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Article

Is Entropy a Force of Nature?

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Abstract: The Universe is treated as an unbound thermodynamic system in which two forces of Nature compete: Attractive gravity loss and repulsive entropic gain. Entropic gain, although not presently considered a force of Nature, is otherwise well known and can be expressed with the gas laws. The Universe at the time of last scattering was all gas. Its instant gain is expressed as gas pressure. When the Gibbs equation is properly applied to this unbound gas, the resulting atomic Hubble parameter is found to be exclusively dependent on baryon density, giving a constant and perpetual two-to-one gain/loss ratio. Instant loss to gravity cannot offset gas pressure in this “thermal equation”. By contrast, the Friedmann equation, a Λ CDM foundation, treats baryon mass as accreted, having little if any entropic gain. These two equations are actually equivalent. Expanding accreted matter at its comoving critical density is herein shown to behave exactly like a freely expanding gas. However, neither equation includes all of today’s Universal kinetic energy, and cannot by themselves fully account for observed stellar movement. The present paper proposes the plausible existence of highly energetic or “suprathermal” free electrons which comprise about half of the intergalactic medium’s kinetic total. It is this suprathermal energy which Λ expresses. The fluid equation, another Λ CDM foundation, eliminates entropic gain from the Gibbs equation. Additionally, rest mass’s Einstein energy $E=Mc^2$ is treated as a thermal variable. These two assumptions, isoentropy and energy conflation, led astronomers to populate Λ as they followed the prior art.

Keywords: dark energy; cosmic background radiation; dark ages; large-scale structure of the Universe

1. INTRODUCTION

After Arno Penzias (b. 1944) and Robert Wilson’s (b. 1936) discovery [1] of the cosmic microwave background (CMB), consensus among astronomers converged around the hot-big-bang concept of the Universe’s origin. The competing “static” or non-expanding Universe was conclusively excluded as a proper model. This discarded model left a legacy: “dark energy” Λ , a repulsive scalar field, which kept a static Universe from collapsing. Dark energy Λ was proposed [2] by no less of an intellect than Albert Einstein (1879-1955) himself, the greatest theoretical physicist who ever lived. For the next 33 years after Penzias and Wilson, Λ was kept in circulation but thought to have no value, as Edwin Hubble’s (1889-1953) linear distance-ladder model [3] was adequate to explain observed stellar movement within the uncertainties of the day. That changed in 1998. Two independent groups of astronomers, in a pair of *tours de force* [4–7], found deviance from the linear distance ladder. The same deviance. They populated Λ , giving rise to today’s “ Λ CDM” model [8].

The Λ CDM model has been applied to the CMB’s tiny anisotropies to calculate the Hubble constant H_0 [9]. These results differ from the distance ladder [10–15]. Explanations have been proffered to resolve this tension [16–26]. Most but not all of these rely on Λ . One alternate, quintessence [27], having a time-dependent scalar field, has been proposed. Quintessence is, like Λ and the energy field first proposed by Peter Higgs (b. 1929)[28], held to exist *in vacuo*. Of these three only the Higgs field has an experimentally consistent *in vacuo* theoretical foundation, and to date neither Λ nor quintessence have been shown to have a clearly defined, energy-conservative source.

Attempts to treat the second law of thermodynamics within a Λ -containing Universe have arisen in numerous papers, many of which (2,000 and counting) cite an original publication [29] by Erik

Verlinde (b. 1962). While Verlinde's proposed entropic "screen" or end state has proven quite popular, it tends to indicate that at some point in time, Universal entropy reaches a final value with no further increase. Inconsistency with the second law remains in such treatments.

The present paper takes an approach to Universal expansion consistent with the second law by treating baryon content as the unbound gas it mostly is. The model that arises from this treatment is called the GCDM model, for gas-cold-dark-matter. The GCDM model does give a repulsive field, plasma kinetic energy, but this field bears little theoretical resemblance to a quintessence, Λ , or Higgs field. Its density is a miniscule fraction of Λ 's purported value, and its ultimate source is nuclear fusion in stars.

The difference between Λ CDM and GCDM is fairly easy to picture. The Λ CDM model treats the Universe as all accreted matter, like rocks all hurtling away from each other. Their mutual gravitational attraction isn't enough to pull them back together. They slow down, but still keep getting further and further apart. Their separation is accelerated by Λ . The GCDM model treats the Universe as an infinitely massive gas. It has some rocks, we live on one of them. But mostly it's a gas. The Universe has no boundary, so the gas is freely expanding. Gas free expansion carries a repulsive force: Thermal pressure. The attractive force of gravity from all kinds of mass can't offset the gas's thermal pressure. Extra force from nonthermal pressure accelerates the expansion.

Gas free expansion gives entropic gain. Herein I propose entropic gain as a force of Nature. When bound, it keeps a balloon inflated. When unbound, it propels rockets. It also pushes distant galaxies apart. The presently defined four forces of Nature cannot account for any of these behaviors. We can use the gas laws to express this force, more properly treating the Universe as an unbound gas and not as expanding accreted matter.

Einstein is widely considered as first among equals in the pantheon of great theoreticians. To this author, J. Willard Gibbs (1839-1903) is a close contender for the crown. It is only through both of these scholars' teachings that we can properly understand Universal behavior.

2. THE GIBBS EQUATION, BOUND AND UNBOUND

Since the Gibbs equation plays such a central role in the development of both Λ CDM and GCDM models, it's important to understand the meaning of its terms, and its expression of the first two thermodynamic laws.

The first and second laws of thermodynamics were developed in the 19th century by many authors and fully quantified by the early 20th, notably by Gibbs.¹ The first law says that energy is neither created nor destroyed:

$$dE/dt = 0 \quad (1)$$

Where E is the *total energy*. Total energy is also called "internal energy".² It's the sum of all forms of energy in the Universe as a whole and in a sufficiently large homogenous and isotropic proxy sphere, generally accepted as not less than 200 megaparsecs (Mpc) or about 600 million light-years (ly) in observed diameter. A sphere this large is said to be "at scale". Long ago, the Universe had a more uniform density, and "at scale" can refer to much smaller comoving volumes for those earlier times.

The second law of thermodynamics says that entropy always increases with time (regardless of scale):

$$dS/dt > 0 \quad (2)$$

An isoentropic process cannot occur in real time. You can slow entropy down in e.g. an insulated vessel, but you can't stop it. At scale, $dS/dt \gg 0$.

Equation (2) can't be directly compared to (1) because they have different units. At scale, (2) can be directly compared to (1) by expression as gain E_s :

¹ Hermann von Helmholtz (1824-1891) is also widely credited.

² "Internal energy" was coined by engineers to describe a gas's thermal energy U_i . Cosmologists added rest energy $E=Mc^2$ to its meaning.

$$d(E_s)/dt > 0 \quad (3)$$

For a gas:

$$d(E_s) = d(TS) \quad (4)$$

In this paper, (3) and (4) refer to an *unbound system*: The Universe at scale. “System” usually means e.g. a bound vessel and its contents, for example, gas in a piston. Herein it also refers to a constant amount of gas at an instant in time, and more generally instant total energy, that isn’t bound. In a bound system, it’s common for $d(TS)/dV < 0$ and $d(S)/dV > 0$ if work is performed. However, for both the system and its surroundings combined, (3) is always true. At scale, system and surroundings are indistinguishable. They’re one and the same thing.

When bound, a gas’s pressure P , volume V , temperature T , entropy S , and internal energy E are connected by the *Gibbs equation*:

$$dE = TdS - PdV \quad (5)$$

If the atomic nuclei aren’t fusing or fissioning, their rest mass doesn’t change, so bound internal energy change dE is just the change in the atoms’ collective kinetic energy $d(U_i)$:

$$dE = d(U_i) \quad (6)$$

The term U_i is also known as “thermal” energy. Practically, for a bound gas:

$$d(U_i) = TdS - PdV \quad (7)$$

In a vessel, we can say the gas’s mass doesn’t change, so its entropy change dS and heat flow dQ into the vessel from its (warmer) surroundings are related:

$$dQ = TdS \quad (8)$$

At a reequilibrium $T_2 = T_1$, all the heat flow $\int dQ$ is converted to work $\int PdV$. The vessel’s internal energy U_i (and PV) doesn’t change, but its entropy S and volume V have increased. If there’s no heat flow into the vessel, both U_i and PV drop as the gas converts thermal energy into work.

Unbound gases behave differently from bound gases: They freely expand *ad infinitum*. There is no reequilibrium. System and surroundings merge, so there is no heat flow, and no change in total energy. However, the thermal energy of the unbound gas does change. It drops without performing work as the atoms collide and repel each other:

$$d(U_i) = d(TS) = -d(PV) \quad (9)$$

Free expansion has kinetic energy which is different from random thermal atomic movement U_i . We will refer to the kinetic energy of free expansion as *kinetic gain*, termed E_k . Kinetic gain E_k is best understood by example. We can shoot a rifle in outer space, using either a blank cartridge or a live round.

With a live round, the gas performs work on the bullet until it exits the barrel. The bullet’s acceleration gives equal-and-opposite recoil. The astronaut holding the gun drifts backwards. The expanding gas’s entropy during recoil doesn’t increase much; most of the PV loss is work. Once the bullet leaves the barrel, the gas freely expands into outer space, and its E_k looks almost constant if we’re holding the gun. The atoms’ individual momentum tensors change relatively little as they move away from our astronaut. From the gas’s reference frame, it’s the gun and bullet that are moving away. The gas is roughly an expanding oblate spheroid of atoms with reduced internal energy U_i . These atoms can be subsumed *in toto* by an imaginary sphere around an atom close to the sphere’s center of gravity. The atoms’ outward (radial) kinetic energy in this sphere is 100% entropic E_k . The expanding sphere’s entropy dS and gain $d(TS)$ continues to increase as the atoms move apart. Collisions approach zero. For an amount of mass this small, E_k doesn’t change after collisions end, and is irreversible when unbound, giving entropic force $d(TS)/dr$. Thermal energy $U_i \rightarrow 0$, as $dS > 0$, and the atoms’ $E_k = d(TS)$ neither changes nor comprises U_i . This is further examined in Appendix A.

With a blank, there’s no bullet, and no work performed. The gas is freely expanding inside the barrel, again giving recoil more like a rocket’s thrust, but now the PV loss in the barrel is 100% kinetic gain E_k . Again, when we take away the gun, we get the subsuming sphere of gas, which is bigger without the bullet. A proper numeric description of this gas’s free expansion requires an extensive finite-element model. If the subsuming sphere maintains a uniform density as it expands, its

conversion of thermal loss into kinetic gain is more simply described, and can be numerically expressed with just a spreadsheet (see Supplemental Material). At scale, gas is well approximated as uniformly dense and peppered with accreted matter: bullets, rocks, planets, stars, and galaxies, all of which can be kinetically connected to the expanding gas from which they formed (section 4.2).

An unbound sphere of gas at scale works internally against its own gravity. Its thermal loss is expressed with the *unbound Gibbs* equation:

$$d(U_i) - d(TS) + d(PV) = 0 \tag{10}$$

Where $d(PV)$ is the classic bound expression of work. Since unbound $d(PV)$ also comprises $-d(TS)$, this may seem confusing, but a classic $d(PV)$ in (10) aligns its meaning with the bound Gibbs equation (7), and (10) will be shown as an accurate expression of gas thermal behavior at scale.

