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Article

Gaussian Versus Mean-Field Model: Contradictory Predictions for the Casimir Force Under Dirichlet–Neumann Boundary Conditions

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Abstract: The mean-field model (MFM) is the workhorse of the statistical mechanics: one normally accepts that it yields results which, despite differing numerically from the correct ones, are not “very wrong”, in that they resemble the actual behavior of the system as eventually obtained by a more advanced treatments. This, for example, turns out to be the case for the Casimir force under, say, Dirichlet–Dirichlet, $(+, +)$ and $(+, -)$ boundary conditions (BC) for which, according to the general expectations the MFM delivers *attractive* for like BC—or *repulsive* for unlike BC—force, with the principally correct position of the maximum strength of the force below, or above the critical point T_c . It turns out, however, that this is *not* the case with Dirichlet–Neumann (DN) BC. In this case, the mean-field approach leads to an attractive Casimir force. This contradiction with the “boundary condition rule” is cured in the case of the Gaussian model under DN BC. Our results, which are mathematically exact, demonstrate that the Casimir force within the MFM is attractive as a function of temperature T and external magnetic field h , while for the Gaussian model it is repulsive for $h = 0$, and can be, surprisingly, both repulsive and attractive for $h \neq 0$. The treatment of the MFM is based on the exact solution of one non-homogeneous nonlinear differential equation of second order. The Gaussian model is analyzed both in its continuum and lattice realization. The obtained outcome teaches us that the mean-field results should be accepted with caution in the case of fluctuation-induced forces and ought to be checked against more precise treatment of the fluctuations within the envisaged system.

Keywords: finite-size effects; exact results; Casimir force; mean-field model; Gaussian model; phase transitions; critical phenomena; phase diagrams

1. Introduction

Currently, the most prominent example of a fluctuation-induced force is the force due to quantum or thermal fluctuations of the electromagnetic field, leading to the so-called QED Casimir effect, named after the Dutch physicist H.B. Casimir who first realized that in the case of two perfectly-conducting, uncharged, and smooth plates parallel to each other in vacuum, at $T = 0$ these fluctuations lead to an *attractive* force between them [1]. Thirty years after Casimir, Fisher and De Gennes [2] showed that a very similar effect exists in critical fluids, today known as critical Casimir effect. A summary of the results available for this effect can be found in the recent reviews [3–6]. The description of the critical Casimir effect is based on finite-size scaling theory [7–9]. Let us envisage a system with a film geometry $\infty^{d-1} \times L$, $L \equiv L_\perp$, and with boundary conditions ζ imposed along the spatial direction of finite extent L . Take $\mathcal{F}_{\text{tot}}^{(\zeta)}$ to be the total free energy of such a system within the grand canonical ensemble (GCE). Then, if $f^{(\zeta)}(T, h, L) \equiv \lim_{A \rightarrow \infty} \mathcal{F}_{\text{tot}}^{(\zeta)} / A$ is the free energy per area A of the system,

one can define the Casimir force for critical systems in the grand-canonical $(T - h)$ -ensemble, see, e.g. Refs. [4,9–11]:

$$\beta F_{\text{Cas}}^{(\zeta)}(L, T, h) \equiv -\frac{\partial}{\partial L} f_{\text{ex}}^{(\zeta)}(L, T, h) \quad (1.1)$$

where

$$f_{\text{ex}}^{(\zeta)}(L, T, h) \equiv f^{(\zeta)}(L, T, h) - L f_b(T, h) \quad (1.2)$$

is the so-called excess (over the bulk) free energy per area and per $\beta^{-1} = k_B T$. Here we suppose a system at temperature T is exposed to an external ordering field h , which couples linearly to its order parameter—such as the number density, the concentration difference, the magnetization etc. Actually, the thermodynamic Casimir force $F_{\text{Cas}}^{(\zeta)}(T, h, L)$ per area is the excess pressure over the bulk one due to the finite size ($L < \infty$) of that system:

$$F_{\text{Cas}}^{(\zeta)}(T, h, L) = P_L^{(\zeta)}(T, h) - P_b(T, h). \quad (1.3)$$

Here $P_L^{(\zeta)}$ is the pressure in the finite system under boundary conditions ζ , while P_b is the one in the infinite, i.e., macroscopically large, system. The above definition is actually equivalent to Eq. (1.1). Note that $f_{\text{ex}}^{(\zeta)}(L, T, h)$ is the excess grand potential per area, $f^{(\zeta)}(L, T, h)$ is the grand canonical potential per area of the finite system, while $f_b(T, h)$ has the meaning of the grand potential per volume V for the macroscopically large system. The equivalence between the definitions in Eqs. (1.1) and (1.3) stems from the observation that for the finite system one has $P_L = -\partial f^{(\zeta)}(L, T, h)/\partial L$, while for the bulk one and $f_b = -P_b$.

When $F_{\text{Cas}}^{(\zeta)}(L, t, h) < 0$ the excess pressure is inward towards the system, i.e., there is an *attraction* of the surfaces of the system towards each other and a *repulsion* if $F_{\text{Cas}}^{(\zeta)}(L, t, h) > 0$.

2. The Casimir Force Within the Continuum Gaussian Model

The continuum version of the Gaussian model with a scalar order parameter consists of the linear and bilinear terms in the Ginzburg-Landau-Wilson formulation of a system in d dimensions that undergoes a continuous symmetry-breaking phase transition at low temperatures. The partition function of this system is the functional integral

$$\mathcal{Z}_G(t, h) = \int \exp[-\mathcal{F}(\psi(\vec{r}))] \mathcal{D}\{\psi(\vec{r})\} \quad (2.1)$$

where

$$\mathcal{F}(\psi(\vec{r})) = \int \left[t\psi(\vec{r})^2 + |\vec{\nabla}\psi(\vec{r})|^2 - h\psi(\vec{r}) \right] d^d r \quad (2.2)$$

In (2.2) t is the reduced temperature, proportional to $T - T_c$, and h is the spatially constant ordering field. Because of the Gaussian nature of the free energy functional $\mathcal{F}(\psi(\vec{r}))$ the partition function resolves into the product

$$\mathcal{Z}_G(t, h) = \mathcal{Z}_{G,I}(t) \times \mathcal{Z}_{G,h}(t, h) \quad (2.3)$$

where $\mathcal{Z}_{G,I}(t)$ is the partition function of the system with $h = 0$. The geometry of the system under consideration is a slab of large—ultimately infinite—cross section and finite thickness L .

With regard to scaling considerations, there are two combinations of parameters that reflect the predictions of finite size scaling. They are

$$x_t = tL^{1/\nu} = tL^2, \quad (2.4)$$

$$x_h = hL^{(d+2-\eta)/2} = hL^{(d+2)/2}, \quad (2.5)$$

where ν , the correlation length exponent, is equal to $1/2$ in the Gaussian model, and as noted above d is the dimensionality of the system. Our end results for the Casimir forces acting upon the systems will depend on the boundary conditions imposed. In all cases, the form of the Casimir force is

$$f_{\text{Cas}}(t, h, L) = L^{-d} \left(w_{\text{Cas}, \text{I}}(x_t) + x_h^2 w_{\text{Cas}, \text{h}}(x_t) \right) \quad (2.6)$$

All results reported in this portion of the article rely on two results, which can be obtained with the use of contour integration techniques; see also [12]. The two results are

$$\sum_{n=-\infty}^{\infty} \frac{1}{an^2 + b} = \frac{\pi \coth(\pi \sqrt{b/a})}{\sqrt{ab}}, \quad (2.7)$$

$$\sum_{n=0}^{\infty} \frac{1}{c(2n+1)^2 + d} = \frac{\pi \tanh(\pi/2 \sqrt{d/c})}{4\sqrt{cd}}. \quad (2.8)$$

In order to carry out the evaluation of the free energy of the Gaussian model we turn to the basis set of functions that will be used to construct the free energy with and without an ordering field. These functions allow us to evaluate the partition function by integrating over the amplitudes of the contributions of each member of the set to the order parameter. Here, we focus on the case of periodic boundary conditions. Ignoring the dependence on position in the “plane” of the slab, the functions are the orthonormal set

$$\psi_c^{(n)}(z) = \sqrt{2/L} \cos(2\pi n z / L) \quad (2.9)$$

$$\psi_s^{(n)}(z) = \sqrt{2/L} \sin(2\pi n z / L) \quad (2.10)$$

$$\psi_0(z) = \sqrt{1/L} \quad (2.11)$$

with n a positive integer. It is straightforward to show that this set is orthonormal as a function of z in that

$$\int_0^L \psi_c^{(n)}(z) \psi_c^{(m)}(z) dz = \delta_{m,n} \quad (2.12)$$

$$\int_0^L \psi_s^{(n)}(z) \psi_s^{(m)}(z) dz = \delta_{m,n} \quad (2.13)$$

$$\int_0^L \psi_0(z)^2 dz = 1 \quad (2.14)$$

The three function types are all mutually orthogonal. In the case of higher dimensions, we construct a new basis set by multiplying the functions (2.9)–(2.11) by suitable functions of the orthogonal position variables. Those functions can be taken to be of the form $e^{i\vec{Q} \cdot \vec{R}}$, where \vec{R} is a $d - 1$ -dimensional position vector in the plane of the slab and \vec{Q} is in its reciprocal space.

We then express the order parameter as follows

$$\psi(z, \vec{R}) = \sum_{\vec{Q}} e^{i\vec{Q} \cdot \vec{R}} \left(\sum_{n=1}^{\infty} a_n^{(c)} \psi_c^{(n)}(z) + \sum_{n=1}^{\infty} a_n^{(s)} \psi_s^{(n)}(z) + a_0 \psi_0(z) \right) \quad (2.15)$$

The free energy for a given configuration of the Gaussian order parameter, in terms of the amplitudes in the expansion of the order parameter in the basis set (2.12)–(2.14), is

$$\sum_{\vec{Q}} \left[\sum_{n=1}^{\infty} (a_n^{(c)2} + a_n^{(s)2}) \left(t + Q^2 + (2\pi n/L)^2 \right) + a_0^2 t - h a_0 \sqrt{L} \right] \quad (2.16)$$

The last term in brackets above reflects the fact that the only basis function that the constant external field couples to is the constant function in (2.11)

The next step is to exponentiate the expression in (2.16), multiply by either $-1/\beta$, or setting $\beta = 1$, by -1 , and, after that, to perform the Gaussian integrals over the $a_n^{(c)}$'s, the $a_n^{(s)}$'s, and a_0 . The resulting partition function is given by

$$\begin{aligned} Z &= \exp \left[\frac{1}{\beta} \left(\frac{h^2 LA}{4t} + \frac{A}{(2\pi)^{d-1}} \int d^{d-1}Q \sum_{n=-\infty}^{\infty} \frac{1}{2} \ln \left(\frac{t + Q^2 + (2\pi n/L)^2}{\pi} \right) \right) \right] \end{aligned} \quad (2.17)$$

The coefficient A in (2.17) is the $d - 1$ dimensional area of the slab.

