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Article

On General Covariance

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Abstract: General covariance is tested controversial after the investigations on gravitational redshift and acceleration. Further inspections on differential geometry indicate the opportunities of inequality of mixed derivatives of bases for the transformations between Riemannian spaces that will then lead to failure of the classical equations of Christoffel symbols, that is the main reason that causes controversies on general covariance. In fact, Christoffel symbols and base derivatives are both valid methodologies for analysis in gravitational fields. Measurable experiments on gravitational redshifts and accelerations have been sponsored to support the theoretical results. Conclusions have been drawn that light speed keeps general covariance in gravitational fields, but light energy momentum would not, while massive matters in gravitational fields do not travel in general covariance. Consequently, inferences on kinematics and relativistic release were put into research, which have got surprising verifications in applications. The problems in classical equations of light ray deflection, time delay of radar echoes and motion of massive matters have been revisited in details, which would help to kick off the errors and help us to find out real kinematics. Relativistic releases reveal the mechanism of evolutions of active galactic nuclei. Relativistic emissions and relativistic absorptions with giant redshifts would have involved with fantastic myths of intrinsic structures of matters that we perhaps know less. Theoretical researches, tremendous experimental and observational verifications would have resulted in comprehensive supports to the conclusions and inferences.

Keywords: general covariance; general relativity; gravitational redshift; gravitational acceleration; differential geometry; covariant differential; Lagrangian; energy momentum conservation; kinematics; relativistic release

1. Preface

Einstein carried out the equivalence principle after discussing the equivalence of gravitational mass and inertial mass, and then he generalized the equivalent principle to create general relativity [1], which predicts same physics in curve space of gravity geometrization as that in no gravity space, that could be called general covariance. Theoretically, general covariance should include but not be limited in the performances of motion inertia, energy and momentum conservations as well as equilibrium states in complicated systems. However, we will find out that quite amount of observations and evidences perform against general covariance.

It is believed that Riemannian geometry has been employed in general relativity for gravity geometrization [2]. But in fact, a transformation of a space does not really determine physics, what on earth the realities do. It is said that it is not the geometry but only the general covariance that matters. In fact, we could find out plenty of contradictions in classical theory of general relativity. It could be verified that even the motions of matter's freefalling on the Earth cannot be well interpreted in the frames of the classical theory. So that it is time to sponsor a series of inspections and perceptions to insight into the topics on general covariance.

The gravitational redshift and gravitational acceleration are of the two typical effects that gravity acts on the light rays and massive matters respectively. Hence researches on these topics would be highly powerful to probe into the investigations and realizations on general covariance.

To discover realities is more important than to carry out new equations and theorems.

2. Contradictions on Gravitational Redshift and Acceleration

2.1. Newtonian Gravitational Redshift

The equation of gravitational redshift can be drawn via Doppler redshift in a thought experiment that a light ray be emitted from the ceiling to the bottom or that is reversely performed from bottom to the ceiling, in a freely falling elevator cabin in a center source field [3]. It could be verified that any observer in the cabin would detect no frequency shift whatever the ways the light emitted. Suppose another observer outside the cabin keeping rest so that to have a relative velocity against the cabin, who will then detect a frequency shift other than the freefalling observer. What I want to say is that whoever freefalling will detect no frequency shift, no matter inside or outside the cabin. It is relative motion that eliminates the gravitational frequency shift. Once the relative velocities catch up to a relativistic velocity, the gravitational frequency shift will then cannot be eliminated anymore.

As a light ray is down ward emitted from a cabin ceiling and at the same time the cabin is released to freely fall down, the light front should spend a time interval to reach the bottom that

$$\Delta t = \frac{\Delta H}{c} \quad (1)$$

where ΔH is the distance that the light travels from the start to the end which may approximately equal to the height of the cabin, c is light speed.

Thus, the cabin velocity increase is

$$\Delta v = a\Delta t \quad (2)$$

where a is gravitational acceleration, which approximately equals to gravity g case velocity does not reach relativistic level. We know it exactly has a minus value as its direction pointing to center source.

Doppler redshift is frequency shift between two observers that one has a relative motion to another, to detect frequency. Here the rest observer outside would have a relative velocity $-\Delta v$ or $+\Delta v$ to the cabin so that when they receive light ray at the same time with the inner, the Doppler redshift could be calculated as

$$z_D = \frac{\Delta v}{c} = \frac{g\Delta H}{c^2} = \frac{\Delta\phi}{mc^2} \quad (3)$$

With the calculation of Doppler redshift we know that the gravitational redshift has happened in the same value. Notwithstanding, I prefer to suggest a new methodology to get an equation for gravitational redshift, in which a proposal should be given that every photon at any position in a center source field could be assumed to have experienced a travel from a farthest point to the current position. This attempt may bring about more physical significances and comprehensive understandings.

A photon in one source field at position r with frequency $\nu_{0(r)}$ is set to have a virtual primary frequency $\nu_{0(\infty)}$ at a farthest point, so that a so called primary dynamic energy is

$$E_{(\infty)} = h\nu_{0(\infty)} \quad (4)$$

where h is Planck's constant. NB, we are not talking about quantum character of photons so that photonic mass momentum we are talking refers to statistic quantities.

The corresponding dynamic mass comes from the mass-energy equation is

$$m = E_{(\infty)}c^{-2} \quad (5)$$

where c is light speed.

Then the gravitational potential at position r , especially as is in weak field with $r \gg r^*$, could be written as

$$\phi_{(r)} = \int_{\infty}^r \frac{GMm_r}{r^2} dr = -\frac{GM\bar{m}}{r} \approx -\frac{GMm}{r} = -\frac{r^*}{2r} mc^2 \quad (6)$$

where G is gravitational constant, M is the mass of the center source, m_r is photonic mass at position r , \bar{m} is the mean mass for the integral and it could be approximately replaced by m case in weak field, and r^* is Schwarzschild radius written as $r^* = \frac{2GM}{c^2}$.

The current dynamic energy is the summation of primary dynamic energy and released potential

$$E_r = h\nu_{0(r)} = E_{(\infty)} + [\phi_{(\infty)} - \phi_{(r)}] \approx mc^2 + \frac{r^*}{2r} mc^2 \approx h\nu_{0(\infty)} (1 + \frac{r^*}{2r}) \quad (7)$$

Case a light ray travels from positions r_1 to r_2 , as is shown in Figure 1, gravitational redshift happens.

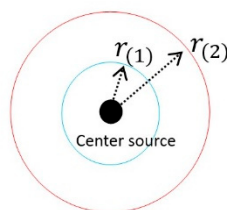


Figure 1. Center source field.

Gravitational redshift will be defined as

$$z_{g\lambda} = \frac{\lambda^1_{(2)} - \lambda^1_{(1)}}{\lambda^1_{(1)}} = \frac{\nu_{0(1)} - \nu_{0(2)}}{\nu_{0(2)}} \approx \frac{r^*}{r_{(1)}} \cdot \frac{r_{(2)} - r_{(1)}}{2r_{(2)} + r^*} \quad (8)$$

Considering weak field effects, the gravitational redshift is

$$z_{g\lambda} \approx r^* \cdot \frac{r_{(2)} - r_{(1)}}{2r_{(1)}r_{(2)}} = \frac{r^*}{2r_{(1)}} - \frac{r^*}{2r_{(2)}} = \frac{\phi_{(2)} - \phi_{(1)}}{mc^2} = \frac{\Delta\phi}{mc^2} \quad (9)$$

They are the forms of frequency shift of wave length based expression in weak fields, called redshift, where λ^1 and ν_0 are wave length and frequency tensors in contra variant space. Of course, a frequency shift could also be defined based on frequency. But we have been used to the forms based on wave length in traditions. In this equation, redshift may go up to more than 1.0, while blueshift must have been limited in -1 to 0. Case in the form of frequency based equations, one could get blueshift greater than 1.0 but redshift limited in -1 to 0.

We could also carry out new forms of frequency shift for conveniences in special discussions, that a differential of wave length based redshift could be defined as

$$dz_{\lambda} = \frac{d\lambda^1_{(r)}}{\lambda^1_{(r)}} = d\ln\lambda^1_{(r)} \quad (10)$$

So that the integral form for $r_{(1)}$ to $r_{(2)}$

$$z_{\lambda} = \ln\lambda^1_{(r)} \Big|_{r_{(1)}}^{r_{(2)}} = \ln \frac{\lambda^1_{(2)}}{\lambda^1_{(1)}} = \ln \frac{\nu_{0(1)}}{\nu_{0(2)}} \quad (11)$$

Or a differential form of frequency based

$$dz_{\nu} = \frac{d\nu_{0(r)}}{\nu_{0(r)}} = d\ln\nu_{0(r)} \quad (12)$$

So that

$$z_{\nu} = \ln\nu_{0(r)} \Big|_{r_{(1)}}^{r_{(2)}} = \ln \frac{\nu_{0(2)}}{\nu_{0(1)}} = \ln \frac{\lambda^1_{(1)}}{\lambda^1_{(2)}} = -\ln \frac{\lambda^1_{(2)}}{\lambda^1_{(1)}} \quad (13)$$

It is said that after the definitions of Eq. (11) and Eq. (13) there is

$$z_{\lambda} = -z_{\nu} \quad (14)$$

We have seen that they have shown difference from traditional equation. But for small frequency shift, both the two integral equations could be use instead of traditional equation.

Moreover, with Eq. (7), the gravitational frequency differential

$$dv_{0(r)} \approx d[v_{0(\infty)} \left(1 + \frac{r^*}{2r}\right)] = -\frac{r^*}{2r^2} v_{0(\infty)} dr \quad (15)$$

Then the differential frequency shift goes

$$dz_{gv} = \frac{dv_{0(r)}}{v_{0(r)}} \approx \frac{-\frac{r^*}{2r^2}}{1 + \frac{r^*}{2r}} dr = -\frac{r^* dr}{r(2r + r^*)} \quad (16)$$

This equation will be also taken into further discussions in the next sections. In next sections, the subscripts of frequency shift symbols will be neglected for convenience and in most cases we use the concept redshift to present frequency shifts.

2.2. Errors in the Equation of so Called Revisit Gravitational Redshift

A thought experiment has ever been employed to present the conception of the so called revisit gravitational redshift [3,4], in which a pulse of photons are supposed to be emitted from position 1, lasting for a time interval $\Delta t_{(1)}$, and be received at position 2 within a time interval $\Delta t_{(2)}$. The world lines of the photons are shown in Figure 2. We then know that the two drawn lines are literally the world lines of the first photon and the final photon of the light pulse.

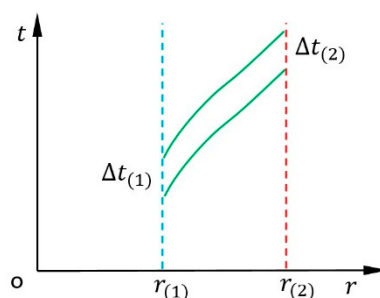


Figure 2. World lines for a pulse of photons.