For a freely expanding sphere of gas with uniform comoving density, its kinetic gain E_k is expressed as:

$$E_k = \frac{d(TS)}{dt} dt = \frac{d(TS)}{dV} dV = \frac{-d(PV)}{dV} dV = -P_{V'} dV \tag{11}$$

Where $P_{V'} = d(PV)/dV$ is the *entropic pressure*. With a rigid boundary, $dV = 0$ and $P_{V'} = P$. At scale, $P_{V'}$ is diminished by the gas's internal gravity. Entropic $P_{V'}$ and thermal P have different meanings for $z < 9$ (see Appendix D). Much of today's $P_{V'}$ cannot exist in a bound state. Although expressable, this extra kinetic gain's theoretical origin is not well understood by the present author. Thermal P is well understood, and comprises about half of today's total $P_{V'}$.

Bound P keeps a balloon inflated. Unbound $P_{V'}$ is pushing the Universe apart. We will now examine this at the time of last scattering.

3. CONSTRUCTION OF THE THERMAL MODEL AT $z = 1089$

3.1. Parameters

The *thermal model*, Eq. (37), expresses the unbound Gibbs equation (10) through a combination of equivalent mass and kinetic energy. We construct the thermal model at the time of last scattering, or more simply "last scatter", with a finite difference method using the radius r of a sphere as the variable. A popular value of the cosmic redshift at last scatter is $z = 1089$ and we use this. A spreadsheet is used for the calculations. This is less satisfactory than an analytic derivation, but it does give solace in that expressions herein describe thermal loss ΔU_i precisely. It is only in the partition of ΔU_i between kinetic gain and gravity loss where error accrues. We treat the gas as a single species with a mean atomic weight \mathcal{K} . The BBN estimate for baryons [30] gives a mass proportion of about 75% hydrogen : 25% helium, so $\mathcal{K} = 1.24 \times 10^{-3}$ kg/mol. Hydrogen was monatomic and nonrecombinant to diatomic form, absent catalysis through aggregation. Atomic collisions are all treated as elastic. The initial parameters for our model are given in Table 1 and were derived by the author from table 6 of Planck 2020 [9]. The baryon temperature T at $z = 1089$ will be set to 2971K, the CMB's extrapolated value. From our chosen H_0 in Table 1, the baryon density ρ_b at last scatter was 5.46×10^{-19} kg/m³. This is very low and we can easily treat the gas as ideal.

Table 1. Values at $z = 0$.

H_0	67.70 km/sec/Mpc
H_0	2.1938×10^{-18} sec ⁻¹
ρ_{crit}	8.6075×10^{-27} kg/m3
Baryons Ω_b	0.04898
Cold dark matter Ω_c	0.26104
Relativistic energy Ω_Λ	0.000091
Dark energy Ω_Λ	0.6908

T_{CMB}	2.6270 K
Note: H_0 and the Ω 's were calculated from table 6 of [9], except $\Omega_\Lambda = 1 - (\Omega_b + \Omega_c + \Omega_\Lambda)$.	

3.2. The Universe at the time of last scattering was Euclidean

The observed tiny wiggles [9], in the otherwise perfect Boltzmann³ distribution of the CMB, are telling. There is a metric, η_{slip} , found from the wiggles. It describes variance in the behavior of light between Einstein's and Newton's⁴ Universes at last scatter. If $\eta_{slip} = 1$, then there was no variance. The value of η_{slip} was found to be 1.004 ± 0.007 . The text in [9] proclaims that such a value of η_{slip} justifies an assumption of minimal gravitational anisotropy. Gravitational stress tensors in Einstein's field matrix vanish. This allows us to say that at last scatter, the entire instant Universe was effectively Euclidean.⁵ Thermal energy U_i is nonrelativistic at 2971K, so atomic momentum tensors were Newtonian, which allows the gas laws to be applied to their collective behavior. The wiggles do indicate minor density variance, attributable in part to baryon acoustic oscillation, i.e. "sonic waves". This variance may have been negligible back then, but these acoustic waves began to overlap within a colder and colder Universe, creating permanent regions of high gas density whose formation was abetted by the cold. Within these regions of higher density, acoustic resonance gave occasional antinodes that were dense enough to cause gravitational collapse into stars via a process first described by James Jeans (1877-1946) [31]. Treatment of baryon mass using all of general relativity becomes necessary after that, and we will accept without reservation the consistency of myriad observations of the behavior of energy in all its forms with the predictions of general relativity. That said, the Euclidian approximation at last scatter remains accurate. We can also treat the intergalactic medium (IGM) today as Euclidean. The Universe is widely regarded as mathematically "flat". A flat Universe is Euclidean *ad infinitum*, and the IGM today comprises about 90% of the Universe's volume. Accreted matter is only a local perturbation within this larger and much more massive volume. Progress of time in the IGM has been linear ever since last scatter. Its density was and is so low that negligible gravitational time dilation has occurred from then to now.

3.3. The thermal model at the time of last scattering

The thermal model is constructed using monatomic gas expressions found in many introductory engineering textbooks and Wikipedia. Most of the z values in the thermal model (9→1089) cover the dark age [31,32]. The effect of ionizing radiation on the thermal model at $z \geq 9$ is practically nil.

3.3.1. Adiabatic energy release

Consider a comoving sphere of initial radius r_1 around a single atom of H_1 , at 2971K and $\rho_b = 5.46 \times 10^{-19} \text{ kg/m}^3$. All the other atoms are surrounded by an r_1 sphere as well. The gas is monatomic, so:

$$U_i = \frac{3}{2}PV \quad (12)$$

There are two competing forces of Nature acting on this sphere: Repulsive entropic push, and attractive gravity pull. We are using a finite difference method, so we define an *increment*: $\frac{(r_2 - r_1)}{r_1} = \frac{\Delta r}{r}$, which must be kept below 10^{-4} for most purposes to minimize the partition error. I use 10^{-9} , as low as the spreadsheet will tolerate. When the gas in the sphere expands, it must do so adiabatically, and there's no void outside the sphere into which free expansion can occur. Under classic bound

³ Ludwig Boltzmann (1844-1906).

⁴ Isaac Newton (1642-1727).

⁵ Euclid of Alexandria (ca. 300 BCE).

conditions, the sphere would then have to lose U_i (through work). From (9) we can postulate that this adiabatic thermal loss occurs in our cosmic setting as well. For monatomic gases it is:

$$U_{i_1} - U_{i_2} = \Delta U_i = \frac{3}{2} P_1 V_1 \left(\left(\frac{V_2}{V_1} \right)^{-\frac{2}{3}} - 1 \right) \quad (13)$$

Where the numeric subscripts refer to the before and after U_i and V values. Volumes V_1 and V_2 are readily found ($4\pi r^3/3$). Thermal pressure P_1 is found from the ideal gas law:

$$P = \frac{\rho RT}{\mathcal{K}} = \frac{MRT}{\mathcal{K}V} = \frac{3MRT}{4\mathcal{K}\pi r^3} \quad (14)$$

If work against gravity is negligible, there is no alternative to free expansion within the sphere that I can find, so the released energy (13) is 100% kinetic gain. The finite differential value ΔE_s is the *increment* kinetic gain $E_{k\Delta r}$:

$$E_{k\Delta r} = \Delta E_s = \int_{V_1}^{V_2} d(E_s) dV = \int_{V_1}^{V_2} d(TS) dV = \int_{V_1}^{V_2} (T dS + S dT) dV \approx (S_2 - S_1) \left(T_2 + \frac{1}{2}(T_1 - T_2) \right) \quad (15)$$

The movement of any one atom is no longer entirely thermally random. It has outward radial momentum and kinetic energy E_k . Volume increase is strictly local to the sphere. In an infinitely large Universe, it all just gets less dense. The pressure gradient which drives this density decrease is temporal, not spatial.

3.3.2. Gravity loss

We now look at the gravitational potential energy U of the sphere:

$$U = \frac{-3GM'^2}{5r} \quad (16)$$

Where G is the gravitational constant. The potential energy U must take into account the *total mass* M' , not just the baryons' thermal mass M . There's evidence for the existence of cold dark matter (CDM) which is held to be about five times as abundant as baryon mass. Its only interaction with nucleons⁶, electrons, or light, is through gravity. CDM is herein treated as scalar in xyz . While it can move relative to accreted baryons like stars, all that occurs within the gravitationally bound network of galaxies known as the "cosmic web", and within Λ CDM, does not affect H . A consistent description of CDM's composition and origin is elusive [34]. Cold dark matter's density evolution is presently treated nonrelativistically; we use this convention. Its density is kept constant with respect to the baryons. Both follow $1/r^3$, as expressed by the scale factor a :

$$a = \frac{r}{r_0} = \frac{1}{(z+1)} = \frac{1}{z'} \quad (17)$$

where r_0 is the comoving radius of a sphere today, and z is the cosmic redshift:

$$z = \frac{\lambda_{ob} - \lambda_{em}}{\lambda_{em}} \quad (18)$$

Where λ_{ob} is the observed wavelength of light of known laboratory value, λ_{em} . There's also relativistic CMB energy density ϵ_{CMB} , which at last scatter fully comprised photon density ϵ_λ . It follows $1/r^4$. The effect of these combined densities on H is expressed in the minimum flat-universe Λ CDM model by (19):

$$H_\Lambda = H_0 \sqrt{\Omega_\lambda a^{-4} + \Omega_b a^{-3} + \Omega_c a^{-3} + \Omega_\Lambda} = H_0 \sqrt{\epsilon_{M'}} \quad (19)$$

Where H_Λ is the Λ CDM Hubble parameter, and H_0 is today's Hubble constant ($z=0$), found with a telescope. The Ω values are energy density ratios $\epsilon/\epsilon_{M'}$ at $z=0$, and $\epsilon_{M'}$ is the total energy density, usually called the "critical energy density" ϵ_{crit} .

Our two models' densities comove with a :

$$\frac{3H^2 c^2}{8\pi G} = \epsilon_{crit} = \epsilon_{M'} \neq \rho_{M'} c^2 \quad (20)$$

⁶ "Nucleon" doesn't include electrons; "Baryon" does.

Where c is the speed of light. The reference density $\epsilon_{M'_0}$ is given from (20) at $H = H_0$.

In Λ CDM, $\epsilon_{M'} = \epsilon_{crit}$ is used. It includes repulsive dark energy Ω_Λ . In GCDM, the total mass density $\rho_{M'}$ is used. It differs from ϵ_{crit} in that dark energy Ω_Λ isn't included. Dark energy is considered an artifact of the fluid equation (Section 6.2). The baryon mass density ρ_b at $z = 0$ is $(\Omega_b \epsilon_{M'_0} / c^2)$. Proper total density $\rho_{M'_0}$ and earlier $\rho_{M'_z}$ values result, as shown below.

The Ω 's in Λ CDM always add up to one at any given a . These contributions to total mass are expressed within GCDM via a *density multiplier* \mathfrak{m} :

$$\mathfrak{m} = \frac{\Omega_\Lambda a^{-4} + \Omega_b a^{-3} + \Omega_c a^{-3}}{\Omega_b a^{-3}} = \frac{\Omega_\Lambda a^{-4} + \Omega_{(b+c)} a^{-3}}{\Omega_b a^{-3}} \quad (21)$$

The “ \mathfrak{m} ” term is dimensionless and expresses the ratio of total mass to baryon mass. It's derived from dark-energy-containing Ω 's but does not depend on dark energy content.

The total mass M' of a sphere is:

$$M' = M\mathfrak{m} = M \left(\frac{\Omega_\Lambda a^{-4} + \Omega_{(b+c)} a^{-3}}{\Omega_b a^{-3}} \right) \quad (22)$$

Total mass $M' = 6.313M$ at $z = 0$ and increases to $M' = 8.336M$ at $z = 1089$.

The energy lost to gravity upon incremental expansion of the sphere is:

$$\Delta U_r = U_1 - U_2 = \frac{-3GM'^2}{5} \left(\frac{1}{r_1} - \frac{1}{r_2} \right) \quad (23)$$

Where U_1 and U_2 are the before and after gravitational potential energies, respectively.