As our next step we evaluate the sum over n on the right hand side of the expression for the partition function. To achieve this, we take the t -derivative of the logarithm of the summand, perform the sum over n and then integrate the resulting expression with respect to t . Taking the derivative of the summand in (2.17) with respect to t leaves us with the sum

$$\frac{1}{2} \sum_{n=-\infty}^{\infty} \frac{1}{t + Q^2 + (2\pi n/L)^2} = \frac{L \coth \left(\frac{1}{2} L \sqrt{Q^2 + t} \right)}{2\sqrt{Q^2 + t}} \quad (2.18)$$

which follows from (2.7). This integrates up to

$$2 \ln \left(\sinh \left(L \sqrt{Q^2 + t} / 2 \right) \right) \quad (2.19)$$

The large- L limit of (2.19) is

$$L \sqrt{t + Q^2} \quad (2.20)$$

To find the contribution to the Casimir force per unit area, we take the L -derivative of the difference between (2.20) and (2.19) and then integrate over \vec{Q} . The derivative yields

$$\frac{1}{2} \sqrt{Q^2 + t} \left(1 - \coth(L \sqrt{Q^2 + t}) \right) = -\sqrt{Q^2 + t} \frac{e^{-L \sqrt{Q^2 + t}}}{e^{L \sqrt{Q^2 + t}} - e^{-L \sqrt{Q^2 + t}}} \quad (2.21)$$

The sum over values of \vec{Q} is expressible as an integral, which takes the form

$$\begin{aligned} & -\frac{K_{d-1}}{(2\pi)^{d-1}} \int_0^\infty Q^{d-2} \sqrt{Q^2 + t} \frac{e^{-L \sqrt{Q^2 + t}}}{e^{L \sqrt{Q^2 + t}} - e^{-L \sqrt{Q^2 + t}}} dQ \\ &= \frac{1}{L^d} \left(-\frac{K_{d-1}}{(2\pi)^{d-1}} x_t^{d/2} \int_0^\infty w^{d-2} \sqrt{1 + w^2} \frac{e^{-\sqrt{x_t(1+w^2)}}}{e^{\sqrt{x_t(1+w^2)}} - e^{-\sqrt{x_t(1+w^2)}}} dw \right) \\ &= \frac{1}{L^d} X_I^{(\text{per},3)}(x_t) \end{aligned} \quad (2.22)$$

where, to get to the last line of (2.22) we defined a new integration variable $w = Q/\sqrt{t}$ and then made use of the definition (2.4) of x_t . The implication of (2.22) is that we can express the $h = 0$ contribution to the Casimir force as L^{-d} times a function of the scaling temperature variable x_t . The coefficient K_d in the equations above is the geometric factor

$$K_d = \frac{2\pi^{d/2}}{\Gamma\left(\frac{d}{2}\right)} \quad (2.23)$$

In the case of three dimensions, further processing of the result (2.22) is possible. We find

$$X_{\text{Cas},I}^{\text{per},3}(x_t) = -\frac{2\sqrt{x_t}\text{Li}_2\left(e^{-2\sqrt{x_t}}\right) + \text{Li}_3\left(e^{-2\sqrt{x_t}}\right) - 2x_t \log\left(1 - e^{-2\sqrt{x_t}}\right)}{8\pi} \quad (2.24)$$

where $\text{Li}_j(x)$ is the polylogarithm function; see [13]. A plot of the function $X_{\text{Cas},I}^{\text{per},3}(x_t)$ is shown in Figure 1.

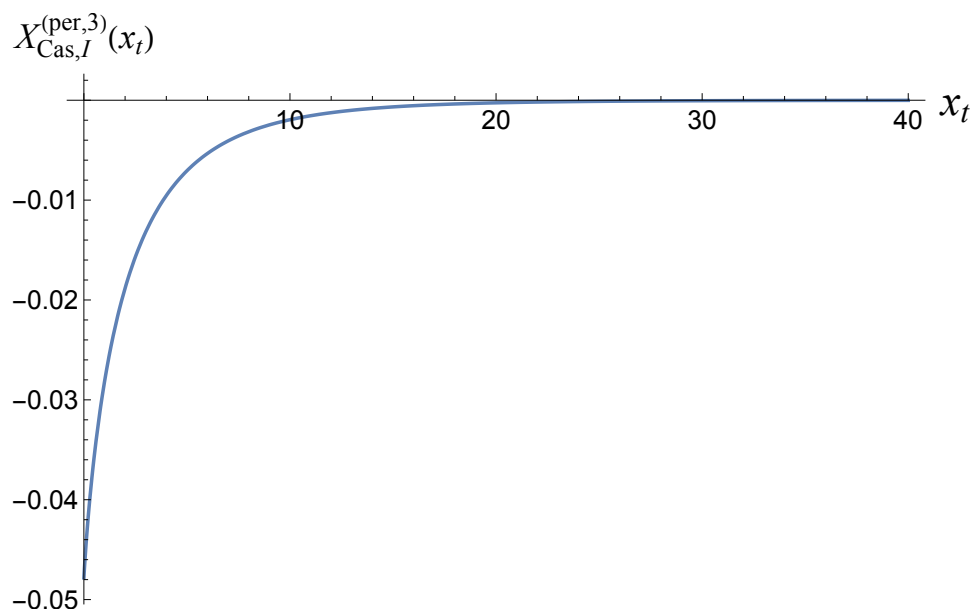


Figure 1. The function $X_{\text{Cas},I}^{\text{per},3}(x_t)$, plotted versus x_t

The first term in parentheses in Eq. (2.17) gives us the h -dependent contribution to the free energy: $-h^2 LA/4t$. This is to be compared to the corresponding free energy of a neighboring bulk phase, which goes as $-h^2(L_0 - L)A/4t$, where L_0 is an extent that will ultimately be taken to go to infinity. If you add the two free energies, the dependence on L , the thickness of the slab, disappears. This means that there is no h -dependent free energy when slab boundary conditions are periodic, and hence no h -dependent contribution to the Casimir force.

The calculations in the case of periodic boundary conditions point the way to evaluating the partition function and the Casimir force of the case of Dirichlet-Neumann boundary conditions.

In this case the (unnormalized) basis functions are, exclusive of their dependence on the in-plane coordinates,

$$\sin((2n+1)\pi z/2L) \quad (2.25)$$

Examples of these functions are shown in Figure 2. Focusing on the h -independent contribution to the partition function, the sum to perform in this case is (see (2.8))

$$\frac{1}{2} \sum_{n=0}^{\infty} \frac{1}{t + Q^2 + ((2n+1)\pi/2L)^2} = \frac{L \tanh\left(L\sqrt{t + Q^2}\right)}{4\sqrt{t + Q^2}} \quad (2.26)$$

Note that in the limit of large L the right hand side goes to the expected asymptotic form. If we subtract that limiting form, and integrate with respect to t , we are left with

$$\frac{1}{2} \left(\log\left(\cosh\left(L\sqrt{t + Q^2}\right)\right) - L\sqrt{t + Q^2} \right) \quad (2.27)$$

Finally, we take minus the derivative of this with respect to L , leaving us with

$$\frac{1}{2} \left(-\sqrt{t+Q^2} \tanh(L\sqrt{t+Q^2}) + \sqrt{t+Q^2} \right) = \sqrt{t+Q^2} \frac{e^{-L\sqrt{t+Q^2}}}{e^{L\sqrt{t+Q^2}} + e^{-L\sqrt{t+Q^2}}} \quad (2.28)$$

Making use of the analysis of previous sections, this leaves us with the following result for the Casimir force in the case of the d -dimensional Gaussian model with Dirichlet-Neumann boundary conditions

$$\begin{aligned} & \frac{K_{d-1}}{(2\pi)^{d-1}} \int_0^\infty Q^{d-2} \sqrt{t+Q^2} \frac{e^{-L\sqrt{t+Q^2}}}{e^{L\sqrt{t+Q^2}} + e^{-L\sqrt{t+Q^2}}} dQ \\ &= \frac{K_{d-1}}{(2\pi)^{d-1}} \frac{1}{L^d} (x_t)^{d/2} \int_0^\infty w^{d-2} \sqrt{1+w^2} \frac{e^{-\sqrt{x_t}\sqrt{1+w^2}}}{e^{\sqrt{x_t}\sqrt{1+w^2}} + e^{-\sqrt{x_t}\sqrt{1+w^2}}} dw \\ &= \frac{1}{L^d} X_{\text{Cas,D/N}}^{(d)}(x_t) \end{aligned} \quad (2.29)$$

When $d = 3$, we have

$$X_{\text{Cas,D/N,I}}^{(3)}(x_t) = -\frac{2\sqrt{x_t} \text{Li}_2(-e^{-2\sqrt{x_t}}) + \text{Li}_3(-e^{-2\sqrt{x_t}}) - 2x_t \log(e^{-2\sqrt{x_t}} + 1)}{8\pi} \quad (2.30)$$

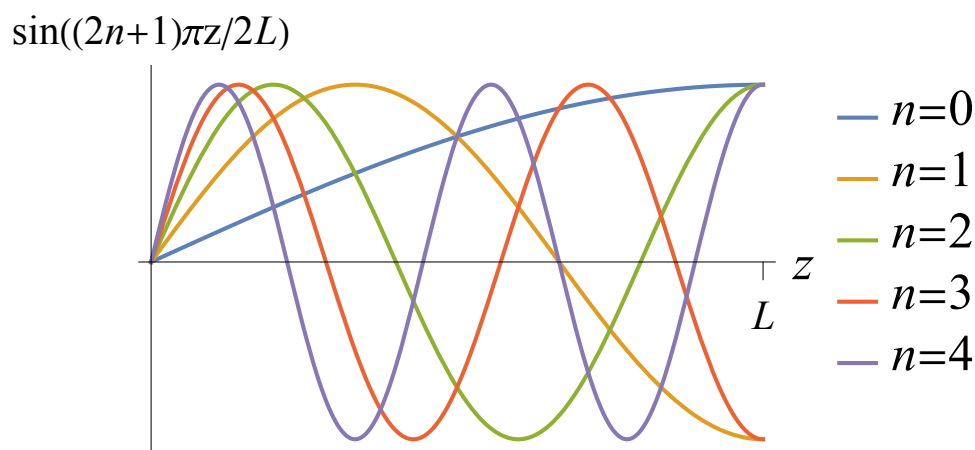


Figure 2. The functions in (2.25)

Figure 3 shows what the function $X_{\text{Cas,D/N,I}}^{(d)}(x_t)$ looks like when $d = 3$.

In order to find the h -dependent contribution to the Casimir force we turn to the normalized the basis set in the case of Dirichlet-Neumann boundary conditions. Assuming that the boundary conditions are Dirichlet at $z = 0$ and Neumann at $z = L$, this basis set is

$$\psi_{DN}^{(n)}(z) = \sqrt{2/L} \sin((n+1/2)\pi z/L) \quad (2.31)$$

with n an integer and

$$0 \leq n < \infty \quad (2.32)$$

It is straightforward to establish that

$$\int_0^L \psi_{DN}^{(n)}(z)^2 dz = 1 \quad (2.33)$$

while

$$\int_0^L \psi_{DN}^{(n)}(z) dL = \frac{2\sqrt{2L}}{(2n+1)\pi} \quad (2.34)$$

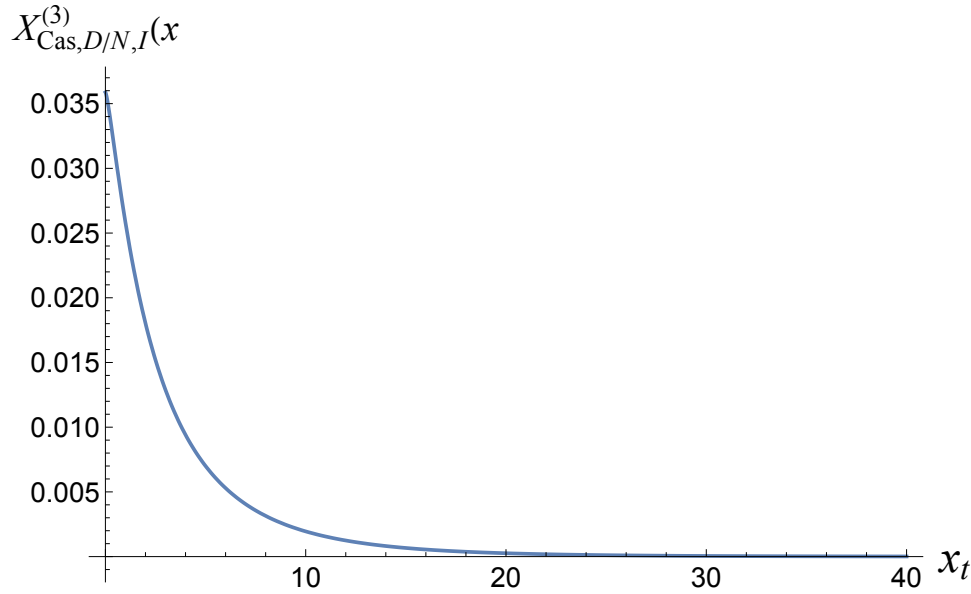


Figure 3. The function $X_{Cas,D/N}^{(3)}(x_t)$, as given in (2.30).

As it turns out there is no need to take into account any dependence of the basis set on coordinates in the plane of the slab. This is because a constant ordering field couples only to order parameter configurations that are independent of those coordinates.