Believing that the two world lines besides the $\Delta t_{(1)}$ and $\Delta t_{(2)}$ are parallel, it is known that the time interval $\Delta t_{(1)}$ equals to that of $\Delta t_{(2)}$. As the light frequency being inversely proportional to proper time interval $\Delta \tau$, the proper forms of frequency ratio were written in some textbooks as [3,4]

$$\frac{v_{(1)}}{v_{(2)}} = \frac{\Delta \tau_{(2)}}{\Delta \tau_{(1)}} = \frac{\sqrt{g_{00(2)} \Delta t_{(2)}}}{\sqrt{g_{00(1)} \Delta t_{(1)}}} = \frac{\sqrt{g_{00(2)}}}{\sqrt{g_{00(1)}}} \approx 1 + \frac{\phi_{(2)} - \phi_{(1)}}{mc^2} \quad (17)$$

where $v_{(1)}$ and $v_{(2)}$ are proper frequencies corresponding to position 1 and position 2, g_{00} is time metric. The right hand side of this equation is an approximate result with Schwarzschild solution of metric g_{00} in condition that $r_2 > r_1 \gg r^*$.

Thus the so called revisit redshift is

$$z_{revisit} = \frac{v_{(1)}}{v_{(2)}} - 1 \approx \frac{\Delta \phi}{mc^2} \quad (18)$$

It is seemingly that the revisit form of equation for gravitational redshift was worked out.

But there are quite many errors in above equations. (i) The two world lines are belong to the first photon and the final photon, thus they might be controlled by emitter, so that the time intervals between the two lines do nothing with any light frequencies. (ii) Any time intervals between neighboring photons could be randomly assigned, so that these intervals also do nothing with any light frequencies. (iii) Frequency of a photon is the reciprocal of its photonic period and is the intrinsic property of a photon, so that it is independent to its positions relating to other photons and any variation of the frequency will not change its position in the pulse of photons. (iv) Detections of

frequencies must involve with wave numbers and time intervals together, this equation has made a mistake by comparing time intervals only.

2.3. Further Investigation into the Revisit Gravitational Redshift

Following the rules in classical frames, the tensor of light wave period is a tensor with upper index

$$T^0 = \frac{dt}{dn} \quad (19)$$

where, T^0 is contra variant period that is described by coordinate time, and dt is coordinate time lasting in which the number of dn waves may have traveled across a specific position. And the proper form of wave period

$$T = \frac{d\tau}{dn} \quad (20)$$

where, $d\tau$ is proper time lasting for dn number of waves to cross the position. So that there is

$$T = e_0 T^0 \text{ or } T^0 = e^0 T \quad (21)$$

where, e_0 is the first component of covariant time base, and it should be noted that the base e_0 is a vector but the component e_0 is not vector even though it is still a tensor. One can get more understandings for these expressions I would sponsor here and in followings.

In this way, we know that different values of the contra variant period and proper period corresponding to a same physical issue. Then it leads to a frequency tensor

$$\nu_0 = \frac{1}{T^0} = e_0 \nu = \frac{dn}{dt} \quad (22)$$

We have seen that, ν_0 is called covariant tensor traditionally, and ν is a proper tensor. This may have brought about confusions in that ν is actually covariant and yet ν_0 has been named the name. I am not going to change the naming methodology thoroughly right now because that may cause more difficulties and sounds more trivial.

Generally, some pure one-order tensors seem to be infinite small quantities such as dr , but for ν_0 , it is dn divided by dt , so that it is not an infinite small quantity. As for velocity tensors, they are really mixed tensors.

Theoretically, the covariant derivative of a frequency in a falling process into a center source can be written as

$$\frac{D\nu_0}{dr} = \frac{\partial \nu_0}{\partial r} - \Gamma_{10}^0 \nu_0 \quad (23)$$

where, $\Gamma_{\nu\mu}^\lambda$ is Christoffel symbols.

We know that, as has been presented in Eq. (15), the contra variant derivative is

$$\frac{\partial \nu_0}{\partial r} = -\frac{r^*}{2r^2} \nu_{0(\infty)} \quad (24)$$

Let us calculate the Christoffel symbols in Eq. (23) that

$$\Gamma_{10}^0 = \frac{1}{2} g^{0\lambda} \left(\frac{\partial g_{1\lambda}}{\partial x^0} + \frac{\partial g_{0\lambda}}{\partial x^1} - \frac{\partial g_{10}}{\partial x^\lambda} \right) \quad (25)$$

The Einstein summation convention has been and will be adopt unless additional declarations.

It is found that only in the condition of $\lambda = 0$ there is a nonvanishing item in the bracket of right hand side of Eq. (25), so that with Schwarzschild metrics it turns to be

$$\Gamma_{10}^0 = \frac{1}{2} g^{00} \frac{\partial g_{00}}{\partial x^1} = \frac{1}{2} \left(-1 + \frac{r^*}{r} \right)^{-1} \left(-1 + \frac{r^*}{r} \right)' = \frac{r^*}{2r^2} \left(1 + \frac{r^*}{r-r^*} \right) \quad (26)$$

The covariant derivative will be calculated to be

$$\frac{D\nu_0}{dr} = -\frac{r^*}{2r^2} \nu_{0(\infty)} - \frac{r^*}{2r^2} \left(1 + \frac{r^*}{r-r^*} \right) \nu_0$$

$$= -\frac{r^*}{2r^2} \nu_{0(\infty)} - \frac{r^*}{2r^2} \left(1 + \frac{r^*}{r-r^*}\right) \left(1 + \frac{r^*}{2r}\right) \nu_{0(\infty)} \quad (27)$$

We could get an approximate solution for weak field that

$$\frac{D\nu_0}{dr} \approx 2 \frac{\partial \nu_0}{\partial r} \quad (28)$$

As we have seen, it shows that the values of covariant redshift doubles that of the contra variant redshift.

2.4. Additional Discussions on the Thought Experiment of Freefalling Elevator Cabin

The thought experiment that observer in freefalling elevator cabin will detect no frequency shift is usually employed to discuss the equivalent principle for the support of general covariance. But in fact, that issue does nothing with covariance. It is just because that the gravitational redshift happens to be offset by Doppler redshift. In fact, this experiment is a comprehensive event that relates both to gravitational redshift and gravitational acceleration.

We know that in a freefalling cabin, freefalling observer will not detect any frequency shift no matter the light ray emitter is on the bottom or on the top to emit up or down to the receiver. Nevertheless, even if the light ray is not vertical, freefalling observer would also observe no frequency shift in the cabin.

For a case that a light ray is emitted to the top with an angle θ to the vertical line

The time interval for light ray from bottom to the top is

$$\Delta t = \frac{1}{c} \frac{\Delta H}{\cos \theta} \quad (29)$$

And the velocity increase of freefall cabin is

$$\Delta v = a \Delta t = \frac{g}{c} \frac{\Delta H}{\cos \theta} \quad (30)$$

Then the velocity increase component at the direction of light ray

$$\Delta v_l = \Delta v \cos \theta = g \frac{\Delta H}{c} \quad (31)$$

So that we get the Doppler redshift again as

$$z_D = \frac{\Delta v_l}{c} = \frac{g \Delta H}{c^2} = \frac{\Delta \phi}{mc^2} \quad (32)$$

In fact, the Doppler redshift for the detector in freefalling cabin could eliminate the gravitational frequency shift, even if the emitter is either outside of the cabin or with any low initial velocity. The real reason for non-detectable frequency shift in freefalling cabin is that the relative motion formed Doppler redshift just has a minus approximate value of gravitational frequency shift.

I prefer to put forward the case that the gravitational frequency shift cannot be eliminated by Doppler redshift. In the case that cabin has a relativistic initial velocity, the geometrical acceleration is not the total gravitational acceleration again that

$$a = \xi g, \xi < 1 \quad (33)$$

Thus, for light rays passing across the cabin there is the Doppler velocity of detector

$$\Delta v_l = a \Delta t = \xi g \frac{\Delta H}{c} \quad (34)$$

so that

$$z_D = \frac{\Delta v_l}{c} = \xi \frac{g \Delta H}{c^2} = \xi \frac{\Delta \phi}{mc^2} \quad (35)$$

While the gravitational frequency is still the form

$$z_g = \frac{g \Delta H}{c^2} = \frac{\Delta \phi}{mc^2} \quad (36)$$

so that in this case

$$z_D \neq z_g \quad (37)$$

We now know that in some situations in freefalling cabin one could detect frequency shift again, so that the thought experiment cannot support general covariance thoroughly. Of course, one can continue to argue that relativistic motion may bring about more sophisticated conditions on frequency shift.

2.5. An Investigation into Gravitational Acceleration

For a freefalling massive matter in gravitational field, the component of the velocity at the direction of radius of center source is

$$V^1 = \frac{dx^1}{d\tau} \quad (38)$$

There is the covariant derivative at one of coordinate direction

$$\frac{DV^1}{d\lambda} = \frac{dV^1}{d\lambda} + \Gamma_{\lambda\mu}^1 V^\mu \quad (39)$$

Because $dx^0 = cdt$, and $dt = e^0 d\tau$, where e^0 is the nonvanishing component of the base e^0 . Thus, the covariant derivative by τ is

$$\frac{DV^1}{d\tau} = \frac{dV^1}{d\tau} + \frac{dx^0}{d\tau} \Gamma_{0\mu}^1 V^\mu = a^1 + ce^0 \Gamma_{0\mu}^1 V^\mu \quad (40)$$

where, $a^1 = \frac{dV^1}{d\tau}$ is the contra variant acceleration of the matter, c is light speed.

NB, accelerations we have and will discuss refer to geometrical accelerations, which will be something different from gravity g , in that the latter sometimes may also be called gravitational accelerations in some writings but in fact matters may not experience accelerating up to that.

With the equation of Christoffel symbols

$$\Gamma_{0\lambda}^1 = \frac{1}{2} g^{1\rho} \left(\frac{\partial g_{0\rho}}{\partial x^\lambda} + \frac{\partial g_{\lambda\rho}}{\partial x^0} - \frac{\partial g_{0\lambda}}{\partial x^\rho} \right) \quad (41)$$

case $\lambda = 1$ in this equation, there is $\Gamma_{01}^1 = 0$. Then considering the condition of $\lambda = 0$, it is

$$\Gamma_{00}^1 = \frac{1}{2} g^{1\rho} \left(\frac{\partial g_{0\rho}}{\partial x^0} + \frac{\partial g_{0\rho}}{\partial x^0} - \frac{\partial g_{00}}{\partial x^\rho} \right) \quad (42)$$

In this condition, only in the case of $\rho = 1$ there is the nonvanishing item in the bracket of right hand side, so that

$$\Gamma_{00}^1 = \frac{1}{2} g^{11} \left(-\frac{\partial g_{00}}{\partial x^1} \right) = \frac{1}{2} \left(1 - \frac{r^*}{r} \right) \left(1 - \frac{r^*}{r} \right)' = \frac{1}{2} \left(1 - \frac{r^*}{r} \right) \frac{r^*}{r^2} \quad (43)$$

And with $V^0 = \frac{dx^0}{d\tau} = \frac{cdt}{d\tau} = e^0 c$, Eq. (40) turns to be

$$\frac{DV^1}{d\tau} = a^1 + ce^0 \Gamma_{00}^1 V^0 = a^1 + e^0 e^0 c^2 \left(1 - \frac{r^*}{r} \right) \frac{r^*}{2r^2} \quad (44)$$

As a matter freefalls on to the earth, its acceleration could be calculated to be

$$a^1 = -e^0 e^0 \frac{GM}{r^2} = -e^0 e^0 c^2 \frac{r^*}{2r^2} \quad (45)$$

Thus, there is the weak field solution

$$\frac{DV^1}{d\tau} \approx 0 \quad (46)$$

2.6. Discussions and Controversies

Now we have got more complex results that the revisit gravitational redshift calculated doubles that of contra variant value while covariant acceleration of freefalling goes to zero. The latter seems to fit with general covariance but the previous does not. And there is still a problem that revisit

gravitational redshift solved in the way of covariant derivation goes contradictory to classical solution. Moreover, the minus form of g_{00} also involves more matters that have cause the result of zero acceleration, deserving further discussions in next sections.