Atoms can freely expand without colliding. At scale they can perform work without colliding. When the atoms of a model sphere move away from the center, they are climbing out of a gravity well caused by the reduced density resulting from their movement, and E_k diminishes accordingly. It is this loss of radial kinetic energy to gravity which gives ΔU_r .

3.3.3. Release equals loss: the adiabatic sphere

Combining (9), (10), (13), and (23) gives a finite expression of thermal loss \rightarrow kinetic gain:

$$E_{k_{\Delta r}} = \left(\frac{3}{2} \right) P_1 V_1 \left(\left(\frac{V_2}{V_1} \right)^{\frac{2}{3}} - 1 \right) - \frac{3GM'^2}{5} \left(\frac{1}{r_1} - \frac{1}{r_2} \right) \quad (24)$$

A finite expression of conserved total energy is given by (25):

$$E_1 - E_2 = \left[(M_b c^2 + M_e c^2 + M_c c^2 + E_{CMB_1} + U_{i_1} + U_1) - (M_b c^2 + M_e c^2 + M_c c^2 + E_{CMB_2} + \Delta E_{S_\lambda} + U_{i_2} + U_2 + E_{k_{\Delta r}}) \right] = \Delta U_i + \Delta U_r - E_{k_{\Delta r}} = 0 \quad (25)$$

Where E_1 and E_2 are the total energies of the before and after sphere. The CMB gain $\Delta E_{S_\lambda} = E_{CMB_1} - E_{CMB_2}$, discussed in section 6.3.2, is decoupled from E_k at $z < 1089$ and separately expressed. The nucleon rest mass M_b , electron mass M_e , and CDM mass M_c are unchanged at last scatter so their rest energies Mc^2 cancel. Eq. (25) is an expanded restatement of the unbound Gibbs equation (10), connected by $\Delta U_r = \Delta(PV)$ and $E_{k_{\Delta r}} = \Delta(TS)$.

When $E_{k_{\Delta r}} \rightarrow 0$ we get an *adiabatic sphere*. Thermal loss is completely taken up by gravity: $\Delta U_i = -\Delta U_r$. Energy is conserved within all of these spheres, and (1) is obeyed. The radius r_e is its adiabatic radius or *endpoint*, found by convergence of r around $-\Delta U_i / \Delta U_r = 1$. If we use $H_0 = 67.70$ km/sec/Mpc, then at last scatter, $r_e = 9.691 \times 10^{16}$ m (about 10 ly). It's very big. The sphere's imaginary boundary is its *adiabatic surface*. In today's variegated Universe, the adiabatic surface around a central atom isn't always spherical due to e.g. anisotropic stress near the cosmic web. In these regions E_k behaves like thrust and must be expressed with entropic stress tensors not found within Einstein's field equation matrix. Deep in today's IGM, Einstein's and Newton's Universes closely converge, and the surfaces are near spherical. At last scatter, it's all spheres.

For an adiabatic sphere, the postulate connecting classic to cosmic gas behavior is clearly seen. The thermal loss in the sphere just balances gravity, like a piston's expansion just holding up a weight. The postulate holds for lesser, *medium* spheres, as their thermal loss creates kinetic gain ($E_k > 0$). These medium spheres aren't adiabatic since kinetic energy is escaping from them.

The adiabatic sphere's mass changes as it expands. We are following a specific amount of gas only instantly, at the differential limit. Over time we follow conserved energy.

The adiabatic sphere at the time of last scattering is a central reference from which both earlier and later values can be derived. We will call its radius the *pure endpoint* and assign it a term, $r_{e_{1090}}$. It can be calculated:⁷

$$r_{e_{1090}} = \frac{6.560754 \cdot 10^{18}}{H_0} \quad (26)$$

Where H_0 is given in km/sec/Mpc.

3.3.4. The expanding adiabatic sphere

The adiabatic sphere contains medium spheres which all have $E_k > 0$. To determine how fast it is expanding, we have to figure out how fast these lesser spheres expand, and add up their combined radial speeds.

The finite kinetic gain $E_{k_{\Delta r}}$ of a medium sphere gives its *increment radial velocity* $v_{s_{\Delta r}}$:

$$v_{s_{\Delta r}} = \sqrt{\frac{2E_{k_{\Delta r}}}{M}} \quad (27)$$

Which is best visualized as each and every atom in the sphere moving away from the center at the same speed. Its instant or *true* velocity v_s gives a true E_k , and vice versa:

$$v_s = \sqrt{\frac{2E_k}{M}} \quad (28)$$

Below a cutoff radius $r_c (= 0.003r_e)$, loss to gravity is negligible, and all these *small spheres* give the same *initial radial velocity* v_i , developed in Appendix B:

$$v_i = \sqrt{\frac{2U_i}{M}} = \sqrt{\frac{3RT}{\mathcal{K}}} = \sqrt{\frac{3(8.3145)(2971)}{(0.00123988)}} = 7731 \text{ m/s} \quad (29)$$

For a medium sphere, its increment and true velocities are connected by v_i :

$$v_s = \frac{v_{s_{\Delta r}}}{(v_{s_{\Delta r}})_0} v_i \quad (30)$$

Where $v_{s_{\Delta r}} = (v_{s_{\Delta r}})_0$ for all $r < r_c$. A 10^{-9} increment does a pretty good job of connecting (27) with (28) via (30). Larger increments can be used, to about $\frac{\Delta r}{r} = 10^{-4}$, but these are less accurate, since the partition error goes up.

The integral radial velocity v_e of the adiabatic sphere is the sum of its medium shells, plus the small core:

$$v_e = (v_i) \left(\frac{r_c}{r_e} \right) + \sum_{r_c}^{r_e} \left(\frac{r}{r_e} v_s \right) = 0.79210 v_i = K v_i \quad (31)$$

Which tells us how fast the sphere expands while conserving energy. The value of K is constant to the 5th decimal place.⁸ It's independent of any change in M' , ρ , T , or \mathcal{K} .

The kinetic gain of an instant sphere is:

$$E_k = \frac{M(v_s)^2}{2} \quad (32)$$

The integral kinetic gain of an adiabatic sphere, also termed E_k , is expressed with v_i and its thermal mass M_e :

⁷ Pure endpoint values were found from (24) for $H_0 = 67.00\text{--}76.00$; $r_{e_{1090}} = 10^{(18.8169537022-0.9999999H_0)}$; correlation = 1.

⁸ Most calculations used 997 steps of linearly increasing r/r_e , beginning at r_c/r_e and ending with $r/r_e = 0.999999$ or 1 at step 997. Derived K values at a 997-point refinement were invariant to 5 decimal places. A 9970-point plot gave $K = 0.792104$ (9969 steps) and 0.792094 (9970 steps); the value $v_e/v_i = K = 0.79210$ was selected.

$$E_k = \frac{M_e K^2 (v_i)^2}{2} \quad (33)$$

Eq. (33) redefines E_k as its r -integral value at the endpoint, rather than its r -differential value (which is zero). This adiabatic redefinition of E_k properly connects it to the Gibbs meanings (7) and (10) through (11). Kinetic gain E_k only exists instantly, as it is continuously transformed into entropic gain $d(TS)$.

A normalized plot, $(v_s/v_i)^2$ vs. (r/r_e) (Figure 1) shows thermal loss distribution within an instant sphere as a function of its radius. Its r -integral value at $(r/r_e) = 1$ is exactly 2/3 kinetic gain:⁹

$$\frac{2}{3} d(U_i) = E_k \quad (34)$$

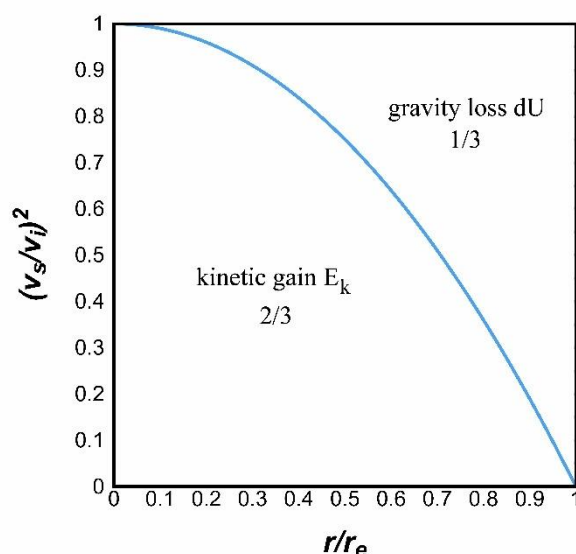


Figure 1. Thermal loss in the adiabatic sphere

Only 1/3 of an adiabatic sphere's thermal loss is taken up by gravity:

$$\frac{1}{3} d(U_i) = -d(U_r) \quad (35)$$

Since the Universe at last scatter was Euclidean, we can construct a line of adiabatic spheres, connected at their tangent points. Anywhere along this line, for any two atoms separated by a distance r , their recession rate v_r is:

$$v_r = K \frac{r}{r_e} v_i \quad (36)$$

Rearrangement gives the fundamental equation:

$$H_g = K \frac{v_i}{r_e} \quad (37)$$

Where $H_g = v_r/r$ is its atomic Hubble parameter. Eq. (37) is the thermal model. Deployment of (37) at varying T from 10K to 100,000K at $z = 1089$, or any other dark z value, gives the same $H_g(z)$ to < 1 ppm every time. The thermal model is zero-order in temperature. Its independence of thermal

⁹ A 9,970-point plot of $y (v_s/v_i)^2$ vs. $x (r/r_e)$ when numerically integrated gave a curve with third-order coefficients $x_0 = -$

0.002999999, $x = 1.0000000$, $x^2 = -1.5 \times 10^{-9}$, and $x^3 = -0.33333333$. Correlation = 1. When cutoff $x = 0.003$ is added to x_0 , $y = 2/3 @ x = 1$.

content is a bit ironic, isn't it? The thermal model is also zero-order in \mathcal{K} . A universe made of xenon atoms (0.131 kg/mole) at the same ρ returns 100.000% of our primordial mix's H_g . The mass density ρ is the only independent variable in the thermal model. This makes possible its direct comparison with Λ CDM.

Although (37) conceals its underlying calculus, the thermal model is simply expressed, and its message is easy to understand: The Universe's entropic pressure perpetually exceeds the attractive force of gravity. Anisotropic mass before last scatter was arguably equally negligible, so by suitable modification of v_i , r_e , and perhaps K , we can more properly understand Universal behavior at scale for its entire post-inflation history. The present paper truncates that history at $z = 1089$, which we further examine.

3.3.5. The Hubble tension

We now compare the thermal and Λ CDM models at the time of last scattering. Since the Hubble constant H_0 is used to get the total density, we must choose from a menu of H_0 estimates. Then we extrapolate today's total density backwards in time with the scale factor to get the total density at last scatter, and use that density to calculate H_{1089} for each model. Most H_0 values derive from distance-ladder measurements of stars. There is another method which relies on the cosmic microwave background [9]. It uses the tiny variances in the CMB's Boltzmann temperature to get an H value at last scatter. This H estimate includes baryon acoustic oscillation, a contribution from sonic effects. The last-scatter H is then extrapolated *forward* in time using Λ CDM to give H_0 . If accurate, this method should give an estimate of H_0 fairly close to the distance ladder estimates. But it doesn't. The distance ladder gives H_0 values around 74 km/sec/Mpc. Last-scatter extrapolation gives an H_0 of around 68 km/sec/Mpc. These values differ too much to be reconciled, and this discrepancy is known as the Hubble tension. There's been a lot of speculation about the Hubble tension's origin [16–26], much of which preserves Λ CDM as a proper method of extrapolation from last scatter to today. This author believes Λ CDM is inaccurate to begin with. The inaccuracy is pronounced at last scatter, as shown in Table 2.

