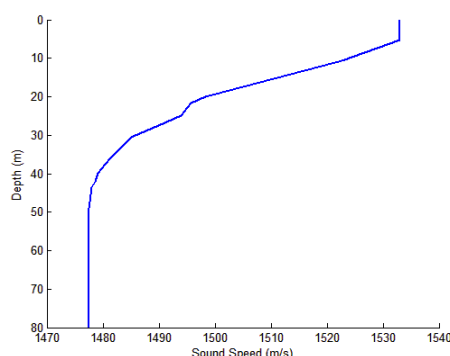
189 4. Numerical Experiments and Performance Analysis

190 This section presents comparisons of the source depth estimation results of the SAB and the
 191 proposed matched-mode autoregressive source depth estimation method (MMAR). Different range
 192 span and SNR should be considered in evaluating the performance of SAB and MMAR in estimating
 193 source depth. For SAB, source depth can be estimated using only a single hydrophone under the
 194 assumption that source moves away from receiver at a constant speed during the observation time.
 195 For convenient comparison, we assume the source of SAB moves along the beam direction of HLA.
 196 Apart from the influence of SNR and range span on the performance of the algorithms, the influence
 197 of the HLA depth on MMAR is also considered. We study this effect from theory and simulation.

198 The pressure field is generated using the KRAKEN program. The SSP for simulation is shown
 199 in Figure 1, and the bottom properties are referred to Reference 7. The consider a HLA possesses N
 200 receivers that are uniformly spaced with d , where $d = \frac{\lambda}{2}$. In studying the difference in depth
 201 estimation performance, we consider two sources depths of 4 and 50 m, which correspond to
 202 shallow and deep sources, respectively. The initial source range is 5010 m, but this information is
 203 assumed to be unknown. The performance influenced by the unknown initial source range has been
 204 discussed in Ref. 7. We assume that each source moves away from the HLA with a speed of 2.5 m/s,
 205 and radiates the narrowband signal with 350 Hz. The SNR in this simulation is defined as:

$$SNR = 10 \lg \frac{P_s}{P_n} \Big|_{r=r_0}. \quad (29)$$

206 According to Equation. (29), when source range is 5010 m, the SNR is described by the ratio of signal
 207 power and noise power at receiver. The array gain is already considered in SNR.



208

209 **Figure 1.** Sound speed profile used in this experiment, which is typical in shallow waveguides.

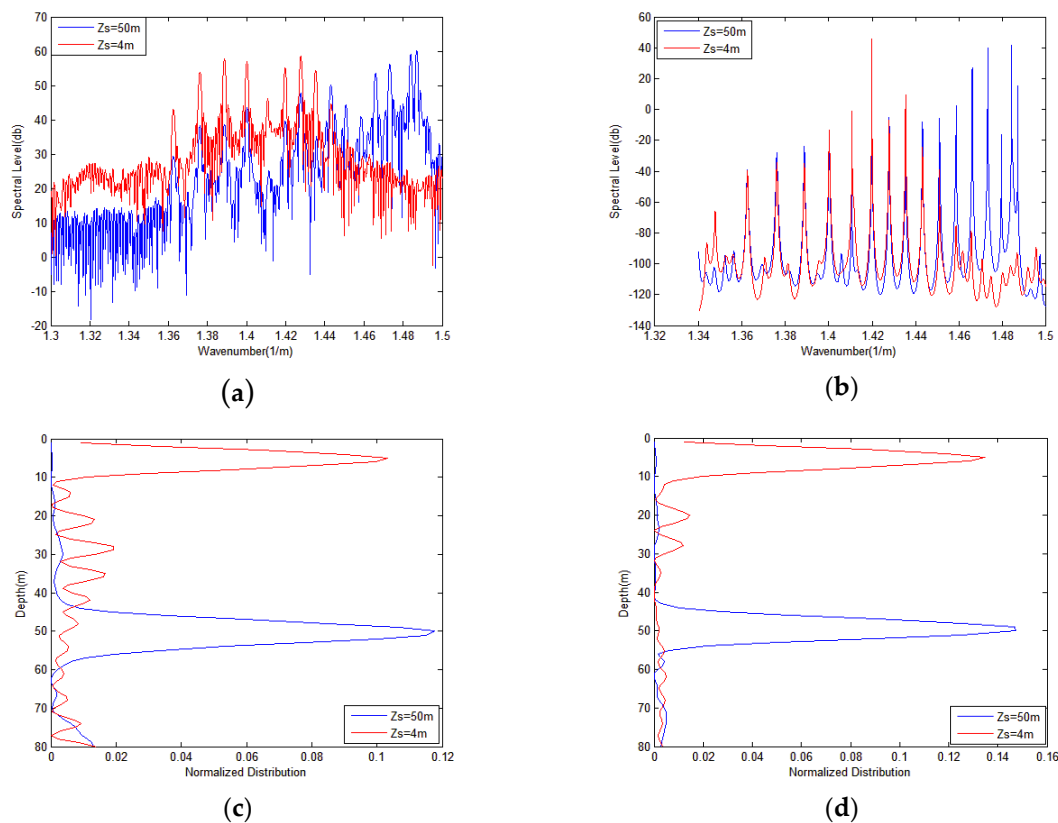
210 4.1. Comparison of SAB and MMAR applied to simulation data