It is extraordinary that the covariant derivative of light frequency goes nonvanishing in the view of general covariance. Even more, it is investigated that covariant derivative analysis shows that the classical solution of revisit redshift may be false in that the frequency has been wrong defined. Some ones may argue that the frequency is not a tensor, but that makes no sense because light frequency is the reciprocal of its wave period which involves with time coordinate.

What I urgently want to say is that these discussions are not enough. The most significant problem is that, the item in original covariant differential in Eq. (44) that matter's velocities multiplied with the base's differential, has been calculated to do nothing with the realistic velocity. We should know that the multiplied item in Eq. (39) originally indicates a base variation ratio multiplied with the very tensors, but the final equation has given up the effects of initial velocity that does lead to contradictions, and the time speed V^0 is a virtual velocity which might have been abused. These contradictions really bothered me until it is occasionally gone through one day, that the real problem is deeply hidden in the equations of Christoffel symbols.

Notwithstanding, a differential of velocity is the differential of that on the trajectory of matter's motion. Thus, the covariant time derivatives cannot be treated directly as ordinary derivatives anymore as in Eq. (40). We are going to carry out detailed discussions on trajectory derivatives in next sections so that to interpret covariant time derivatives correctly.

3. Theoretical Investigations on Christoffel Symbols

3.1. Classical Equations of Christoffel Symbols

Christoffel symbols have been defined as

$$\frac{\partial e^\mu}{\partial x^\nu} = -\Gamma_{\nu\lambda}^\mu e^\lambda \quad (47)$$

There is nothing wrong with the definition in that the derivative of a base must have a direction and so that to be written as a linear combination of the total bases. The key problem is what the Christoffel symbols are.

In this and following sections, all symbols of vectors and matrix would be bold written while their components and that of other quantities may be simplified written, no matter they are tensors or not.

In most conditions, Christoffel symbols could be discussed by the derivation of metrics as

$$\frac{\partial g_{\mu\nu}}{\partial x^\lambda} = \frac{\partial}{\partial x^\lambda} (e_\mu \cdot e_\nu) = \frac{\partial e_\mu}{\partial x^\lambda} \cdot e_\nu + e_\mu \cdot \frac{\partial e_\nu}{\partial x^\lambda} = \Gamma_{\mu\lambda}^\rho e_\rho \cdot e_\nu + \Gamma_{\nu\lambda}^\rho e_\rho \cdot e_\mu \quad (48)$$

For the derivative forms with alternative indexes mathematically, there will be

$$\frac{\partial g_{\mu\nu}}{\partial x^\lambda} - \Gamma_{\mu\lambda}^\rho g_{\rho\nu} - \Gamma_{\nu\lambda}^\rho g_{\rho\mu} = 0 \quad (49)$$

$$\frac{\partial g_{\lambda\nu}}{\partial x^\mu} - \Gamma_{\lambda\mu}^\rho g_{\rho\nu} - \Gamma_{\nu\mu}^\rho g_{\rho\lambda} = 0 \quad (50)$$

$$\frac{\partial g_{\mu\lambda}}{\partial x^\nu} - \Gamma_{\mu\nu}^\rho g_{\rho\lambda} - \Gamma_{\lambda\nu}^\rho g_{\rho\mu} = 0 \quad (51)$$

In the case that the so called torsions $S_{\mu\lambda} = \Gamma_{\mu\nu}^\rho - \Gamma_{\nu\mu}^\rho$ is set zero, the summation of the previous two equations minus to the last one that

$$\frac{\partial g_{\mu\nu}}{\partial x^\lambda} + \frac{\partial g_{\lambda\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\lambda}}{\partial x^\nu} - 2\Gamma_{\mu\lambda}^\rho g_{\rho\nu} = 0 \quad (52)$$

Thereafter the equations of Christoffel symbols will be solved as [2,3,5]

$$\Gamma_{\mu\lambda}^{\rho} = \frac{1}{2} g^{\rho\nu} \left(\frac{\partial g_{\mu\nu}}{\partial x^{\lambda}} + \frac{\partial g_{\lambda\nu}}{\partial x^{\mu}} - \frac{\partial g_{\mu\lambda}}{\partial x^{\nu}} \right) \quad (53)$$

Generally, Eq. 49 - Eq. 51 could also be rewritten as the original forms

$$\frac{\partial g_{\mu\nu}}{\partial x^{\lambda}} - \frac{\partial e_{\mu}}{\partial x^{\lambda}} \cdot e_{\nu} - \frac{\partial e_{\nu}}{\partial x^{\lambda}} \cdot e_{\mu} = 0 \quad (54)$$

$$\frac{\partial g_{\lambda\nu}}{\partial x^{\mu}} - \frac{\partial e_{\lambda}}{\partial x^{\mu}} \cdot e_{\nu} - \frac{\partial e_{\nu}}{\partial x^{\mu}} \cdot e_{\lambda} = 0 \quad (55)$$

$$\frac{\partial g_{\mu\lambda}}{\partial x^{\nu}} - \frac{\partial e_{\mu}}{\partial x^{\nu}} \cdot e_{\lambda} - \frac{\partial e_{\lambda}}{\partial x^{\nu}} \cdot e_{\mu} = 0 \quad (56)$$

We will find that in some conditions the torsions do not always equal to zero. It is said that the mixed derivatives of bases $\frac{\partial e_{\mu}}{\partial x^{\nu}}$ and $\frac{\partial e_{\nu}}{\partial x^{\mu}}$ do not always equal, and then the Christoffel symbols with mixed subscripts $\Gamma_{\mu\nu}^{\rho}$ and $\Gamma_{\nu\mu}^{\rho}$ do not always equal, so that the Eq. (53) might be invalid in those conditions.

3.2. The Inequality of Christoffel Symbols of Mixed Subscripts

In differential geometry, the equality of Christoffel symbols of mixed subscripts is usually adopted the doctrine. But no forceful researches could provide reliable supports. The truth is that the problem of mixed derivatives of bases in a Riemannian space are far different from the problem of normal mixed derivatives of a 3-dimensional surface in Euclidean geometry.

We could find out the truth that in the deduction of Γ_{00}^1 in Eq. (43), $\frac{\partial g_{00}}{\partial x^1}$ has been used instead of $\frac{\partial g_{11}}{\partial x^0}$ so that to gain $\Gamma_{00}^1 = \frac{1}{2} g^{11} \left(-\frac{\partial g_{00}}{\partial x^1} \right)$. But we can easily calculate that $\frac{\partial e_0}{\partial x^1}$ and $\frac{\partial e_1}{\partial x^0}$ are not equal. We might have found out the problems.

3.2.1. Coordinate Transformations and Bases

Any points in a Riemannian space of Riemannian manifold of dimension n has a neighborhood homeomorphic to a subset of Euclidean space of dimension n , so that there must be probable maps between the neighborhoods and the corresponding subsets. It is just to say that the coordinates of any points in Riemannian space could be expressed with the coordinates of corresponding points of Euclidean space, and reversely. If a part or the entire of a Riemannian space are continuous and differentiable, Euclidean coordinate lines could be drawn in the part or the entire of the Riemannian space. On the other side, coordinate lines of Riemannian space could also be drawn in the corresponding Euclidean space. For convenience, the Riemannian space could be called covariant space, and the corresponding Euclidean space could be called contra variant space. A contra variant space is curved in the view of its covariant space, and the covariant space is also curved in the view of contra variant space.

It is obvious that transformations of spaces are actually coordinate transformations. These transformations could happen between covariant space and contra variant space, as well as they could happen among homeomorphic Riemannian spaces. Coordinate transformations may perform in the way with unequal metrics as well as the way with equal metrics.

The more effective method for coordinate transformation is to define bases and distances for spaces. Two examples would be presented firstly for definitions and for following discussions.

Example 1: Bases of Riemannian manifold of super surface

The derivative vectors of 3-dimensional surfaces were usually employed to form bases in classical differential geometry. The curve space has an extra dimension than a plane space that could be called the super surface. The Riemannian manifold of the super surface in the 3-dimensional space (u, v, w) would have a homeomorphic Euclidean space (x, y) in the space (x, y, z) . The coordinate lines x and y in contra variant space could be transformed to be $\xi(x)$ and $\eta(y)$ in covariant space,

while u , v and w in covariant space be transformed to be $\alpha(u)$, $\beta(v)$ and $\gamma(w)$ in the space (x, y, z) as shown in Figures 3 and 4.

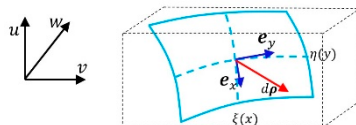


Figure 3. A super surface as covariant space in 3-dimensional space.

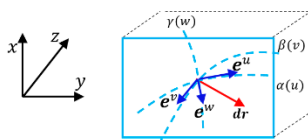


Figure 4. A coordinate plane (x, y) as contra variant space in 3-dimensional space.

The super surface could be determined by a vector function ρ

$$\rho = \rho(u, v, w)|_{2 \text{ of variables independent}} \quad (57)$$

And the function could also be written as

$$\rho = \rho(\xi, \eta) \quad (58)$$

or

$$\rho = \rho(x, y) \quad (59)$$

At the same time, the contra variant space could be defined by r

$$r = r(x, y, z)|_{z=\text{const.}} = r(\alpha, \beta, \gamma)|_{2 \text{ of variables independent}} = r(x, y) = r(\xi, \eta) \quad (60)$$

There will be varieties of available ways to develop the expressions of bases and distances. I prefer to put forward the followings might as well.

The way in super space:

In super space, the differential $d\rho$ has 3 components

$$d\rho = \begin{pmatrix} du \\ dv \\ dw \end{pmatrix} \quad (61)$$

That of differential dr could be simplified to be 2 dimensional because it just locates in the space (x, y)

$$dr = \begin{pmatrix} dx \\ dy \end{pmatrix} \quad (62)$$

To define a set of covariant bases for a position in covariant space by

$$e_x = \frac{\partial \rho}{\partial x}, \quad e_y = \frac{\partial \rho}{\partial y} \quad (63)$$

It should be pointed out that in some publications coordinate and vector symbols have been used reversely, which is just a kind of treatment, but in the theory of relativity they may bring about confusions.