Table 2. The effect of relativistic (CMB) energy on the Hubble parameter.

$z = 0$		$z = 1089$				
H_0	H_0	H_g	H_Λ	H_g/H_0	H_Λ/H_0	H_g/H_Λ
(km/sec/Mpc)	(sec ⁻¹)	(sec ⁻¹)	(sec ⁻¹)			
BOTH GCDM AND Λ CDM CONTAIN CMB ENERGY (Einstein's Universe)						
67.70	2.194×10^{-18}	6.319×10^{-14}	5.045×10^{-14}	28,805	22,995	1.253
74.40	2.411×10^{-18}	6.945×10^{-14}	5.544×10^{-14}	28,805	22,995	1.253
GCDM DOES NOT CONTAIN CMB ENERGY, Λ CDM DOES						
67.70	2.194×10^{-18}	4.784×10^{-14}	5.045×10^{-14}	21,807	22,995	0.948
74.40	2.411×10^{-18}	5.257×10^{-14}	5.544×10^{-14}	21,807	22,995	0.948
NEITHER GCDM NOR Λ CDM CONTAIN CMB ENERGY (Newton's Universe)						
67.70	2.194×10^{-18}	4.784×10^{-14}	4.389×10^{-14}	21,807	20,008	1.090
74.40	2.411×10^{-18}	5.257×10^{-14}	4.824×10^{-14}	21,807	20,008	1.090
NEITHER Λ CDM NOR GCDM CONTAIN CMB ENERGY, AND GCDM'S GAS DENSITY IS REDUCED BY 15.8%: $\rho_g = 0.8418 \rho_b$						
67.70	2.194×10^{-18}	4.389×10^{-14}	4.389×10^{-14}	20,008	20,008	1.000
74.40	2.411×10^{-18}	4.824×10^{-14}	4.824×10^{-14}	20,008	20,008	1.000

All of Table 2's calculations start with H_0 and are extrapolated backward in time to find H_{1089} . The top two rows give the Hubble parameters with all energy included. The rows below that show what happens when we remove energy.

On the far-left column of Table 2 are estimates of H_0 . The upper number in each set of rows, 67.70 km/sec/Mpc, is a CMB-derived value from [9]. The topmost row of the table recalculates its last-scatter H_Λ . The result is $5.045 \times 10^{-14} \text{ sec}^{-1}$ which presumably re-expresses the authors' starting point; I couldn't find it anywhere in their paper. The lower number, 74.40 km/sec/Mpc, is one of many distance-ladder estimates.

The second column converts km/sec/Mpc \rightarrow (sec) $^{-1}$, which gives the proportionate increase in adiabatic sphere radius, per second. The next two columns give the Hubble parameters at last scatter for each model. The final three columns give H/H ratios for the four different input sets.

The ratio H_{1089}/H_0 is constant for each pair of H_0 's. It's always the same within any of the eight pairs shown in the table. Other H_0 's give the same result. The accuracy of this H_{1089}/H_0 ratio depends on the accuracy of the model, and the accuracy of H_{1089} depends on both H_0 and the model. I believe the distance-ladder H_0 and the thermal model, in Einstein's Universe, gives the best result: $H_{1089} = 6.945 \times 10^{-14} \text{ sec}^{-1}$, or 2.14 million km/sec/Mpc.

The far-right column gives the relative values of the models at last scatter. When CMB energy is included in both models, Λ CDM gives a value that is 125% of Λ CDM in a single row. That tension increases to 138% if we compare H_{1089} 's between the top two rows. When we remove CMB energy from Λ CDM, its value drops to 95% within a single row. This is because relativistic mass shrinks the adiabatic sphere, increases its baryon density, and increases H . When this mass is removed from the sphere, it gets larger and its baryon density drops, which reduces H . Density is covariant with entropy in a freely expanding gas, and an adiabatic sphere's increase in density gives an increase in its entropic pressure. Entropic pressure is presently neglected by acoustic oscillation calculations. The below excerpt from Wikipedia is succinct:

"Without the photo-baryon pressure driving the system outwards, the only remaining force on the baryons was gravitational."¹⁰

The above assertion does not account for entropic pressure. Gaseous baryons comprise this pressure. They are the force-carrying particles. Entropic force $d(TS)/dr$ plausibly propels expansion at last scatter.

Acoustic oscillation is only a small part of the CMB picture. When it's excluded from the Planck 2020 calculations [9], H_0 moves to 67.32 km/sec/Mpc, not a very large drop. Maybe inclusion of entropic force bolsters oscillation enough to bring H_{1089} up to 2.14 million km/sec/Mpc, maybe not. I can't comment intelligently about that, or about how H_{1089} is derived from thermal variance in the CMB's Boltzmann curve. It's the ratio H_{1089}/H_0 where the models' predictions differ the most. They give different results at last scatter, and when acoustic oscillation is included, some of the variance credibly arises from entropic neglect. This author believes that all of the variance, and the Hubble tension, can be attributed to Λ CDM's treatment of baryon mass as accreted rather than gaseous. This difference between the models is developed for $z < 1089$ in the next section.

4. VARIANCE BETWEEN THE Λ CDM AND GCDM MODELS AT $z = 1089 \rightarrow 9$

4.1. The thermal model in z'

The thermal model (37) can be expressed with the inverse scale factor z' :

$$H_g = \frac{K}{r_{e1}} \eta' \sqrt{\frac{3RT' (z'_2)^3}{\mathcal{K} z'_1}} \quad (38)$$

¹⁰ https://en.wikipedia.org/wiki/Baryon_acoustic_oscillations

Eq. (38) is developed in Appendix C. The term $T' = 0.0025\text{K}$ is a constant, expressed in Kelvins.¹¹ The term η' expresses relative mass density, has a range of $0.757 \leq \eta' \leq 1$, and is given by (39):

$$\eta' = \frac{\eta'_{z_2}}{\eta'_{z_1}} = \frac{(\Omega_\lambda(z'_2)^4 + \Omega_{(b+c)}(z'_2)^3)(z'_1)^3}{(\Omega_\lambda(z'_1)^4 + \Omega_{(b+c)}(z'_1)^3)(z'_2)^3} \quad (39)$$

The Ω 's are given in Table 1. When $z'_1 = 1090$, we get the *pure thermal* model:

$$H_g = \frac{K}{r_{e1090}} \eta' \sqrt{\frac{3RT'(z')^3}{\mathcal{K} \cdot 1090}} \quad (40)$$

Eq. (40) gives < 0.1 ppm deviance from (37)'s manually calculated H_g 's for all $z = 0 \rightarrow 1089$ (see Supplemental Material). It's exact for any input H_0 and is a perfect expression of gas thermal behavior at scale. We just need to find the pure endpoint from (26) to get thermal H values at lower z from (40).

When we eliminate CMB energy from (40) we get Newton's Universe. The term $\eta' = 1.000$ for all z , giving H'_g :

$$H'_g = \frac{K}{r_{e1090}} \sqrt{\frac{3RT'(z')^3}{\mathcal{K} \cdot 1090}} \quad (41)$$

Here, the "pure endpoint" r_{e1090} is Newtonian. Eq. (41) also gives < 0.1 ppm deviance from (37). Since " r_{e1090} " has increased, $H'_g < H_g$.

We can compare (41) with a modified ΛCDM H'_Λ , where relativistic mass and dark energy have been removed:

$$H'_\Lambda = H_0 \sqrt{\Omega_{(b+c)} a^{-3}} = H_0 \sqrt{\Omega_{(b+c)} (z')^3} \quad (42)$$

This is the Newtonian version of the *Friedmann*¹² equation. It describes the behavior of expanding rocks and CDM in a vacuum, in Newton's Universe, at their comoving critical density. The thermal (41) and Friedmann (42) equations give an invariant $H'_g/H'_\Lambda = 1.09$ for all z . We can bring H'_g/H'_Λ down to 1.00 with an accretion term.

4.2. Mass accretion

Accretion is expressed with a *mass partition* $\rho^* \leq 1$, giving accreted thermal models. The mean gas density ρ_g is divided by the mean baryon density ρ_b :

$$\rho^* = \frac{\rho_g}{\rho_b} \quad (43)$$

The mass partition removes accreted mass, its kinetic energy, and its proportion of CDM mass from (24)'s endpoint calculation. CMB energy, if present, is unaffected. This twinned "Universe" is both unaccreted and has a lower baryon density, so it expands more slowly. We connect it with its denser, mass-partitioned twin through ρ^* :

$$H'_{g^*} = \frac{K}{r_{e1090}} \sqrt{\frac{3RT'(z')^3 \rho^*}{\mathcal{K} \cdot 1090}} \quad (44)$$

The pure endpoint must be used for ρ^* to be properly expressed as a second independent variable. Eq. (44) again gives < 0.1 ppm deviance from its less-dense twin's manually found values (which have larger endpoints). Since we use the pure endpoint, the accreted baryons aren't really missing from (44). They're coasting alongside the expanding gas from which they formed, and exhibit Friedmann behavior.

When $\rho^* = 0.8418$, for the entire domain $z = 0$ to 1089, the thermal and Friedmann equations give identical results:

¹¹ $T' = 0.002500631$. This is the *root temperature* for a last-scatter 2971K. It's derived in appendix C.

¹² Alexander Friedmann (1888-1925).

$$\frac{H'_{g^*}}{H'_\Lambda} = \frac{\frac{K}{r_{e1090}} \sqrt{\frac{3RT'(z')^3 \rho^*}{\mathcal{K} \cdot 1090}}}{H_0 \sqrt{\Omega_{(b+c)}(z')^3}} = 1.000 \text{ for all } z = 0 \text{ to } 1089 \quad (45)$$

In other words, expanding rocks which are forever slowing to an eventual halt behave exactly a freely expanding gas forever pushing itself apart. The only thing the CMB does is increase the push, and that mostly happens way back near last scatter. The “rocks”, of course, are stars. In today’s intergalactic medium, entropic pressure makes its gas expand, separating the tendrils of the cosmic web. About 90% of our Universe’s present volume is occupied by this expanding gas, which contains 84% of all baryon mass. After around $z = 9$ or so, accretion into the cosmic web seems to have stabilized at 16% of the total. This 84/16 proportion probably hasn’t changed much since then; however, mass loss from nuclear fusion remains unaddressed.

Eqs. (41) – (45) deal with Newton’s Universe, where the CMB’s equivalent mass isn’t included. In Einstein’s Universe, they underestimate H_g . Also, CMB mass is unaffected by ρ^* .

The thermal model in Einstein’s Universe, with accretion, is termed H_{g^*} :

$$H_{g^*} = \frac{K}{r_{e1090}} \mathfrak{m}' \sqrt{\frac{3RT'(z')^3 \rho^*}{\mathcal{K} \cdot 1090}} \quad (46)$$

The pure endpoint r_{e1090} re-includes CMB energy. The \mathfrak{m}' term now includes ρ^* :

$$\mathfrak{m}' = \frac{\left(\frac{\Omega_\lambda(z'_2)^4}{\rho^*} + \Omega_{(b+c)}(z'_2)^3 \right) (z'_1)^3}{\left(\Omega_\lambda(z'_1)^4 + \Omega_{(b+c)}(z'_1)^3 \right) (z'_2)^3} \quad (47)$$

Where $z'_1 = 1090$. For $\rho^* = 0.8418$, its \mathfrak{m}' range is $0.757 \leq \mathfrak{m}' \leq 1.046$. We again find that (46) gives < 1ppm deviance from the manually found values (37) for our input H_0 ’s (see Supplemental Material).