With this in mind, we expand the order parameter as follows

$$\Psi(z) = \sum_{n=0}^{\infty} a_n^{(DN)} \psi_{DN}^{(n)}(z) \quad (2.35)$$

The Gaussian integrations over the $a_n^{(DN)}$'s leaves us with the summation over n for the h -dependent contribution to the partition function

$$\begin{aligned} & \exp \left[h^2 \sum_{n=0}^{\infty} \left(\frac{2\sqrt{2L}}{(2n+1)\pi} \right)^2 \frac{1}{4((\pi(n+1/2)/L)^2 + t)} \right] \\ &= \exp \left[h^2 \left(\frac{L}{4t} - \frac{\tanh(L\sqrt{t})}{4t^{3/2}} \right) \right] \\ &= \exp \left[\frac{h^2}{4t^{3/2}} (L\sqrt{t} - \tanh(L\sqrt{t})) \right] \end{aligned} \quad (2.36)$$

where the evaluation of the sum over n in (2.36) is accomplished with the use of (2.8) and a partial fraction decomposition of the summand. The first term in parentheses on the last line of (2.36) gives us exactly the same expression as the h -dependent contribution to the partition function of the slab with

periodic boundary conditions. Its influence on the Casimir force is exactly canceled by the influence of the bulk. What remains is

$$\begin{aligned}
 -h^2 \frac{\partial}{\partial L} \tanh(L\sqrt{t})/4t^{3/2} &= -\frac{h^2}{4t} \operatorname{sech}^2(L\sqrt{t}) \\
 &= -\frac{h^2 L^2}{4x_t} \operatorname{sech}^2(\sqrt{x_t}) \\
 &= -\frac{1}{L^d} \frac{x_h^2}{4x_t} \operatorname{sech}^2(\sqrt{x_t})
 \end{aligned} \tag{2.37}$$

where we have made use of the definition of the scaling combination x_h in (2.5). The scaling form of the contribution to the Casimir force is, then

$$X_{D/N}^{(3)}(x_t, x_h) = \frac{-x_h^2}{4x_t} \operatorname{sech}^2(\sqrt{x_t}) \tag{2.38}$$

This function is shown in Figure 4. Note that this function is *always attractive*.

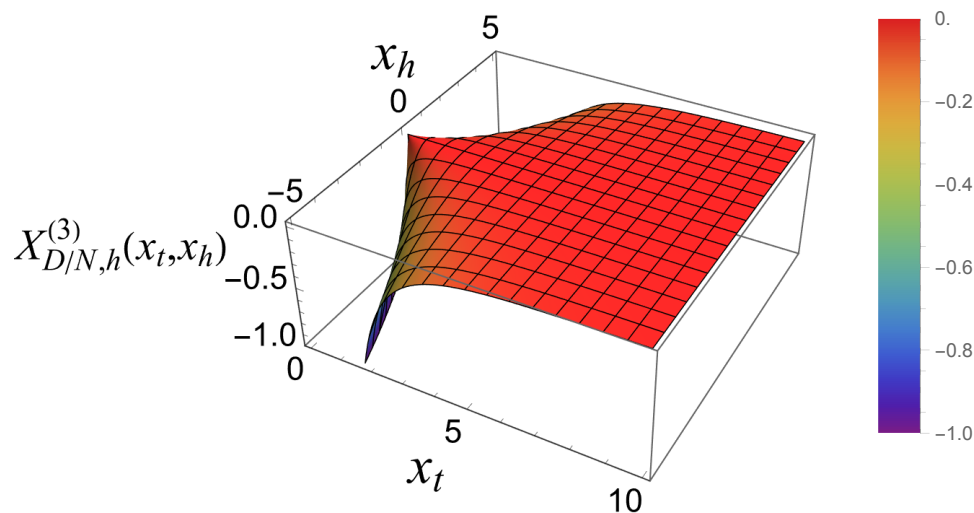


Figure 4. The function $X_{D/N,h}^{(3)}(x_t, x_h)$, as given by (2.38).

The total scaling function $X_{D/N}^{(3)}(x_t, x_h)$ is given by

$$X_{D/N,h}^{(3)}(x_t, x_h) = -\frac{2\sqrt{x_t} \operatorname{Li}_2(-e^{-2\sqrt{x_t}}) + \operatorname{Li}_3(-e^{-2\sqrt{x_t}}) - 2x_t \log(e^{-2\sqrt{x_t}} + 1)}{8\pi} - \frac{x_h^2}{4x_t} \operatorname{sech}^2(\sqrt{x_t}). \tag{2.39}$$

Figure 5 shows what this function looks like.

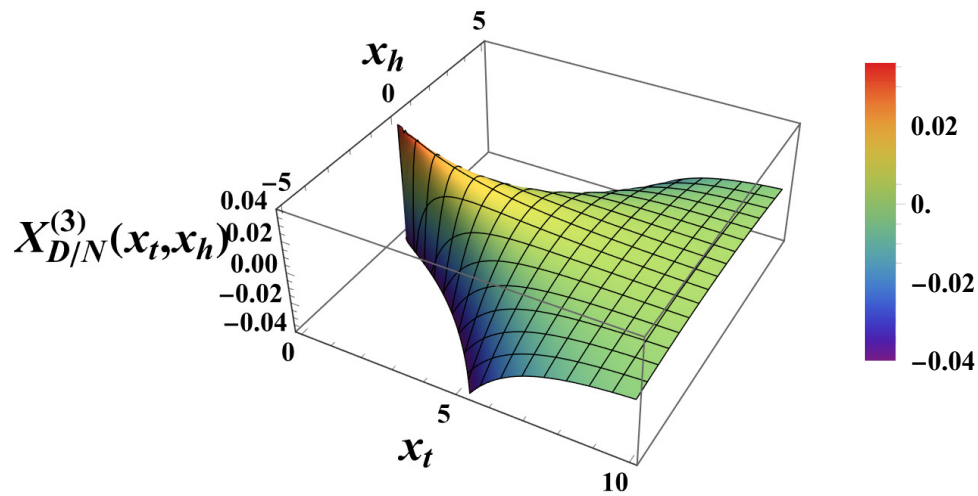


Figure 5. The total scaling contribution to the Casimir force for Dirichlet-Neumann boundary conditions in the three dimensional Gaussian model with a scalar order parameter, $X_{D/N}^{(3)}(x_t, x_h)$. Note that this function can be both positive (repulsive) and negative (attractive).

Another depiction of the scaling contribution to the Casimir force for Dirichlet-Neumann boundary conditions in the three dimensional Gaussian model with a scalar order parameter, $X_{D/N}^{(3)}(x_t, x_h)$, Figure 6, highlights the regions in which the function is attractive and repulsive.

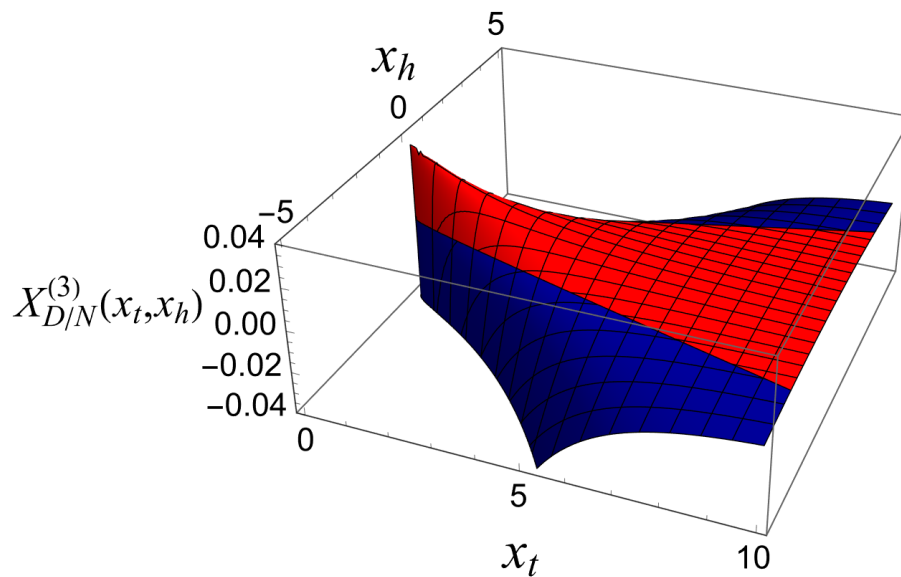


Figure 6. The total scaling contribution to the Casimir force for Dirichlet-Neumann boundary conditions in the three dimensional Gaussian model with a scalar order parameter, $X_{D/N}^{(3)}(x_t, x_h)$. The red region in the figure corresponds to a repulsive force, and the blue region corresponds to an attractive force.

3. The Casimir Force Within the Lattice Gaussian Model

We consider a ferromagnetic model with nearest-neighbor interactions on a fully finite d -dimensional hypercubic lattice $\Lambda \in \mathbb{Z}^d$ of $|\Lambda|$ sites. Let us take $\Lambda \in \mathbb{Z}^d$ to be the parallelepiped $\Lambda = \mathcal{L}_1 \times \cdots \times \mathcal{L}_d$, where \times denotes the direct (Cartesian) product of the finite sets $\mathcal{L}_\nu = \{1, \dots, L_\nu\}$.

It is convenient to consider the configuration space $\Omega_\Lambda = R^{|\Lambda|}$ as an Euclidean vector space in which each configuration is represented by a column-vector S_Λ with components labeled according to the lexicographic order of the set $\{\mathbf{r} = (r_1, \dots, r_d) \in \Lambda\}$. Let S_Λ^\dagger be the corresponding transposed row-vector and let the dot (\cdot) denote matrix multiplication. Then, for given boundary conditions $\tau = (\tau_1, \dots, \tau_d)$, specified for each pair of opposite faces of Λ by some τ_v takes the form

$$\beta \mathcal{H}_\Lambda^{(\tau)}(S_\Lambda | K) = -\frac{1}{2} K S_\Lambda^\dagger \cdot Q_\Lambda^{(\tau)} \cdot S_\Lambda. \quad (3.1)$$

Here $K = \beta J$, where J is the interaction constant (to be set to $J = 1$ in the remainder), and the $|\Lambda| \times |\Lambda|$ interaction matrix $Q_\Lambda^{(\tau)}$ can be written as

$$Q_\Lambda^{(\tau)} = (\Delta_1^{(\tau_1)} + 2E_1) \times \dots \times (\Delta_d^{(\tau_d)} + 2E_d), \quad (3.2)$$

where $\Delta_v^{(\tau_v)}$ is the one-dimensional discrete Laplacian defined on the finite chain \mathcal{L}_v under boundary conditions τ_v , and E_v is the $L_v \times L_v$ unit matrix.

By using the results of [9, Chapter 7], we can write down the eigenfunctions of the interaction matrix (3.2) in the form

$$u_\Lambda^{(\tau)}(\mathbf{r}, \mathbf{k}) = u_{L_1}^{(\tau_1)}(r_1, k_1) \cdots u_{L_d}^{(\tau_d)}(r_d, k_d), \quad \mathbf{k} = (k_1, \dots, k_d) \in \Lambda, \quad (3.3)$$

and obtain the corresponding eigenvalues of it

$$\mu_\Lambda^{(\tau)}(\mathbf{k}) = 2 \sum_{v=1}^d \cos \varphi_{L_v}^{(\tau_v)}(k_v), \quad \mathbf{k} \in \Lambda. \quad (3.4)$$

Obviously, $\max_{\mathbf{k} \in \Lambda} \mu_\Lambda^{(\tau)}(\mathbf{k}) = 2d$. Note that the interaction Hamiltonian (3.1) has negative eigenvalues, which makes necessary the inclusion of a positive-definite quadratic form in the Gibbs exponent, to ensure the existence of the corresponding partition function. Thus, we consider the Hamiltonian

$$\beta \mathcal{H}_\Lambda^{(\tau)}(S_\Lambda | \beta, h_\Lambda; s) = -\frac{1}{2} \beta S_\Lambda^\dagger \cdot Q_\Lambda^{(\tau)} \cdot S_\Lambda + s S_\Lambda^\dagger \cdot S_\Lambda - h_\Lambda^\dagger \cdot S_\Lambda. \quad (3.5)$$

Here $h_\Lambda = \{h(\mathbf{r}), \mathbf{r} \in \Lambda\}$ is a column-vector representing (in units of $k_B T$) the inhomogeneous magnetic field configuration acting upon the system, and let h_Λ^\dagger be the transposed row-vector.

