# Satellite and Ground Remote Sensing Techniques to Trace the Hidden Growth of a Lava Flow Field: The 2014-15 Effusive Eruption at Fogo Volcano (Cape Verde)

Sonia Calvari<sup>1\*</sup>, Gaetana Ganci<sup>1</sup>, Sónia Silva Victória<sup>2</sup>, Pedro Hernandez<sup>3,4</sup>, Nemesio Perez<sup>3,4</sup>, Vera Alfama<sup>2</sup>, José Barrancos<sup>3</sup>, Jeremias Cabral<sup>5</sup>, Nadir Cardoso<sup>2</sup>, Samara Dionis<sup>3</sup>, Paulo Fernandes<sup>2</sup>, Gladys Melian<sup>3</sup>, José Pereira<sup>2</sup>, Hélio Semedo<sup>5</sup>, Germán Padilla<sup>3,4</sup>, Fátima Rodriguez<sup>4</sup>

<sup>1</sup>Istituto Nazionale di Geofisica e Vulcanologia, Osservatorio Etneo (INGV-OE), Catania, 95125, Italy, 
<sup>2</sup>University of Cabo Verde (UNICV), CP 279 Praia, Santiago, Cape Verde, 
<sup>3</sup>Instituto Tecnológico de Energias Renovábles de Tenerife (ITER), 38611 Granadilla de Abona, Tenerife, 
Spain.

<sup>4</sup>Instituto Volcanológico de Canarias (INVOLCAN), 38400 Puerto de la Cruz, Tenerife, Spain, <sup>5</sup>National Protection Service of Cabo Verde, Praia, Santiago, Cape Verde

\*Corresponding author:

Sonia Calvari

e-mail: sonia.calvari@ingv.it

tel.: +39 095 7165862

ORCID ID: 0000-0001-8189-5499

# **Key Points:**

- Remote sensing techniques allowed us to trace the time averaged effusion rate (TADR) and magma supply rate (SR) during the Fogo 2014-15 eruption
- The lava flow field was fed by a ~5.9 km long lava tube that formed in ~10 days
- The magma intruded within the lava flow field and causing its inflation is estimated at ~31% of the total erupted volume

**Keywords:** Remote sensing monitoring, Fogo volcano, effusive eruption, lava flow inflation, lava tubes, time averaged effusion rate (TADR), magma supply rate (SR)

Running Title: Lava Flows and Lava Tubes at Fogo Volcano

Satellite and Ground Remote Sensing Techniques to Trace The Hidden Growth of a Lava Flow Field: The 2014-15 Effusive Eruption at Fogo Volcano (Cape Verde)

Sonia Calvari¹¹, Gaetana Ganci¹, Sónia Silva Victória², Pedro Hernandez³, Nemesio Perez³, Vera Alfama², José Barrancos³, Jeremias Cabral⁵, Nadir Cardoso², Samara Dionis³, Paulo Fernandes², Gladys Melian³, José Pereira², Hélio Semedo⁵, Germán Padilla³, Fátima Rodriguez⁴

¹Istituto Nazionale di Geofisica e Vulcanologia, Osservatorio Etneo (INGV-OE), Catania, 95125, Italy, ²University of Cabo Verde (UNICV), CP 279 Praia, Santiago, Cape Verde, ³Instituto Tecnológico de Energias Renovábles de Tenerife (ITER), 38611 Granadilla de Abona, Tenerife, Spain,

¹Instituto Volcanológico de Canarias (INVOLCAN), 38400 Puerto de la Cruz, Tenerife, Spain,

⁵National Protection Service of Cabo Verde, Praia, Santiago, Cape Verde

**Keywords:** Remote sensing monitoring, Fogo volcano, effusive eruption, lava flow 18 inflation, lava tubes, time averaged effusion rate (TADR), magma supply rate (SR)

# **Abstract**

Fogo volcano erupted in 2014-15 producing an extensive lava flow field in the summit caldera that destroyed two villages, Portela and Bangaeira. The eruption started with powerful explosive activity, lava fountains, and a substantial ash column accompanying the opening of an eruptive fissure. Lava flows spreading from the base of the eruptive fissure produced three arterial lava flows. By a week after the start of the eruption, a master lava tube had already developed within the eruptive fissure and along the arterial flow. In this paper, we analyze the emplacement processes on the basis of observations carried out directly on the lava flow field, remote sensing measurements carried out with a thermal camera, SO<sub>2</sub> fluxes, and satellite images, in order to unravel the key factors

leading to the development of lava tubes. These were responsible for the rapid expansion of lava for the ~7.9 km length of the flow field, as well as the destruction of the Portela and Bangaeira villages. The key factors leading to the development of tubes were the low topography and the steady magma supply rate along the arterial lava flow. Comparing time-averaged effusion rates (TADR) obtained from satellite and Supply Rate (SR) derived from SO<sub>2</sub> flux data, we estimate the amount and timing of the lava flow field endogenous growth, with the aim of developing a tool that could be used for hazard assessment and risk mitigation at this and other volcanoes.

# Introduction

When an effusive eruption starts, the maximum distance that a flow can travel can be easily estimated on the basis of the measured peak effusion rate [1-4]. However, when lava tubes develop within a flow, these prevent lava from cooling, increasing its ability to cover longer distances [3,5]. Lava tubes are normally hidden below tens to hundreds of meters of lava and thus it's not easy to detect them [3,6]. The increasing use of remote sensing techniques allows a more accurate and faster detection of lava tube formation [7], and is essential for hazard assessment and risk mitigation [4,8].

The general shape of a complex lava flow field is defined by a few arterial lava flows generally displaying *aa* texture, with its outline modified by secondary lava flows normally having a *pahoehoe* surface [2,9-10]. When lava tubes develop within complex lava flows, their hidden path is revealed by the distribution of ephemeral vents or breakouts [3,11-12]. On Etna, there have been distinguished three types of ephemeral vents [3] on the basis of age and position with respect to the arterial flows and the tube system: (1) first-order ephemeral vents open at the flow front and normally feed secondary flows having *aa* 

morphology, playing a major role in the formation and propagation of the lava tube because they are continuously supplied by lava; (2) second-order ephemeral vents open on top of lava tubes and act as pressure release valves during periods of tube blockage, increase in the supply rate or inflation, forming aa or pahoehoe flow lobes; (3) third-order ephemeral vents are located at the flow margins around flows that are no longer fed, draining their interior and producing small pahoehoe or toothpaste flows. Thus, pahoehoe and toothpaste morphologies on Etna are common at the margins of aa lava flows and master tubes or in the proximal portion of aa flows that produced extensive lava tube systems [3,6,13].

The process of endogenous growth or inflation in *pahoehoe* lava flows was described [14] and quantified [15] a long time ago, as also its role in the formation of lava tubes within sheet flows. Since then, other studies have been devoted to recognizing inflation in active and past lava flows [5,16], describing its features [17-19], its importance in the growth of *aa* lava flow fields [3,13] and in the emplacement of flood basalts [20-21].

Inflation of a lava flow is often related to the formation of lava tubes. Preferred pathways develop in the older portions of the liquid-cored flow as the flow advances, and these pathways can evolve into lava tube systems within a few weeks [15]. Inflation combined with flow overlapping cause a slow and mostly undetected expansion of a complex lava flow field, and can result in sudden propagation of previously stagnating lava flow fronts [3,6,15,22-24].

Quantitative measurements of the amount of magma intruded within a lava flow field and causing inflation are scant and normally carried out only on a portion of a compound lava flow field [21,23-28]. On pahoehoe flows from Kilauea it has been documented a volumetric growth of ten times the initial flow lobe by inflation [15]. However, to the best of

84

85

86

87

88

89

90

91

92

93

94

95

96

97

98

99

100

101

102

103

104

105

106

107

108

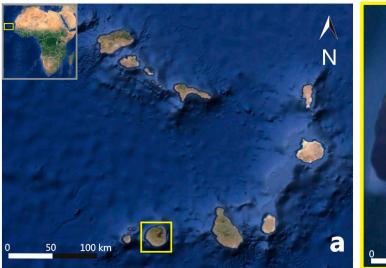
our knowledge, there are no published data offering an estimate of the amount of endogenous growth for a whole lava flow field, although a previous study carried out on Kilauea evaluated the partitioning of lava between surface and tube flows but was limited by lava spilling into the ocean [29]. In this paper, we present for the first time an estimation of the total volume endogenously intruded within the lava flow field of the Fogo 2014-15 eruption. We obtained this estimate by comparing results of time averaged discharge rate (TADR) from satellite measurements [30] and of magma supply rate (SR) from the source as retrieved from daily measurements of SO<sub>2</sub> flux, which is related to the amount of degassing magma intruded within the system [31-32]. Given that satellite images only detect surface emplacement of lava, and that SO2 flux instead is related to the total magma volume stored in the source region [32] and feeding the eruption, the difference between the two should give an estimation of the magma intruded within the lava flow field and responsible for its inflation and supplying lava tubes. We compare these results with satellite and field data on the lava flow field growth as well as published erupted volumes, and explain the fast growth of the lava flow field at its northern end that caused the destruction of the villages of Portela and Bangaeira. Provided that routine SO<sub>2</sub> flux measurements and fast calculation of TADR from satellite can be carried out, these results could be applied during future eruptions to detect lava flow field inflation. This would help predict, hopefully well in advance, the formation of lava tubes that extend lava flow fields at this and other basaltic volcanoes.