4.3. Progress of accretion at cosmic dawn

The term $(1-\rho^*)$ is the *accretion parameter*. It’s the fraction of (baryon + CDM) mass that is gravitationally bound: Stars, planets, black holes, bound gases, etc. At scale, these accreted baryons act like rocks and comprise the cosmic web of galaxies with all of its gravity-bound behavior. How did accretion evolve? We may be able to answer this question through observation of stars at cosmic dawn.

For $z > 9$, “dark” energy is minimal ($\Omega_\Lambda < 0.002$), and ρ^* (43) might be estimable from the H values of any luminous bodies we are fortunate enough to see. The accuracy of ρ^* depends on a proper value of H_0 for which the range of distance-ladder estimates remains large. What we can do is determine if observed H/H_Λ values deviate upwards as we go back in time from $z = 9 \rightarrow 50$ or so. This in turn depends on whether or not star formation and accretion are coevolving phenomena. If accretion ends before starlight begins, then $\rho^* = 0.84$ applies and the value of H/H_Λ will remain close to one. However, if accretion and starlight coevolve, then $H/H_\Lambda > 1$ may be significant enough to measure. For example, the newly found galaxy JADES-GS-z13-0 [35,36] has a spectroscopic redshift of 13.2, so $\Omega_\Lambda = 0.0007$, a minimal dark energy value. If $\rho^* = 0.84$ at this redshift then $H_{g^*}/H_\Lambda = 1.002$, not significant. If $\rho^* = 0.9$ then $H_{g^*}/H_\Lambda = 1.037$ which if true ought to be detectable. The theoretical upper limit of H_{g^*}/H_Λ at $z = 13.2$ is 1.09, shown in Figure 2. These ratios don’t change in z for any H_0 , but if $H_0 = 67.70$ km/sec/Mpc is in fact a low estimate then observed ratios using it should start high ($H/H_\Lambda = 1.139$ for $\rho^* = 0.9$) and get higher as z increases (if the Universe coevolves). This author believes that star formation and accretion do coevolve, and I predict upwardly-trending deviance of H from H_Λ in the $z > 9$ domain now accessible from the James Webb space telescope. If we can get reliable luminosity-based distance estimates from these faint bodies, and can find the proper H_0 , we should be able to follow the progress of accretion through $(1-\rho^*)$.

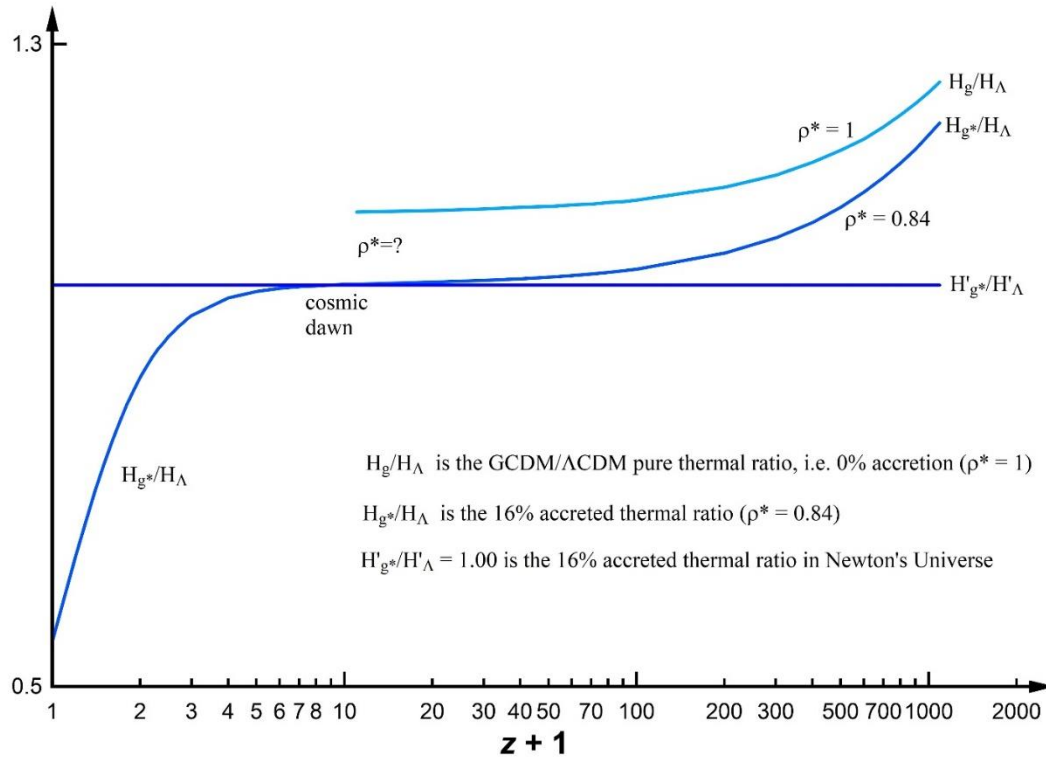


Figure 2. Effects of CMB and dark energy on the accreted thermal model at a proper H_0

Figure 2 gives a complete picture of the thermal model over the entire domain $z = 0$ to 1089. There are two domains of variance, overlapping at cosmic dawn, where both dark and CMB energies aren't large. For $z > 9$, predicted variances H_g^*/H lie somewhere in the area between the pure and accreted curves in Figure 2. For $z < 9$, the variance of H_g^*/H due to added repulsive force is clearly shown and discussed in section 5 below.

Two numbers are worthy of mention. The dark-matter-to-baryon ratio Ω_c/Ω_b is 5.311. The accretion ratio $\rho_g/(1 - \rho_g)$ for $z < 9$ is 5.321. These differ by less than half a percent. Is this just a coincidence, or is there a causal connection?

5. SUPRATHERMAL ENERGY

5.1. The suprathermal model, $z = 9$ to 0

None of the above expressions come any closer to explaining the source of the repulsive “dark energy” term Ω_Λ in the Λ CDM model. I propose that suprathermal kinetic energy in the IGM is responsible. It arises mostly from Compton¹³ scattering, reliant in turn on energetic photon flux E_γ . There are direct partial flux estimates available at the high end [37,38] but the process of connecting these and other sources to produce a definitive suprathermal model is an undertaking of considerable magnitude. The present paper is merely an introduction.

The suprathermal model is developed in Appendix D and expressed by (48):

$$H = H_{g^*} \sqrt{\left(1 + \frac{U_{\beta_s}}{U_b}\right)} \quad (48)$$

Where U_{β_s} is electron suprathermal energy in the adiabatic sphere, U_b is nucleon thermal energy, and U_{β_s}/U_b is the *suprathermal ratio*.

Eq. (48) presupposes thermal plasma, a safe bet for a reionized Universe. Accretion is presumed complete after $z = 9$ so we use $\rho^* = 0.84$. The suprathermal ratio U_{β_s}/U_b is, like H_{g^*} ,

¹³ Arthur Compton (1892-1962).

zero-order in T ,¹⁴ so (48) is completely temperature-independent. We fit the suprathreshold model to the Λ CDM model by convergence of U_{β_s}/U_b around $H/H_\Lambda = 1$ for each datum. These results are shown in Figure 3. A ln-ln line is found¹⁵ for $z = 0 \rightarrow 2$, giving (49):

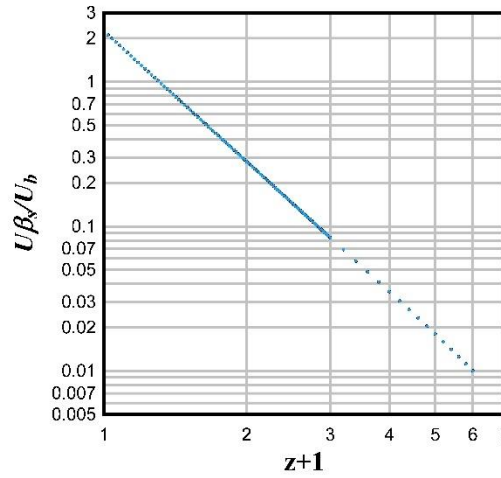


Figure 3. Suprathreshold ratio vs. redshift

$$\frac{U_{\beta_s}}{U_b} = e^{(0.8045 - 2.9915 \ln(z'))} \approx \frac{2.2397}{(z')^3} \quad (49)$$

At $z = 0$, $U_{\beta_s}/U_b = 2.2397$ gives $H/H_\Lambda = 1$; the regressed value is 2.236. This is close to the ratio $\Omega_\Lambda/(\Omega_b + \Omega_c)$ in the Λ CDM model, 2.235, and a simple restatement of the source of “dark energy” repulsion. A modest deviance from linearity in the plot occurs above $z = 2$; these are shown in Figure 3 but not included in the regression. The line is still very good out to $z = 4$. At $z = 5$ there’s deviance and U_{β_s}/U_b drops to 0.01. The crossover to U_{β_s} dominance is found at $z = 0.309$. Eq. (49) derives from Λ CDM and accordingly gives the suprathreshold ratio as proportionate to $(z')^{-3}$.

Inserting (49) into (48) gives:

$$H = H_g^* \sqrt{\left(1 + \frac{2.2397}{(z')^3}\right)} = H_g^* \sqrt{1 + 2.2397 a^3} \quad (50)$$

Using the same data, a ln-ln regression of r_e vs. z' gives (51):¹⁶

$$r_e = r_0 e^{[0.00006 - 1.5006 \ln(z')]} \approx r_0 (z')^{-\frac{3}{2}} \quad (51)$$

Where $r_0 = r_e$ at $z = 0$. From (51) we see that the adiabatic volume V_e/V_0 varies as $(z')^{-\frac{9}{2}}$, not $(z')^{-3}$. A constant “dark energy density” expressed as $[U_{\beta_s}/U_b]/(z')^3$ might be accurate, but within Λ CDM, isn’t proper.

From (48) – (51) we arrive at an expression of H for $z = 0$ to 2:

$$H = H_{g_0}^* (z')^{\frac{3}{2}} \sqrt{1 + \frac{2.2397}{(z')^3}} = H_{g_0}^* a^{-\frac{3}{2}} \sqrt{1 + 2.2397 a^3} \quad (52)$$

When $H_0 = 74.40$ km/sec/Mpc, the thermal $H_{g_0}^* = 1.339 \times 10^{-18}$ sec⁻¹ (at $z = 0$). Eq. (52) gives 100.00% → 99.90% of the Λ CDM value (19) for $z = 0$ to 2. The exponents in (49) – (51) are rounded to the nearest fraction. The rounding excises CMB density from r_e/r_0 (51), giving some of the observed

¹⁴ $T = 4,000-50,000$ K gave the same results for all z .

¹⁵ For the ln-ln regression of $H(z)$ from $z = 0$ to 2, 101 data points were used with 3+ significant figures for all calculated U_{β_s}/U_b . Found: $y = 0.80455 - 2.99154x$; Correlation 0.99999996; std. error 0.0007. The y intercept gives $z = 0.30852$.

¹⁶ Found for ln-ln: $y = 0.00006 - 1.500563x$; correlation 0.99999999. The endpoint is temperature-dependent so a constant T must be used.

-0.1% deviance at $z = 2$. When (50) is properly expressed with (46), the error drops considerably, only reaching 0.1% at $z = 9$. (see Supplemental Material).