211 The HLA depth is 70 m without special instructions. The effectiveness of our algorithm is
 212 analyzed in the following three cases and compared with that of SAB. A unified description is given
 213 to avoid duplication, that is, the deep source is expressed by the blue line and the shallow source is
 214 expressed by the red line. In Figure 2-4, subplot (a) shows the wavenumber spectrum obtained by
 215 the generalized Hankel transform; subplot (b) shows the AR wavenumber spectrum obtained by AR
 216 model, which is used to estimate the mode wavenumbers in MMAR; the source depths estimated
 217 using SAB and MMAR are shown in subplot (c) and (d). We determine the performance of source
 218 depth estimation under three cases.

219 1. Range span is sufficient and SNR is high

220 The source moves 4990 m during the observation time. The SNR is 40 dB. From Figure 2a,b, the
 221 peaks of wavenumber spectrum obtained by two methods are found to be clear. Therefore, we can
 222 estimate wavenumbers accurately using the two methods under this simulation condition. Both
 223 methods yield good depth estimation results as shown in Figure 2c,d. Comparing Figure 2c,d shows
 224 that the sidelobes of the depth estimation in Figure 2d are smaller than those in Figure 2c, especially
 225 for the shallow source. The number of mode wavenumbers estimated is 9 in Figure 2a and the

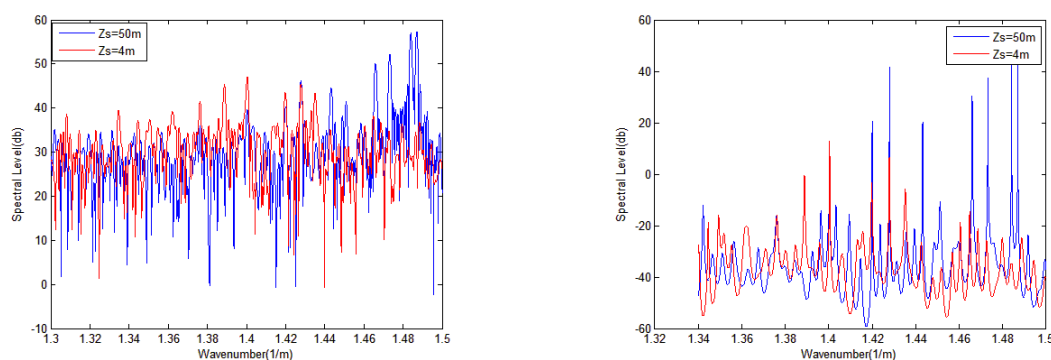
226 number is 11 in Figure 2b. The more mode wavenumbers are estimated correctly, the more
 227 information can be taken into depth estimation. The performance of depth estimation can be
 228 improved by increasing the number of mode wavenumbers estimated. Small sidelobes can also be
 229 obtained. Thus, when the range span is sufficient and the SNR is high, the performance of depth
 230 estimation using MMAR is slightly better than the performance of that using SAB.

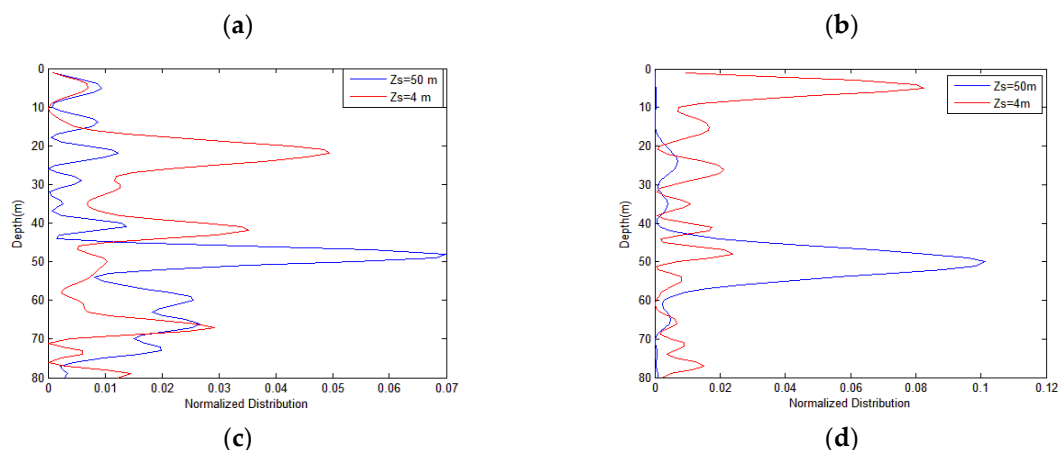


231 **Figure 2.** Comparison of SAB and MMAR in terms of source depth estimation by using simulated
 232 pressure field data covering the range span of 4990 m with an SNR of 40 dB. (a) Mode wavenumber
 233 extracted using the modified Hankel transform. (b) Wavenumber estimated using the AR model. (c)
 234 Depth estimation using SAB. (d) Depth estimation using MMAR.