In covariant space, the differential $d\rho$ expressed by dr with covariant bases

$$d\rho = dx e_x + dy e_y \quad (64)$$

Differential distance could be defined as

$$ds^2 = du^2 + dv^2 + dw^2 = d\rho \cdot d\rho = (dx e_x + dy e_y) \cdot (dx e_x + dy e_y) \quad (65)$$

If the bases are orthogonal, there is

$$ds^2 = e_x \cdot e_x dx^2 + e_y \cdot e_y dy^2 \quad (66)$$

We have seen that the covariant bases are defined in covariant space to help contra variant coordinates to form covariant distances.

There are more complexities for a transformation between a super surface in 3-dimensional space and R^2 that the 2 covariant bases would have 3 components

$$\mathbf{e}_x = \begin{pmatrix} \frac{\partial u}{\partial x} \\ \frac{\partial v}{\partial x} \\ \frac{\partial w}{\partial x} \end{pmatrix}, \mathbf{e}_y = \begin{pmatrix} \frac{\partial u}{\partial y} \\ \frac{\partial v}{\partial y} \\ \frac{\partial w}{\partial y} \end{pmatrix} \quad (67)$$

Define the contra variant bases for a point in the plane that

$$\mathbf{e}^u = \frac{\partial \mathbf{r}}{\partial u}, \mathbf{e}^v = \frac{\partial \mathbf{r}}{\partial v}, \mathbf{e}^w = \frac{\partial \mathbf{r}}{\partial w} \quad (68)$$

The 3 contra variant bases all have 2 components as

$$\mathbf{e}^u = \begin{pmatrix} \frac{\partial x}{\partial u} \\ \frac{\partial y}{\partial u} \end{pmatrix}, \mathbf{e}^v = \begin{pmatrix} \frac{\partial x}{\partial v} \\ \frac{\partial y}{\partial v} \end{pmatrix}, \mathbf{e}^w = \begin{pmatrix} \frac{\partial x}{\partial w} \\ \frac{\partial y}{\partial w} \end{pmatrix} \quad (69)$$

The differential $d\mathbf{r}$ expressed by $d\boldsymbol{\rho}$ in contra variant space

$$d\mathbf{r} = du\mathbf{e}^u + dv\mathbf{e}^v + dw\mathbf{e}^w \quad (70)$$

Of course, one can create transformation matrix to perform the relationship between $d\mathbf{r}$ and $d\boldsymbol{\rho}$ directly.

The distance could be defined as

$$d\zeta^2 = dx^2 + dy^2 = d\mathbf{r} \cdot d\mathbf{r} = (du\mathbf{e}^u + dv\mathbf{e}^v + dw\mathbf{e}^w) \cdot (du\mathbf{e}^u + dv\mathbf{e}^v + dw\mathbf{e}^w) \quad (71)$$

If the bases are orthogonal

$$d\zeta^2 = \mathbf{e}^u \cdot \mathbf{e}^u du^2 + \mathbf{e}^v \cdot \mathbf{e}^v dv^2 + \mathbf{e}^w \cdot \mathbf{e}^w dw^2 \quad (72)$$

One could imagine that this condition you cannot give the relationship of metrics that g_{ii} equals to $1/g^{ii}$, in that the covariant bases have 3 components and contra variant bases have 2.

The way in tangent space:

Consequently, the issues could be simplified in tangent spaces. At a position $\boldsymbol{\rho}$ in the covariant space, there is a neighborhood which will be labeled with coordinate lines of $\xi(x)$ and $\eta(y)$, at the same time at the position \mathbf{r} , there is a corresponding neighborhood in contra variant space labeled with coordinate lines of x and y , as shown in Figures 5 and 6. Generally, coordinate lines could be set orthogonal. In most of publications, $\xi(x)$ and $\eta(y)$ were seen as x and y , but one should realize that the difference really matters.

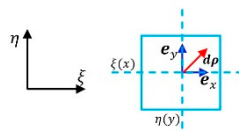


Figure 5. Covariant tangent space .

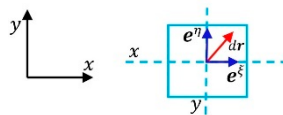


Figure 6. Contra variant tangent space.

One could define the differential vector in covariant space

$$d\boldsymbol{\rho} = \begin{pmatrix} d\xi \\ d\eta \end{pmatrix} \quad (73)$$

As a result, the bases

$$\mathbf{e}_x = \frac{\partial \mathbf{p}}{\partial x} = \begin{pmatrix} \frac{\partial \xi}{\partial x} \\ \frac{\partial \eta}{\partial x} \end{pmatrix}, \quad \mathbf{e}_y = \frac{\partial \mathbf{p}}{\partial y} = \begin{pmatrix} \frac{\partial \xi}{\partial y} \\ \frac{\partial \eta}{\partial y} \end{pmatrix} \quad (74)$$

Again, there is the distance

$$ds^2 = d\xi^2 + d\eta^2 = d\mathbf{p} \cdot d\mathbf{p} = \mathbf{e}_x \cdot \mathbf{e}_x dx^2 + \mathbf{e}_y \cdot \mathbf{e}_y dy^2 \quad (75)$$

The differential $d\mathbf{r}$ keep the form as Eq. (62), so that the contra variant bases could be defined as

$$\mathbf{e}^\xi = \frac{\partial \mathbf{r}}{\partial \xi} = \begin{pmatrix} \frac{\partial x}{\partial \xi} \\ \frac{\partial y}{\partial \xi} \end{pmatrix}, \quad \mathbf{e}^\eta = \frac{\partial \mathbf{r}}{\partial \eta} = \begin{pmatrix} \frac{\partial x}{\partial \eta} \\ \frac{\partial y}{\partial \eta} \end{pmatrix} \quad (76)$$

Also, there is

$$d\zeta^2 = dx^2 + dy^2 = d\mathbf{r} \cdot d\mathbf{r} = \mathbf{e}^\xi \cdot \mathbf{e}^\xi d\xi^2 + \mathbf{e}^\eta \cdot \mathbf{e}^\eta d\eta^2 \quad (77)$$

In this case, the relationship of metrics go harmonized that g_{ii} equals to $1/g^{ii}$. It should be pointed out that the ways of expressions of bases are all equivalent except that the substitutions of coordinate lines ξ and η might have hidden some information of super surface, so that I would like to make analysis within super space in most cases.

Example 2: Bases of Riemannian manifold of equal dimension

As a Riemannian manifold has equal dimension with its contra variant space, it could be called equal dimension manifold. A plane space (u, v) maps to a plane space (x, y) could be taken for granted, as shown in Figures 7 and 8.



Figure 7. Covariant space.

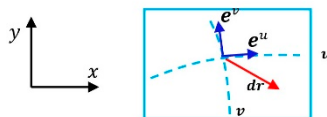


Figure 8. Contra variant space.

A differential vector in covariant space is

$$d\mathbf{r} = \begin{pmatrix} dx \\ dy \end{pmatrix} \quad (78)$$

The differential vector in contra variant space is

$$d\mathbf{p} = \begin{pmatrix} du \\ dv \end{pmatrix} \quad (79)$$

Thus, the definition of contra variant bases could be

$$\mathbf{e}^u = \frac{\partial \mathbf{r}}{\partial u}, \quad \mathbf{e}^v = \frac{\partial \mathbf{r}}{\partial v} \quad (80)$$

To express $d\mathbf{r}$ with $d\mathbf{p}$

$$d\mathbf{r} = du\mathbf{e}^u + dv\mathbf{e}^v \quad (81)$$

Also, there is the definition of covariant bases

$$\mathbf{e}_x = \frac{\partial \rho}{\partial x}, \mathbf{e}_y = \frac{\partial \rho}{\partial y} \quad (82)$$

So that the expression of $d\rho$ by $d\mathbf{r}$ should be

$$d\rho = dx\mathbf{e}_x + dy\mathbf{e}_y \quad (83)$$

Something different is that a covariant base has 2 components

$$\mathbf{e}_x = \begin{pmatrix} \frac{\partial u}{\partial x} \\ \frac{\partial v}{\partial x} \end{pmatrix}, \mathbf{e}_y = \begin{pmatrix} \frac{\partial u}{\partial y} \\ \frac{\partial v}{\partial y} \end{pmatrix} \quad (84)$$

And a contra variant base also has 2 components

$$\mathbf{e}^u = \begin{pmatrix} \frac{\partial x}{\partial u} \\ \frac{\partial y}{\partial u} \end{pmatrix}, \mathbf{e}^v = \begin{pmatrix} \frac{\partial x}{\partial v} \\ \frac{\partial y}{\partial v} \end{pmatrix} \quad (85)$$

In the case that bases are orthogonal, the distance

$$ds^2 = du^2 + dv^2 = d\rho \cdot d\rho = \mathbf{e}_x \cdot \mathbf{e}_x dx^2 + \mathbf{e}_y \cdot \mathbf{e}_y dy^2 \quad (86)$$

and

$$d\zeta^2 = dx^2 + dy^2 = d\mathbf{r} \cdot d\mathbf{r} = \mathbf{e}^u \cdot \mathbf{e}^u du^2 + \mathbf{e}^v \cdot \mathbf{e}^v dv^2 \quad (87)$$

3.2.2. Inequalities of Mixed Derivatives of Bases

Now it is the time to carry out the first discussion on the inequality of mixed derivatives of bases. The mixed derivatives of bases are just special defined for bases alternative derivations. As transformation from contra variant space to covariant space is concerned, the covariant bases could be considered to be derivated by the coordinate lines in chain rule

$$\mathbf{e}_x = \frac{\partial \rho}{\partial x} = \frac{\partial \rho}{\partial \xi} \frac{\partial \xi}{\partial x}, \mathbf{e}_y = \frac{\partial \rho}{\partial y} = \frac{\partial \rho}{\partial \eta} \frac{\partial \eta}{\partial y} \quad (88)$$

where, $\frac{\partial \rho}{\partial \xi}$ and $\frac{\partial \rho}{\partial \eta}$ are the direction derivatives along the coordinate lines $\xi(x)$ and $\eta(y)$ in covariant space, and $d\xi$ and $d\eta$ are their differential lengths in covariant space, which could be called the covariant lengths. And there will be a setting that Einstein summation convention does not act on double $d\xi$ and double $d\eta$.

It should be pointed out that in most mathematics and physics, mixed derivatives being confirmed to be equal is because in the Eq. (88) $d\mathbf{x}$ and $d\mathbf{y}$ is incorrectly understood to be the differential length in covariant space $(\mathbf{u}, \mathbf{v}, \mathbf{w})$, but they are really the lengths in contra variant space $(\mathbf{x}, \mathbf{y}, \mathbf{z})$. That is the reason we have carried out the concept of covariant length $d\xi$ and $d\eta$.

Thus, the mixed derivatives will be

$$\frac{\partial \mathbf{e}_x}{\partial y} = \frac{\partial^2 \rho}{\partial \xi \partial \eta} \frac{\partial \xi}{\partial x} \frac{\partial \eta}{\partial y} + \frac{\partial \rho}{\partial \xi} \frac{\partial^2 \xi}{\partial x \partial y} \quad (89)$$

and

$$\frac{\partial \mathbf{e}_y}{\partial x} = \frac{\partial^2 \rho}{\partial \eta \partial \xi} \frac{\partial \eta}{\partial y} \frac{\partial \xi}{\partial x} + \frac{\partial \rho}{\partial \eta} \frac{\partial^2 \eta}{\partial y \partial x} \quad (90)$$

Consequently 3 conditions could be focused on:

Condition 1:

If there is an equality between the first items of the two equations that

$$\frac{\partial^2 \rho}{\partial \xi \partial \eta} = \frac{\partial^2 \rho}{\partial \eta \partial \xi} \quad (91)$$

For example, in the super surface, the mixed derivatives of course have the equality just as the equality of normal mixed derivatives of a 3-dimensional surface in a Euclidean space.