From (26), (46), (47), and (50), a full expression of H can be given:

$$H = \frac{KH_0(1090)^3 a^{1.5} \left(\frac{\Omega_\lambda}{\rho^* a^4} + \frac{\Omega_{(b+c)}}{a^3} \right) \sqrt{\frac{3RT^* \rho^*}{1090K} (1 + 2.2397 a^3)}}{(6.5608 \times 10^{18}) (\Omega_\lambda (1090)^4 + \Omega_{(b+c)} (1090)^3)} \quad (53)$$

Which contains two independent variables: $a (=1/[z+1])$, and $\rho^* \leq 1$. For $z < 9$, $\rho^* = 0.84$, and Eq. (53) agrees with Λ CDM (19) ($\pm 0.1\%$). For $z > 9$, ρ^* rises, giving tension with (19). At $z = 1089$, $\rho^* = 1$, and (53)/(19) = 1.25.

At $z = 0$, the models converge:

$$\frac{U_{\beta_s}}{(U_b + U_{\beta_s})} = \Omega_\Lambda = 0.691 \quad (54)$$

If we include thermal free electron energy U_{β_t} in the denominator of (54), we get $\frac{U_{\beta_s}}{(U_b + U_{\beta_s} + U_{\beta_t})} = 0.53$, still more than half of all kinetic energy in the IGM today.

The suprathermal model gives a “pumped Universe” scenario. In a pumped Universe, the intergalactic medium is continually fed with suprathermal energy U_{β_s} . This energy persists and accumulates. Most of it comes from electrons which in turn derive their energy from photons produced by nuclear fusion within the cosmic web of galaxies. Compton scattering imparts kinetic energy to the electrons. Much of that energy is well above the nucleons’ thermal range given by the Boltzmann curve, so the Universe isn’t just getting hotter. Its kinetic energy is increasing above and beyond simple thermal heat. This adds a tremendous amount of entropic force $d(TS)/dr$.

We should be able to connect suprathermal energy generation with the small number of known sources which are capable of producing E_γ photons: Type “O” stars and active galactic nuclei. Maybe additional scattering of free electrons by the lower-energy photons emitted by all stars today gives net addition enough to account for the “dark energy” effect. If U_{β_s} persists indefinitely, then $d(U_{\beta_s})/dt$ can be estimated from stellar photon flux and lookback. Those calculations involve several other assumptions not covered here and lie outside the scope of the present paper.

5.2. Origin of suprathermal electrons in the IGM

By $z \approx 6$ the Universe was fully reionized, so Compton scattering from neutral IGM atoms isn’t a significant source of suprathermal energy after that. A second source is Compton scattering of free electrons in the IGM, which would taper off to a steady state if the energy profiles of the electrons and photons were to converge. Cosmic radiation U_{b_s} from the web is a third source. The reader is asked to consider a fourth source, kinetic energy of electron escape from the web, as a present-day contributor to the IGM’s suprathermal content. Any resulting IGM charge buildup is presently untreated. This author estimates these electrostatic effects as minor compared to U_{β_s} .

5.3. Suprathermal effects on r_e and K

Suprathermal electrons do not behave like thermal baryons, and their kinetic gain hasn’t been properly addressed. Thermal free electrons are treated as a gas (appendix D) which seems proper, but I’m not so sure if it’s accurate. This author is unaware of any treatments involving electron kinetic gain. Electron thermal energy may partly comprise U_{β_s} . What we can say is that electron kinetic energy in general only minimally affects nucleon endpoint values. The presumption that $v_e/v_i = K$ remains constant with added suprathermal energy is conjecture, but it appears to work well. For special effects, any relativistic mass increase would cause a decrease in r_e . The entropic partition (2/3) and K may also change with introduction of highly relativistic kinetic energy. I believe such changes in K and r_e aren’t large. Even if K/r_e is theoretically shown to change by as much as 1%, the suprathermal model still allows us to use known and conserved energy sources, in compliance with (1), to account for H .

6. DISCUSSION: GCDM VERSUS Λ CDM

The Λ CDM model (19) is a benchmark, giving the most accurate empirical fit to date. It closely converges with the GCDM model at $z = 0$. The models have different theoretical foundations and their predictions diverge at last scatter.

The Λ CDM model combines three formulas:

- 1) The Friedmann equation gives a relation between H and total energy density $\epsilon_{M'}$.
- 2) The fluid equation adds a (kinetic) mass equivalence term " P " and describes covariant ($\epsilon_{M'} + "P"$) vs. H .
- 3) The equation of state divides " P " into three different constituents.

We use $H(t)$ below. It's connected to $H(a)$ by lookback $\int (dt/da)da$.

6.1. The Friedmann equation

The question of Universal curvature is historically important and extensively treated in modern introductory texts [39,40]. That debate is largely settled now. Most of us believe in a flat Universe, so the Friedmann equation can be simply expressed:

$$H = \sqrt{\frac{8\pi G \epsilon_{M'}}{3c^2}} \quad (55)$$

The Friedmann equation overlooks the fact that most Universal baryon mass is gaseous. Instead of the thermal model's perpetual entropic pressure, we have Λ CDM's rocks at a "miraculous" critical density. While the Friedmann equation (43) may not properly express entropic gain, it can be shown as identical to thermal behavior (45) through ρ^* .

Friedmann employed Einstein's field equation which makes no provision for entropic gain. Einstein invented the Λ force to offset gravity, as the Universe was thought to be static at the time, and he had to do something to prevent its collapse. Both Friedmann's and Einstein's entropic omissions were benign. It just didn't occur to either author to include entropy. Einstein's later doubts about Λ are well known [41].

6.2. The fluid and acceleration equations

The fluid equation is a different story. It actively excludes entropic gain. We start, as do the texts, with the bound Gibbs equation (5) which describes e.g. gas in a vessel:

$$dE = TdS - PdV \quad (5)$$

Over time, this is:

$$\frac{dE}{dt} = T \frac{dS}{dt} - P \frac{dV}{dt} \quad (56)$$

When applied to the unbound Universe, $dE/dt = 0$. We include rest mass equivalence $E = Mc^2$. Nuclear fusion. The unbound Gibbs equation (10) is now:

$$dE = d(Mc^2) + d(U_i) + d(PV) - d(TS) = 0 \quad (57)$$

At last scatter there was no fusion, so M was unchanged, $d(Mc^2) = 0$, and (57) = (10). Photon energy from the CMB isn't included here and is discussed in more detail below (section 6.3.2). It's unimportant for now.

The fluid equation's development continues with heat flow:

$$\frac{dQ}{dt} = T \frac{dS}{dt} \quad (58)$$

This relates thermal change dQ with the entropy change dS in a vessel. If there's no vessel, we get the Universe. We will assume there's no heat flow in or out of the Universe. This assumption gives rise to a philosophical discussion about the meaning of the first law vs. unbound thermal loss. We will sidestep that entire debate and can properly treat lost kinetic energy as entropic gain.

The second law (2) is clear about entropy: $dS/dt > 0$ always. At scale, we can get a rough estimate of gas dS/dt from the thermal model (37) at a constant T and P by setting $dE/dt = 0$ in (56). For $H_0 = 74.40$ km/sec/Mpc and $z = 1089$, this gives:

$$\left(\frac{\partial S}{\partial t}\right)_{T,P} = \frac{4\pi r_e^2 PK v_i}{TM_e} = 106 \text{ (J/K)/sec/kg} \quad (59)$$

The fluid equation sets Universal $dS/dt = 0$ despite the second law. Baryon mass is treated as a pressureless perfect fluid, called a “dust”, which is inconsistent with its actual existence as a gas having entropic pressure. Entropic gain is eliminated, and all energy change is isoentropic. Eq. (56) now looks like this:

$$PdV/dt = -dE/dt \quad (60)$$

Which describes gas in an adiabatic vessel. The term $-dE/dt = -d(U_i)/dt$ describes its rate of thermal loss. Setting $dS = 0$ means that all the loss is reversibly stored, and in the unbound Universe, there’s only one place this author knows of to store it: work against gravity. I’ve shown that only a third of this thermal loss is stored that way. The remaining two-thirds vanishes, and requires entropic power $d(TS)/dt$ to properly describe its fate. Entropic power gives over time an irreversible loss of thermal energy, and a reduction in density which acts as a thermodynamic repository or “sink”. If entropy is excised from the Gibbs equation, the lost thermal energy cannot be accounted for, giving inconsistency with (1).

There is another significant issue with the fluid equation which only became relevant after 1998. Internal energy change dE gets improperly redefined in the fluid equation’s development. It isn’t thermal anymore, it’s total, and now includes rest mass $dE = d(Mc^2)$. Eq. (60) is then used to treat the energy change of rest mass $d(Mc^2)$ as a thermal variable. Rest mass does not change thermally, it stays constant.¹⁷ From (1), dE/dt in (60) should be zero. Conflation of the dE terms can thus appear to create an enormous amount of what is fictitious repulsive energy from what is actually a much smaller amount of suprathermal energy. In this author’s opinion, that’s what happened when the distance ladder line was found to be a curve by the seminal observations of the two research groups who first published their findings in 1998 [4],[6]. They accounted for the distance-ladder curvature by populating Λ with fictitious negative mass. The reader may wish to consider the alternative, suprathermal energy, and the possibility that the source of Λ cannot be identified because it doesn’t exist.

Further development of the fluid equation is in the texts. The result is the same for both Einsteinian and Newtonian versions, which differ only by c^2 . The Newtonian expression of the fluid equation is:

$$\frac{d(\rho_{M'})}{dt} + 3H(\rho_{M'} + P/c^2) = 0 \quad (61)$$

In (61), $\rho_{M'}$ includes ϵ_Λ/c^2 . The other terms in P/c^2 are equivalent mass density of energy not at rest, so if we exclude Λ , (61) is an attractive term. The energy density term P is thus also attractive and labelled as “pressure”:

$$P = w_b \epsilon_b + w_\Lambda \epsilon_\Lambda + w_\Lambda \epsilon_\Lambda \quad (62)$$

The w terms are dimensionless numbers whose values are expressed using a : $w_b \ll 1$, $w_\Lambda = 1/3$, and $w_\Lambda = -1$. This definition of “pressure P ” is known as the equation of state. It’s the third leg of Λ CDM and is discussed below. The equation of state inverts the meaning of pressure. Most of us think of positive pressure as repulsive, like in a balloon. Not here. Repulsive energy density $w_\Lambda \epsilon_\Lambda$ in (62) is called “negative pressure”.

When we talk about the meaning of pressure, the Jeans model of star formation [30] is relevant. Both (62) and the Jeans model operate concurrently within any given volume. The Jeans model treats P as repulsive, an offset against gravitational collapse. The fluid and state equations treat P as attractive, which is inconsistent with the Jeans model. The Λ CDM model treats P as repulsive, which is consistent with the Jeans model’s treatment of P .

6.3. The equation of state

¹⁷ Mass loss from nuclear fusion is neglected here, but its added kinetic energy isn’t enough to account for Λ .

The equation of state (62) is combined with the Friedmann (55) and fluid (61) equations to complete the Λ CDM model (19).

6.3.1. Baryonic mass

Baryonic matter comprises stars, rocks, and helium balloons, and is treated by Λ CDM as 100% accreted and therefore 100% attractive. In GCDM, baryon mass is mostly repulsive, a gas. At last scatter, baryon mass was 100% gas. Presently, the repulsive/attractive ratio of baryon mass in the Universe is about 84:16, or 5¼:1.

The Λ CDM term $w_b \epsilon_b$ is expressed as:

$$w_b \epsilon_b \approx \left(\frac{kT}{\mu c^2} \right) \epsilon_b = \left(\frac{kT}{\mu c^2} \right) (\rho_b c^2) = \frac{kT \rho_b}{\mu} \quad (63)$$

Where μ is the mean atomic mass and k is Boltzmann's constant. Eq. (63) is simply thermal pressure. Its equivalent mass density, and suprathermal pressure, are both negligible compared to Λ CDM's ϵ_{crit} , whose conflated term Ω_Λ and rest term $\Omega_{(b+c)}$ together comprise almost 100% of ϵ_{crit} today. In GCDM, there is no Ω_Λ , and $\Omega_{(b+c)} \approx 0.9997$. Although only an infinitesimal fraction of total energy today, entropic pressure $P_{V'}$ is ample, and provides both thermal and suprathermal repulsive force.