235 2. Range span is sufficient and SNR is low

236 In this simulation, the range span is 4990 m and the SNR is 5 dB. Only the deep source can be
 237 estimated correctly as shown in Figure 3c. By contrast, the results in Figure 3d show good depth
 238 estimation for the shallow and deep sources. SAB presents difficulty in estimating the mode
 239 wavenumbers correctly as shown in Figure 3a. This difficulty is due to the influence of sidelobe
 240 mask. MMAR based on AR model can reduce the influence of sidelobe mask. Thus, the spectral
 241 peaks can be determined and the mode wavenumbers can be estimated accurately. When the range
 242 span is sufficient and the SNR is low, the performance of depth estimation using MMAR improves
 243 greatly compared with the performance of that using SAB.

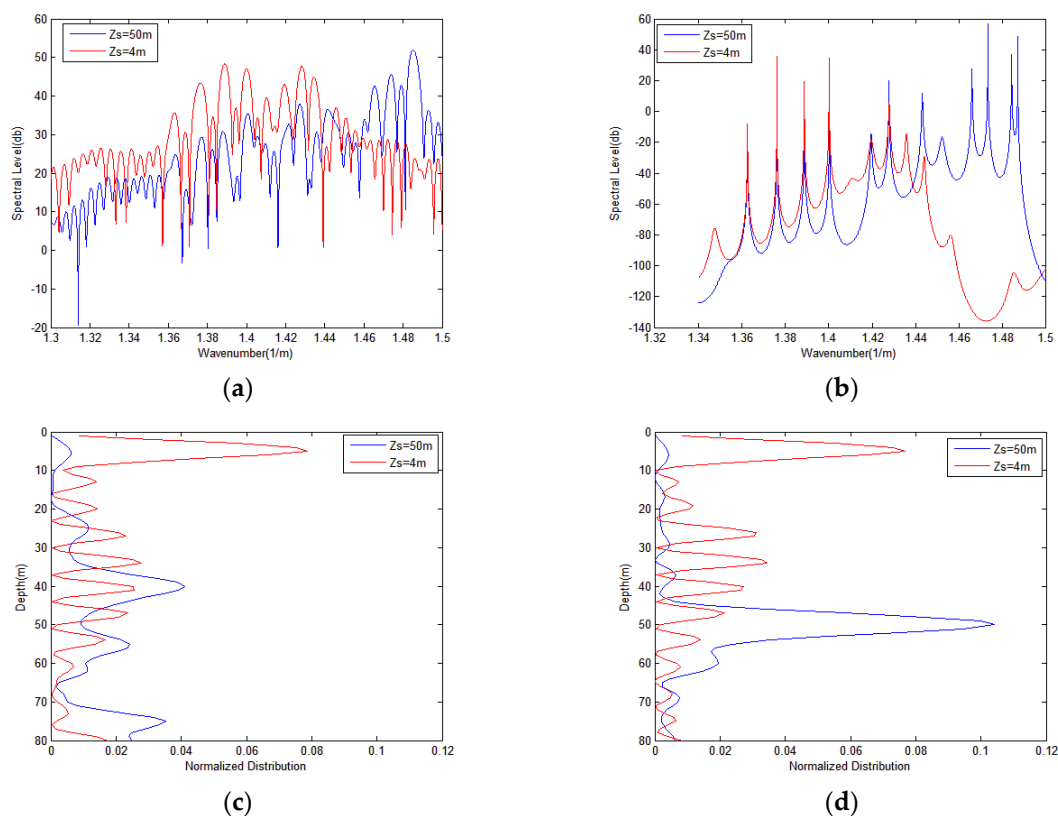




244 **Figure 3.** Comparison of SAB and MMAR in terms of source depth estimation by using simulated
 245 pressure field data covering the range span of 4990 m with an SNR of 5 dB. (a) Mode wavenumber
 246 extracted using the modified Hankel transform of SAB. (b) Wavenumber estimated by the AR model
 247 of MMAR. (c) Depth estimation using SAB. (d) Depth estimation using MMAR.

248 3. Range span is insufficient and SNR is high

249 The source moves 1990 m and the SNR is 40 dB in this simulation. Both methods can estimate
 250 the shallow source depth effectively under this condition as shown in Figure 4c,d. SAB fails whereas
 251 MMAR succeeds to estimate deep source depth. SAB mistakenly selects the sidelobe peaks as
 252 estimations of the mode wavenumbers, and the false wavenumbers are shown in Figure 4a. By
 253 contrast, the peaks of the main lobe can be extracted effectively and the miscarriage of justice can be
 254 reduced using AR model. When the range span is insufficient and the SNR is high, the performance
 255 of MMAR is better than the SAB. In particular, MMAR performs obviously better than SAB in terms
 256 of estimation of deep source depth.



257 **Figure 4.** Comparison of SAB and MMAR in terms of source depth estimation by using simulated
 258 pressure field data covering the range span of 4990 m with an SNR of 40 dB. (a) Mode wavenumber

259 extracted using the modified Hankel transform of SAB. (b) Wavenumber estimated by the AR model
 260 of MMAR. (c) Depth estimation using SAB. (d) Depth estimation using MMAR.