In order to ensure the existence of the partition function, all the eigenvalues $-\frac{1}{2} \beta \mu_\Lambda^{(\tau)}(\mathbf{k}) + s$, $\mathbf{k} \in \Lambda$, of the quadratic form in $\beta \mathcal{H}_\Lambda^{(\tau)}(S_\Lambda | \beta, h_\Lambda; s)$, ought to be positive. Hence, the field $s^{(\tau)}$ must satisfy the inequality

$$s > \frac{1}{2} \beta \max_{\mathbf{k} \in \Lambda} \mu_\Lambda^{(\tau)}(\mathbf{k}) \equiv \frac{1}{2} \beta \mu_\Lambda^{(\tau)}(\mathbf{k}_0), \quad (3.6)$$

with

$$\beta_{c,L} = \frac{1}{2} \mu_\Lambda^{(\tau)}(\mathbf{k}_0) \quad (3.7)$$

defining the critical temperature of the *finite* system. Since, as stated above $\max_{\mathbf{k} \in \Lambda} \mu_\Lambda^{(\tau)}(\mathbf{k}) = 2d$, it is clear that for the infinite system

$$\beta_c = d. \quad (3.8)$$

The free energy density of a *finite* system in a region Λ is

$$\beta f_\Lambda^{(\tau)}(\beta, h_\Lambda) = \frac{1}{2} \left\{ \ln(\beta/2\pi) - 2s + \mathcal{U}_\Lambda^{(\tau)}(\beta, s) - P_\Lambda^{(\tau)}(\beta, h_\Lambda, s) \right\}. \quad (3.9)$$

In Eq. (3.9) the first two terms do not depend on the size of the system, i.e., they are the same in both finite and infinite systems. The other two terms do depend, however on the size of the system. The function $\mathcal{U}_\Lambda^{(\tau)}(\beta, s)$ is due to the spin-spin interaction (and will be called "interaction term"); it depends on s , but does not depend on h . It is equal to

$$\mathcal{U}_\Lambda^{(\tau)}(\beta, s) = |\Lambda|^{-1} \sum_{\mathbf{k} \in \Lambda} \ln \left[\frac{2s}{\beta} - \mu_\Lambda^{(\tau)}(\mathbf{k}) \right], \quad (3.10)$$

and is obtained after performing the corresponding Gaussian integrals in the free energy of the finite system. The dependence of the free energy on the field variables h is given by the "field term"

$$P_\Lambda^{(\tau)}(\beta, h_\Lambda; s) = \frac{1}{\beta |\Lambda|} \sum_{\mathbf{k} \in \Lambda} \frac{|\hat{h}_\Lambda^{(\tau)}(\mathbf{k})|^2}{2s/\beta - \mu_\Lambda^{(\tau)}(\mathbf{k})}. \quad (3.11)$$

Here $\hat{h}_\Lambda^{(\tau)}(\mathbf{k})$ denotes the projection of the magnetic field configuration h_Λ on the eigenfunction $\{\bar{u}_\Lambda^{(\tau)}(\mathbf{r}, \mathbf{k}), \mathbf{k} \in \Lambda\}$ (by \bar{u} we denote the complex conjugate of $u \in \mathbb{C}$):

$$\hat{h}_\Lambda^{(\tau)}(\mathbf{k}) = \sum_{\mathbf{r} \in \Lambda} h(\mathbf{r}) \bar{u}_\Lambda^{(\tau)}(\mathbf{r}, \mathbf{k}). \quad (3.12)$$

Defining β_c so, that

$$\frac{2s}{\beta} = 2d \frac{\beta_c}{\beta}, \quad (3.13)$$

the above expressions can be rewritten in the form

$$\mathcal{U}_\Lambda^{(\tau)}(\beta) = |\Lambda|^{-1} \sum_{\mathbf{k} \in \Lambda} \ln \left[2d(\beta_c/\beta - 1) + 2d - \mu_\Lambda^{(\tau)}(\mathbf{k}) \right], \quad (3.14)$$

and

$$P_\Lambda^{(\tau)}(\beta, h_\Lambda) = \frac{1}{\beta |\Lambda|} \sum_{\mathbf{k} \in \Lambda} \frac{|\hat{h}_\Lambda^{(\tau)}(\mathbf{k})|^2}{2d(\beta_c/\beta - 1) + 2d - \mu_\Lambda^{(\tau)}(\mathbf{k})}. \quad (3.15)$$

Using the notations of [9, Chapter 7], below we give a list of the complete sets of orthonormal eigenfunctions, $\{u_L^{(\tau)}(r, k), k = 1, \dots, L\}$, of the one-dimensional discrete Laplacian under the Neumann - Dirichlet (ND) boundary conditions:

- periodic (p) boundary conditions

$$u_L^{(p)}(r, k) = L^{-1/2} \exp[-ir\varphi_L^{(p)}(k)]; \quad (3.16)$$

- Neumann - Dirichlet (ND) boundary conditions

$$u_L^{(\text{ND})}(r, k) = 2(2L+1)^{-1/2} \cos(r-1/2)\varphi_L^{(\text{ND})}(k). \quad (3.17)$$

The quantities $\varphi_L^{(\tau)}, k = 1, \dots, L$, are defined as follows

$$\varphi_L^{(p)}(k) = 2\pi k/L, \quad \varphi_L^{(\text{ND})}(k) = \pi(2k-1)/(2L+1). \quad (3.18)$$

Now we are ready to find the finite-size behavior of the Gaussian model under the Dirichlet-Neumann boundary conditions. According to Eq. (3.17), $S(0) = S(1)$, i.e., one has there realization of Neumann boundary conditions, while $L+1 = 0$, which corresponds to Dirichlet boundary

conditions. Thus, in the envisaged one-dimensional chain one has L independent spin variables $\{S(1), S(2), \dots, S(L)\}$.

We start with the consideration of $d = 3$ dimensional system. Note that:

- under fully periodic (p) boundary conditions, $\tau = (p, p, p)$, one has $\mathbf{k}_0 = (L_1, L_2, L_3)$, hence $\mu_{\Lambda}^{(p,p,p)}(\mathbf{k}_0) = 6$.
- under Neumann-Dirichlet boundary conditions along z direction, i.e., $\tau = (p, p, \text{ND})$, one has $\mathbf{k}_0 = (L_1, L_2, 1)$, hence $\mu_{\Lambda}^{(p,p,\text{ND})}(\mathbf{k}_0) = 4 + 2 \cos[\pi / (2L + 1)]$.

3.1. The Gaussian Model on a Lattice for the Case $d = 3$

We recall that for this model $\alpha = 1/2$, $\gamma = 1$ and $\nu = 1/2$ [9,14].

The behavior of the interaction term $\mathcal{U}_{\Lambda}^{(\tau)}(\beta)$

We set $\tau = (p, p, \text{ND})$ and use the short-hand notation $\tau = \text{ND}$ for these boundary conditions. Then, we perform in Eq. (3.10) the limits $L_1, L_2 \rightarrow \infty$, keeping $L_3 = L$ fixed. For the interaction term one then obtains

$$\mathcal{U}_{L,3}^{(\text{ND})}(\beta) = \lim_{L_1, L_2 \rightarrow \infty} \mathcal{U}_{\Lambda}^{(p,p,\text{ND})}(\beta) = \frac{1}{L} \sum_{k=1}^L \mathcal{V}_2 \left[6(\beta_c / \beta - 1) + 2 \left(1 - \cos \pi \frac{2k-1}{2L+1} \right) \right], \quad (3.19)$$

where

$$\mathcal{V}_d(z) := \frac{1}{(2\pi)^d} \int_{-\pi}^{\pi} d\theta_1 \cdots \int_{-\pi}^{\pi} d\theta_d \ln \left[z + 2 \sum_{\nu=1}^d (1 - \cos \theta_{\nu}) \right]. \quad (3.20)$$

The behavior of the interaction term in the bulk system

In accord with Eq. (3.20), one has

$$\mathcal{U}_{\infty,3}(\beta) = \mathcal{V}_3[6(\beta_c / \beta - 1)]. \quad (3.21)$$

The behavior of the interaction term in the film system with Neumann-Dirichlet boundary conditions

Explicitly, from Eq. (3.19) one obtains

$$\mathcal{U}_{L,3}^{(\text{ND})}(\beta) = \frac{1}{L} \sum_{k=1}^L \mathcal{V}_2 \left[6(\beta_c / \beta - 1) + 2 \left(1 - \cos \pi \frac{2k-1}{2L+1} \right) \right] = \frac{1}{(2\pi)^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 S^{(\text{ND})}(\beta, L | \theta_1, \theta_2), \quad (3.22)$$

with

$$S^{(\text{ND})}(\beta, L | \theta_1, \theta_2) = \frac{1}{L} \sum_{k=1}^L \ln \left[6(\beta_c / \beta - 1) + 2 \sum_{\nu=1}^2 (1 - \cos \theta_{\nu}) + 2 \left(1 - \cos \pi \frac{2k-1}{2L+1} \right) \right]. \quad (3.23)$$

This sum is of the form

$$S^{(\text{ND})}(x, L) = \frac{1}{L} \ln \prod_{k=0}^{L-1} 2 \left[\cosh(x) - \cos \pi \frac{2k+1}{2L+1} \right], \quad (3.24)$$

where $x = x(\beta | \theta_1, \theta_2)$ is defined as

$$\cosh x = 1 + 3(\beta_c / \beta - 1) + \sum_{\nu=1}^2 (1 - \cos \theta_{\nu}). \quad (3.25)$$

The summations in Eq. (3.24) can be performed using [12] the identity

$$\frac{\cosh[(L+1/2)x]}{\cosh(x/2)} = \prod_{k=0}^{L-1} 2 \left[\cosh x - \cos \pi \frac{2k+1}{2L+1} \right]. \quad (3.26)$$

With the help of the identity one derives

$$S^{(\text{ND})}(x, L) = \frac{1}{L} \ln \frac{\cosh[(L+1/2)x]}{\cosh(x/2)}. \quad (3.27)$$

Obviously $\lim_{L \rightarrow \infty} S^{(\text{ND})}(x, L) = x$. Thus, the part of the excess free energy under Neumann – Dirichlet boundary conditions that depends only on the interaction term is

$$\begin{aligned} \beta \Delta f_{\text{ex},3}^{(\text{ND})}(\beta, h=0) &= \frac{1}{2} L \left[\mathcal{U}_{L,3}^{(\text{ND})}(\beta) - \mathcal{U}_{\infty,3}(\beta) \right] \\ &= \frac{L}{8\pi^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 \left[\frac{1}{L} \ln \frac{\cosh[(L+1/2)x]}{\cosh(x/2)} - x \right] \\ &= \frac{1}{8\pi^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 \left[\ln \left(\frac{e^{-(2L+1)x} + 1}{e^{-x} + 1} \right) \right]. \end{aligned} \quad (3.28)$$

Thus, $\beta \Delta f_{\text{ex},3}^{(\text{ND})}(\beta, h=0)$ can be decomposed in the sum of $g_1(L, \phi)$ and $g_2(L, \phi)$ where

$$g_1(L, \beta) = \frac{1}{8\pi^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 \ln(e^{-(2L+1)x} + 1), \quad (3.29)$$

and

$$g_2(L, \beta) = -\frac{1}{8\pi^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 \ln(e^{-x} + 1). \quad (3.30)$$

Let us consider the behavior of g_1 and g_2 in the scaling regime

$$x_t = 6(\beta_c/\beta - 1)(2L+1)^2 = \mathcal{O}(1). \quad (3.31)$$

Let us first start with the function $g_1(L, \phi)$. Obviously, if $x = \mathcal{O}(1)$ then g_1 will be exponentially small. Thus, we need to consider the regime $(2L+1)x = \mathcal{O}(1)$. It follows that $x \ll 1$. From Eq. (3.25) we obtain

$$1 + \frac{1}{2}x^2 = 3(\beta_c/\beta - 1) + \frac{1}{2}(\theta_1^2 + \theta_2^2). \quad (3.32)$$

It follows that

$$x^2 = 6(\beta_c/\beta - 1) + (\theta_1^2 + \theta_2^2) = 6(\beta_c/\beta - 1) + r^2, \quad (3.33)$$

where we have introduced polar coordinates. In terms of them $g_1(L, \beta)$ becomes

$$\begin{aligned} g_1(L, \beta) &\simeq \frac{1}{4\pi} \int_0^R \ln(e^{-(2L+1)x} + 1) dr^2 \simeq \frac{1}{4\pi} \int_{\sqrt{6(\beta_c/\beta-1)}}^{\infty} \ln(e^{-(2L+1)x} + 1) dx^2 \\ &= -\frac{1}{4\pi} \frac{\sqrt{x_t} \text{Li}_2(-e^{-\sqrt{x_t}}) + \text{Li}_3(-e^{-\sqrt{x_t}})}{(2L+1)^2}, \end{aligned} \quad (3.34)$$

where R can be defined from the constraint $(2\pi) \times (2\pi) = 4\pi^2 = \pi R^2$, i.e., $R = 2\sqrt{\pi}$.