In this case and if there is another equality for the last items of the two equations that

$$\frac{\partial \rho}{\partial \xi} \frac{\partial^2 \xi}{\partial x \partial y} = \frac{\partial \rho}{\partial \eta} \frac{\partial^2 \eta}{\partial y \partial x} \quad (92)$$

Then that must come to the conclusion

$$\frac{\partial e_x}{\partial x} = \frac{\partial e_y}{\partial y} \quad (93)$$

Otherwise, that depends.

It should be pointed out that $\frac{\partial \rho}{\partial \xi}$ and $\frac{\partial \rho}{\partial \eta}$ do not equal in general conditions because they have different directions and in most cases they are usually set orthogonal, so that if that equality Eq. (92) happens, it asks for

$$\frac{\partial^2 \xi}{\partial x \partial y} = \frac{\partial^2 \eta}{\partial y \partial x} = 0 \quad (94)$$

We will see that in some cases it is really well satisfied.

Condition 2:

Most special if

$$\frac{\partial^2 \rho}{\partial \xi \partial \eta} \neq \frac{\partial^2 \rho}{\partial \eta \partial \xi} \quad (95)$$

that indicate the first items of the two equations are not equal, but at the same time if the total equations are still equal that

$$\frac{\partial^2 \rho}{\partial \xi \partial \eta} \frac{\partial \xi}{\partial x} \frac{\partial \eta}{\partial y} + \frac{\partial \rho}{\partial \xi} \frac{\partial^2 \xi}{\partial x \partial y} = \frac{\partial^2 \rho}{\partial \eta \partial \xi} \frac{\partial \eta}{\partial y} \frac{\partial \xi}{\partial x} + \frac{\partial \rho}{\partial \eta} \frac{\partial^2 \eta}{\partial y \partial x} \quad (96)$$

We will still obtain the equality that

$$\frac{\partial e_x}{\partial x} = \frac{\partial e_y}{\partial y} \quad (97)$$

Condition 3:

This is the condition after the previous two conditions and else to them. Generally if

$$\frac{\partial^2 \rho}{\partial \xi \partial \eta} \frac{\partial \xi}{\partial x} \frac{\partial \eta}{\partial y} + \frac{\partial \rho}{\partial \xi} \frac{\partial^2 \xi}{\partial x \partial y} \neq \frac{\partial^2 \rho}{\partial \eta \partial \xi} \frac{\partial \eta}{\partial y} \frac{\partial \xi}{\partial x} + \frac{\partial \rho}{\partial \eta} \frac{\partial^2 \eta}{\partial y \partial x} \quad (98)$$

No matter the first items of the two equations are equal or not, the mixed derivatives will perform inequality

$$\frac{\partial e_x}{\partial x} \neq \frac{\partial e_y}{\partial y} \quad (99)$$

Then, turn to the issue of geometrical influence that the inequality of mixed derivatives will cause closure errors [6,7]. I prefer to give a brief presentation. Consider a differential in a curve line coordinate system expressed by bases along different coordinate paths as shown in Figure 9 that in path 1,

$$(d\rho)_1 = \int_x^{x+dx} \mathbf{e}_x dx + \int_y^{y+dy} (\mathbf{e}_y + \frac{\partial \mathbf{e}_y}{\partial x} dx) dy \quad (100)$$

The irregular expressions of same symbols of integral variables and integral range could be adopted in special cases.

By Taylor's approximation, it could be written as

$$(d\rho)_1 \approx \mathbf{e}_x dx + \frac{1}{2} \frac{\partial \mathbf{e}_x}{\partial x} dx^2 + \mathbf{e}_y dy + \frac{\partial \mathbf{e}_y}{\partial x} dx dy + \frac{1}{2} (\frac{\partial \mathbf{e}_y}{\partial y} + \frac{\partial \mathbf{e}_y}{\partial x \partial y} dx) dy^2 \quad (101)$$

We could also get the differential in path 2,

$$(d\rho)_2 \approx \mathbf{e}_y dy + \frac{1}{2} \frac{\partial \mathbf{e}_y}{\partial y} dy^2 + \mathbf{e}_x dx + \frac{\partial \mathbf{e}_x}{\partial y} dy dx + \frac{1}{2} (\frac{\partial \mathbf{e}_x}{\partial x} + \frac{\partial \mathbf{e}_x}{\partial y \partial x} dy) dx^2 \quad (102)$$

Trimming off the 3-order infinite small quantities, the difference of $(d\rho)_1$ and $(d\rho)_2$ is

$$(d\rho)_1 - (d\rho)_2 \approx (\frac{\partial \mathbf{e}_y}{\partial x} - \frac{\partial \mathbf{e}_x}{\partial y}) dy dx \quad (103)$$

There will be a closure error in close path if the mixed derivatives of bases do not equal.

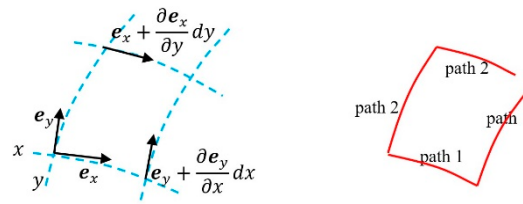


Figure 9. Bases vary in different paths.

3.2.3. Verifications and Discussions

Example 1: Polar coordinate system

Polar coordinate system that we are familiar with is a transformation from its contra variant space (r, θ) , as shown in Figures 10 and 11.

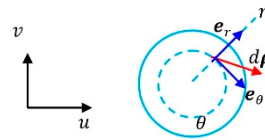


Figure 10. Covariant space.

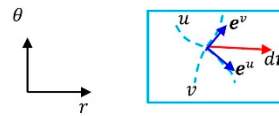


Figure 11. Contra variant space.

A position in contra variant space could be expressed by vector

$$\mathbf{r} = \mathbf{r}(r, \theta) \quad (104)$$

and the differential is

$$d\mathbf{r} = \begin{pmatrix} dr \\ d\theta \end{pmatrix} \quad (105)$$

Corresponding position in covariant space, will be expressed by

$$\boldsymbol{\rho} = \boldsymbol{\rho}(u, v) \quad (106)$$

and the differential is

$$d\boldsymbol{\rho} = \begin{pmatrix} du \\ dv \end{pmatrix} \quad (107)$$

In the contra variant space, the differential distance between two positions could be defined as

$$d\zeta^2 = d\mathbf{r} \cdot d\mathbf{r} = dr^2 + d\theta^2 \quad (108)$$

The system we have used to is the one that has experienced transformation from space (r, θ) with

$$u = r \cos \theta, \quad v = r \sin \theta \quad (109)$$

The bases

$$\mathbf{e}_r = \frac{\partial \boldsymbol{\rho}}{\partial r} = \begin{pmatrix} \cos \theta \\ \sin \theta \end{pmatrix} \quad (110)$$

$$\mathbf{e}_\theta = \frac{\partial \boldsymbol{\rho}}{\partial \theta} = \begin{pmatrix} -r \sin \theta \\ r \cos \theta \end{pmatrix} \quad (111)$$

so that

$$d\boldsymbol{\rho} = \mathbf{e}_r dr + \mathbf{e}_\theta d\theta \quad (112)$$

Thus, there is the covariant distance

$$ds^2 = du^2 + dv^2 = d\boldsymbol{\rho} \cdot d\boldsymbol{\rho} = \mathbf{e}_r \cdot \mathbf{e}_r dr^2 + \mathbf{e}_\theta \cdot \mathbf{e}_\theta d\theta^2 = dr^2 + r^2 d\theta^2 \quad (113)$$

The mixed derivatives of bases

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = \frac{\partial}{\partial \theta} \begin{pmatrix} \cos\theta \\ \sin\theta \end{pmatrix} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \quad (114)$$

$$\frac{\partial \mathbf{e}_\theta}{\partial r} = \frac{\partial}{\partial r} \begin{pmatrix} -r\sin\theta \\ r\cos\theta \end{pmatrix} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \quad (115)$$

We have seen the mixed derivatives got equal

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = \frac{\partial \mathbf{e}_\theta}{\partial r} \quad (116)$$

It could also be verified in Eq. (89) and Eq. (90) that

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \xi_{(r)} \partial \eta_{(\theta)}} \frac{\partial \xi_{(r)}}{\partial r} \frac{\partial \eta_{(\theta)}}{\partial \theta} + \frac{\partial \boldsymbol{\rho}}{\partial \xi_{(r)}} \frac{\partial^2 \xi_{(r)}}{\partial r \partial \theta} \quad (117)$$

and

$$\frac{\partial \mathbf{e}_\theta}{\partial r} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \xi_{(r)}} \frac{\partial \eta_{(\theta)}}{\partial \theta} \frac{\partial \xi_{(r)}}{\partial r} + \frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} \frac{\partial^2 \eta_{(\theta)}}{\partial \theta \partial r} \quad (118)$$

The vector $\boldsymbol{\rho}$ is

$$\boldsymbol{\rho} = \begin{pmatrix} u \\ v \end{pmatrix} = \begin{pmatrix} r\cos\theta \\ r\sin\theta \end{pmatrix} \quad (119)$$

Because $d\xi_{(r)}$ is radius length dr , and $d\eta_{(\theta)}$ is arc length $r d\theta$, then

$$\frac{\partial \boldsymbol{\rho}}{\partial \xi_{(r)}} = \frac{\partial}{\partial r} \begin{pmatrix} r\cos\theta \\ r\sin\theta \end{pmatrix} = \begin{pmatrix} \cos\theta \\ \sin\theta \end{pmatrix} \quad (120)$$

$$\frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} = \frac{\partial}{\partial \theta} \begin{pmatrix} r\cos\theta \\ r\sin\theta \end{pmatrix} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \quad (121)$$

Thus, the first item of Eq. (117) is

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \xi_{(r)} \partial \eta_{(\theta)}} \frac{\partial \xi_{(r)}}{\partial r} \frac{\partial \eta_{(\theta)}}{\partial \theta} = \frac{\partial}{\partial \theta} \begin{pmatrix} \cos\theta \\ \sin\theta \end{pmatrix} \frac{\partial r}{\partial r} \frac{\partial \theta}{\partial \theta} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \quad (122)$$

The second item is

$$\frac{\partial \boldsymbol{\rho}}{\partial \xi_{(r)}} \frac{\partial^2 \xi_{(r)}}{\partial r \partial \theta} = \begin{pmatrix} \cos\theta \\ \sin\theta \end{pmatrix} \frac{\partial}{\partial \theta} \frac{\partial r}{\partial r} = \begin{pmatrix} 0 \\ 0 \end{pmatrix} \quad (123)$$

And the first item of Eq. (118)

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \xi_{(r)}} \frac{\partial \eta_{(\theta)}}{\partial \theta} \frac{\partial \xi_{(r)}}{\partial r} = \frac{\partial}{\partial r} \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \frac{\partial \theta}{\partial \theta} \frac{\partial r}{\partial r} = \begin{pmatrix} 0 \\ 0 \end{pmatrix} \quad (124)$$

The second item

$$\frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} \frac{\partial^2 \eta_{(\theta)}}{\partial \theta \partial r} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \frac{\partial}{\partial r} \frac{\partial \theta}{\partial \theta} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \quad (125)$$

so that

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} + \begin{pmatrix} 0 \\ 0 \end{pmatrix} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \quad (126)$$

and

$$\frac{\partial \mathbf{e}_\theta}{\partial r} = \begin{pmatrix} 0 \\ 0 \end{pmatrix} + \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} = \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \quad (127)$$

Obviously there is

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \xi_{(r)} \partial \eta_{(\theta)}} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \xi_{(r)}} \quad (128)$$

and

$$\frac{\partial \rho}{\partial \xi_{(r)}} \frac{\partial^2 \xi_{(r)}}{\partial r \partial \theta} = \frac{\partial \rho}{\partial \eta_{(\theta)}} \frac{\partial^2 \eta_{(\theta)}}{\partial \theta \partial r} \quad (129)$$

and then

$$\frac{\partial e_r}{\partial \theta} = \frac{\partial e_\theta}{\partial r} \quad (130)$$

Again, we have got the equality of mixed derivatives. But to our surprise is that this solution really subject to condition 2. It is said that the first items of Eq. (117) and Eq. (118) do not equal. One of the reasons in this case, is that there is no super surface.