261 The above mentioned analysis is based on the qualitative analysis of a single experiment, and
 262 the following two types of algorithms are analyzed quantitatively from the statistical point of view.
 263 In quantifying the estimated performance in an experiment, the correct index (CI) and output peak
 264 sidelobe ratio (OPSLR) should be defined. CI can be expressed as

$$CI = \begin{cases} 1 & |z_e - z_s| \leq g \\ 0 & |z_e - z_s| > g \end{cases} \quad (30)$$

265 where z_e is the result of source depth estimation and g is the tolerance of the depth estimation.
 266 The OPSLR is defined as follows

$$OPSLR = \begin{cases} \max(P(z)) - \max(P(z')) & CI = 1 \\ 0 & CI = 0 \end{cases} \quad (31)$$

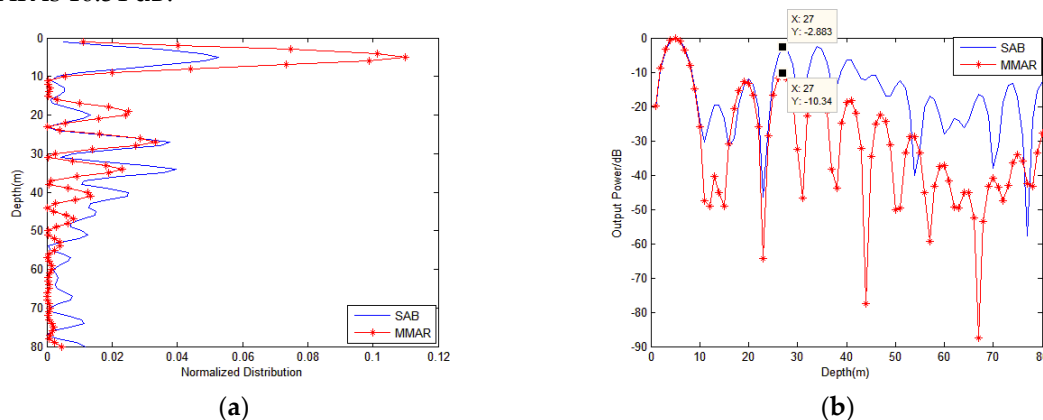
267 where

$$P(z) = 20 \lg \frac{D(z)}{\max(D(z))}, \quad (32)$$

$$z \subseteq [z_{\min}, z_{\max}], \quad (33)$$

$$z' \subseteq [z_{\min}, z_s - g] \cup [z_s + g, z_{\max}], \quad (34)$$

268 We adopt the condition in Figure 4 and consider only the shallow source to intuitively
 269 understand the above-mentioned definition. The depth estimation results of the two algorithms
 270 are shown in Figure 5a. In this case, the two methods can estimate the source depth correctly. The
 271 OPSLR is marked for the two methods in Figure 5b. The OPSLR of SAB is 2.883 dB and that of
 272 MMAR is 10.34 dB.



273 **Figure 5.** Depth estimation. (a) Normalized depth ambiguity functions, (b) Output power at all
 274 depths.

275 The correct probability (CP) and average peak sidelobe ratio (APSLR) are proposed to evaluate
 276 the statistical performance of source depth estimation in multiple experiments. CP can be expressed

277 as $CP = \frac{1}{N_0} \sum_{n=1}^{N_0} CI$, and the APSLR is defined as $APSLR = \frac{1}{N_0} \sum_{n=1}^{N_0} OPSLR$, where

278 number of experiments in this simulation. The CP and APSLR with different range spans and SNRs
 279 for source depth estimation is shown in Tables 1 and 2. The source depths are 4 and 50 m in Tables 1
 280 and 2, respectively. From Tables 1 and 2, the performance of source depth estimation by MMAR is

281 found to better than the performance of that by SAB regardless of the source depth. The source
 282 depth influences the algorithm as shown in Table 1 and 2. For effective comparison of performance
 283 of two algorithms, the source depth is varied from 1 m to 80 m in Figure 6. The source moves 3990 m
 284 and SNR is 20 dB. The similar regulation shown in Figure 6 can be found in other cases shown in
 285 Tables 1 and 2. Figure 6 shows that MMAR presents an improvement in depth estimation for
 286 different source depths compared with SAB.

287

Table 1. The results for the source depth of 4 m.