Next, we deal with $g_2(L, \phi)$. Taking into account that x_L is small we derive

$$\begin{aligned} g_2(L, \beta) &= -\frac{1}{8\pi^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 \ln(e^{-x} + 1) \simeq -\frac{1}{8\pi^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 \left[\ln 2 - \frac{1}{2}x \right] \\ &= -\frac{1}{2} \ln 2 + \frac{1}{16\pi^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 x \simeq -\frac{1}{2} \ln 2 + \frac{1}{8\pi} \int_{\sqrt{6(\beta_c/\beta-1)}}^R x dx^2 \\ &= -\frac{1}{2} \ln 2 + \frac{1}{12\pi} \left\{ R^3 - \left(\frac{\sqrt{x_t}}{2L+1} \right)^3 \right\}. \end{aligned} \quad (3.35)$$

Note that for $x_t = \mathcal{O}(1)$ one has that for the L -dependent part $\Delta g_2(L, \beta)$ of g_2 one has $\Delta g_2(L, \phi) \propto L^{-3}$, i.e., Δg_2 is one order of magnitude *smaller* than g_1 . Because of that, g_2 contributes only sub-leading contributions to L -dependent part of the excess free energy and, therefore, to the Casimir force. Based on the above, we will no longer be interested in the function g_2 .

Summarizing the above, we conclude that the excess free energy can be written in a scaling form

$$\beta \Delta f_{\text{ex},3}^{(\text{ND})}(\beta, h = 0) = -\frac{1}{L^2} X_{\text{ex}}(a_{\beta} t L^{1/\nu}) \quad (3.36)$$

where a_{β} is a non-universal constants, and X_{ex} is an universal scaling function, $t = (T - T_c)/T_c$, where T has the meaning of the temperature of the system, and T_c is its bulk temperature. From Eq. (3.34), taking into account that with $\nu = 1/2$ one has $(2L+1)^2 \simeq 4L^2 \simeq L^{1/\nu}$, we identify that

$$X_{\text{ex}}(x_t) = \frac{1}{16\pi} \left[\sqrt{x_t} \text{Li}_2(-e^{-\sqrt{x_t}}) + \text{Li}_3(-e^{-\sqrt{x_t}}) \right]. \quad (3.37)$$

The behavior of the field term $P_{\Lambda}^{(\tau)}(\beta, h_{\Lambda})$

The dependence of the free energy on the field variable is given by the "field term", given by Eq. (3.15). For a homogeneous field h and for (per) $\equiv (p, p, p)$ and ND $\equiv (p, p, \text{ND})$ boundary conditions, it is easy to obtain that

- for (p, p, p) boundary conditions

$$\hat{h}_{\Lambda}^{(\text{per})}(\mathbf{k}) = \sum_{\mathbf{r} \in \Lambda} h(\mathbf{r}) \bar{u}_{\Lambda}^{(\text{per})}(\mathbf{r}, \mathbf{k}) = \sqrt{L_1 L_2 L_3} \delta_{k_1,0} \delta_{k_2,0} \delta_{k_3,0} h, \quad (3.38)$$

and

$$P_L^{(\text{per})}(K, h; \phi) = \frac{h^2}{6\beta(\beta_c/\beta - 1)}. \quad (3.39)$$

Obviously

$$P_{\infty}(\beta, h) = \lim_{L \rightarrow \infty} P_L^{(\text{per})}(K, h; \phi) = \frac{h^2}{6\beta(\beta_c/\beta - 1)}. \quad (3.40)$$

- for (p, p, ND) boundary conditions

$$\hat{h}_{\Lambda}^{(\text{ND})}(\mathbf{k}) = \sum_{\mathbf{r} \in \Lambda} h(\mathbf{r}) \bar{u}_{\Lambda}^{(\text{ND})}(\mathbf{r}, \mathbf{k}) = 2\sqrt{\frac{L_1 L_2}{2L_3 + 1}} \delta_{k_1,0} \delta_{k_2,0} h \sum_{r=1}^{L_3} \cos\left[(r-1/2) \varphi_{L_3}^{(\text{ND})}(k_3)\right], \varphi_{L_3}^{(\text{ND})}(k_3) = \pi \frac{2k_3 - 1}{2L_3 + 1}. \quad (3.41)$$

Thus, setting $k_3 = k, r_3 = r$ and $L_3 = L$, for a film geometry we arrive at

$$P_L^{(\text{ND})}(\beta, h) = \frac{4h^2}{\beta L(2L+1)} \sum_{k=1}^L \frac{\left| \sum_{r=1}^L \cos\left[(r-1/2) \pi \frac{2k-1}{2L+1}\right] \right|^2}{6(\beta_c/\beta - 1) + 2\left(1 - \cos \pi \frac{2k-1}{2L+1}\right)}. \quad (3.42)$$

It is easy to show that

$$2 \sum_{r=1}^L \cos\left(\pi(r-1/2) \frac{(2k-1)}{(2L+1)}\right) = \frac{\sin\left(\pi \frac{2k-1}{2L+1} L\right)}{\sin\left(\pi \frac{2k-1}{2L+1}\right)}. \quad (3.43)$$

Thus, one has

$$P_L^{(\text{ND})}(\beta, h) = \frac{h^2}{\beta L(2L+1)} \sum_{k=1}^L \frac{\cot^2\left(\pi \frac{2k-1}{2L+1}\right)}{6(\beta_c/\beta - 1) + 2\left(1 - \cos \pi \frac{2k-1}{2L+1}\right)}. \quad (3.44)$$

Let us consider the small k behavior of the above sum. One derives

$$\begin{aligned} P_L^{(\text{ND})}(\beta, h) &\simeq \frac{h^2}{\beta} \frac{1}{L(2L+1)} \sum_{k=1}^L \frac{1}{\left[\frac{\pi(2k-1)}{2(2L+1)}\right]^2 \left\{6(\beta_c/\beta - 1) + \left[\frac{\pi(2k-1)}{2L+1}\right]^2\right\}} \\ &\simeq \frac{4}{\pi^2} \frac{h^2}{\beta} \frac{(2L+1)^3}{L} \sum_{k=1}^L \frac{1}{(2k-1)^2 [x_t + \pi^2(2k-1)^2]} \\ &= \frac{h^2}{\beta} \frac{(2L+1)^3}{L} \left\{ \frac{1}{2x_t} \left[1 - \frac{\tanh(\sqrt{x_t}/2)}{\sqrt{x_t}/2}\right] + \mathcal{O}(L^{-3}) \right\}. \end{aligned} \quad (3.45)$$

In the limits $x_t \rightarrow 0$ and $x_t \rightarrow \infty$ for the behavior of the field term one obtains

$$P_L^{(\text{ND})}(\beta, h) \simeq \frac{h^2}{\beta} \frac{(2L+1)^3}{L} \begin{cases} 1/24 + \mathcal{O}(x_t), & x_t \rightarrow 0; \\ 1/(2x_t) + \mathcal{O}[\exp(-\sqrt{x_t})], & x_t \gg 1. \end{cases} \quad (3.46)$$

When $L \rightarrow \infty$, then $x_t \rightarrow \infty$, we obtain that

$$\lim_{L \rightarrow \infty} P_L^{(\text{ND})}(\beta, h) = \frac{h^2}{6\beta(\beta_c/\beta - 1)}, \quad (3.47)$$

which indeed equals the bulk expression - see Eq. (3.40).

From Eq. (3.45) for the behavior of the susceptibility in the finite system we derive

$$\chi_L^{(\text{ND})}(\beta, h) = \frac{1}{\beta} \frac{(2L+1)^3}{L} \frac{1}{x_t} \left[1 - \frac{\tanh(\sqrt{x_t}/2)}{\sqrt{x_t}/2}\right]. \quad (3.48)$$

According to the finite-size scaling theory [9,15]

$$\chi_L^{(\zeta)}(t) = a_h L^\gamma X_\chi(a_\beta t L^{1/\nu}), \quad (3.49)$$

where a_h and a_β are non-universal constants, and X_χ is an universal scaling function, $t = (T - T_c)/T_c$, where T has the meaning of the temperature of the system, and T_c is its bulk temperature. From Eq. (3.48), taking into account that $(2L+1)^3/L \simeq 8L^2$, we identify that

$$\gamma = 2, \quad \nu = 1/2, \quad \text{and} \quad tL^2 = x_t. \quad (3.50)$$

It is clear that the field term in the free energy of the finite system will be of the same order as the field term, i.e., $\propto L^{-3}$ if $h \propto L^{-5/2}$. In order to achieve that, we define a field dependent scaling variable

$$x_h = \beta^{-1/2}(2L+1)^{3/2}L h. \quad (3.51)$$

In terms of it, Eq. (3.45) becomes

$$P_L^{(\text{ND})}(x_h, x_t) = \frac{x_h^2}{L^3} X_\chi(x_t), \text{ where } X_\chi(x_t) = \frac{1}{2x_t} \left[1 - \frac{\tanh(\sqrt{x_t}/2)}{\sqrt{x_t}/2} \right]. \quad (3.52)$$

The behavior of the scaling function $X_\chi(x_t)$ is given in Figure 7.

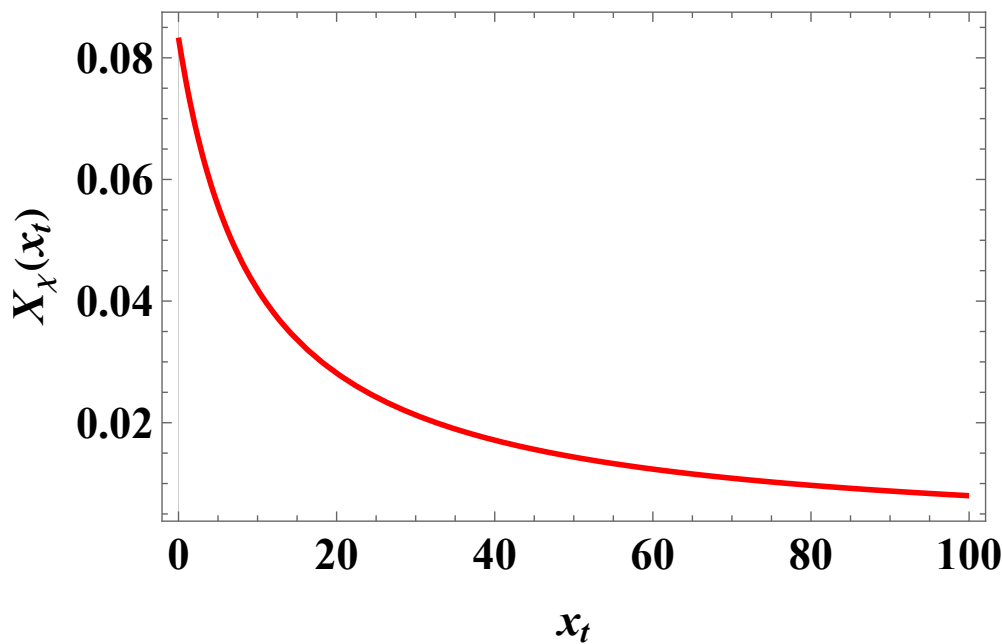


Figure 7. The behavior of the scaling function $X_\chi(x_t)$.