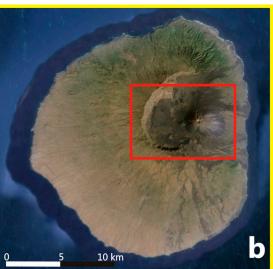
# Geologic background

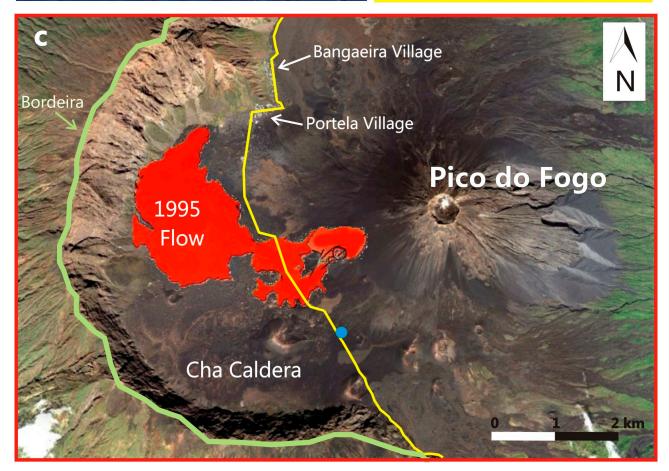
Cape Verde Islands comprise a ~50 Ma old volcanic archipelago located ~700 km off W Africa (Fig. 1a). The archipelago lies on the African plate, which drifts at a velocity of ~0.9 cm/yr above a hot spot mantle plume [33-34]. Fogo Island is located at the SW end of the archipelago (Fig. 1a), and has a conical shape, with a diameter of ~25 km and an

Peer-reviewed version available at Remote Sens. 2018, 10, 1115; doi:10.3390/rs10071115

elevation of ~2829 m (Fig. 1b). The summit is truncated at ~1700 m a.s.l. by the Cha Caldera (Fig. 1c), a 9 km wide depression open to E by a flank collapse scar [35] formed ~73 ka [36]. The upper part of this depression has a flat bottom, confined to the N, W and S by the 1000 m high vertical cliff known as Bordeira (Fig. 1c). The eastern boundary of the depression is occupied by the ~1100-m-high active volcano named Pico do Fogo (Fig. 1c), which gave rise between 1500 and 1750 CE to several eruptions from its summit. After this lapse of time, ~9 eruptive events occurred from fissures at its base, with the last episode in 1995 [34; Table SI1 and Fig. 1c]. Eruptive activity at Fogo is fed directly from the mantle plume at depths greater than 16 km, with an estimated magma SR of  $1.7 \times 10^6$  m<sup>3</sup> yr<sup>-1</sup> during the last three centuries [34]. No shallow magma storage has been found during the last effusive eruptions, and an estimated 16-24 km depth has been inferred for the magma storage that fed the 1995 eruption [37]. Erupted lavas range in composition from basanites, tephrites and nephelinites [38], with the 1995 eruption producing cogenetic basanites and phonotephrites [37].







**Figure 1 –** Google Maps with location of (a) Fogo Island (yellow square/rectangle) at Cape Verde Archipelago, western Africa, (b) Cha Caldera and Pico do Fogo volcano (inside the red rectangle), and (c) Cha Caldera depression and the paved road (yellow line) that crosses it from SE to NW, Pico do Fogo volcano, and the Portela and Bangaeira villages. The blue dot shows the fixed camera location. The 1995 lava flow field is from [39].

**Methods** 

133

134

135

136

137

138

139

140

141

142

143

144

145

146

147

148

149

150

151

152

153

154

155

156

157

Fogo volcano, and the fumaroles located within the summit crater of Pico do Fogo (Fig. 1c), have been monitored since 2008 by the Instituto Volcanológico de Canarias [40-41]. The 2014-15 eruption was jointly monitored by researchers from the University of Cape Verde (UNICV) and from the Instituto Volcanológico de Canarias (Canary Islands) directly in field, and remotely by the Istituto Nazionale di Geofisica e Vulcanologia - Osservatorio Etneo (INGV-OE) of Italy through satellite images [30]. The amount and quality of data collected during the eruption allowed us to reconstruct the expansion of the lava flow field even during the phases of inflation and endogenous growth, leading to the identification of a master lava tube that was eventually responsible for the destruction of the villages of Portela and Bangaeira. The chronology of the eruption described in this paper has been obtained through almost daily field surveys carried out by several authors of this manuscript, and using photos and thermal images, as well as through satellite images (i.e. Landsat-8 OLI, EO-1 ALI, Pléiades) and time-lapse frames recorded from a fixed monitoring camera. The fixed camera was installed by INVOLCAN on 27 November 2014 on a 4-m-high pole located at the S entrance of the Cha Caldera valley (Fig. 1c), allowing a view from the S and from a distance of ~2 km away from the eruptive fissure and proximal lava flow field. It recorded frames every 4-5 seconds from 30 November 2014 to 27 January 2015, turning to infrared at dusk. Camera failure caused a lack of image acquisition between 1 and 12 December 2014. The morphology of the lava flow field has been analyzed using the images freely available in Google Earth and acquired on 2 March 2016 (Image c 2016 CNES / Astrium). With the aim of identifying the hidden path of lava tubes within the lava flow field, we analyzed the lava flow morphology from the images freely available in Google Earth. Plume SO<sub>2</sub> flux measurements were performed along road traverses using a mini-DOAS. The mini-DOAS instrument is based on an Ocean Optics USB2000 UV spectrometer, which collects the UV radiation via an optical fiber coupled to a vertically pointing

159

160

161

162

163

164

165

166

168

169

170

171

172

173

174

175

176

177

178

179

180

181

182

telescope [42-43]. The instrument position was tracked using a handheld Global Positioning System (GPS) receiver. The mini-DOAS measures the  $SO_2$  column density in parts per million per mass at every measurement point along the plume transect. Integrated path values were obtained by adding the products of the  $SO_2$  column density and perpendicular displacement for each segment along the path. The  $SO_2$  emission rate was obtained by multiplying the integrated path values by the average wind speed. Using the daily  $SO_2$  flux estimations, we calculated magma degassing rates and volumes of magma degassed in the period between 28 November 2014 and 7 February 2015. The total volume of degassed magma ( $V_d$ ) was calculated using [32]:

$$V_d = \frac{V_S}{2[S]\rho(1-x)} \tag{1}$$

where Vs is the volume of elemental sulphur, [S] is the weight fraction of sulphur degassed per unit of magma, x is the crystal fraction and  $\rho$  is the density of magma. We used a value of 10% [44] as mean crystal fraction, 0.3 wt% as mean original sulphur content, and 2600 kg/m<sup>3</sup> as density of magma, assimilating Fogo's magma to Etna's basalt [45]. Plume SO<sub>2</sub> measurements were undertaken daily, allowing calculating the degassed magma flux (in m<sup>3</sup> s<sup>-1</sup>), and from the daily degassed magma volume we have obtained a time-averaged (daily) magma supply rate (SR) [46]. Given SR, we can integrate it through time to calculate the volume of degassed magma over the measurement period. Errors associated with the SO<sub>2</sub> method have been shown to range between 20 and 30% [47]. The thermal activity at Fogo volcano was observed and quantified from multispectralsatellite data via the HOTSAT system [48]. This volcano monitoring system processes data acquired by the Moderate Resolution Imaging Spectroradiometer (MODIS) as well as Spinning Enhanced Visible and Infrared Imager (SEVIRI) sensors in order to: (i) locate thermal anomalies, (ii) compute the associated radiant heat flux, and (iii) provide an estimation of the time averaged effusion rate (TADR; [49]). Clouds in satellite data can

alter or mask the thermal activity, thus it is retrieved by using a texton-based cloud detection algorithm [50], and a cloud index is provided to interpret the radiant heat flux signal. The hotspot detection algorithm is based on a contextual approach [51] and the radiant heat flux is computed following [52]. TADR is computed from radiant heat flux by using:

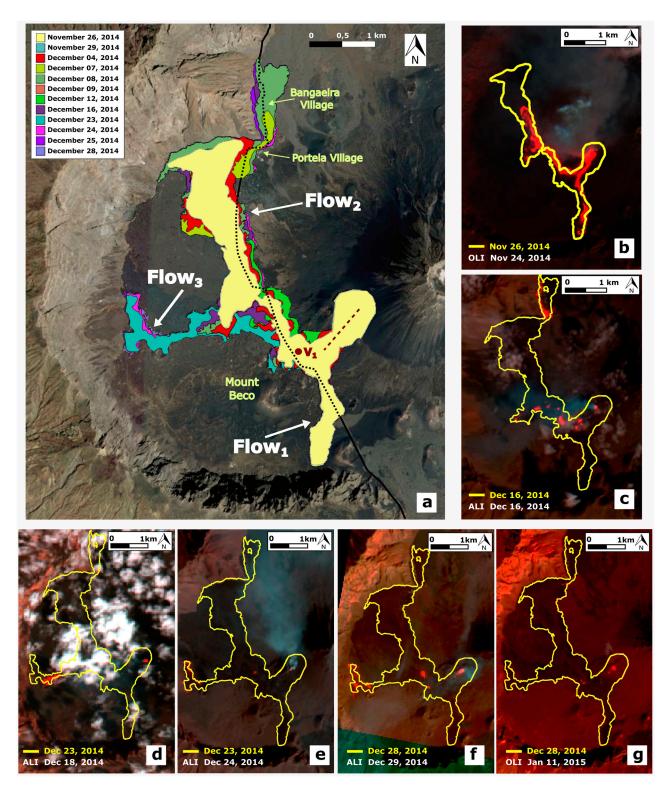
$$TADR = \frac{Q}{\rho(C_p\Delta T + C_L\Delta\varphi)}$$
 (2)

where Q is the total thermal flux obtained summing up the radiant heat flux computed for each hot spot pixel,  $\rho$  is the lava density,  $C_p$  is the specific heat capacity,  $\Delta T$  is the eruption temperature minus temperature at which flow stops,  $C_L$  is the latent heat of crystallization, and  $\Delta \varphi$  is the volume percent of crystals that form while cooling through  $\Delta T$  [53]. TADR were calculated by considering, in the variability range of each lava parameter, the largest values [30]. Moreover, for the percentage of crystals and temperature variations, we used the recent constrained values [44]. HOTSAT uses satellite images as soon as they are available, i.e. every 15 minutes in the case of SEVIRI full disk, up to five minutes in the case of SEVIRI Rapid Scanning for the

# Chronology of the eruption

northern third of Meteosat disc.