Example 2: Spherical surface coordinate system

A spherical surface coordinate system is also a transformation of the corresponding contra variant space as in Figures 12 and 13, in which

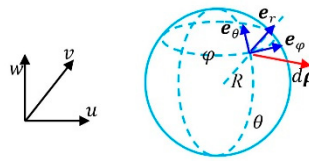


Figure 12. Covariant space.

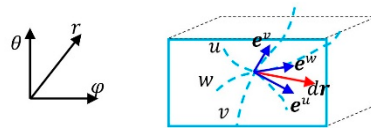


Figure 13. Contra variant space.

$$\mathbf{r} = \mathbf{r}(r, \theta, \varphi)|_{r=R} = \begin{pmatrix} \theta \\ \varphi \end{pmatrix} \Big|_{r=R} \quad (131)$$

Differential distance is

$$d\zeta^2 = d\mathbf{r} \cdot d\mathbf{r} = d\theta^2 + d\varphi^2 \quad (132)$$

And the coordinates of covariant space will be expressed with

$$\boldsymbol{\rho} = \boldsymbol{\rho}(u, v, w) = \begin{pmatrix} u \\ v \\ w \end{pmatrix} \quad (133)$$

The spherical coordinates could be transformed to Cartesian coordinates,

$$u = R \sin \theta \cos \varphi, \quad v = R \sin \theta \sin \varphi, \quad w = R \cos \theta \quad (134)$$

The bases could be defined as

$$\mathbf{e}_\theta = \frac{\partial \boldsymbol{\rho}}{\partial \theta} = \begin{pmatrix} R \cos \theta \cos \varphi \\ R \cos \theta \sin \varphi \\ -R \sin \theta \end{pmatrix}, \quad \mathbf{e}_\varphi = \frac{\partial \boldsymbol{\rho}}{\partial \varphi} = \begin{pmatrix} -R \sin \theta \sin \varphi \\ R \sin \theta \cos \varphi \\ 0 \end{pmatrix} \quad (135)$$

and

$$d\boldsymbol{\rho} = \mathbf{e}_\theta d\theta + \mathbf{e}_\varphi d\varphi \quad (136)$$

Thus, there is the covariant distance

$$ds^2 = d\boldsymbol{\rho} \cdot d\boldsymbol{\rho} = \mathbf{e}_\theta \cdot \mathbf{e}_\theta dr^2 + \mathbf{e}_\varphi \cdot \mathbf{e}_\varphi d\varphi^2 = R^2 d\theta^2 + R^2 \sin^2 \theta d\varphi^2 \quad (137)$$

The derivatives

$$\frac{\partial \mathbf{e}_\theta}{\partial \varphi} = \begin{pmatrix} -R\cos\theta\sin\varphi \\ R\cos\theta\cos\varphi \\ 0 \end{pmatrix}, \quad \frac{\partial \mathbf{e}_\varphi}{\partial \theta} = \begin{pmatrix} -R\cos\theta\sin\varphi \\ R\cos\theta\cos\varphi \\ 0 \end{pmatrix} \quad (138)$$

so that

$$\frac{\partial \mathbf{e}_\theta}{\partial \varphi} = \frac{\partial \mathbf{e}_\varphi}{\partial \theta} \quad (139)$$

It could also be verified in Eq. (89) and Eq. (90) that

$$\frac{\partial \mathbf{e}_\theta}{\partial \varphi} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \zeta_{(\varphi)}} \frac{\partial \eta_{(\theta)}}{\partial \theta} \frac{\partial \zeta_{(\varphi)}}{\partial \varphi} + \frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} \frac{\partial^2 \eta_{(\theta)}}{\partial \theta \partial \varphi} \quad (140)$$

and

$$\frac{\partial \mathbf{e}_\varphi}{\partial \theta} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \zeta_{(\varphi)} \partial \eta_{(\theta)}} \frac{\partial \zeta_{(\varphi)}}{\partial \varphi} \frac{\partial \eta_{(\theta)}}{\partial \theta} + \frac{\partial \boldsymbol{\rho}}{\partial \zeta_{(\varphi)}} \frac{\partial^2 \zeta_{(\varphi)}}{\partial \varphi \partial \theta} \quad (141)$$

The vector $\boldsymbol{\rho}$ is

$$\boldsymbol{\rho} = \begin{pmatrix} u \\ v \\ w \end{pmatrix} = \begin{pmatrix} R\sin\theta\cos\varphi \\ R\sin\theta\sin\varphi \\ R\cos\theta \end{pmatrix} \quad (142)$$

Because $d\eta_{(\theta)}$ is arc length $Rd\theta$, and $d\zeta_{(\varphi)}$ is arc length $Rd\varphi$, then

$$\frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} = \frac{\partial}{R\partial\theta} \begin{pmatrix} R\sin\theta\cos\varphi \\ R\sin\theta\sin\varphi \\ R\cos\theta \end{pmatrix} = \begin{pmatrix} \cos\theta\cos\varphi \\ \cos\theta\sin\varphi \\ -\sin\theta \end{pmatrix} \quad (143)$$

$$\frac{\partial \boldsymbol{\rho}}{\partial \zeta_{(\varphi)}} = \frac{\partial}{R\partial\varphi} \begin{pmatrix} R\sin\theta\cos\varphi \\ R\sin\theta\sin\varphi \\ R\cos\theta \end{pmatrix} = \begin{pmatrix} -\sin\theta\sin\varphi \\ \sin\theta\cos\varphi \\ 0 \end{pmatrix} \quad (144)$$

Thus, the first item of Eq. (140) is

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \zeta_{(\varphi)}} \frac{\partial \eta_{(\theta)}}{\partial \theta} \frac{\partial \zeta_{(\varphi)}}{\partial \varphi} = \frac{\partial}{R\partial\theta} \begin{pmatrix} \cos\theta\cos\varphi \\ \cos\theta\sin\varphi \\ -\sin\theta \end{pmatrix} \frac{R\partial\theta}{\partial\theta} \frac{R\partial\varphi}{\partial\varphi} = \begin{pmatrix} -R\cos\theta\sin\varphi \\ R\cos\theta\cos\varphi \\ 0 \end{pmatrix} \quad (145)$$

The second item is

$$\frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} \frac{\partial^2 \eta_{(\theta)}}{\partial \theta \partial \varphi} = \begin{pmatrix} \cos\theta\cos\varphi \\ \cos\theta\sin\varphi \\ -\sin\theta \end{pmatrix} \frac{\partial}{\partial\theta} \frac{R\partial\varphi}{\partial\varphi} = \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix} \quad (146)$$

And the first item of Eq. (141)

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \zeta_{(\varphi)} \partial \eta_{(\theta)}} \frac{\partial \zeta_{(\varphi)}}{\partial \varphi} \frac{\partial \eta_{(\theta)}}{\partial \theta} = \frac{\partial}{R\partial\theta} \begin{pmatrix} -\sin\theta\sin\varphi \\ \sin\theta\cos\varphi \\ 0 \end{pmatrix} \frac{R\partial\varphi}{\partial\varphi} \frac{R\partial\theta}{\partial\theta} = \begin{pmatrix} -R\cos\theta\sin\varphi \\ R\cos\theta\cos\varphi \\ 0 \end{pmatrix} \quad (147)$$

The second item

$$\frac{\partial \boldsymbol{\rho}}{\partial \zeta_{(\varphi)}} \frac{\partial^2 \zeta_{(\varphi)}}{\partial \varphi \partial \theta} = \begin{pmatrix} -\sin\theta\sin\varphi \\ \sin\theta\cos\varphi \\ 0 \end{pmatrix} \frac{\partial}{\partial\varphi} \frac{R\partial\theta}{\partial\theta} = \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix} \quad (148)$$

With Eq. (145) to Eq. (148) we found that

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \zeta_{(\varphi)}} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \zeta_{(\varphi)} \partial \eta_{(\theta)}} \quad (149)$$

and

$$\frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} \frac{\partial^2 \eta_{(\theta)}}{\partial \theta \partial \varphi} = \frac{\partial \boldsymbol{\rho}}{\partial \zeta_{(\varphi)}} \frac{\partial^2 \zeta_{(\varphi)}}{\partial \varphi \partial \theta} \quad (150)$$

so that there is

$$\frac{\partial \mathbf{e}_\theta}{\partial \varphi} = \frac{\partial \mathbf{e}_\varphi}{\partial \theta} \quad (151)$$

One can see that this is of condition 1.

Example 3: Spherical coordinate system

A spherical coordinate system is also a transformation of the corresponding contra variant space as in Figures 14 and 15, in which

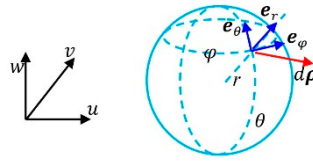


Figure 14. Covariant spac.