Then for the excess free energy related to the field term, see Eq. (3.9), one derives

$$\beta \Delta f_{\text{ex},3}^{(\text{ND})}(\beta, h) = -\frac{1}{2}L \left[P_L^{(\text{ND})}(h; \beta) - P_\infty(h; \beta) \right] = \frac{x_h^2}{L^2} \frac{\tanh(\sqrt{x_t}/2)}{2x_t^{3/2}}. \quad (3.53)$$

3.2. The Behavior of the Casimir Force

Let us determine the contributions of the interaction term $\Delta F_{\text{Cas}}^{(\text{ND})}(\beta, h = 0)$ and of the field term $\Delta F_{\text{Cas}}^{(\text{ND})}(\beta, h \neq 0)$. Obviously, one has

$$\Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h) = \Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h = 0) + \Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h \neq 0). \quad (3.54)$$

We start with determining the behavior of $\Delta F_{\text{Cas}}^{(\text{ND})}(\beta, h = 0)$. By definition, it is equal to

$$\Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h = 0) \equiv -\frac{\partial}{\partial L} \beta \Delta f_{\text{ex},3}^{(\text{ND})}(\beta, h = 0). \quad (3.55)$$

From Eq. (3.28) we derive the *exact* expression

$$\beta \Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h = 0) = \frac{1}{4\pi^2} \int_{-\pi}^{\pi} d\theta_1 \int_{-\pi}^{\pi} d\theta_2 \frac{x}{e^{(2L+1)x} + 1}. \quad (3.56)$$

Here we did not make any assumption about L . Naturally, we will obtain a scaling form of $\beta\Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h=0)$ only for $L \gg 1$. Then Eq. (3.32) is valid and, after performing the integration, we arrive at

$$\begin{aligned}\beta\Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h=0) &= -\frac{1}{(2(L+1)^3)} \frac{1}{\pi} \left\{ \text{Li}_3(-e^{-\sqrt{x_t}}) + \sqrt{x_t} \text{Li}_2(-e^{-\sqrt{x_t}}) - \frac{1}{2} x_t \log(e^{-\sqrt{x_t}} + 1) \right\} \\ &= \frac{1}{(L+1/2)^3} X_{\text{Cas},3}(x_t),\end{aligned}\quad (3.57)$$

where

$$X_{\text{Cas},3}(y) = -\frac{1}{8\pi} \left\{ \text{Li}_3(-e^{-\sqrt{x_t}}) + \sqrt{x_t} \text{Li}_2(-e^{-\sqrt{x_t}}) - \frac{1}{2} x_t \log(e^{-\sqrt{x_t}} + 1) \right\}. \quad (3.58)$$

The behavior of the scaling function $X_{\text{Cas}}(x_t, h=0)$ is given in Figure 8. Obviously, the function is positive, which means that the Casimir force is repulsive when the external field is zero. For the Casimir amplitude we obtain

$$\Delta_{\text{Cas},3}^{(\text{ND})} \equiv X_{\text{Cas},3}(x_t=0, h=0)/2 = \frac{3}{64\pi} \zeta(3). \quad (3.59)$$

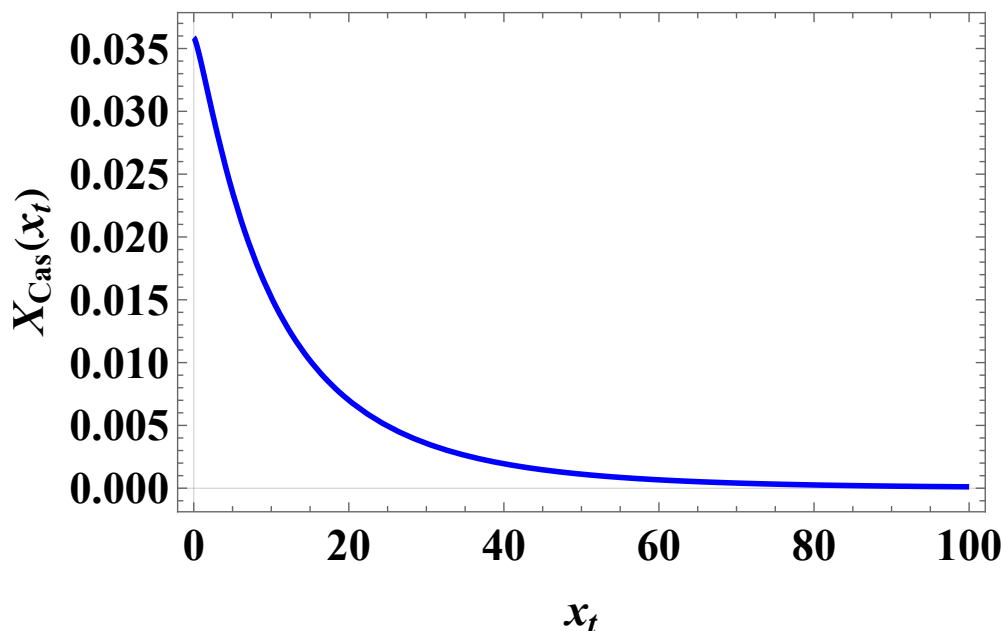


Figure 8. The behavior of the scaling function $X_{\text{Cas}}(x_t)$ when $h=0$.

Obviously, Eq. (3.59) coincides with the corresponding result for the Gaussian model obtained via studying the $O(n)$, $n=1, d=3$ Φ^4 model — see [4, Eq. (6.99)]. Analogically, after proper renaming of the scaling variable the expression Eq. (3.58) of the scaling function of the force coincides with the corresponding one for the $O(n)$, $n=1, d=3$ Φ^4 model — see [4, Eq. (6.104)].

Let us now determine the h -dependent part of the Casimir force. By definition, one has

$$\Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h) \equiv -\frac{\partial}{\partial L} \beta \Delta f_{\text{ex},3}^{(\text{ND})}(\beta, h). \quad (3.60)$$

Then, from Eq. (3.53) one obtains

$$\Delta F_{\text{Cas},3}^{(\text{ND})}(\beta, h) = -\frac{\partial}{\partial L} \left[\frac{x_h^2 \tanh(\sqrt{x_t}/2)}{L^2 2x_t^{3/2}} \right] \simeq -\frac{x_h^2}{2L^2(1+2L)} \left[\frac{\text{sech}^2(\sqrt{x_t}/2)}{x_t} \right] \quad (3.61)$$

$$= \frac{1}{L^2(L+1/2)} X_{\text{Cas},3}(x_t, x_h), \quad (3.62)$$

where

$$X_{\text{Cas},3}(x_t, x_h) = -\frac{x_h^2}{4} \left[\frac{\text{sech}^2(\sqrt{x_t}/2)}{x_t} \right] < 0. \quad (3.63)$$

A visualization of $X_{\text{Cas},3}(y, x_h)$ as a function of y for $x_h = 1$ is shown in Figure 9.

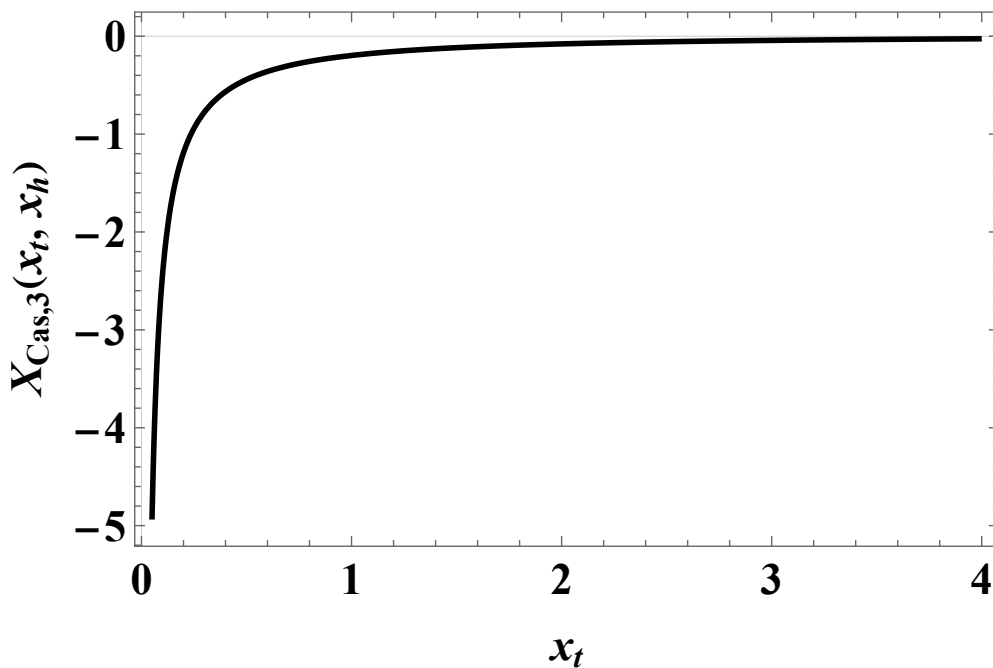


Figure 9. The behavior of the scaling function $X_{\text{Cas}}(x_t, x_h = 1)$. We observe that the force is *attractive*.

The total Casimir force is a sum of $X_{\text{Cas},3}(x_t)$, see Eq. (3.58), and $X_{\text{Cas},3}(x_t, x_h)$ given by Eq. (3.63). The plot of the result as a function of x_t for $x_h = 0.05$ is shown in Figure 10. As we see, the force can be both positive and negative, i.e., *repulsive* and *attractive*.

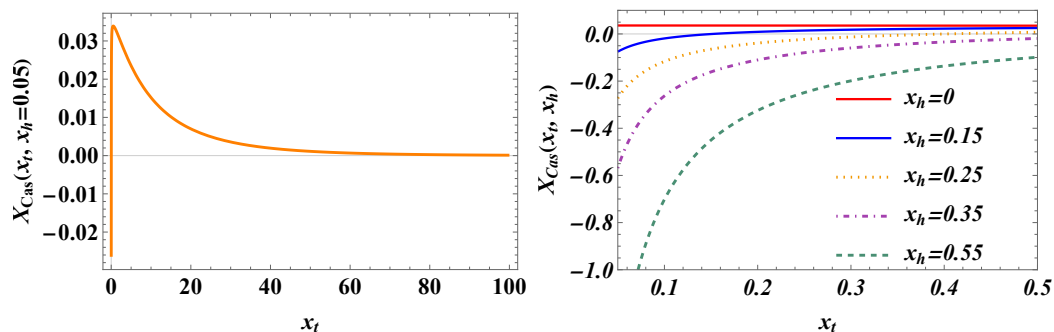


Figure 10. The behavior of the scaling function of the total Casimir force as a function of y for several values of x_h . **Left panel:** We see that for $x_h = 0.05$ the force is attractive very near the critical temperature, then becomes repulsive with increase of x_t (i.e., of T). **Right panel:** It is clear, that for zero field the force is repulsive, then — for small values of x_h — the force changes sign from attractive to repulsive with the increase of x_t (i.e., of the temperature), while for large values of x_h the force becomes attractive for all values of T (i.e., x_t).

The overall 3D behavior of the force as a function both on x_t and x_h is given in Figure 11.

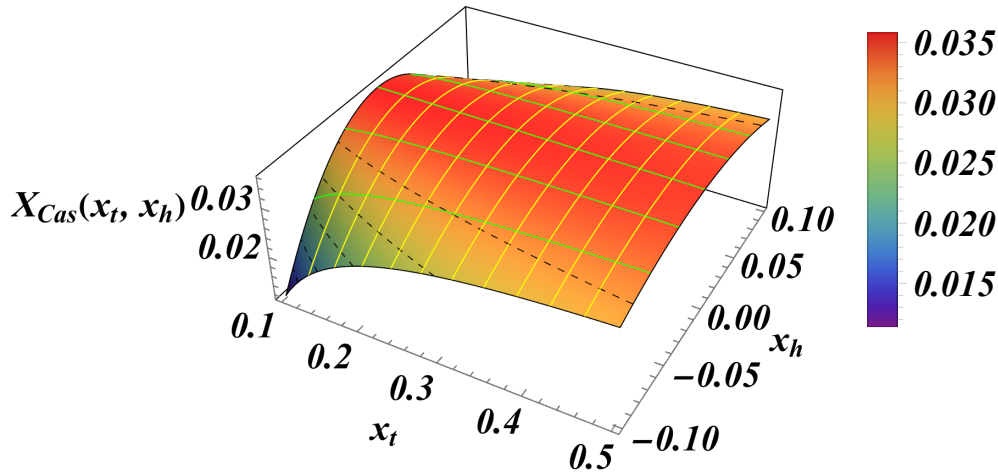


Figure 11. The behavior of the scaling function $X_{Cas}(x_t, x_h)$. Here $x_t \in [0.1, 0.5]$ and $x_h \in [-0.1, 0.1]$.