The first few days of the eruption were characterized by sudden changes in the eruptive activity and fast growth of the lava flows. The main events are summarized in Table SI1, and Figure 2a displays the key phases of growth of the lava flow field. Figures 2b-g show the active portion of the lava flow field (in red) as observed by satellites, compared to the nearest lava flow field map (yellow outline).



**Figure 2 –** (a) Map displaying the most important stages of growth of the lava flow field. The legend shows in different colours the stages of growth of the lava flow field, as well as the eruptive fissure (red dotted line), Mount Beco, Portela, Bangaeira, Vent 1, and the names of the 3 arterial flows: Flow 1, Flow 2, and Flow 3. (b-g) satellite images showing in red the active portion of the lava flow field, with the outline of the lava flow field in yellow. The destroyed (dotted black line) and residual portion (black line) of the access road to the Cha Caldera is also shown.

217 The eruption started on 23 November 2014 at 09.45 LT (LT= local time; all times in this 218 paper are LT) within the Cha Caldera depression. It began with the opening of a ENE-219 WSW, ~700-m-long eruptive fissure at the SW base of the Pico do Fogo cone between 220 2200 m and 1800 m above sea level (a.s.l.), close to the previous 1995 fissure (Figs. 1, 2). 221 Four eruptive vents opened along the fissure, giving rise to powerful Strombolian explosive 222 activity that within a few hours became lava fountains, forming a ~6 km high ash plume 223 (estimated by the Capo Verde Express pilots). At 10.00, lava flows erupted from a vent at 224 ~2100 m a.s.l., spreading at first SW along the fissure, then splitting at its base in two 225 directions: S (Flow 1) and NNW (Flow 2), being divided by the hill of Mount Beco (Figs. 226 2a-b), and then following the paved road that crossed the Cha Caldera from S to N (Fig. 227 1c), which was the only access to the villages of Portella and Bangaeira. Flow 1 travelled 228 ~1.2 km during the first 2 hours, and by 12.00 covered the Cha Caldera paved road (Fig. 229 2a). Flow 2 was initially flowing parallel to Flow 1 along the fissure, expanding SW towards 230 Mount Beco, and then it moved NNW, bounded by the high topography of Mount Beco and 231 by the 1995 lava flow field (Fig. 2a). 232 On 24 Nov, the eruptive fissure continued to open upslope, with up to seven active 233 explosive vents along it (Table SI1, Fig. 2b). The two channel-fed arterial lava flows (Flow 234 1 and Flow 2, Fig. 2b), had aa surface and reached ~2 km in length. Flow 1 slowed down 235 and became stagnated by dusk, while Flow 2 was spreading NNW. On 25 Nov morning, 236 Flow 1 was again slowly spreading S, whereas Flow 2 was expanding N and approaching 237 the Portela village. By the afternoon Flow 2 became thicker and faster, and when on 26 238 Nov it encountered the N cliff, the flow front split into two branches directed E to Portela 239 and W (Fig. 2a, b). The flow front directed E began covering the building of the Fogo 240 Natural Park, completed just a few months earlier. On 27 November, a new vent (Vent 1, 241 Figs. 2a, 3a) opened at the base of the eruptive fissure, feeding a small ~30-40 m wide

243

244

245

246

247

248

249

250

251

252

253

254

255

256

257

258

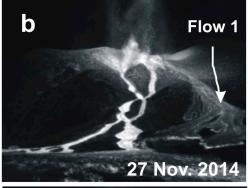
259

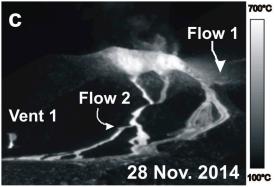
260

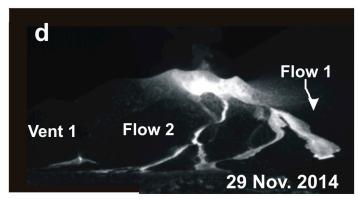
261

lava flow and expanding above Flow 2. While powerful Strombolian explosions from the upper part of the fissure were still building up two cinder cones (Figs. 3a-b), several overflows occurred from the upper S rim of the fissure, breaching and eroding the still growing cinder cones and forming several lava branches overlapping Flow 1 (Fig. 3b). This activity also continued the following days (Fig. 3c-d), and lava emerging from Vent 1 gradually built a tumulus (Fig. 3d). Vent 1 drained lava from the cinder cones on the upper fissure for three days before a significant decline of the explosive activity was observed at the summit vents, accompanied by a reduction of the surface temperature at the cones along the fissure and in the upper lava flow field (Figs. 3d-e). On 30 Nov, most of the cinder cones along the eruptive fissure were already built up, and the uppermost tumulus and lava channel fed by Vent 1 began sealing over, forming several skylights along its path, and thus revealing the presence of a proximal lava tube (Fig. 3e). Explosions and pulsating ash emission occurred along the uppermost fissure, indicating a lower magma level when compared to the previous lava fountain activity [54]. On 30 Nov, the time-lapse images from the fixed monitoring camera installed at the S end of the Cha Caldera became available (Fig. 1c), showing a fast and continuous sealing of the upper lava tube and occasional overflows from Vent 1. In general, when explosive activity at the summit vents along the fissure was increasing, we observed a decrease in the explosive and degassing activity at the skylights along the upper tube, and vice-versa, indicating that the two systems were connected at a shallow depth.









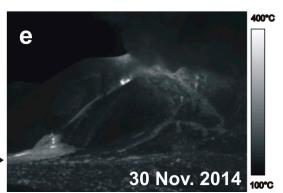


Figure 3 - Lava tube formation and growth during the early stages of the eruption shown by photo and thermal images. (a) Photo taken from SW (Mt. Beco, see Fig. 1) on 27 Nov 2014 showing the opening of Vent 1 at the base of the eruptive fissure, and (b) thermal image collected on the same day, showing lava overflows from the S crater rim feeding several branches overlapping Flow 1. Powerful Strombolian activity occurred from 7 vents along the fissure. (c) Thermal image displaying a decreased explosive activity from two active vents within the cinder cone, several overflows from the S crater rim overlapping Flow 1, and Vent 1 at the base of the eruptive fissure; (d) composition of two overlapped thermal images, displaying a significant erosion of the S crater rim by lava overflows (right of the image), and a small tumulus formed by lava accumulation around Vent 1 (left of the image), (e) Thermal image collected on 30 Nov 2014, displaying the inactive overflows from the crater rim, and Vent 1 at the base of the fissure forming a much higher tumulus than the previous day, with several skylights (white dots) displaying the path of the proximal lava tube. Note the much smaller range of the thermal scale (100-400°C) to the right of this image when compared to the scales of previous thermal images (100-700°C), suggesting a much cooler upper lava flow field.

280

281

282

283

284

285

286

287

288

289

290

291

292

293

294

295

296

263

264265

266

267268

269

270

271272

273

274

275

276277

278

279

In early December, the front of Flow 2 reached the N cliff of Bordeira, where it slowed down and inflated (Fig. 2a). At that time, most of the master tube within Flow 2 had already formed, and only ephemeral vents along the uppermost lava flow field were feeding small lava lobes on the flow surface (Fig. 2c). At this stage the lava flow field began expanding W through a second-order ephemeral vent, giving rise to initial Flow 3, whereas Flow 1 was already inactive, and most of the surface of Flow 2 between Vent 1 and Portela was crusted over (Fig. 2c). On 7 December, two first-order ephemeral vents [3] opened at the E margin of the Flow 2 front and at the exit of the lava tube that was already ~ 5.7 km long. These vents were discharging lava towards Portela and Bangaeira at a high rate, thus expanding the lava flow field a further ~1 km by 16 December (Fig. 2c). From this date to the end of December, the lava flow field widened by westward expansion of Flow 3 (Fig. 2d-e). Also this flow stopped, widened and inflated as soon as it reached the W cliff, but did not expand further because its supply was eventually cut off. From the end of December (Fig. 2f) to the end of the eruption (Fig. 2g), only small third-order ephemeral vents [3] were observed along the lava flow field, causing thickening of the proximal lava flow field [55], drainage of the main flows, and slight changes in the outline of the lava flow

298

299

300

301

302

303

304

305

306

307

308

309

310

311

312

313

314

315

316

317

318

319

320

321

322

field. The eruption eventually ended on 8 February 2015, after a gradual decline of the output rate [30].