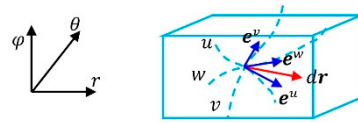


Figure 15. Contra variant space.

$$\mathbf{r} = \mathbf{r}(r, \theta, \varphi) = \begin{pmatrix} r \\ \theta \\ \varphi \end{pmatrix} \quad (152)$$

Differential distance is

$$d\zeta^2 = d\mathbf{r} \cdot d\mathbf{r} = dr^2 + d\theta^2 + d\varphi^2 \quad (153)$$

And the coordinates of covariant space will be expressed with

$$\boldsymbol{\rho} = \boldsymbol{\rho}(u, v, w) = \begin{pmatrix} u \\ v \\ w \end{pmatrix} \quad (154)$$

The spherical coordinates could be transformed to Cartesian coordinates,

$$u = r \sin \theta \cos \varphi, \quad v = r \sin \theta \sin \varphi, \quad w = r \cos \theta \quad (155)$$

The bases could be defined as

$$\mathbf{e}_r = \frac{\partial \boldsymbol{\rho}}{\partial r} = \begin{pmatrix} \sin \theta \cos \varphi \\ \sin \theta \sin \varphi \\ \cos \theta \end{pmatrix}, \quad \mathbf{e}_\theta = \frac{\partial \boldsymbol{\rho}}{\partial \theta} = \begin{pmatrix} r \cos \theta \cos \varphi \\ r \cos \theta \sin \varphi \\ -r \sin \theta \end{pmatrix}, \quad \mathbf{e}_\varphi = \frac{\partial \boldsymbol{\rho}}{\partial \varphi} = \begin{pmatrix} -r \sin \theta \sin \varphi \\ r \sin \theta \cos \varphi \\ 0 \end{pmatrix} \quad (156)$$

and

$$d\boldsymbol{\rho} = \mathbf{e}_r dr + \mathbf{e}_\theta d\theta + \mathbf{e}_\varphi d\varphi \quad (157)$$

Thus, there is the covariant distance

$$ds^2 = d\boldsymbol{\rho} \cdot d\boldsymbol{\rho} = \mathbf{e}_r \cdot \mathbf{e}_r dr^2 + \mathbf{e}_\theta \cdot \mathbf{e}_\theta d\theta^2 + \mathbf{e}_\varphi \cdot \mathbf{e}_\varphi d\varphi^2 = dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\varphi^2 \quad (158)$$

The derivatives

$$\begin{aligned} \frac{\partial \mathbf{e}_r}{\partial \theta} &= \begin{pmatrix} \cos \theta \cos \varphi \\ \cos \theta \sin \varphi \\ -\sin \theta \end{pmatrix}, \quad \frac{\partial \mathbf{e}_\theta}{\partial r} = \begin{pmatrix} \cos \theta \cos \varphi \\ \cos \theta \sin \varphi \\ -\sin \theta \end{pmatrix}, \quad \frac{\partial \mathbf{e}_\theta}{\partial \varphi} = \begin{pmatrix} -r \cos \theta \sin \varphi \\ r \cos \theta \cos \varphi \\ 0 \end{pmatrix} \\ \frac{\partial \mathbf{e}_\varphi}{\partial \theta} &= \begin{pmatrix} -r \cos \theta \sin \varphi \\ r \cos \theta \cos \varphi \\ 0 \end{pmatrix}, \quad \frac{\partial \mathbf{e}_\varphi}{\partial r} = \begin{pmatrix} -\sin \theta \sin \varphi \\ \sin \theta \cos \varphi \\ 0 \end{pmatrix}, \quad \frac{\partial \mathbf{e}_r}{\partial \varphi} = \begin{pmatrix} -\sin \theta \sin \varphi \\ \sin \theta \cos \varphi \\ 0 \end{pmatrix} \end{aligned} \quad (159)$$

so that

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = \frac{\partial \mathbf{e}_\theta}{\partial r}, \quad \frac{\partial \mathbf{e}_\theta}{\partial \varphi} = \frac{\partial \mathbf{e}_\varphi}{\partial \theta}, \quad \frac{\partial \mathbf{e}_\varphi}{\partial r} = \frac{\partial \mathbf{e}_r}{\partial \varphi} \quad (160)$$

It could also be verified in Eq. (89) and Eq. (90) that

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \xi_{(r)} \partial \eta_{(\theta)}} \frac{\partial \xi_{(r)}}{\partial r} \frac{\partial \eta_{(\theta)}}{\partial \theta} + \frac{\partial \boldsymbol{\rho}}{\partial \xi_{(r)}} \frac{\partial^2 \xi_{(r)}}{\partial r \partial \theta} \quad (161)$$

and

$$\frac{\partial \mathbf{e}_\theta}{\partial r} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \xi_{(r)}} \frac{\partial \eta_{(\theta)}}{\partial \theta} \frac{\partial \xi_{(r)}}{\partial r} + \frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} \frac{\partial^2 \eta_{(\theta)}}{\partial \theta \partial r} \quad (162)$$

The vector $\boldsymbol{\rho}$ is

$$\boldsymbol{\rho} = \begin{pmatrix} u \\ v \\ w \end{pmatrix} = \begin{pmatrix} r \sin \theta \cos \varphi \\ r \sin \theta \sin \varphi \\ r \cos \theta \end{pmatrix} \quad (163)$$

Because $d\xi_{(r)}$ is radius length dr , $d\eta_{(\theta)}$ is arc length $r d\theta$, and $d\zeta_{(\varphi)}$ is arc length $r d\varphi$, then

$$\frac{\partial \boldsymbol{\rho}}{\partial \xi_{(r)}} = \frac{\partial}{\partial r} \begin{pmatrix} r \sin \theta \cos \varphi \\ r \sin \theta \sin \varphi \\ r \cos \theta \end{pmatrix} = \begin{pmatrix} \sin \theta \cos \varphi \\ \sin \theta \sin \varphi \\ \cos \theta \end{pmatrix} \quad (164)$$

$$\frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} = \frac{\partial}{r \partial \theta} \begin{pmatrix} r \sin \theta \cos \varphi \\ r \sin \theta \sin \varphi \\ r \cos \theta \end{pmatrix} = \begin{pmatrix} \cos \theta \cos \varphi \\ \cos \theta \sin \varphi \\ -\sin \theta \end{pmatrix} \quad (165)$$

$$\frac{\partial \boldsymbol{\rho}}{\partial \zeta_{(\varphi)}} = \frac{\partial}{r \partial \varphi} \begin{pmatrix} r \sin \theta \cos \varphi \\ r \sin \theta \sin \varphi \\ r \cos \theta \end{pmatrix} = \begin{pmatrix} -\sin \theta \sin \varphi \\ \sin \theta \cos \varphi \\ 0 \end{pmatrix} \quad (166)$$

Thus, the first item of Eq. (161) is

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \xi_{(r)} \partial \eta_{(\theta)}} \frac{\partial \xi_{(r)}}{\partial r} \frac{\partial \eta_{(\theta)}}{\partial \theta} = \frac{\partial}{r \partial \theta} \begin{pmatrix} \sin \theta \cos \varphi \\ \sin \theta \sin \varphi \\ \cos \theta \end{pmatrix} \frac{\partial r}{\partial r} \frac{\partial \theta}{\partial \theta} = \begin{pmatrix} \cos \theta \cos \varphi \\ \cos \theta \sin \varphi \\ -\sin \theta \end{pmatrix} \quad (167)$$

The second item is

$$\frac{\partial \boldsymbol{\rho}}{\partial \xi_{(r)}} \frac{\partial^2 \xi_{(r)}}{\partial r \partial \theta} = \begin{pmatrix} \sin \theta \cos \varphi \\ \sin \theta \sin \varphi \\ \cos \theta \end{pmatrix} \frac{\partial}{r \partial \theta} \frac{\partial r}{\partial r} = \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix} \quad (168)$$

And the first item of Eq. (162)

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \xi_{(r)}} \frac{\partial \eta_{(\theta)}}{\partial \theta} \frac{\partial \xi_{(r)}}{\partial r} = \frac{\partial}{\partial r} \begin{pmatrix} \cos \theta \cos \varphi \\ \cos \theta \sin \varphi \\ -\sin \theta \end{pmatrix} \frac{r \partial \theta}{\partial \theta} \frac{\partial r}{\partial r} = \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix} \quad (169)$$

The second item

$$\frac{\partial \boldsymbol{\rho}}{\partial \eta_{(\theta)}} \frac{\partial^2 \eta_{(\theta)}}{\partial \theta \partial r} = \begin{pmatrix} \cos \theta \cos \varphi \\ \cos \theta \sin \varphi \\ -\sin \theta \end{pmatrix} \frac{\partial}{\partial r} \frac{r \partial \theta}{\partial \theta} = \begin{pmatrix} \cos \theta \cos \varphi \\ \cos \theta \sin \varphi \\ -\sin \theta \end{pmatrix} \quad (170)$$

With Eq. (167) and Eq. (170) we found that

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \xi_{(r)} \partial \eta_{(\theta)}} \neq \frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \xi_{(r)}} \quad (171)$$

but there is

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = \frac{\partial \mathbf{e}_\theta}{\partial r} \quad (172)$$

One can also calculate that

$$\frac{\partial^2 \boldsymbol{\rho}}{\partial \eta_{(\theta)} \partial \zeta_{(\varphi)}} = \frac{\partial^2 \boldsymbol{\rho}}{\partial \zeta_{(\varphi)} \partial \eta_{(\theta)}}, \quad \frac{\partial^2 \boldsymbol{\rho}}{\partial \zeta_{(\varphi)} \partial \xi_{(r)}} \neq \frac{\partial^2 \boldsymbol{\rho}}{\partial \xi_{(r)} \partial \zeta_{(\varphi)}} \quad (173)$$

At the end we can obtain

$$\frac{\partial \mathbf{e}_\theta}{\partial \varphi} = \frac{\partial \mathbf{e}_\varphi}{\partial \theta}, \quad \frac{\partial \mathbf{e}_\varphi}{\partial r} = \frac{\partial \mathbf{e}_r}{\partial \varphi} \quad (174)$$

One can see that one of them is of condition 1 and the others of them are of condition 2.

Additional discussion: deformed bases of example 3

If one of the bases in example 3 be set deformed as

$$\mathbf{e}_r = f(r) \begin{pmatrix} \sin\theta\cos\varphi \\ \sin\theta\sin\varphi \\ \cos\theta \end{pmatrix} \quad (175)$$

where, $f(r) \neq 1$ is a function of coordinate r .

One will find that the derivatives

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = f(r) \begin{pmatrix} \cos\theta\cos\varphi \\ \cos\theta\sin\varphi \\ -\sin\theta \end{pmatrix} + \frac{\partial f(r)}{\partial \theta} \begin{pmatrix} \sin\theta\cos\varphi \\ \sin\theta\sin\varphi \\ \cos\theta \end{pmatrix} = f(r) \begin{pmatrix} \cos\theta\cos\varphi \\ \cos\theta\sin\varphi \\ -\sin\theta \end{pmatrix} \quad (176)$$

and

$$\frac{\partial \mathbf{e}_r}{\partial \varphi} = f(r) \begin{pmatrix} -\sin\theta\sin\varphi \\ \sin\theta\cos\varphi \\ 0 \end{pmatrix} + \frac{\partial f(r)}{\partial \varphi} \begin{pmatrix} \sin\theta\cos\varphi \\ \sin\theta\sin\varphi \\ \cos\theta \end{pmatrix} = f(r) \begin{pmatrix} -\sin\theta\sin\varphi \\ \sin\theta\cos\varphi \\ 0 \end{pmatrix} \quad (177)$$

while $\frac{\partial \mathbf{e}_\theta}{\partial r}$ and $\frac{\partial \mathbf{e}_\varphi}{\partial r}$ will still keep the results as in Eq. (159), that will cause

$$\frac{\partial \mathbf{e}_r}{\partial \theta} \neq \frac{\partial \mathbf{e}_\theta}{\partial r}, \quad \frac{\partial \mathbf{e}_\varphi}{\partial r} \neq \frac{\partial \mathbf{e}_r}{\partial \varphi} \quad (178)$$

This result reminds us that the same performance would have happened in gravitational space time that will be put into discussions in next section.