4. The Casimir Force Within the Mean-Field Model

We start by defining the mean-field model used in the current study.

4.1. The Ginzburg-Landau Functional

In the present work we consider the standard ϕ^4 Ginzburg-Landau functional

$$\mathcal{F}[\phi|\tau, h, L] = \int_0^L \mathcal{L}(\phi, \phi'|\tau, h) dz \quad (4.1)$$

with

$$\mathcal{L}(\phi, \phi'|\tau, h) = \frac{1}{2}\phi'^2 + \frac{1}{2}\tau\phi^2 + \frac{1}{4}g\phi^4 - h\phi. \quad (4.2)$$

Here $L, g \in \mathbb{R}^+$, while $\tau, h \in \mathbb{R}$, $z \in (0, L)$ and $\phi = \phi(z)$ are the independent and dependent variables, respectively, and the prime indicates differentiation with respect to the variable z .

The functional (4.1) describes a critical system of Ising type in a film geometry $\infty^2 \times L$, where the film thickness L is supposed to be along the z axis. In Equation (4.1), $\phi(z|\tau, h, L)$ is the order parameter of the system, which is assumed to depend on the perpendicular position $z \in (0, L)$ only, g is the bare coupling constant, and, as before, $\tau = (T - T_c)/T_c$ is the bare reduced temperature and h is the external ordering field. Given τ, h and L , the physical state of the regarded system is described by the minimizer of the respective Ginzburg-Landau functional $\mathcal{F}[\phi; \tau, h, L]$ given above whose extremals are determined by the solutions of the corresponding Euler-Lagrange equation

$$\frac{d}{dz} \frac{\partial \mathcal{L}}{\partial \phi'} - \frac{\partial \mathcal{L}}{\partial \phi} = 0. \quad (4.3)$$

In case the Lagrangian density \mathcal{L} is defined by Equation (4.2), Equation (4.3) reads

$$\phi'' - \phi[\tau + g\phi^2] + h = 0. \quad (4.4)$$

Multiplying Equation (4.4) by ϕ' and integrating once over z one obtains that

$$P[\phi] \equiv \frac{1}{2}\phi'^2 - \frac{1}{2}\tau\phi^2 - \frac{1}{4}g\phi^4 + h\phi \quad (4.5)$$

is a first integral of Equation (4.4), cf., e.g., [4,16,17]. This means that P is a constant on any smooth solution $\phi(z|\tau, h, L)$ of the Euler-Lagrange equation (4.4).

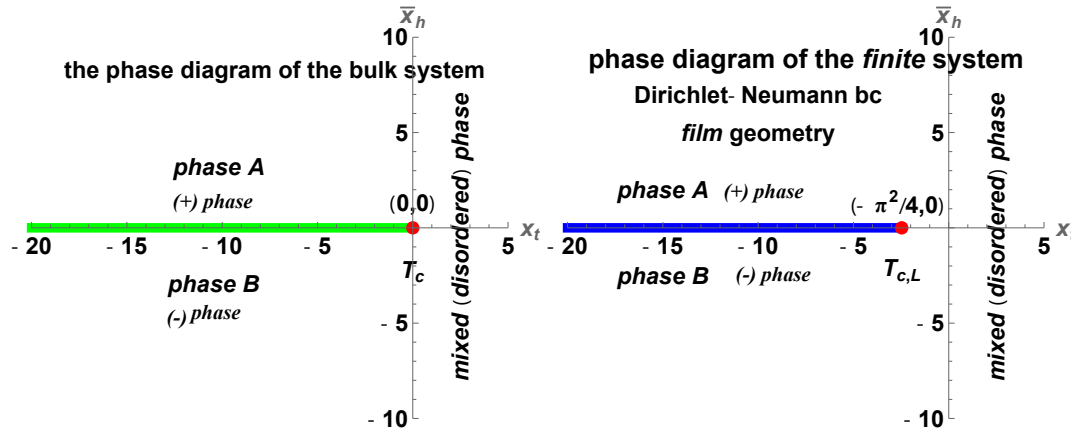


Figure 12. Phase diagrams. **Left panel:** The phase diagram of the bulk system. **Right panel:** The phase diagram of the finite system with Dirichlet-Neumann boundary conditions. In the bulk system a phase transition of first order happens when crossing the phase coexistence line that is at $\bar{x}_h = 0$ and spans for $T \in (0, T = T_c)$. At $T = T_c$ the system exhibits a second order phase transition. In the finite system the coexistence line is at $\bar{x}_h = 0$ and spans for $T \in (0, T = T_{c,L})$. The second order phase transition happens at $T = T_{c,L} \equiv (-\pi^2/4, 0)$. Note the change with Dirichlet-Dirichlet boundary conditions where the critical point is at $T_{c,L} = (-\pi^2, 0)$.

In general, the thermodynamic Casimir force $F_{\text{Cas}}(\tau, h, L)$ in such a system is the *excess pressure*, over the bulk one, acting on the boundaries of the finite system, which is due to the finite size of that system, i.e.,

$$F_{\text{Cas}}(\tau, h, L) = P_L(\tau, h) - P_b(\tau, h). \quad (4.6)$$

Here P_L is the pressure in the finite system, while P_b is that one in the infinite system.

Now, assuming that the thickness L of the film is free to move, the variation $\delta\mathcal{F}$ of the free energy $\mathcal{F}[\phi|\tau, h, L]$ of the finite system is given as follows

$$\delta\mathcal{F} = \int_0^L \left(\frac{\partial\mathcal{L}}{\partial\phi} - \frac{d}{dz} \frac{\partial\mathcal{L}}{\partial\phi'} \right) \delta\phi dz + \mathcal{L}_{\phi'} \delta\phi \Big|_0^L - \left(\phi' \mathcal{L}_{\phi'} - \mathcal{L} \right) \delta z \Big|_0^L \quad (4.7)$$

(see, e.g., [18, p. 54], [19, p. 260] and [20]), where δz and $\delta\phi$ are the variations of the independent and dependent variables, while

$$T_{zz} = \phi' \mathcal{L}_{\phi'} - \mathcal{L} \quad (4.8)$$

is the one-dimensional counterpart of the stress tensor (see, e.g., [21,22]). Relation (4.7) estimates the change of the finite-size contribution to the free energy of the system corresponding to a small variation of the variables including the variation of the film thickness L . In this sense, $T_{zz}|_L$ can be interpreted (see, e.g., [17,23]) as the pressure in the finite system, that is

$$P_L = T_{zz}|_L. \quad (4.9)$$

On the other hand, taking into account Equations (4.2) and (4.8), one can see that

$$T_{zz} = P[\phi] \quad (4.10)$$

and hence T_{zz} is a constant on any smooth solution $\phi(z|\tau, h, L)$ of the Euler-Lagrange equation (4.4) including the minimizer of the Ginzburg–Landau functional (4.1). Thus, the pressure in the finite system is

$$P_L(\tau, h) = \frac{1}{2}\phi_{min}'^2 - \frac{1}{4}g\phi_{min}^4 - \frac{1}{2}\tau\phi_{min}^2 + h\phi_{min} \quad (4.11)$$

where ϕ_{min} is the foregoing minimizer.

As for the bulk system, it is easy to see following the same way of reasoning that the corresponding pressure is

$$P_b(\tau, h) = -\frac{1}{4}g\phi_b^4 - \frac{1}{2}\tau\phi_b^2 + h\phi_b. \quad (4.12)$$

Here, the value ϕ_b of the order parameter of the bulk system is determined as the constant solution of Equation (4.4), i.e., as the root of the cubic equation

$$-\phi_b \left[\tau + g\phi_b^2 \right] + h = 0, \quad (4.13)$$

that minimizes

$$\mathcal{L}_b = \frac{1}{2}\tau\phi_b^2 + \frac{1}{4}g\phi_b^4 - h\phi_b. \quad (4.14)$$

Of course, ϕ_b does not depend on the boundary conditions at all. Let us note that $P_b = -\mathcal{L}_b$, i.e., P_b has its *maximum* over the solution ϕ_b of the cubic equation for (4.13).

Obviously, the relation (4.11) does *not* depend on the boundary conditions applied on the finite system too. This dependence arises solely from the dependency of the order parameter profile that minimizes the particular boundary value problem considered.

In the light of the above it is evident that once the order parameter profile ϕ_{min} and its bulk value ϕ_b are known in analytic form for given values of the parameters τ and h , then the respective Casimir force is determined in an exact manner by Equation (4.6).

In the current article we consider the Dirichlet-Neumann boundary conditions meaning that

$$\phi(z=0|\tau, h, L) = 0 \quad \text{and} \quad \frac{\partial}{\partial z}\phi(z|\tau, h, L)\Big|_{z=1} = 0. \quad (4.15)$$

In addition, ν is a critical exponent characterizing the behavior of the correlation length, while Δ is another exponent related to the behavior of, say, order parameter as a function of the external field h .

It is convenient to introduce new parameters

$$x_t = \frac{\tau L^{1/\nu}}{[\xi_0^+]^{1/\nu}}, \quad x_h = \frac{\sqrt{2g}hL^{\Delta/\nu}}{[\xi_{0,h}]^{\Delta/\nu}} \quad (4.16)$$

and variables

$$\zeta = z/L, \quad \phi(z) = \sqrt{\frac{2}{g}} L^{-\beta/\nu} X_m(\zeta|x_t, x_h), \quad (4.17)$$

where $\beta = \nu = 1/2$ and $\Delta = 3/2$, while ξ_0^+ and $\xi_{0,h}$ are the respective amplitudes of the correlation length along the τ and h axes (see, e.g., [22,24,25]). In terms of these new parameters and variables, Equations (4.1), (4.2), (4.4) and (4.5) becomes

$$\mathcal{F}[X_m|x_t, x_h] = \frac{1}{gL^4} \int_0^1 \mathcal{L}[X_m, X_m'|x_t, x_h] d\zeta, \quad (4.18)$$

$$\mathcal{L}[X_m, X_m'|x_t, x_h] = X_m'^2(\zeta) + X_m^4(\zeta) + x_t X_m^2(\zeta) - x_h X_m(\zeta), \quad (4.19)$$

$$X_m''(\zeta) = X_m(\zeta) \left[x_t + 2X_m^2(\zeta) \right] - \frac{x_h}{2}, \quad (4.20)$$

and

$$P[X_m(\zeta)] = X_m'^2(\zeta) - X_m^4(\zeta) - x_t X_m^2(\zeta) + x_h X_m(\zeta), \quad (4.21)$$

respectively. The primes here and hereafter indicate differentiation with respect to the variable $\zeta \in [0, 1]$. Then, according to Equations (4.6), (4.11) and (4.12), the expression for the Casimir force $X_{\text{Cas}}(x_t, x_h)$ written by means of the new parameters (4.16) and variables (4.17) reads

$$X_{\text{Cas}}(x_t, x_h) = \hat{X}_m'^2 - (\hat{X}_m^4 - X_b^4) - x_t(\hat{X}_m^2 - X_b^2) + x_h(\hat{X}_m - X_b), \quad (4.22)$$

where \hat{X}_m and X_b are the minimizers of the functional (4.18) and its “bulk counterpart” corresponding to x_t and x_h .

As mentioned above, in the present article we assume that the system is subject to Dirichlet-Neumann boundary conditions, that is

$$X_m(\zeta = 0|x_t, x_h) = 0 \quad \text{and} \quad X_m'(\zeta = 1|x_t, x_h) = 0. \quad (4.23)$$

In other words, we are interested in the solution of Equation (4.20) that meet the conditions (4.23). It should be remarked that exact results associated with the Casimir effect have been derived in Refs. [17,23,26–30] in the cases of $(+, +)$, $(+, -)$ and Dirichlet-Dirichlet boundary conditions.