SNR(dB)	Range(m)	CP		APSLR(dB)	
		SAB	MMAR	SAB	MMAR
40	1990	0.772	1.000	1.440	10.737
30		0.598	1.000	1.121	9.788
20		0.338	0.998	1.018	8.201
15		0.260	0.898	0.841	6.856
20	2990	0.608	1.000	5.199	10.027
10		0.316	0.832	1.246	5.962
20	3990	0.934	1.000	7.924	10.804
10		0.422	0.942	1.604	8.012
10		4990	0.458	0.970	2.142

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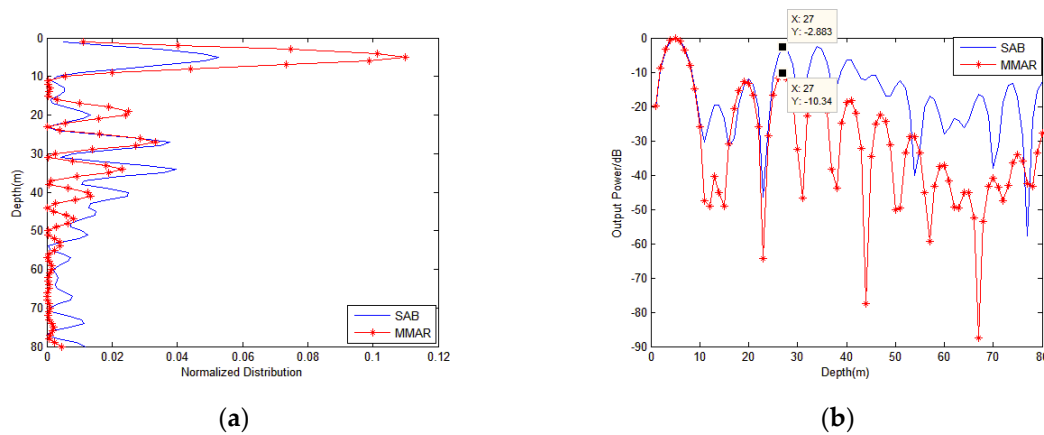
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Table 2. The results for the source depth of 50 m.

SNR(dB)	Range(m)	CP		ASPLR(dB)	
		SAB	MMAR	SAB	MMAR
40	1990	0.000	1.000	0.000	25.907
30		0.000	1.000	0.000	20.491
20		0.000	1.000	0.000	20.788
15		0.002	1.000	0.007	19.192
20	2990	0.670	1.000	2.345	24.510
10		0.484	1.000	2.602	19.420
20	3990	0.680	1.000	5.804	25.830
10		0.820	1.000	6.017	21.508
10		4990	0.704	1.000	8.901

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294 **Figure 6.** Depth estimation performance for the various source depths by using (a) CP and (b)
 295 APSLR.

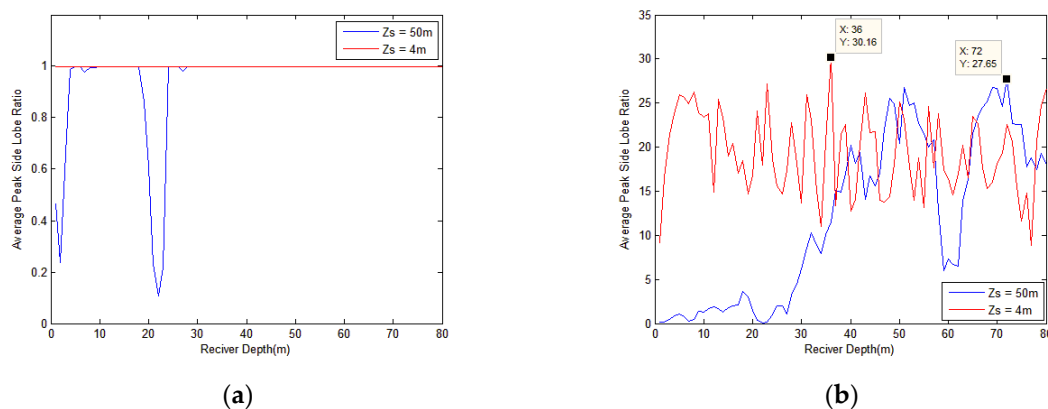
296 4.2. Selection of the HLA depth

297 After substituting Equation. (23) into Equation.(21), we can express the depth ambiguity
 298 function as

$$D(z) = \left| \sum_{m=1}^M \phi_m(z) \frac{b_m}{1 + \frac{\Delta^2}{\phi_m^2(z_r)}} \right|^2, \quad (35)$$

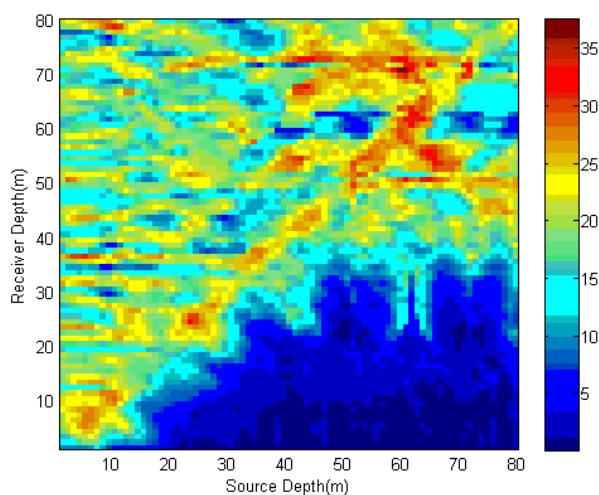
299 From Equation. (35), the mode is found to not contribute to depth estimation when the HLA depth
 300 z_r coincides with the zero-crossing points of the mode depth function, that is , $\phi_m(z_r) \approx 0$. The
 301 performance is influenced not only by the range span and the SNR, but also by the HLA depth.