6.3.2. Relativistic mass; entropy of a photon

Relativistic mass, expressed as $w_\lambda \epsilon_\lambda$ in the Λ CDM model, is attractive in both models and arises from photon and neutrino energy.¹⁸ Its effect on the Hubble parameter at last scatter gives rise to the Hubble tension.

We now examine photon energy more closely. An expanding sphere of CMB light has an r^4 dependence of energy density. Volume increases as r^3 , so there appears to be a $1/r$ loss of CMB energy upon expansion. During the dark age there was minimal CMB energy transfer to the IGM's baryons [41]. Most of the CMB's energy vanished; we get inconsistency with (1). I see no escape from this conundrum except to apply (3): CMB light yields gain through wavelength stretch $\Delta\lambda$. Any one CMB photon's wavelength increases with time and their combined lost energy becomes entropic gain ΔE_{S_λ} :

$$\Delta E_{S_\lambda} = E_{CMB_1} - E_{CMB_2} = \sum_{\lambda_1 \approx 0}^{\infty} n_\lambda h c \left(\frac{1}{\lambda_1} - \frac{1}{\lambda_2} \right) \quad (64)$$

where E_{CMB_1} and E_{CMB_2} are the before and after photon energies during the dark age, h is Planck's¹⁹ constant, λ_1 and λ_2 are the before and after wavelengths of the stretched photon, and n_λ is the number of photons at a wavelength λ_1 . The distribution n_λ vs. λ is observed in the CMB as a blackbody curve, which gives n_λ vs. λ at earlier times.

The above analysis gives an individual photon's entropy S_λ as equal to Planck's constant:

$$S_\lambda = h \quad (65)$$

Entropy is expressed as J/Hz rather than the more conventional J/K.²⁰ Unbound photon energy E_λ is potentially 100% entropic, in that all of it is eventually lost to time. From (65), photons are isentropic, so the second law as applied to them is expressed with gain $d(Es)$ rather than simple entropy increase dS . Gain $d(Es)$ is connected to E_k , hence volume. The rate of volume increase dV_λ/dt of radial light in a model sphere is:

$$\frac{dV_\lambda}{dt} = \frac{4\pi c^3}{3} \quad (66)$$

which far outpaces nonrelativistic E_k .

¹⁸ Neutrinos are believed to have had relativistic kinetic energy at last scatter but became nonrelativistic in the dark age.

This affects their temporal mass density dependence, which is untreated in the present paper.

¹⁹ Max Planck (1858-1947).

²⁰ An alternate treatment of photon entropy using J/K instead of J/Hz is given by Kirwan [43].

Current treatment of CMB energy also begins isentropically, again from the bound Gibbs equation (5) [39,40]. While the present paper concurs with isentropic photon treatment, photon gain ΔE_{S_λ} is neglected, and light energy is purported to expand more slowly than baryonic matter. A different result might be found if ΔE_{S_λ} is included in an *ab initio* derivation.

Cosmic free electron gain resembles that of photons, as they both have wavelike character. Electron kinetic gain has yet to be properly expressed. Before last scatter, when $z > 1089$, free electrons coupled with photons. The coupling gives an increase in H . The present author believes that the thermal model (37) can be applied to give H over the entire domain $a \rightarrow 0$. This interesting subject lies beyond the scope of the present paper.

6.3.3. Dark energy

The remaining term, $w_A \epsilon_A$, describes repulsion. In the Λ CDM model, Λ is used to account for distance-ladder curvature, e.g. [4–7]. Its predominance, $\epsilon_A = 0.69\epsilon_{crit}$ at $z = 0$, arises from the fluid equation's isentropic and variable total energy. The behavior described by $w_A \epsilon_A$ is herein proposed to arise instead from suprathermal pressure $P_{V_s'}$ in the IGM, mostly carried by electrons. Electron and nucleon pressure does create a repulsive *in toto* scalar field, but its Ω_{U_i} is more than ten orders of magnitude smaller than Ω_A , is variable, and can be locally nonscalar. A noncovariant Λ means a constant ϵ_A . This creates more and more energy as the Universe expands, which is inconsistent with (1). If (1) is obeyed, the field must have a conserved source. Suprathermal energy U_{i_s} meets this requirement.

7. CONCLUDING REMARKS

The present paper proposes a fundamental change in the way the Universe is viewed: As an unbound thermodynamic system in which a freely expanding gas has partly accreted into stars. The gas comprises a repulsive field, IGM kinetic energy, scalar in xyz at last scatter. It's now locally variable in xyz but still behaves *in toto* as a time-variant scalar in the flat Universe we see. Thermal behavior is shown to be identical to the Friedmann equation's predictions. Suprathermal behavior causes "dark energy" Λ . The Λ CDM model predicts variance from Λ CDM at cosmic dawn, and the Hubble tension at the time of last scattering.

This author has herein made the case that Λ is an artifact arising from inconsistency with the thermodynamic laws, and that kinetic density, rather than total density, is the primary metric through which H is properly expressed. While not the first to suggest that entropy is a force of Nature, I hope that the present paper will be persuasive enough to convince the reader.²¹

Supplemental Material: An .XLSX workbook containing the model and its output is available.

APPENDIX A: Entropic development at the atomic level

We connect atomic movement to kinetic gain on a small scale, where work against gravity $d(PV)$ can be neglected.

A1.1 Bound, equilibrium free expansion

Take a spherical helium balloon, of radius $r_1 = 10$ cm, at a temperature $T = 300$ K and pressure $P = 1$ atmosphere, and place it in the center of a perfectly rigid, insulated, spherical vacuum chamber of radius $r_2 = 50$ cm. The insulation and rigidity of the chamber means any gas expansion from r_1 to r_2 will be adiabatic. The gas's internal kinetic energy U_i is 100% thermal. Helium is monatomic, so:

$$U_i = \frac{3}{2}PV = \frac{3}{2}nRT = \frac{3MRT}{2\mathcal{K}} \quad (\text{A1})$$

²¹ This paper is deposited with Physical Review D; code number DN13643.

Where n is the number of moles of gas, \mathcal{K} is the gas's atomic weight, and R is the gas constant.

The U_i in the chamber is the instant sum of its atoms' individual kinetic energies:

$$U_i = \sum_{r=0}^{r_2} \sum_{\theta=0}^{\pi} \sum_{\varphi=0}^{2\pi} \left\{ \frac{1}{2} m [(v \sin[\theta'])^2 + (v \cos[\theta'])^2] \right\} \quad (A2)$$

Where the tensor v is the atom's kinetic energy, m is its atomic rest mass, r is its distance from the center, θ is its conic angle of latitude, φ is its angle of longitude, and θ' is the conic angle of v 's deviance from radial. These are shown in two dimensions in Figure A1. The balloon is an *idle* sphere with a constant r_1 . The void between r_1 and r_2 makes no contribution to U_i as long as the balloon is intact.

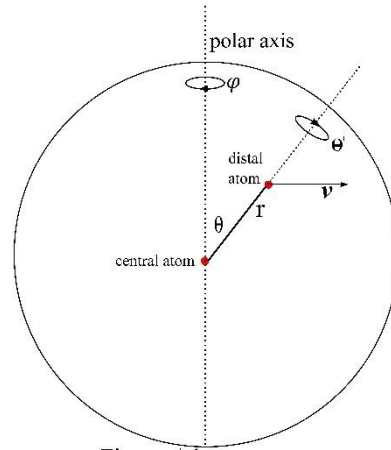


Figure A1.

We pop the balloon. The thermal energy U_i is temporarily and partly transformed into (radial) kinetic gain E_k :

$$E_k = \sum_{r=0}^{r_2} \sum_{\theta=0}^{\pi} \sum_{\varphi=0}^{2\pi} \sum_{\theta'=0}^{\pi/2} \left\{ \frac{1}{2} m [(v \cos[\theta'])^2] \right\} - \sum_{r=0}^{r_2} \sum_{\theta=0}^{\pi} \sum_{\varphi=0}^{2\pi} \sum_{\theta'=\pi/2}^{\pi} \left\{ \frac{1}{2} m [(v \cos[\theta'])^2] \right\} \quad (A3)$$

Which is the scalar difference between the outward and inward radial components of the atoms' tensors. For an idle sphere, the inward and outward radials of v in (A3) are equal. We can double (A3)'s inward radial and replace the radial expression in (A2), giving U'_i :

$$U'_i = \sum_{r=0}^{r_2} \sum_{\theta=0}^{\pi} \sum_{\varphi=0}^{2\pi} \frac{1}{2} m [(v \sin[\theta'])^2] + 2 \sum_{r=0}^{r_2} \sum_{\theta=0}^{\pi} \sum_{\varphi=0}^{2\pi} \sum_{\theta'=\pi/2}^{\pi} \left\{ \frac{1}{2} m [(v \cos[\theta'])^2] \right\} \quad (A4)$$

Combining (A2), (A3) and (A4) gives:

$$U_i = U'_i + E_k \quad (A5)$$

When idle, $U_i = U'_i$. When expanding, $U_i > U'_i$.

Since loss to gravity is negligible, $E_k = (U_i - U'_i)$ for the r_2 sphere throughout expansion, which lasts for a second or two. Initially U'_i drops and E_k rises. The atoms quickly bounce off the wall and E_k drops back down. At reequilibrium, $E_k \rightarrow 0$, the terms of (A3) again cancel, and U_i is restored: " $U_i \rightarrow E_k \rightarrow U'_i$ ". The gain ΔE_s from volume increase is:

$$\Delta E_s = T(S_2 - S_1) = nRT \ln \left(\frac{V_2}{V_1} \right) \quad (A6)$$

A1.2 Bound, nonequilibrium free expansion

Take that same balloon, put it in the center of a large vacuum chamber ($r_2 = 10^8$ m) and pop it. A helium atom at $T = 300$ K has a root mean square speed $v_{rms} = 1368$ m/s. Those atoms will take about 20 hours to reach the wall if their tensor of movement is perfectly radial. As they expand, they stop colliding with each other at any meaningful rate. After that happens, atomic movement is in an unbound regime which is best considered when the atoms have stopped colliding, but haven't hit the wall yet. In this regime, the atomic Hubble parameter H_g is simple:

$$H_g = v_r / r = 1/t \quad (A7)$$

Once the atoms stop colliding, $E_k \gg U_i'$ at 300K and $U_i \rightarrow E_k$ almost entirely. Eventually the atoms bounce off the wall, $E_k \rightarrow U_i$, and the regime slowly ends.

A1.3 Unbound, nonequilibrium free expansion

What if there's no boundary? There's no equilibrium to be reached, so for a freely expanding gas, $U_i \rightarrow E_k$ only and there's no $E_k \rightarrow U_i$. One can approach Universal conditions at the time of last scattering by looking only at the comoving core of a popped sphere with initial $r_{core} = 10^{-6} r_{sphere}$. The atoms in that core would be slow, cold, and nearly isodense ($dp/dr \approx 0$) with a uniform U_i which obeys (11). This sort of treatment could be accurate at scales large enough to include gravity. I'm not suggesting that the Universe has finite mass, only that it can be so modeled.