4.2. The Casimir Force for Zero External Field

In Ref. [31] it has been shown that for $x_t \in (-\infty, -\pi^2/4]$ there are two order parameter profiles that minimize the functional (4.18) in the case of Dirichlet-Neumann boundary conditions and zero external field. They can be expressed using an auxiliary parameter $k \in [0, 1]$ as follows

$$\hat{X}_m(\zeta) = \pm k K(k) \text{sn}(\zeta K(k)|k) \quad (4.24)$$

at

$$x_t = -(k^2 + 1)K(k)^2, \quad (4.25)$$

where $K(\cdot)$ is the complete elliptic integral of the first kind and $\text{sn}(\cdot|\cdot)$ is the sine Jacobi elliptic function. Simultaneously, it is easy to see that in this case

$$X_b = \frac{1}{\sqrt{2}} \sqrt{(k^2 + 1)K(k)^2}. \quad (4.26)$$

Now, substituting Equations (4.24) and (4.26) into Equation (4.22) one obtains

$$X_{\text{Cas}}(x_t, x_h = 0) = -\frac{1}{4}(k^2 - 1)^2 K(k)^4, \quad (4.27)$$

for the Casimir force at $x_t \in (-\infty, -\pi^2/4]$ given by Equation (4.25).

If $x_t \in (-\pi^2/4, 0]$, then $\hat{X}_m = 0$, $X_b = -\sqrt{-x_t/2}$ and hence, according to Equation (4.22), the expression for the Casimir force reads

$$X_{\text{Cas}}(x_t, x_h = 0) = -\frac{x_t^2}{4}. \quad (4.28)$$

Finally, if $x_t \in (0, \infty)$, then $X_{\text{Cas}}(x_t, x_h = 0) = 0$. Combining these results one can write down

$$X_{\text{Cas}}(x_t, x_h = 0) = \begin{cases} -\frac{1}{4}(k^2 - 1)^2 K(k)^4, & x_t \in (-\infty, -\frac{\pi^2}{4}], \\ -x_t^2/4, & x_t \in (-\frac{\pi^2}{4}, 0], \\ 0, & x_t \in (0, \infty). \end{cases} \quad (4.29)$$

The behavior of the scaling function $X_{\text{Cas}}(x_t, x_h = 0)$ for $x_t \in [-30, 30]$ is depicted in Figure (13)

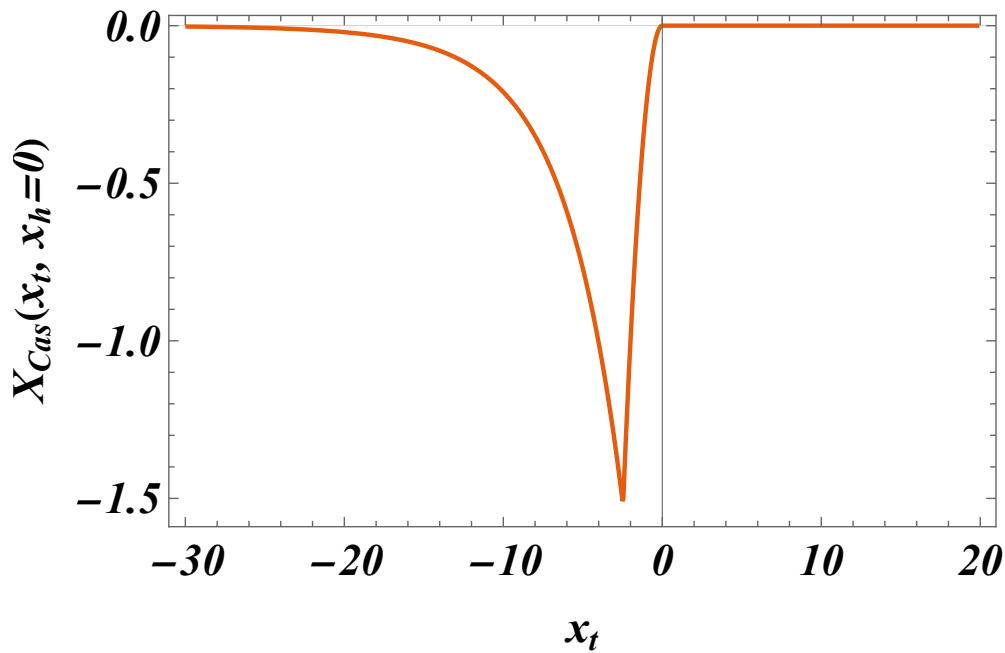


Figure 13. The behavior of the scaling function $X_{\text{Cas}}(x_t, x_h = 0)$ for $x_t \in [-30, 30]$. We observe that the force is *attractive*, contrary to the corresponding result for the Gaussian model.

4.3. The Casimir Force for Nonzero External Field

In Ref. [31] it has been shown, following [32, p. 454], that each solution of Equation (4.20) that meets Dirichlet-Neumann boundary conditions can be written in the form

$$X_m(\zeta|x_t, x_h, X_{m,r}) = X_{m,r} + \frac{6X_{m,r}(x_t + 2X_{m,r}^2) - 3x_h}{12\wp(\zeta - 1; g_2, g_3) - (x_t + 6X_{m,r}^2)}, \quad (4.30)$$

where $X_{m,r} = X_{m,r}(x_t, x_h)$ is a real number that depends only on the values of the parameters x_t and x_h . Here $\wp(v; g_2, g_3)$ is the Weierstrass elliptic function corresponding to the invariants g_2 and g_3 given as follows

$$\begin{aligned} g_2 &= \frac{1}{12}x_t^2 - X_{m,r}(X_{m,r}^3 + x_t X_{m,r} - x_h), \\ g_3 &= -\frac{1}{432} \left[27x_h^2 + 2x_t^3 + 72x_t X_{m,r}(X_{m,r}^3 + x_t X_{m,r} - x_h) \right]. \end{aligned} \quad (4.31)$$

It is easy to see that $X_{m,r}$ is the value of the order parameter at the right end of the system since $\wp(v; g_2, g_3)$ tends to infinity when v tends to zero. It is also easy to see that $X'_m(\zeta \rightarrow 1|x_t, x_h, X_{m,r}) = 0$, i.e., each function of the form (4.30) meets the boundary condition imposed on the right end of the system. The only remaining requirement that $X_m(\zeta \rightarrow 0|x_t, x_h, X_{m,r}) = 0$ leads to a transcendental equation from where we have to determine $X_{m,r}$. Usually, one obtains several solution of this equation. However, the one that corresponds to the physical reality is the one that minimizes the energy given by Equations (4.18) and (4.19). In this way we find the order parameter profile \hat{X}_m as a function of the parameters x_t and x_h . We also obtain X_b as a function of x_t and x_h

Finally, using Equation (4.22) we obtain the Casimir force $X_{\text{Cas}}(x_t, x_h)$.

5. Conclusion

As reported above, we have obtained exact results for the Casimir force in two basic statistical mechanical models: the Gaussian and the mean field model. In the case of the Gaussian model we performed the calculations for two realizations: a continuum version, see Sec. 2, and a lattice version see Sec. 3 realizations. The mean-field model is considered in Sec. 4. The models are considered under Neumann-Dirichlet boundary conditions in the presence of an external magnetic field h .

We summarize our main results as follows:

- (I) We derived exact closed form expression for the free energy of the Gaussian model in both the continuum version (CGM) and the lattice formulation of the model (LGM). The results for the Casimir force can be written as a sum of
 - i) expressions pertinent to the $h = 0$ case - see Eq. (2.30) for CGM and Eq. (3.58) for the LGM.
 - ii) equations for the field-dependent parts of the force — see Eq. (2.38) for the CGM, and Eq. (3.63) for the LGM.

We observe that these expression are identical, as is to be expected on the ground of the universality hypothesis, provided proper definitions of the scaling variables are used.

- (II) The behavior of the Casimir force in the CGM is shown in Figures 3 and 5, and the behavior for the LGM - in Figures 8 – 11. We observe that for $h = 0$ the force is repulsive and, depending on magnitude of h , it can be both repulsive or attractive for $h \neq 0$. Contrary to this behavior, we observe that the force in the MFM is *always* attractive - both for $h = 0$, see Figure 13, as well as for $h \neq 0$ – see Figure 14 and 15.

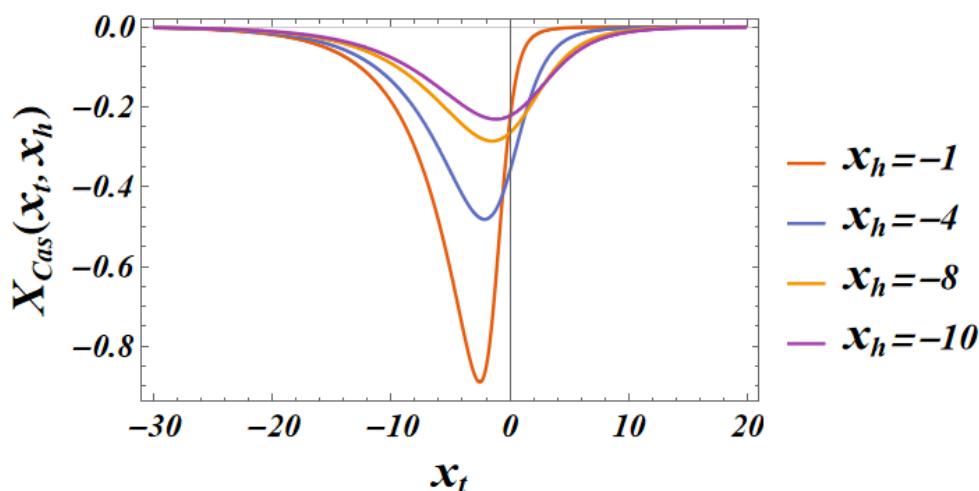


Figure 14. The behavior of the scaling function $X_{\text{Cas}}(x_t, x_h)$, $x_t \in [-30, 30]$ for several values of x_h . We observe that the force is *attractive*.

From all of the above one can, at the very least, conclude the following:

- (*) The sign of the Casimir force for the GM is not necessarily the same for $h = 0$, for which case they are very well known - see, e.g., [4,9,33,34], as it is for $h \neq 0$.
- (**) The predictions of the “workhorse” of statistical mechanics — the mean-field approach sometimes—in particular in the studies of the Casimir force—can be wrong even with respect to the predicted sign of the force.

The results presented in the current article are based on exact analytical expressions for both the Gaussian and mean-field model.

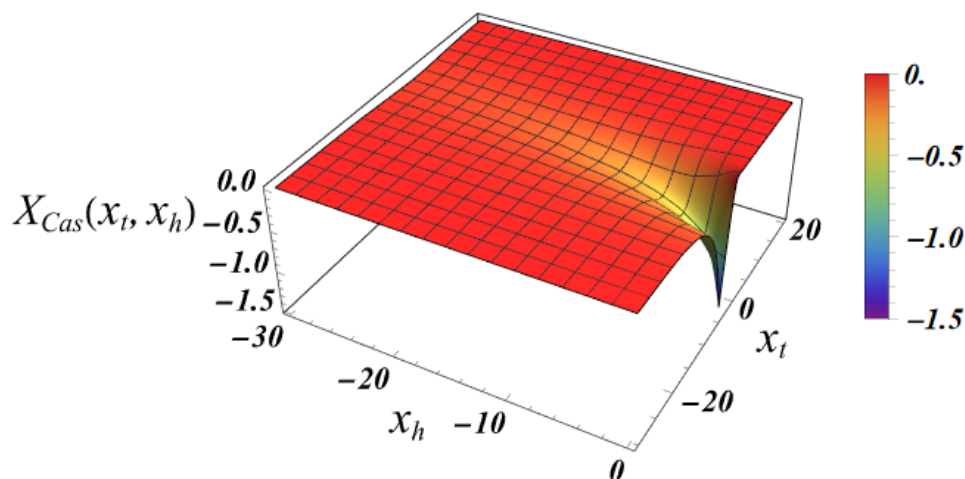


Figure 15. The behavior of the scaling function $X_{Cas}(x_t, x_h)$, $x_t \in [-30, 30]$, $x_h \in [-30, 0]$. We observe that the force is *attractive*.

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Abbreviations

The following abbreviations are used in this manuscript:

| | |
|------|--------------------------|
| CCF | critical Casimir force |
| BC's | boundary conditions |
| DN | Dirichlet–Neumann |
| GCE | grand canonical ensemble |
| LGM | lattice Gaussian model |
| CGM | continuum Gaussian model |

Appendix A On the evaluation of the field term

I am not sure that this is necessary, but just in case an improvement of the evaluation of the field term is necessary some probably helpful details are collected.

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