# Field surveys and observations of lava tubes

A survey along the 2014-15 inactive lava flow field was carried out on 23 November 2015, one year after the start of the eruption. The first lava flow (Flow 1) emerging from the fissure had aa morphology (observed in the field while spreading), but this flow is not exposed. The topographic surface was then covered by a surge of lava that gave rise to a sheet flow of shining blue pahoehoe with a glassy surface, very similar to those observed at Kilauea [56]. This was covered by an aa flow that formed a channel but not a lava tube. The levees were ~1 m high outside the channel and above the pahoehoe sheet flow, and ~2 m high inside the channel. The aa surface inside the channel was 2 m lower than the levees, indicating drainage of the channel and possibly erosion of the base. In the lower portion of the channel there was a break in slope where the inner coating was detached from the walls and collapsed within the channel. This inner coating was a mere 10-cmthick layer, suggesting that the channel was not active for long (only 1 coating) and that lava feeding it was rather viscous (indicated by the thickness of the coating; [13]). In fact, Flow 1 was only active for 6 days and intermittently fed by overflows from the S rim of the fissure (Figure 3b-d). Further down along the channel and a few meters below this site, the inner coating was jammed in the middle of the channel forming a significant obstruction, and the channel disappeared below aa flows coming from overflows both from the channel and from the fissure. The intermittent supply and fast drainage of the channel were apparently responsible for its instability and blockage, deactivation, and for the lack of a lava tube within this flow. As already observed [3,5] a steady supply is the key requirement for lava tube formation.

324

325

326

327

328

329

330

331

332

333

334

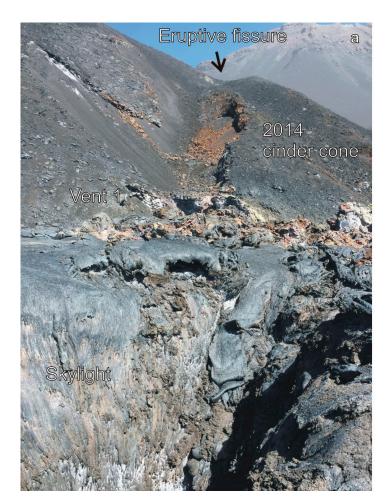
335

336

337

338

Further N, the eruptive fissure gave rise to several lava benches along the fissure crossing the cone. These benches were formed by lava stagnating at progressively lower levels while the fissure was propagating along the cone and down slope (Fig. 7a). During this propagation, at the base of the cones it formed 3-4 layers of overlapped bubbly grey lava ~30-50 cm thick suggesting a temporary lava pond, just where Vent 1 was located (Fig. 7a). This pond eventually broke when the eruptive fissure propagated down slope. confirmed by a fracture crossing it. The dike emptied from the base of the fracture, giving rise to pahoehoe sheet flows spreading like a fan. We have found a lava pond about 20 m down slope from the base of the fissure that eventually drained completely forming a skylight (Fig. 7a-b). At the base of this skylight there was a lava tube ~2-3 m in diameter and located ~20 m below the topographic surface (Fig. 7c). About 50 m down slope from this pit a larger lava pond, ~50 m wide, drained forming a circular collapsed structure ("shatter ring"; [5]). Along the path of the lava tube and further W, a 50 m wide tumulus elongated E-W suggested the prolonged feeding to a lava tube in this direction, possibly supplied by lava during the final stages of the eruption [55]. The last pahoehoe flows erupted in Portela were very black, compact and poorly vesiculated.





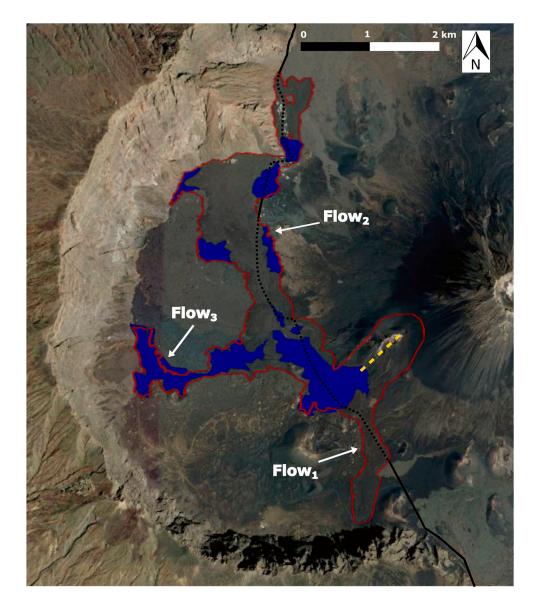


**Figure 7 -** Photos of Fogo: (a) 2014 eruptive fissure and cinder cone, taken from West on 23 November 2015, with at its base a frozen lava pond where Vent 1 was located. In the foreground is a skylight with a tube (visible in c) at its base, about 20 m below the surface. Pico do Fogo is in the background; (b) view of the skylight as taken from East and looking down flow, with the *pahoehoe* surface coating the inner walls of the skylight, displaying final drainage of lava within the depression; (c) lava tube, about 2-3 m wide, at the bottom of the skylight shown in *a* and *b*.

# Lava flow field morphology

With the aim of identifying the hidden path of lava tubes within the lava flow field, we analyzed the lava flow morphology from the images freely available in Google Earth. Figure 4 displays the portion of the lava flow field of mainly *aa* flow surfaces with the red outline, and with blue color the portions that are mainly *pahoehoe* flows. Flow 1 is essentially made up of *aa* flows, while Flow 3 is mostly *pahoehoe*. Flow 2, on the other hand, is the longest and most complex, characterized by *pahoehoe* surface at the base of

the eruptive fissure and at the boundary of the flow, and by aa surface in the middle of the flow and at the flow front. Flow 1 comprises mainly overlapped sheet flows that were intermittently supplied by lava through overflows from the eruptive fissure (Fig. 3b-d). It did not show any significant amount of inflation and its outline is not surrounded by *pahoehoe* flows. The proximal portion of the lava flow field at the base of the eruptive fissure and around Vent 1 is mainly of *pahoehoe* flows because this portion emplaced during the final stages of the eruption by drainage of degassed lava at low output rate through the skylight that replaced Vent 1 at the base of the eruptive fissure (Fig. 2a, 3e). Most of the surface of Flow 2 is made up of *aa* clinkers. It was emplaced as an arterial flow surrounded by late third-order ephemeral vents causing the partial drainage of the inflated flow interior and the emplacement of *pahoehoe* flow lobes all around it and especially at the exit of the two first-order ephemeral vents that covered the village of Portela with lava. Flow 3 is mostly formed by *pahoehoe* surface. It took place from a second-order ephemeral vent draining the middle portion of the inflated flow field when the front of Flow 2 was stopped and inflating against the N Bordeira wall.



**Figure 4 –** Map with outline of the whole lava flow field (red line) mainly comprising aa lava flow surfaces, with the blue area indicating the distribution of *pahoehoe* lava flows. See text for further explanation. (Flow 1 made of *aa* (arterial flow volume-controlled); Flow 3 made of *pahoehoe* (secondary flow fed by drainage of Flow 2); Flow 2 complex flow made of *aa* (initial arterial flow) + *pahoehoe* (at the flow margins, mainly by emptying of the flow interior) + proximal *pahoehoe* by late stage (degassed) flow emplacement. The yellow dotted line indicates the eruptive fissure, the black line the access road to Cha Caldera (dotted where covered by the lava flows).

# Results

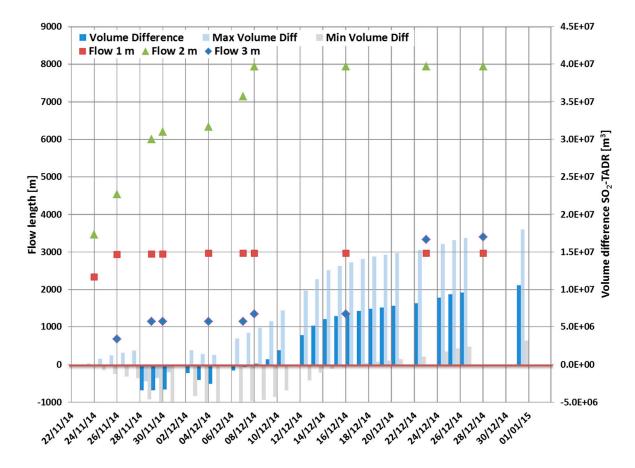
#### Pyroclastic cone

During the first week of the eruption, the explosive activity initially built up a spatter rampart and then above it two cinder cones elongated NE-SW along the eruptive fissure

(Fig. 2a). They eventually merged, forming a unique cinder cone that extended between 1950 m and 2200 m a.s.l., that was ~125 m high and with a ~500 m wide base, and with a crater width of ~165 m (measured from the map [57]). This results in a vesiculated volume of ~ 11.8 × 10<sup>6</sup> m³. Considering ~20% vesiculation for the deposit and ~30% for the magma typical of basaltic compositions [28,49,59], a volume of ~5.9 × 10<sup>6</sup> m³ dense rock equivalent (DRE) is obtained for the pyroclastic cone built up during the explosive phase of the eruption.

# Lava flow field growth

We have extracted the parameters of lava flow field growth, taking the maximum length of the three arterial flows measured along their main axes at different stages of the eruption. Figure 5 shows the expansion of Flow 1, Flow 2 and Flow 3 with time. Flow 1 displayed a significant growth only during the first two days of the eruption, expanding more than 2.0 km on the first day, but its length remained stable at 2.9 km after 26 November, when the lava was directed NNW along Flow 2. Flow 2 was the longest, reaching its maximum length of ~ 7.9 km on 8 December. Flow 3 formed after 26 November by a second-order ephemeral vent opened at the W margin of a partially roofed over lava channel, but its main expansion occurred mostly after 16 December when Flow 2 stopped, and caused a westward widening of the lava flow field. Flow 3 reached its maximum length of 3.4 km on 28 December. From December until the end of the eruption, only small secondary pahoehoe flows changed the outline of the lava flow field, causing a significant thickening and widening of the middle portion, but no more increase in length.