302 In determining the influence of the HLA depth, we consider the two above-mentioned sources
 303 with a sufficient range span (4990 m) and high SNR. The HLA depth is varied from 1 m to 80 m to
 304 obtain the performance of depth estimation for different HLA depths. The number of Monte Carlo
 305 experiments is 500. Figure 7a shows the performance evaluated by CP with various HLA depths. For
 306 the deep source, the HLA depth significantly influences the correction of depth estimation.
 307 However, for the shallow source, CPs are all ones no matter where the HLA is located. This finding
 308 means that the HLA depth slightly influences the correctness of depth estimation. Consequently, as
 309 the source locates at different depths in the ocean, the HLA depth affects the source depth estimation
 310 differently. The performance evaluated by APSLR is shown in Figure 7b. For the deep source, when
 311 the HLA is located at 72 m, the largest value of APSLR occurs. For the shallow source, the APSLR
 312 presents the largest value when the HLA depth is 36 m. We can conclude from the above-mentioned
 313 observation that the most suitable HLA depth for depth estimation depends on the source depth.
 314 Thus, performance for various source depths and HLA depths should be studied. On the basis of the
 315 analysis present above, a suitable HLA depth span can be obtained for all source depths.



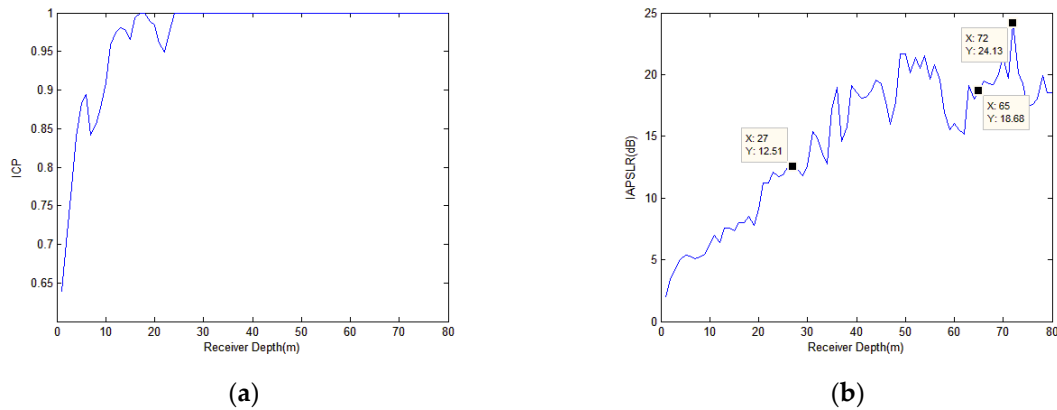
316 **Figure 7.** Depth estimation performance for shallow and deep sources from a single receiver at
 317 various depth using (a) CP, and (b) APSLR.

318 A common conclusion for HLA depth span is obtained after evaluating the performance for all
 319 probable source depth. The performance evaluated by the APSLR for depth estimation under
 320 various source and HLA depths is shown in Figure 8. As shown in the figure, when the HLA is
 321 located near the water surface, the depth is difficult to be estimated correctly for some deep sources.
 322 The optimum placement depth of HLA differs for dissimilar source depths. The general trend is that,
 323 with the increasing of the source depth, the optimal placement depth of HLA also increases. In other
 324 words, the optimum placement depth of HLA is determined by the source depth. Therefore, we
 325 synthesize the depth estimation performance of the source with all possible depths and obtain the
 326 HLA depth span that is suitable for depth estimation. The distribution of the source depth is
 327 expressed by $p(z)$, and is assumed to be uniform. The result of the integrated correct estimation
 328 probability (ICP) for different HLA depths is present in Figure 9. The figure shows that, when HLA
 329 is located below 25 m, the correctness of depth estimation can be guaranteed for all possible source
 330 depths. The integrated average peak sidelobe ratio (IAPSLR) as the placement depth of HLA
 331 increases is shown in Figure 9b. IAPSLR can be calculated by $p(z) \cdot APSLR$. Synthesizing all
 332 possible source depths shows that, when the HLA depth is smaller than 30 m, the IAPSLR increases
 333 with the placement depth of HLA. The general trend of the curve increases from 30 m to near the
 334 boundaries, with an oscillatory plateau in the local part. Furthermore, the IAPSLR presents higher
 335 value at this region than at the region of a source depth smaller than 30 m. When the HLA is located
 336 at 72 m, the IAPSLR possesses the largest value. These curves are valid for the given environment
 337 and frequency, but similar plateau can be observed with other downward-refracting SSPs.
 338 Consequently, when the HLA is located below the transitional layer, we can obtain a satisfying
 339 result of the source depth estimation. The conclusion is studied on the basis of the normal mode
 340 theory.