APPENDIX B: Initial radial velocity in the adiabatic sphere

We start with the increment radial velocity $v_{s\Delta r}$:

$$v_{s\Delta r} = \sqrt{\frac{2E_{k\Delta r}}{M}} \quad (27)$$

We follow $v_{s\Delta r}$ as a function of r , using a fixed increment $\frac{\Delta r_i}{r} = 10^{-9}$. Below the cutoff radius $r_c = 0.003 r_e$, loss to gravity is negligible, and all these small spheres have the same $E_{k\Delta r}/M$ value to within 5 ppm (Figure B1):

$$\frac{E_{k\Delta r}}{M} = \frac{dE}{dM} = \frac{dV}{dM} \frac{dE}{dV} = \left(\frac{RT}{\mathcal{K}P}\right) \frac{dE}{dV} = \frac{RT}{\mathcal{K}} \left(\frac{dE}{PdV}\right) = \frac{RT}{\mathcal{K}} \quad (B1)$$

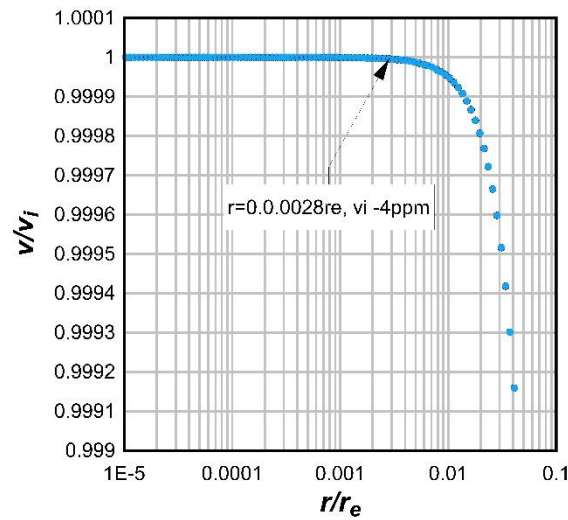


Figure B1. small sphere cutoff

Here we used $dE = PdV$. Combining (27) and (B1) gives an initial radial velocity v_i :

$$v_i = \sqrt{\frac{2RT}{\mathcal{K}}} \quad (B2)$$

We can determine if energy is conserved within (B2) by examining the partition error (B3):

$$\frac{\frac{1}{2}M \left[(v_{i(T_1)})^2 - (v_{i(T_2)})^2 \right] - \Delta E}{\Delta E} \quad (B3)$$

We increment a small sphere ($r_1 = 1 \times 10^{12}$ m) giving P_2 and T_2 . The pressure drop from thermal loss of our gas is given by:

$$P_2 = P_1 \left(\frac{v_2}{v_1} \right)^{-\frac{5}{3}} \quad (B4)$$

And the temperature drop by:

$$T_2 = T_1 \left(\frac{P_2}{P_1} \right)^{2/5} \quad (B5)$$

Which for (B2) gives a large error. Development in (B1) uses the ideal gas law and not thermal energy. We rearrange (A1):

$$\frac{U_i}{M} = \frac{3RT}{2\mathcal{K}} \quad (\text{B6})$$

Substituting $\frac{U_i}{M}$ for $\frac{E_k}{M}$ in (B2) gives (29):

$$v_i = \sqrt{\frac{2U_i}{M}} = \sqrt{\frac{3RT}{\mathcal{K}}} = \sqrt{\frac{3(8.3145)(2971)}{(0.00123988)}} = 7731 \text{ m/s} \quad (29)$$

By use of (29) the partition error (B3) is only 2×10^{-8} .

APPENDIX C: Expression of the thermal model in z'

The redshift z_2 can be obtained from any starting value z_1 by enlarging the increment $\frac{\Delta r_i}{r}$:

$$z_2 = \frac{z_1+1}{\frac{\Delta r_{i+1}}{r}+1} - 1 \quad (\text{C1})$$

The temperature is found from (B4) and (B5) using (C1) to get V_2 . Given $T_{(z'=1090)} = 2971\text{K}$, we find from the spreadsheet that $T_{z'} = T_{(z+1)}$ is exactly:²²

$$T_{z'} = T'(z')^2 \quad (\text{C2})$$

Where the *root temperature* $T' = 0.002500631$ is expressed as degrees K. If treated as an abrupt decoupling at $z = 1089$, then at $z = 60$, $T = 9.3\text{K}$; at $z = 10$, $T = 0.3\text{K}$. These are lower than other estimates [42] but since H_g is temperature-independent, dark thermal heating by the CMB doesn't affect H . We just need the root temperature.

We insert (C2) into (29), giving:

$$v_i = \sqrt{\frac{3RT'z'^2}{\mathcal{K}}} \quad (\text{C3})$$

The endpoint r_e is adjusted for both T and ρ .

For the T adjustment, the thermal radius change r_{e2}/r_{e1} vs. T at constant ρ and \mathcal{K} is found exactly:

$$r_{e2} = r_{e1} \sqrt{\frac{T_2}{T_1}} \quad (\text{C4})$$

For the ρ adjustment, the thermal radius change r_{e2}/r_{e1} vs. ρ at constant T and \mathcal{K} is found exactly:

$$r_{e2} = r_{e1} \sqrt{\frac{\rho_1}{\rho_2}} \quad (\text{C5})$$

For nonrelativistic mass,

$$\frac{\rho_1}{\rho_2} = \frac{(z'_1)^3}{(z'_2)^3} \quad (\text{C6})$$

Combining these gives:

$$r_{e2} = r_{e1} \sqrt{\frac{z'_1}{z'_2}} \quad (\text{C7})$$

Inserting (C3) and (C7) into the thermal model (37) gives Newton's thermal model (C8):

$$H_g = K \frac{v_{i2}}{r_{e2}} = \frac{K}{r_{e2}} \sqrt{\frac{3RT'(z'_2)^2}{\mathcal{K}}} = \frac{K}{r_{e1}} \sqrt{\frac{3RT'(z'_2)^3}{\mathcal{K} z'_1}} \quad (\text{C8})$$

We adjust for relativistic mass with η' (39):

²² 210 points from $z' = 1090$ to 10; median $z' = 350$. Found: $T = 2 \times 10^{-9} + 5 \times 10^{-11}(z') + 0.002500631 (z')^2$; correlation = 1.

$$\eta' = \frac{\eta_{z'_2}}{\eta_{z'_1}} = \frac{(\Omega_\lambda(z'_2)^4 + \Omega_{(b+c)}(z'_2)^3)(z'_1)^3}{(\Omega_\lambda(z'_1)^4 + \Omega_{(b+c)}(z'_1)^3)(z'_2)^3} \quad (39)$$

A linear dependence of H on η' is found from the spreadsheet, giving the thermal model in Einstein's Universe (38):

$$H_g = \frac{K}{r_{e1}} \eta' \sqrt{\frac{3RT'(z'_2)^3}{\mathcal{K} z'_1}} \quad (38)$$

When $z'_1 = 1090$, the pure endpoint r_{e1090} applies, and we get the pure thermal model (40).

APPENDIX D: Expression of the suprathreshold model in z'

Suprathreshold kinetic energy adds entropic pressure to the IGM:

$$P_{V'} = P_{V'_t} + P_{V'_s} \quad (D1)$$

Where $P_{V'_t}$ and $P_{V'_s}$ are thermal and suprathreshold pressures respectively.

The thermal model (37) has three terms: K , v_i , and r_e . If we want to express suprathreshold content within that framework, we need to increase v_i or K , decrease r_e , or some combination. We express v_i (29) as a sum:

$$v_i = \sqrt{\frac{2(U_{i_t} + U_{i_s})}{M}} \quad (D2)$$

where U_{i_t} and U_{i_s} are thermal and suprathreshold kinetic energies in the adiabatic sphere.

The total nucleon kinetic energy U_b is:

$$U_b = U_{b_t} + U_{b_s} \quad (D3)$$

Where U_{b_t} is the thermal value of U_b , and U_{b_s} is cosmic radiation.

The total electron kinetic energy U_β is:

$$U_\beta = U_{\beta_b} + U_{\beta_t} + U_{\beta_s} \quad (D4)$$

Where U_{β_b} is the thermal energy of atomically bound electrons, U_{β_t} is the thermal energy of free electrons, and U_{β_s} is their suprathreshold energy.

Inserting (D3) and (D4) into (D2) gives:

$$v_i = \sqrt{\frac{2[(U_{b_t} + U_{\beta_b} + U_{\beta_t}) + (U_{b_s} + U_{\beta_s})]}{M}} \quad (D5)$$

We neglect cosmic radiation U_{b_s} for now as its omission doesn't substantially affect the logic of the following expressions. This means $U_{b_t} = U_b$, so:

$$v_i = \sqrt{\frac{2(U_b + U_{\beta_b} + U_{\beta_t} + U_{\beta_s})}{M}} \quad (D6)$$

We examine the thermal energies U_{β_b} (bound) and U_{β_t} (free). Thermal free electrons are held to behave at very low densities as a monatomic gas. Treatment as such reduces the mean atomic weight \mathcal{K} . The thermal model is independent of both \mathcal{K} and T and dependent only on the mean gas density ρ_g . The result of thermal ionization is thus an increase in both v_i and r_e without affecting H or K . If v_i is doubled, so is r_e , as is the case with pure hydrogen plasma which will serve as our example.

In a thermal system with no ionized H₁:

$$U_i = U_b + U_{\beta_b} \approx 1.0005 U_b \quad (D7)$$

so $U_i = U_b$ is reasonably accurate. When H₁ is 100% ionized at e.g. 4000K, the number of gas particles is doubled, the atomic weight halved, and energy equipartitioned: $U_{\beta_b} = 0$, $U_{\beta_t} = U_b$, and $\mathcal{K}' = \frac{1}{2}\mathcal{K}$, where \mathcal{K}' is the mean atomic weight of the plasma. Making these substitutions into (D6) with no U_{β_s} gives:

$$v_i = \sqrt{\frac{2(U_b + U_{\beta_t})}{M}} = \sqrt{\frac{2(2U_b)}{M}} \approx \sqrt{\frac{4U_i}{M}} = \sqrt{\frac{6RT}{\mathcal{K}'}} = \sqrt{\frac{6RT}{\mathcal{K}/2}} = \sqrt{\frac{12RT}{\mathcal{K}}} = 2\sqrt{\frac{3RT}{\mathcal{K}}} \quad (\text{D8})$$

The added $U_{\beta_t} = U_b$ gives twice the old value of v_i from (29); more generally, added U_{β_t} gives a linear increase in v_i and we can expect the same for r_e . This means that for thermal plasmas, the thermal model (37) is more properly expressed using the nucleon kinetic energy U_b alone:

$$H_g = K \frac{v_i \left(\frac{U_b + U_{\beta_t}}{U_b} \right)}{r_e \left(\frac{U_b + U_{\beta_t}}{U_b} \right)} = K \frac{\sqrt{\frac{2U_b}{M}}}{r_e} \quad (\text{D9})$$

Where v_i and r_e have their non-ionized values. The denominator term associated with r_e in (D9) is inserted to comply with the thermal model's zero-order dependencies. In (D9), v_i remains close to (29): $U_b \approx 0.9995U_i$, so the effect of thermal ionization on H is minimal, and we can exclude U_{β_t} from the thermal model.

We proceed by assuming that suprathermal energy U_{β_s} has no effect on either K or r_e . It may have some effect but we will say it doesn't. Removal of U_{β_b} and U_{β_t} from (D6) gives:

$$v_{i(b+\beta_s)} = \sqrt{\frac{2(U_b + U_{\beta_s})}{M}} = v_i \sqrt{\left(1 + \frac{U_{\beta_s}}{U_b}\right)} \quad (\text{D10})$$

When accretion $\rho^* = 0.84$ is included we arrive at the suprathermal model (48):

$$H = H_{g^*} \sqrt{\left(1 + \frac{U_{\beta_s}}{U_b}\right)} \quad (\text{48})$$

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