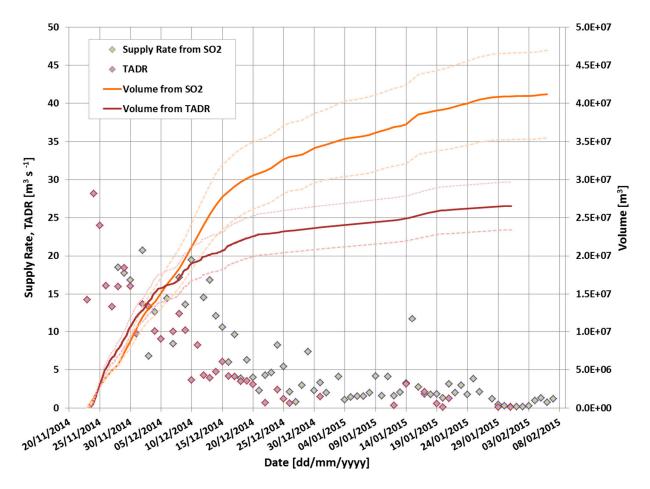
**Figure 5 –** Graph showing maximum lava flow lengths (in meters) vs. time (dd/mm/yyyy) for the three arterial lava flows that developed during the Fogo 2014-15 effusive eruption. L1 is length for Flow 1, L2 for Flow 2, and L3 for Flow 3. Note that Flow 1 increased its length only until the end of November; Flow 2 had two major growing phases, at the end of November and in early December; Flow 3 was mainly active during the second half of December, when the other two flows had largely halted. The difference in volume between the TADR measured by satellite and the Supply Rate calculated on the basis of the SO<sub>2</sub> flux measurements (blue bar), with uncertainty bars (gray and pale blue bars), is also shown, with values increasing especially from early December onwards.

# Lava flow field volume

Figures 5 and 6 show the difference in volume between the TADR measured by satellite (data from [30]) and the SR calculated here from  $SO_2$  flux measurements. It is worth noting that  $SO_2$  flux measurements started on 27 November, thus the SR estimates are lacking for the first three days of eruption. The difference between intruded and erupted magma volume becomes greater than 2 ×  $10^6$  m<sup>3</sup> in the second half of December, when the lava tube within Flow 2 was mostly sealed and started draining, thus feeding Flow 3 with mostly

pahoehoe lava. The tube efficiency caused the lack of large surface aa flows (Fig. 2 c-g), thus the lava flow field was dotted with only small pahoehoe lobes. Figure 6 displays the difference between supplied (SR) and erupted (TADR) magma volumes during the whole eruption. The difference between SR and TADR increases from mid December 2014 until the end of the eruption. The final difference between supplied and erupted magma is 14.7  $\pm$  8.8  $\times$  10<sup>6</sup> m³ (Fig. 6), and we suggest that this volume represents the magma causing inflation and endogenous growth of the lava flow field, undetected by satellite since it did not expand on the surface.



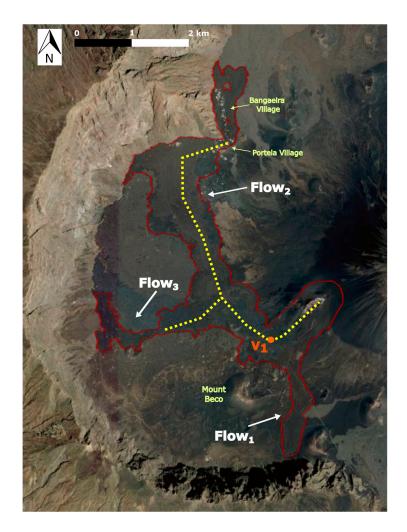


**Figure 6 –** Graph showing Time Averaged Discharge Rate (TADR, purple diamonds) retrieved by HOTSAT and Supply Rate derived from SO<sub>2</sub> data (green diamonds) vs. time (dd/mm/yyyy) during the Fogo 2014-15 effusive eruption. Red and orange solid lines and relative errors (dashed lines) show volumes obtained by integrating TADR and Supply Rate curves, respectively.

Discussion

Comparing this episode with lava fountains observed at Mt Etna, we assume that the cinder cone was mainly formed by proximal ballistic fallout during the lava fountain phase of the eruption [59-61]. Given that the cone built up in 7 days (between 23 and 30 November 2014), the DRE TADR [49] for its growth is ~9.7 m³ s⁻¹. This is much higher than the DRE 0.12 m³ s⁻¹ measured for the growth of the cinder cone within Etna's Bocca Nuova summit crater in 2012 [28] and than the DRE 1.5 m³ s⁻¹ of the 2002 Laghetto cone on Etna's S flank [59]. It is also lower than the 50 m³ s⁻¹ that gave rise to the New SE Crater during powerful explosive episodes in 2011-2013 [60].

The direct observation of the eruption compared with thermal images collected from both the ground and satellite, as well as the morphology map distinguishing *pahoehoe* from *aa* surfaces, has enabled recognizing the stages of formation and growth of the lava tubes within the lava flow field, their position and extension (Fig. 8). The first sector of lava tube to become sealed was the uppermost portion within the eruptive fissure and up to Vent 1. This was ~1.1 km long and sealed in ~ 7 days. In fact, as shown in Figure 3e, on 30 November the eruptive fissure was mostly roofed over, there was a tumulus at its base above Vent 1, and several skylights along the upper tube. This caused a significant cooling of the flow surface, as shown by the comparison between Figures 3b-d and Figure 3e. Tubes on steep slopes and along eruptive fissures form more rapidly, because local changes in discharge rate can easily trigger inward accretion of levees and roofing over of lava channels [3,5].



**Figure 8 -** Interpretative model. Outline of the final lava flow field (red line) with path of lava tubes (yellow dotted line). The position of V1 at the base of the eruptive fissure is also shown.

The second sector of lava tube to form was the longest, extending on the low slope between Vent 1 and the N cliff of Bordeira, in the middle portion of Flow 2. This tube was the down slope continuation of the proximal portion; it was about ~4.8 km long and formed in about 10 days (Fig. 2c). It led to the destruction of the two villages of Portela and Bangaeira by fast-spreading lava flows erupted from first-order ephemeral vents at the E margin of Flow 2 front. Indeed, on 16 December (Fig. 2c), only small *pahoehoe* lobes were spreading along the upper flow field, most of the lava being carried within the ~5.9 km master tube to the N end of the lava flow field.

483 The third sector of lava tube to form was the ~1.1 km branch that fed Flow 3. It formed 484 much more slowly than the previous, because it started on 26 November but then stopped between 29 November and 7 December. It was only after the halting of the Flow 2 front 485 486 that the middle portion of the lava field inflated, and a second-order ephemeral vent 487 opened on the W margin of the lava flow field, acting as a pressure-release valve [3] 488 Calvari and Pinkerton, 1998]. It drained the flow field interior feeding Flow 3 with mostly 489 pahoehoe lava. The supply to Flow 3 ended by 23 December, and no further expansion of 490 Flow 3 was observed after this date. No evidence of lava tubes has been found within 491 Flow 1. 492 The distribution of pahoehoe and aa flow surfaces has important implications for 493 recognizing the hidden path of lava tubes. In fact, Flow 1, which is mainly made up of aa 494 surfaces, did not show any evidence of lava tubes in the field. It was characterized by a 495 drained channel obstructed by its collapsed inner coating, indicating complete drainage of 496 the channel and successive lack of supply. Flow 1 was intermittently fed by overflows from 497 the fissure, a condition that does not enable lava tube development [3,5]. Flow 2, which 498 mainly comprises aa surfaces and pahoehoe at the margins of the flow, has been 499 characterized by a long lava tube extending from the eruptive fissure down to V1 (~ 1.1 500 km), and from here all along the flat surface of Cha Caldera up to the N boundary of 501 Bordeira (~ 4.8 km), for a total of ~5.9 km length. Another small sector of lava tube, ~1.1 502 km long, formed during the second half of December as a lateral branch of the master tube 503 that fed Flow 2. This fed Flow 3, made up almost entirely by pahoehoe surface. The 504 presence of the tube system was confirmed in the proximal lava flow field by the skylight 505 observed during the field survey (Fig. 7c), displaying a ~2-3 m wide lava tube located ~20 506 m below the flow surface. The skylight inner wall was coated by pahoehoe lava congealed 507 while dribbling inwards and showing final drainage of fluid lava within the upper tube (Fig. 508 7a-b).