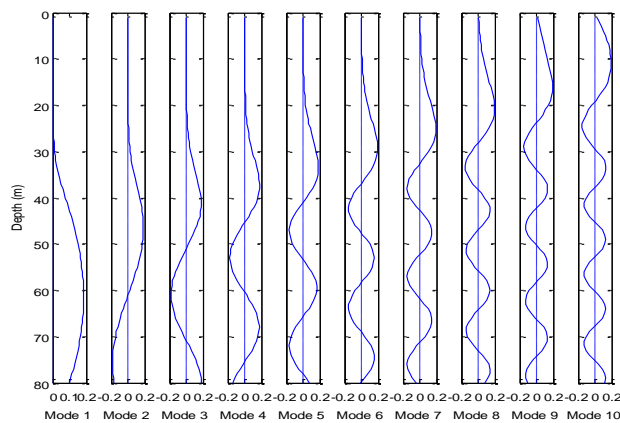
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342 **Figure 8.** The performance of the depth estimation is evaluated by APSLR. The figure shows APSLR
 343 versus HLA and source depths.



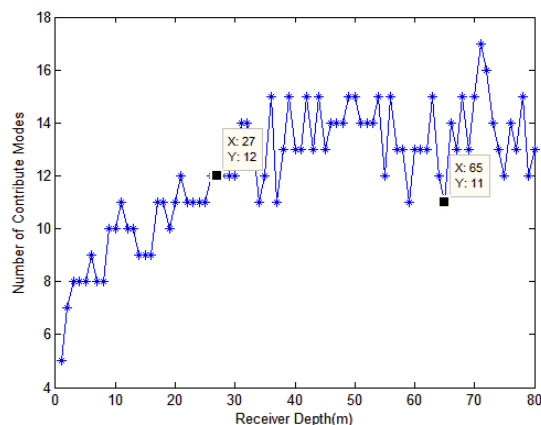
344 **Figure 9.** Curve for depth estimation performance with different HLA depths by using (a) ICP and
 345 (b) IAPSLR.

346 Figure 10 shows the 10 modes excited by the 350 Hz narrowband signal, which is radiated by
 347 the source located at 50 m. The mode excitation amplitudes are the function of depth. For example,
 348 for the modes wherein the HLA is located at a range of 5010 m and a depth of 30 m, several mode
 349 amplitudes (1th, 2nd, 8th, and 10th orders) is close to zero. The proposed method in this paper
 350 estimates the wavenumber before depth estimation. The wavenumbers are difficult to estimate
 351 when $|\phi_m(z_r)\phi_m(z_s)| \approx 0$. The missing information degrades the performance of depth estimation.



352
 353 **Figure 10.** Ten order normal modes excited by the 350 Hz cw. Several mode amplitudes (1th, 2nd,
 354 8th, and 10th orders) is close to zero are shown in this figure.

355 We define the modes that contribute to the depth estimation as the modes with amplitudes
 356 larger than $\Delta/2$. The number of modes that contribute to estimate source depth at various HLA
 357 depths is shown in Figure 11. Compared with other depths for HLA, 71 m and 72 m present a large
 358 number of the modes that contribute to depth estimation. This conclusion is the same as that for
 359 Figure 9b. In addition, the trend observed in Figure 11 is similar to that found in Figure 9b.



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Figure 11. Number of modes that contribute to the source depth estimation at various receiver depths.

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The HLA is located at several depths wherein the IAPSLR is high whereas the number of modes that contribute to depth estimation is small. For example, the number of modes is 11 when the HLA is located at 65 m and the number is 12 when the HLA is located at 27 m (Figure 12). However, IAPSLR at 37 m is high as shown in Figure 7b. The phenomenon can be explained by normal mode theory. Figure 11 shows the amplitudes of modes when the HLA is located at 27 and 65 m. As shown in Figure 10a, the lower order mode amplitudes are approximately zero when the HLA depth is 25 m. Figure 10b shows that the lower order mode amplitudes possess large values when the HLA depth is 65 m. Using Equation. (21), the proposed depth estimation method is made equivalent to

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weighted matched-mode depth estimation with coefficients of $c_m = \frac{2e^{-\alpha_m r_0}}{\alpha_m k_m} \sinh\left(\frac{\alpha_m \Delta R}{2}\right)$. Figure

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13 shows that the weighting coefficients is monotonously decrease with the increase in mode order. This finding indicates that the lower order modes provide high contribution for the depth estimation. The number of modes is smaller but lower order mode amplitudes are higher when HLA is located at 65 m than when HLA is located at 27 m. Consequently, the performance of depth estimation when HLA is located at 65 m is good.

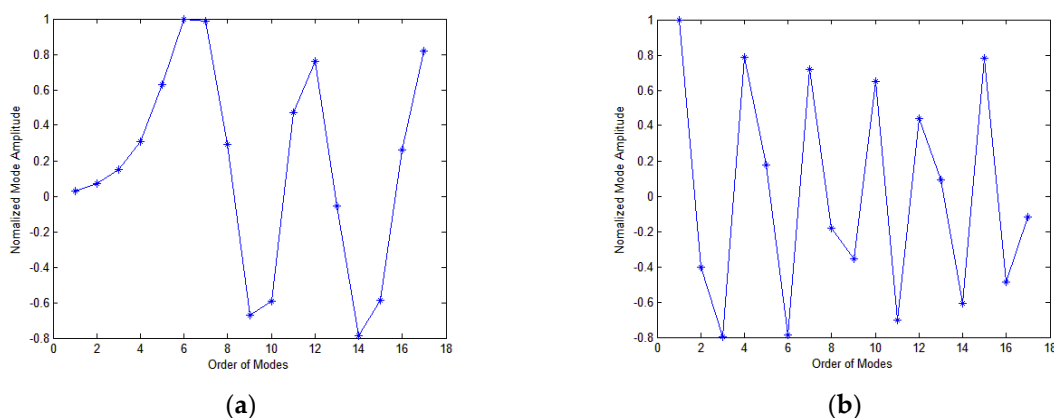
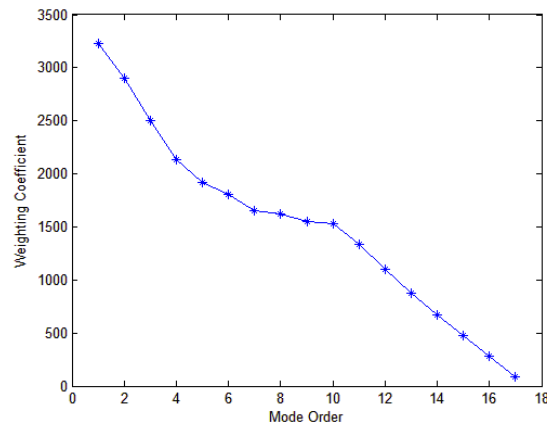
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Figure 12. Variation curve of mode amplitudes and mode orders of modes when HLA is located at (a) 27m and (b) 65m.



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Figure. 13. Weighting coefficients decrease with the increasing mode order.

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On the basis of the discussion presented above, the performance of depth estimation is determined by the number and the amplitudes of contributing modes, which are related with the HLA depth. The combined contribution of both to depth estimation can be expressed as follows

$$B = \sum_{m=1}^M c_m X_m, \quad (36)$$

$$X_m = \begin{cases} 1 & \phi_m(z_r) \geq \Delta/2 \\ 0 & \phi_m(z_r) < \Delta/2' \end{cases} \quad (37)$$

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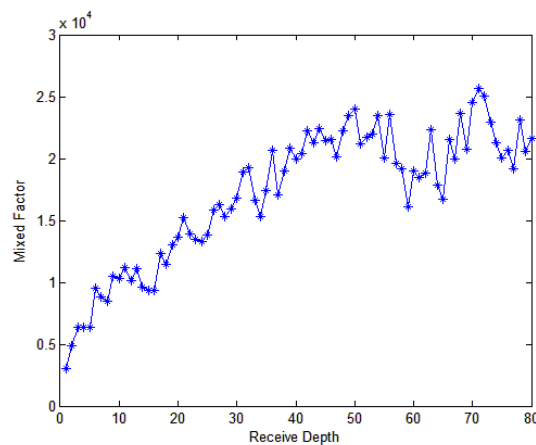
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Figure 14 shows the contribution to depth estimation. Compared with Figure 11 and Figure 9b, a higher similarity is obtained between Figure 14 and Figure 9b. In summary, when the HLA is located below the transitional layer, the number of contributing modes is large and the lower order modes are dominant. Therefore, the combined contribution is larger in this case. Thus, a good depth estimation can be obtained when the HLA is located below the transitional layer.



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Figure. 14. Contribution to depth estimation.

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5. Conclusions

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A matched-mode method based on AR using an HLA to estimate the moving source depth is developed in this study. The modal wavenumber spectrum is obtained using generalized Hankel transform with the data from the moving source. The mode amplitudes can be extracted by combining the information based on AR modal wavenumber spectrum and the FT wavenumber spectrum. The amplitudes contain the information of source depth. The source depth is estimated by

397 matching the mode estimation with the mode depth function calculated by the KRAKEN. The
398 method is robust for estimating source depth as a result of the robustness of its mode depth function.

399 The method proposed in this study is evaluated using the simulated data. For the data with a
400 small range span or low SNR, the proposed method can achieve source depth estimation with better
401 performance than SAB. The depth estimation performance is influenced not only by the range span
402 and SNR but also by the HLA depth. The selection for HLA depth is studied and the result is
403 explained by normal mode theory. In summary, when the HLA is located below the transitional
404 layer, good depth estimation result can be expected. This conclusion is valid for the shallow
405 waveguides with downward-refracting SSP.

406 Compared with that in SAB, the requirement of the moving source traveling range decreases in
407 proposed method. The proposed method can also be applied to the data with low SNR. However,
408 this method is limited because it assumes the source depth fixed during the observation time. The
409 effectiveness of this method degrades sharply, when the source depth varies rapidly. Thus,
410 additional work is needed for its future application in source depth tracking.

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416 Yi-Feng and Liang Guo-Long performed the experiments; Zhang Yi-Feng, Liang Guo-Long and Zou Nan
417 analyzed the data; Wang Jin-Jin contributed reagents/materials/analysis tools; Zhang Yi-Feng wrote the paper.

418 **Conflicts of Interest:** The authors declare no conflict of interest.

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