509 Thus, on the basis of the data here described, a master tube developed within the eruptive 510 fissure and along Flow 2 after just one week from the start of the eruption, and its growth completed within about 10 days from the start of the eruption. This tube was responsible 511 512 for the fast lengthening of Flow 2, and formed a short branch to the W as soon as Flow 2 513 slowed down when encountering the N cliff of Bordeira. The magma inside the tube then 514 accumulated upslope inflating the lava field and eventually feeding Flow 3. The path of 515 Flow 2 was dictated by the line of maximum steepness [55], but its expansion was 516 probably favored also by the smooth surface of the paved road crossing the Cha Caldera 517 from S to N, given that most of the road had been covered by lava, as already happened 518 also during the 1995 eruption [39]. 519 In order to estimate the quantity of magma intruding the lava flow field and causing its 520 inflation, we have compared the measurements of SR derived from SO<sub>2</sub> flux 521 measurements, and the TADR derived from satellite [30]. Although the measurements of 522 SO<sub>2</sub> flux, and then of SR, are lacking for the first three days of the eruption, we can 523 assume that at that time there was no difference between supplied and erupted magma, 524 given that the lava flows were spreading with no significant surface cooling (Fig. 2b), and 525 that there was no inflation or endogenous growth of the lava flow field. The difference 526 between SR and TADR increased significantly after 9 December, as soon as Flow 2 527 stopped advancing (Figs. 5-6), and this is when Flow 3 increased its length. Once Flow 2 528 reached its maximum length allowed by surface crust cooling [9], the magma flowing within 529 the master tube accumulated to the rear and upslope, inflating the middle portion of the 530 flow field. This triggered an expansion of the tube towards the W for ~1.1 km, feeding Flow 531 3. Flow 3 was mostly pahoehoe, and fed by part of the degassed interior causing the 532 inflation of the lava flow field. 533 Another significant increase in the difference between SR and TADR occurred after 21 534 December when also Flow 3 stopped advancing (Figs. 5-6), and only small pahoehoe flow

536

537

538

539

540

541

542

543

544

545

546

547

548

549

550

551

552

553

554

555

556

557

558

559

lobes were observed on the proximal lava flow field (Fig. 2e). Flow 3 stopped after coming up against the W Bordeira wall and following a significant decline in the output rate [30]. At this stage, the supply of lava was not sufficient to keep the ~5.9 km long tube active. Lava thus poured out from the base of the eruptive fissure [Richter et al., 2016], thickening the proximal lava flow field with a pahoehoe fan of lava from the skylight that eventually drained at the end of the effusion (Figure 7). This is the second major inflation stage of the lava flow field, not followed by major lava flow expansion. This is probably due to the confining effect of the lava crustal growth [9]. By the end of the eruption, the erupted volume estimated by satellite measurements is  $26.5 \pm 3.2 \times 10^6$  m<sup>3</sup> DRE (Figure 6) corresponding only to surface lava, while mass volume obtained from SO<sub>2</sub> flux measurements is  $46.9 \pm 5.74 \times 10^6$  m<sup>3</sup>. Considering the respective uncertainty associated to each estimation, a total difference of 14.7 ± 8.8 × 10<sup>6</sup> m<sup>3</sup> DRE was obtained between SR and TADR. We argue that this differential volume, that is about 31% of the total erupted volume, caused inflation of the lava flow field and endogenous growth. Published bulk values obtained from differential digital elevation models (DEMs; 45.83 ±  $0.02 \times 10^6 \text{ m}^3$  [62]; 43.7 ± 5.2 × 10<sup>6</sup> m<sup>3</sup>, [55]) are in agreement with mass estimation derived from SO<sub>2</sub> flux measurements. Differences between supply rates and TADR can be also ascribed to voids from the tubes. drained flows and empty spaces between crust and flow. The high amount of voids within the flow field is testified by the significant vertical displacement by cooling and contraction measured ~9 months after the end of the eruption [62], and observed also at other volcanoes [63]. It is noteworthy that the greatest contraction affected the thickest zones of the flow field, these being the eruptive fissure and the middle line of Flow 2, following the path of the master tube that fed the lava flow field (Figure 7).

561

562

563

564

565

566

567

568

569

570

571

572

573

574

575

576

577

578

579

580

581

582

583

584

585

The comparison between SR and TADR highlighted a growing difference between the two values from early December 2014 onwards. This is consistent with surface cooling and inflation of the lava field, and with the endogenous accumulation of lava. It is worth noting that the main destruction of Portela and Bangaeira occurred at this very stage of the eruption, when there was enough magma available within the inflated flow field to feed fast-spreading secondary flows from the exit of the master tube. The development of lava tubes within the Fogo lava flow field was favored by a stable supply of lava along Flow 2, and by the flat topography on which lava was spreading. We suggest that any time lava flow spreads on a flat surface, the early comparison between TADR estimated by satellite and SR obtained from SO<sub>2</sub> flux measurements might allow recognizing if endogenous growth is occurring, leading to lava tube expansion and to potentially much greater length and destructive power of the lava flow field. The 2014-15 effusive eruption at Fogo was very similar to the previous 1995 eruption. The effusive vents of both eruptions were located in the same zone, i.e. at the W base of Pico do Fogo cone (Fig. 1), sharing the same NE-SW orientation. The duration and erupted volume of both eruptions is similar, with 54 days of eruption and 46 × 10<sup>6</sup> m<sup>3</sup> lava erupted in 1995 [33 Amelung and Day, 2002] and 77 days and ~46 × 10<sup>6</sup> m<sup>3</sup> erupted magma in 2014-15. Given that the last two lava flow fields affected the same flat ground of Cha Caldera, and that this area is the most prone to lava invasion [55], should a new eruption occur in the same zone, we suggest measuring and comparing the calculated SR from SO<sub>2</sub> flux measurements and the TADR obtained by satellite as soon as possible. This would enable detecting major phases of inflation of the lava flow field in good time to prevent damage.

# **Conclusive remarks**

586

587

588

589

590

591

592

593

594

595

596

597

598

599

600

601

602

603

604

605

606

607

608

609

610

611

The 2014-15 eruption at Fogo volcano developed a ~5.9 km long lava tube system that started from the eruptive fissure that became sealed in just six days after the start of the eruption. The master tube soon propagated within the arterial Flow 2, following the line of maximum slope and probably also the smooth paved road of access to the Cha Caldera. When Flow 2 stopped against the N Bordeira cliff, the lava flow field inflated and eventually led to vent opening at the E margin of its front with lava flows that destroyed the villages of Portela and Bangaeira. At this stage, further lengthening of the lava flow field was probably prevented by surface crust cooling. Thus the lava feeding the eruption accumulated upslope, inflating the proximal flow field and causing a fast spreading of Flow 3 during the second half of December. This is also when the difference between SR and TADR increased significantly (Figs. 5-6), causing inflation and endogenous growth of the lava flow field. When also Flow 3 halted, we recorded a growing difference between SR and TADR and a significant inflation of the lava flow field. However, the low SR of this stage and the containment offered by the cooling crust did not allow further expansion of the lava flow field, and only proximal minor pahoehoe lobes increased the thickness of the lava flow field. We have estimated that by the end of the eruption, ~31% of the lava emplaced endogenously. We suggest that early detection of flow inflation might help recognize on time the destructive power of a lava flow field in time to help prevent further devastation.

# **Acknowledgements**

We would like to thank the following institutions that gave us support and help for the monitoring of the eruptive activity and field work at Fogo: Universidade de Cabo Verde (UNICV), Praia, Santiago, Cape Verde; Instituto Tecnológico de Energias Renovábles (ITER); Instituto Volcanológico de Canarias (INVOLCAN); Serviço Nacional de Proteção

- 612 Civil, Cabo Verde; Ministério de Desenvolvimento Rural, Cabo Verde; Câmara Municipal 613 de São Filipe, Ilha do Fogo, Cabo Verde; Parque Natural do Fogo, Direcção Nacional do Ambiente, Cabo Verde; Associação dos Guias Turísticos da Ilha do Fogo, Cabo Verde; 614 615 Nuno Coelho, ESRI Portugal; Cabildo Insular de Tenerife-Acción Exterior; Forças 616 Armadas de Cabo Verde; Policia Nacional de Cabo Verde. Thanks are due to European 617 Organisation for the Exploitation of Meteorological Satellites (EUMETSAT) for SEVIRI data 618 (www.eumetsat.int) and to National Aeronautics and Space Administration (NASA) for 619 MODIS data (modis.gsfc.nasa.gov). Landsat 8 OLI and Eo-1 ALI images are courtesy of 620 the U.S. Geological Survey (earthexplorer.usgs.gov). We are grateful to the Copernicus 621 emergency management service (emergency.copernicus.eu/mapping/list-of-622 components/EMSR111) for mapping the actual lava flow field by Cosmo-SkyMed and 623 Pleiades images. The authors would like to thank M. James and M. Patrick who produced 624 thoughtful reviews that helped clarify and improve a previous version of the manuscript. 625 The English style has been reviewed by Stephen Conway. Data are available from the 626 figures and references.
- 628 References
- 1. Walker, G. P. L. Lengths of lava flows. *Phil. Trans. R. Soc. Lond.* **1973**, *A 274*, 107–
- 630 118.

- 631 2. Kilburn, C.R.J.; Lopes, R.M.C. The growth of aa lava flow fields on Mount Etna, Sicily,
- 632 J. Geophys. Res. Solid Earth 1988, 93, 14759-14772.
- 633 3. Calvari, S.; Pinkerton, H. Formation of lava tubes and extensive flow field during the
- 634 1991-93 eruption of Mount Etna, *J. Geophys. Res. Solid Earth* **1998**, *103 (B11)*.
- 635 27291-27302.
- 4. Bonaccorso, A.; Calvari, S.; Boschi, E. Hazard mitigation and crisis management
- during major flank eruptions at Etna volcano: reporting on real experience. In:

- 638 Detecting, Modelling and Responding to Effusive Eruptions; Harris, A. J. L., De
- Groeve, T., Garel, F., Carn, S. A., Eds.; Geological Society, London, Special
- Publications, London, U.K., 2015; Volume 426, pp. 447-461; ISBN 978-1-86239-736-
- 1. http://doi.org/10.1144/SP426.4
- 5. Kauahikaua, J.; Cashman, K.V.; Mattox, T.N.; Heliker, C.C.; Hon, K.A.; Mangan, M.T.;
- Thornber, C.R. Observations on basaltic lava streams in tubes from Kilauea Volcano,
- island of Hawaii. *J. Geophys. Res. Solid Earth* **1998**, *103*(*B11*), 27303-27323.
- 6. Calvari, S.; Coltelli, M.; Neri, M.; Pompilio, M.; Scribano, V. The 1991-93 Etna
- eruption: chronology and lava flow field evolution. *Acta Vulcanol.* **1994**, *4*, 1-14.
- 7. Harris, A.J.L. *Thermal Remote Sensing of Active Volcanoes*, Cambridge University
- 648 Press: Cambridge, U.K., 2013; pp. 728. ISBN 978-0-521-85945-5.
- 8. Solana, M.C.; Calvari, S.; Kilburn, C.R.J.; Gutierrez, H.; Chester, D.; Duncan, A.
- Supporting the Development of Procedures for Communications During Volcanic
- Emergencies: Lessons Learnt from the Canary Islands (Spain) and Etna and Stromboli
- 652 (Italy). Advances in Volcanology **2017**, DOI 10.1007/11157\_2016\_48.
- 9. Guest, J.E.; Kilburn, C.R.J.; Pinkerton, H.; Duncan, A.M. The evolution of lava flow
- fields: Observations of the 1981 and 1983 eruptions of Mount Etna, Sicily. *Bull.*
- 655 *Volcanol.* **1987**, *4*9, 527-540.
- 10. Kilburn, C.R.J.; Lopes, R.M.C. General Patterns of Flow Field Growth: Aa and Blocky
- 657 Lavas. J. Geophys. Res. Solid Earth **1991**, 96(B12), 19721-19732.
- 11. Guest, J.E.; Underwood, J.R.; Greeley, R. Role of lava tubes in flows from the
- Observatory Vent, 1971 eruption on Mount Etna. *Geol. Mag.* **1980**, *117*, 601-606.
- 12. Mattox, T.N.; Heliker, C.; Kauahikaua, J.; Hon, K. Development of the 1990 Kalapana
- Flow Field, Kilauea Volcano, Hawaii. *Bull. Volcanol.* **1993**, *55*, 407-413.
- 13. Calvari, S.; Pinkerton, H. Lava tube morphology on Etna and evidence for lava flow
- emplacement mechanisms. J. Volcanol. Geotherm. Res. 1999, 90, 263-280.

- 14. Walker, G.P.L. Structure, and origin by injection under surface crust, of tumuli, "lava
- rises," "lava-rise pits," and "lava inflation clefts" in Hawaii. Bull. Volcanol. 1991, 53,
- 666 546– 58.
- 15. Hon, K.; Kauahikaua, J.; Denlinger, R.; Mackay K. Emplacement and inflation of
- pahoehoe sheet flows: Observations and measurements of active lava flows on
- 669 Kilauea Volcano, Hawaii. *Geol. Soc. Am. Bull.* **1994**, *106*, 351-370.
- 16. Rossi, M.J.; Gudmundsson A. The morphology and formation of flow-lobe tumuli on
- 671 Icelandic shield volcanoes. *J. Volcanol. Geotherm. Res.* **1996**, *72(11996)*, 291-308.
- 17. Cashman, K.V.; Kauahikaua, J.P. Reevaluation of vesicle distributions in basaltic lava
- 673 flows. *Geology* **1997**, 25, 419–22.
- 18. Keszthelyi, L.; Self, S. Some physical requirements for the emplacement of long
- basaltic lava flows. *J. Geophys. Res. Solid Earth* **1998**, 103(B11), 27447-27464.
- 19. Self, S.; Keszthelyi, L.; Thordarson, T. The importance of pahoehoe. *Annu. Rev. Earth*
- 677 Planet. Sci. **1998**, 26, 81-110.
- 20. Self, S.; Thordarson, T.; Keszthelyi, L.; Walker, G.P.L.; Hon, K.; Murphy, M.T. A new
- model for the emplacement of Columbia River basalts as large, inflated pahoehoe lava
- 680 flow fields. Geophys. Res. Lett. **1996**, 23(19), 2689-2692.
- 21. Thordarson, T.; Self, S. The Roza Member, Columbia River Basalt Group: A gigantic
- pahoehoe lava flow field formed by endogenous processes? J. Geophys. Res. Solid
- 683 *Earth* **1998**, *103(B11)*, 27411-27445.
- 22. Applegarth, L.J.; Pinkerton, H.; James, M. R.; Calvari, S. Lava flow superposition: the
- reactivation of flow units in compound flow fields. J. Volcanol. Geotherm. Res. 2010,
- 686 194, 100-106, doi:10.1016/j.jvolgeores.2010.05.001.
- 687 23. Favalli, M.; Harris, A.J.L.; Fornaciai, A.; Pareschi, M.T.; Mazzarini, F. The distal
- segment of Etna's 2001 basaltic lava flow. Bull. Volcanol. 2010, 72, 119–127, doi:
- 689 10.1007/s00445-009-0300-z.

- 690 24. Favalli, M.; Fornaciai, A.; Mazzarini, F.; Harris, A.J.L.; Neri, M.; Behncke, B.; Pareschi,
- 691 M.T.; Tarquini, S.; Boschi, E. Evolution of an active lava flow field using a
- 692 multitemporal LIDAR acquisition. J. Geophys. Res. Solid Earth 2010, 115,B11203,
- 693 doi:10.1029/2010JB007463.
- 694 25. Umino, S.; Nonaka, M.; Kauahikaua, J. Emplacement of subaerial pahoehoe lava
- sheet flows into water: 1990 Kūpaianaha flow of Kilauea volcano at Kaimu Bay,
- 696 Hawaii. Bull. Volcanol. **2006**, 69, 125-139, doi: 10.1007/s00445-006-0059-4.
- 697 26. James, M.R.; Applegarth, L.J.; Pinkerton, H. Lava channel roofing, overflows,
- breaches and switching: insights from the 2008–2009 eruption of Mt. Etna. *Bull.*
- 699 *Volcanol.* **2012**, 74, 107-117, doi: 10.1007/s00445-011-0513-9.
- 700 27. Hamilton, C.W.; Glaze, L.S.; James, M.R.; Baloga, S.M. Topographic and stochastic
- influences on pahoehoe lava lobe emplacement. *Bull. Volcanol.* **2013**, *75*, 756,
- 702 doi:10.1007/s00445-00013-00756-00448.
- 703 28. Slatcher, N.; James, M.R.; Calvari, S.; Ganci, G.; Browning, J. Quantifying effusion
- rates at active volcanoes through integrated time-lapse laser scanning and
- 705 photography. *Remote Sens.* **2015**, 7, 14967-14987, doi:10.3390/rs71114967.
- 706 29. Koeppen, W.C.; Patrick, M.; Orr, T.; Sutton, A.J.; Dow, D.; Wright, R. Constraints on
- the partitioning of Kīlauea's lavas between surface and tube flows, estimated from
- infrared satellite data, sulfur dioxide emission rates, and field observations. *Bull.*
- 709 *Volcanol.* **2013**, 75:716, doi:10.1007/s00445-013-0716-3.
- 30. Cappello, A.; Ganci, G.; Calvari, S.; Perez, N.M.; Hernandez, P.A.; Silva, S.V.; Cabral,
- J.; Del Negro, C. Lava Flow Hazard Modeling during the 2014-2015 Fogo eruption,
- 712 Cape Verde. J. Geophys. Res. Solid Earth **2016**, 121, 2290-2303, doi:
- 713 10.1002/2015JB012666.

- 31. Sutton, A.J.; Elias, T.; Kauahikaua, J. Lava-Effusion Rates for the Puu Oo-Kupaianaha
- Fruption derived from SO2 Emissions and Very Low Frequency (VLF) Measurements.
- 716 USGS Prof. Paper **2003**, 1676, 137-148.
- 32. Allard, P. Endogenous magma degassing and storage at Mount Etna. *Geophys. Res.*
- 718 Lett. **1997**, 24(17), 2219-2222.
- 33. Courtney, R.C.; White, R.S. Anomalous heat flow and geoid across the Cape Verde
- Rise: evidence for dynamic support from a thermal plume in the mantle. *Geophys. J.*
- 721 R. Astr. Soc. 1986, 87, 815-867.
- 34. Amelung, F.; Day, S. InSAR observations of the 1995 Fogo, Cape Verde, eruption:
- 723 Implications for the effects of collapse events upon island volcanoes. *Geophys. Res.*
- 724 Lett. **2002**, 29(12):1606, doi: 10.1029/2001GL013760.
- 35. Day, S.J.; Heleno da Silva, S.I.N.; Fonseca, J.F.B.D. A past giant lateral collapse and
- present-day flank instability of Fogo, Cape Verde Islands. J. Volcanol. Geotherm. Res.
- 727 **1999**, 94, 191–218.
- 728 36. Ramalho, R.S.; Winckler, G.; Madeira, J.; Helffrich, G.R.; Hipolito, A.; Quartau, R.;
- Adena, K.; Schaefer, J.M. Hazard potential of volcanic flank collapses raised by new
- 730 megatsunami evidence. *Sci. Adv.* **2015**, *1*(9)e1500456, doi:10.1126/sciadv.1500456.
- 731 37. Hildner, E.; Klügel, A.; Hauff, F. Magma storage and ascent during the 1995 eruption
- of Fogo, Cape Verde Archipelago. Contrib. Mineral. Petrol. 2011, 162, 751–772,
- 733 doi:10.1007/s00410-011-0623-6.
- 38. Escrig, S.; Doucelance, R.; Moreira, M.; Allègre, C.J. Os isotope systematics in Fogo
- 735 Island: Evidence for lower continental crust fragments under the Cape Verde Southern
- 736 Islands. *Chem. Geol.* **2005**, *219*, 93–113.
- 737 39. Texier-Teixeira, P.; Chouraqui, F.; Perrillat-Collomb, A.; Lavigne, F.; Cadag, J.R.;
- Grancher, D. Reducing volcanic risk on Fogo Volcano, Cape Verde, through a

- participatory approach: which outcome? Nat. Hazards Earth Syst. Sci., 2014, 14,
- 740 2347–2358, doi:10.5194/nhess-14-2347-2014.
- 40. Dionis, S.M.; Melian, G.; Rodriguez, F.; Hernandez, P.A.; Padron, E.; Perez, N.M.;
- Barrancos, J.; Padilla, G.; Sumino, H.; Fernandes, P.; Bandomo, Z.; Silva, S.; Pereira,
- J.M.; Semedo, H. Diffuse volcanic gas emission and thermal energy release from the
- summit crater of Pico do Fogo, Cape Verde. *Bull. Volcanol.* **2015**, 77:10, doi:
- 745 10.1007/s00445-014-0897-4.
- 41. Dionis, S.M.; Perez N.M., Hernandez P.A., Melian G., Rodriguez F., Padron E.,
- Sumino H., Barrancos J., Padilla G., Fernandes P., Bandomo Z., Silva S., Pereira
- J.M., Semedo H., Cabral J. Diffuse CO2 degassing and volcanic activity at Cape
- 749 Verde islands, West Africa. *Earth Plan. Space* **2015**, *67:48*, doi: 10.1186/s40623-015-
- 750 0219-x.
- 42. Galle, B.; Oppenheimer, C.; Geyer, A.; McGonigle, A.; Edmonds, M.; Horrocks, L.A. A
- miniaturized ultraviolet spectrometer for remote sensing of SO2 fluxes: a new tool for
- volcano surveillance. J. Volcanol. Geotherm. Res. 2002, 119, 241–254.
- 43. Barrancos, J.; Roselló, J.I.; Calvo, D.; Padrón, E.; Melián, G.; Hernández, P.A.; Pérez,
- N.M.; Millán, M.M.; Galle, B. SO2 emission from active volcanoes measured
- simultaneously by COSPEC and mini-DOAS. *Pure Appl. Geophys.* **2008**, *165*, 115–
- 757 133.
- 44. Mata, J.; Martins, N.; Mattielli, N.; Madeira, J.; Faria, B.; Ramalho, R.S.; Silva, P.;
- Moreira, M.; Caldeira, R.; Rodrigues, J.; Martins, L. The 2014–15 eruption and the
- short-term geochemical evolution of the Fogo volcano (Cape Verde): Evidence for
- small-scale mantle heterogeneity. *Lithos* **2017**, *288-289*, 91-107,
- 762 doi:10.1016/j.lithos.2017.07.001.
- 45. Spilliaert, N.; Allard, P.; Metrich, N.; Sobolev, A.V. Melt inclusion record of the
- conditions of ascent, degassing, and extrusion of volatile-rich alkali basalt during the

- powerful 2002 flank eruption of Mount Etna (Italy). J. Geophys. Res. Solid Earth 2006,
- 766 *111(B043203)*, doi:10.1029/2005JB003934.
- 46. Steffke, A.M.; Harris, A.J.L.; Burton, M.; Caltabiano, T.; Salerno, G.G. Coupled use of
- 768 COSPEC and satellite measurements to define the volumetric balance during effusive
- eruptions at Mt. Etna, Italy. J. Volcanol. Geotherm. Res. 2010, 205, 47–53.
- 47. Caltabiano, T.; Burton, M.; Giammanco, S.; Allard, P.; Bruno, N.; Muré, F.; Romano,
- R. Volcanic gas emissions from the summit craters and flanks of Mt. Etna, 1987–2000.
- In: Mt. Etna: Volcano Laboratory. Bonaccorso, A., Calvari, S., Coltelli, M., Del Negro,
- 773 C., Falsaperla, S. (Eds.). *AGU Geophys. Monograph* 2004; Volume 143, pp. 111–128,
- 774 doi: 10.1029/143GM08, ISBN 0-87590-408-4...
- 48. Ganci, G.; Bilotta, G.; Cappello, A.; Hérault, A.; Del Negro, C. HOTSAT: a
- multiplatform system for the satellite thermal monitoring of volcanic activity. In:
- 777 Detecting, Modelling and Responding to Effusive Eruptions; Harris, A. J. L., De
- Groeve, T., Garel, F., Carn, S. A., Eds.; Geological Society, London, Special
- 779 Publications, London, U.K., 2015; Volume 426, pp. 207-222; ISBN 978-1-86239-736-
- 780 1.
- 49. Harris, A.J.L.; Dehn, J.; Calvari, S. Lava effusion rate definition and measurement: a
- 782 review. Bull. Volcanol. 2007, 70, 1-22, doi: 10.1007/s00445-007-0120-y.
- 783 50. Ganci, G.; Vicari, A.; Bonfiglio, S.; Gallo, G.; Del Negro, C. A texton-based cloud
- detection algorithm for MSG-SEVIRI multispectral images. *Geomatics Nat. Haz. Risk*
- 785 **2011**, 2, 279-290, doi: 10.1080/19475705.2011.578263.
- 51. Ganci, G.; Vicari, A.; Fortuna, L.; Del Negro, C. The HOTSAT volcano monitoring
- system based on combined use of SEVIRI and MODIS multispectral data. *Ann.*
- 788 *Geophys.* **2011**, *54*, 544-550, doi:10.4401/ag-5338.

- 789 52. Wooster, M.; Zhukov, B.; Oertel, D. Fire radiative energy release for quantitative study
- of biomass burning: derivation from the BIRD experimental satellite and comparison to
- 791 MODIS fire products. *Rem. Sens. Env.* **2003**, *86*, 83–107.
- 792 53. Harris, A.; Blake, S.; Rothery, D.; Stevens, N. A chronology of the 1991 to 1993 Mount
- Etna eruption using advanced very high resolution radiometer data: implications for
- real-time thermal volcano monitoring. J. Geophys. Res. Solid Earth 1997, 102, 7985-
- 795 8003.
- 54. Patrick, M.R.; Harris, A.J.L.; Ripepe, M.; Dehn, J.; Rothery, D.A.; Calvari, S.
- 797 Strombolian explosive styles and source conditions: insights from thermal (FLIR)
- 798 video. *Bull. Volcanol.* **2007**, *69*, 769-784, doi: 10.1007/s00445-006-0107-0.
- 799 55. Richter, N.; Favalli, M.; Zeeuw-van Dalfsen, E.; Fornaciai, A.; da Silva Fernandes,
- R.M.; Perez Rodriguez, N.; Levy, J.; Silva, S.V.; Walter, T.R. Lava flow hazard at Fogo
- Volcano, Cape Verde, before and after the 2014-2015 eruption. *Nat. Hazards Earth*
- 802 Syst. Sci. Discuss. **2016**, *16*, 1925-1951, doi:10.5194/nhess-16-1925-2016.
- 56. Oze, C.; Winter, J.D. The occurrence, vesiculation, and solidification of dense blue
- glassy pahoehoe. *J. Volcanol. Geotherm. Res.* **2005**, *142*, 285-301.
- 57. Gonzales, A.R.; Pérez Torrado, F.J.; Carracedo Gómez, J.C.; Medina, C.J.M.; Garcia,
- A.B.; de la Torre, E.G.; Cigala, A.N.; Paris, R.; Rodrigues, A.N.; Dinis, H.A.; Andrade,
- J.P. Carta Geológica Ilha do Fogo, Cabo Verde. Mercurio Editorial 2015, ISBN: 978-
- 808 84-943863-8-1. D.L.: GC 372-2015.
- 809 58. Houghton, B.F.; Schmincke, H.-U. Rothenberg scoria cone, East Eifel: a complex
- Strombolian and phreatomagmatic volcano. *Bull. Volcanol.* **1989**, *52*, 28-48.
- 59. Calvari, S.; Pinkerton, H. Birth, growth and morphologic evolution of the "Laghetto"
- cinder cone during the 2001 Etna eruption. J. Volcanol. Geotherm. Res. 2004, 132,
- 813 225-239, doi:10.1016/S0377-0273(03)00347-0.

- 814 60. Behncke, B.; Branca, S.; Corsaro, R.A.; De Beni, E.; Miraglia, L.; Proietti, C. The
- 2011–2012 summit activity of Mount Etna: Birth, growth and products of the new SE
- 816 crater. *J. Volcanol. Geotherm. Res.* **2014**, 270, 10–21.
- 817 61. Bonaccorso, A.; Calvari, S.; Linde, A.; Sacks, S. Eruptive processes leading to the
- most explosive lava fountain at Etna volcano: The 23 November 2013 episode.
- 819 Geophys. Res. Lett. **2014**, 41, 4912–4919, doi:10.1002/2014GL060623.
- 820 62. Bagnardi, M.; Gonzales, P.J.; Hooper, A. High-resolution digital elevation model from
- tri-stereo Pleiades-1 satellite imagery for lava flow volume estimates at Fogo Volcano.
- 822 Geophys. Res. Lett. **2016**, 43, doi:10.1002/2016GL069457.
- 823 63. Stevens, N.F.; Wadge, G.; Williams, C.A.; Morley, J.G.; Muller, J.-P.; Murray, J.B.;
- Upton, M. Surface movements of emplaced lava flows measured by synthetic aperture
- radar interferometry. *J. Geophys. Res. Solid Earth* **2001**, *106,B6*, 11293-11313.