4. Metrics and Covariant Derivatives in Space Time

4.1. Metrics in Pseudo Riemannian Space

Pseudo Riemannian space is raised for the description of space time for general relativity, after Minkowski space for special relativity [8,9]. For Minkowski space, invariant distance for flat space time could be written as

$$ds^2 = -c^2 dt^2 + dx^2 + dy^2 + dz^2 \quad (179)$$

We know that there is a minus signal in the equation. The minus signal does not come from transformations of spaces or coordinates. It is a kind of mathematics and physics setting. Former researchers have made efforts on this topic, for example, the concept of plural employed to reform the base \mathbf{e}_0 [6]. But plural bases for relativity is not a good idea. Another treatment is to define $x^0 = ict$, which looks like more reasonable [7]. The more important is that the metrics for this condition should be carefully treated so that the metric g_{00} will not be minus.

In general relativity, spherical coordinates are usually suggested for one source problem. So that there will be contra variant space (t, r, θ, φ) and covariant space $(\tau, \rho, \theta, \varphi)$. The invariant distance of Schwarzschild solution is

$$ds^2 = d\boldsymbol{\rho} \cdot d\boldsymbol{\rho} = -(1 - \frac{r^*}{r})(cdt)^2 + (1 - \frac{r^*}{r})^{-1} dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\varphi^2 \quad (180)$$

Most of publications have set the metric [3,4,10,11,12] to be

$$g_{00} = -(1 - \frac{r^*}{r}) \quad (181)$$

So that there could be a brief expression of invariant distance

$$ds^2 = g_{ij} dx^i dx^j \quad (182)$$

As we have discussed, that will cause plural bases. Anyway, minus metrics is improper. In fact, it is one of the reasons that cause the wrong result of acceleration calculation in Eq. (43).

Considering the above discussions, I prefer to give the invariant distance of one source field as

$$ds^2 = -g_{00}(cdt)^2 + g_{11}dr^2 + g_{22}d\theta^2 + g_{33}d\varphi^2 \quad (183)$$

Thus, we will prevent from plural items. That doesn't hurt the expressions of relativity. Metrics and bases defined are just employed for the transformation from Minkowski space time to pseudo Riemannian space time. Even if one persists the coordinate $x^0 = ict$ the invariant distance will still keep the expression. Additionally, I still suggest $x^0 = ct$ in use that will bring about convenience in most cases.

It should be highlighted that the bases still would have sophisticated forms that

$$\mathbf{e}_0 = \begin{pmatrix} (1 - \frac{r^*}{r})^{1/2} \\ 0 \\ 0 \\ 0 \end{pmatrix}, \mathbf{e}_1 = \begin{pmatrix} 0 \\ \sin\theta\cos\varphi(1 - \frac{r^*}{r})^{-1/2} \\ \sin\theta\sin\varphi(1 - \frac{r^*}{r})^{-1/2} \\ \cos\theta(1 - \frac{r^*}{r})^{-1/2} \end{pmatrix}, \mathbf{e}_2 = \begin{pmatrix} 0 \\ r\cos\theta\cos\varphi \\ r\cos\theta\sin\varphi \\ -r\sin\theta \end{pmatrix}, \mathbf{e}_3 = \begin{pmatrix} 0 \\ -r\sin\theta\sin\varphi \\ r\sin\theta\cos\varphi \\ 0 \end{pmatrix} \quad (184)$$

That is because the spherical space we have discussed has already experienced coordinates transformations. The transformed coordinates are not real spherical coordinates, they are the Cartesian coordinates (u, v, w) expressed by parameters of r, θ, φ , which could be called parameterized Cartesian coordinates. One can learn from previous section for the reasons.

It could be calculated that

$$\frac{\partial \mathbf{e}_1}{\partial t} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}, \frac{\partial \mathbf{e}_1}{\partial \theta} = \begin{pmatrix} 0 \\ \cos\theta\cos\varphi(1 - \frac{r^*}{r})^{-1/2} \\ \cos\theta\sin\varphi(1 - \frac{r^*}{r})^{-1/2} \\ -\sin\theta(1 - \frac{r^*}{r})^{-1/2} \end{pmatrix}, \frac{\partial \mathbf{e}_1}{\partial \varphi} = \begin{pmatrix} 0 \\ -\sin\theta\sin\varphi(1 - \frac{r^*}{r})^{-1/2} \\ \sin\theta\cos\varphi(1 - \frac{r^*}{r})^{-1/2} \\ 0 \end{pmatrix}$$

$$\frac{\partial \mathbf{e}_0}{\partial r} = \begin{pmatrix} \frac{r^*}{2r^2}(1 - \frac{r^*}{r})^{-1/2} \\ 0 \\ 0 \\ 0 \end{pmatrix}, \frac{\partial \mathbf{e}_2}{\partial r} = \begin{pmatrix} 0 \\ \cos\theta\cos\varphi \\ \cos\theta\sin\varphi \\ -\sin\theta \end{pmatrix}, \frac{\partial \mathbf{e}_3}{\partial r} = \begin{pmatrix} 0 \\ -\sin\theta\sin\varphi \\ \sin\theta\cos\varphi \\ 0 \end{pmatrix} \quad (185)$$

there are

$$\frac{\partial \mathbf{e}_1}{\partial t} \neq \frac{\partial \mathbf{e}_0}{\partial r}, \frac{\partial \mathbf{e}_1}{\partial \theta} \neq \frac{\partial \mathbf{e}_2}{\partial r}, \frac{\partial \mathbf{e}_1}{\partial \varphi} \neq \frac{\partial \mathbf{e}_3}{\partial r} \quad (186)$$

In general relativity, it could be suggested to simplify the presentation that we only focus on the gravitational transformation. So that the contra variant space is no longer of real spherical coordinates, which could be set as parameterized Cartesian coordinates, and the covariant space could also be the form of parameterized Cartesian. Thus, the invariant distance could be expressed as

$$ds^2 = -g_{00}(cdt)^2 + g_{11}dr^2 + g_{22}r^2d\theta^2 + g_{33}r^2\sin^2\theta d\varphi^2 \quad (187)$$

In which, $g_{22} = g_{33} = 1$.

These metrics could be called the gravitational metrics, while the metrics previous could be called total metrics and the metrics before gravitational transformation could be called original metrics or spherical transformation metrics.

For gravitational metrics of Schwarzschild solution, we have the bases

$$\mathbf{e}_0 = \begin{pmatrix} (1 - \frac{r^*}{r})^{1/2} \\ 0 \\ 0 \\ 0 \end{pmatrix}, \mathbf{e}_1 = \begin{pmatrix} 0 \\ (1 - \frac{r^*}{r})^{-1/2} \\ 0 \\ 0 \end{pmatrix}, \mathbf{e}_2 = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}, \mathbf{e}_3 = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix} \quad (188)$$

It could be calculated that

$$\frac{\partial \mathbf{e}_0}{\partial r} = \begin{pmatrix} \frac{r^*}{2r^2}(1 - \frac{r^*}{r})^{-1/2} \\ 0 \\ 0 \\ 0 \end{pmatrix}, \frac{\partial \mathbf{e}_1}{\partial t} = 0 \quad (189)$$

so that

$$\frac{\partial \mathbf{e}_0}{\partial r} \neq \frac{\partial \mathbf{e}_1}{\partial t} \quad (190)$$

This is of the condition 3 that we have discussed in section 3.2.2. One can also calculate the inequality of mixed derivatives of bases of total metrics.

That reminds us that the inequality of mixed derivatives of bases will cause closure errors in space time, which would be left for more discussions elsewhere.

4.2. Discussions on Bases, Tensors and Their Derivatives

Tensors could be recognized as the quantities relating to coordinates in space time. Case a tensor varies in space time, the variation rate could be inspected by derivation. The simplest tensor is position vector $\boldsymbol{\rho}(x_0, x_1, x_2, x_3,)$. You have seen that we are going to use middle subscriptions to express coordinates and tensors in covariant space time, though they are rarely mentioned in most of references. To study its variation in space time, one could define the distance variation ratio to form the bases

$$\mathbf{e}_i = \frac{\partial \boldsymbol{\rho}}{\partial x^i} \quad (191)$$

For any tensors involved with bases, such as a proper tensor

$$\mathbf{A} = A^i \mathbf{e}_i \quad (192)$$

where, A^i is number i component of the total contra variant tensor.

The derivatives of the tensor are

$$\frac{\partial \mathbf{A}}{\partial x^j} = \mathbf{e}_i \frac{\partial A^i}{\partial x^j} + A^i \frac{\partial \mathbf{e}_i}{\partial x^j} \quad (193)$$

The differential of a proper tensor $d\mathbf{A}$ could be defined to be covariant differential labeled as $D\mathbf{A}$, and then the derivative $\frac{\partial \mathbf{A}}{\partial x^j}$ to be covariant derivative $\frac{D\mathbf{A}}{\partial x^j}$

$$\frac{D\mathbf{A}}{\partial x^j} = \frac{\partial \mathbf{A}}{\partial x^j} = \mathbf{e}_i \frac{\partial A^i}{\partial x^j} + A^i \frac{\partial \mathbf{e}_i}{\partial x^j} \quad (194)$$

In these equations, the middle subscriptions of proper tensors maybe neglected conventionally so that it is expressed as \mathbf{A} . And the tensor \mathbf{A} could be called proper tensor because A_i has already been named covariant tensor conventionally. Because A^i or A_i is just a component, we could imagine that there must be the total quantity. That will be expressed to be $\mathbf{A}^\dagger = (A^0, A^1, A^2, A^3)$ or $\mathbf{A}_\dagger = (A_0, A_1, A_2, A_3)$ for convenience.

Bases sometimes look like one-order tensors since they have a single index in expressions. But in fact, they really are two-order mixed tensors. For example, a component of contra variant base of collinear transformation, i.e. coordinate lines of contra variant space and covariant space coinciding, could be written as

$$e^\mu \nu = \frac{\partial x^\mu}{\partial x^\nu} \quad (195)$$

where the contra variant base and proper coordinate differential are all labeled with middle index ν .

So that the base really is

$$\mathbf{e}^\mu = \begin{pmatrix} e^{\mu 0} \\ e^{\mu 1} \\ e^{\mu 2} \\ e^{\mu 3} \end{pmatrix} \quad (196)$$

If $e^\mu \nu = 0$ with $\mu \neq \nu$, we could use e^μ instead of $e^\mu \mu$ for convenience, as it is the only nonvanishing component.

For any one order tensors there is a transformation

$$A^\mu = \mathbf{e}^\mu \cdot \mathbf{A}, \text{ or } A_\mu = \mathbf{e}_\mu \cdot \mathbf{A} \quad (197)$$

But for two order tensors, there are some things different. For example, a component of contra variant velocity could be transformed from proper velocity

$$V^\mu = \frac{dx^\mu}{d\tau} = \mathbf{e}^\mu \cdot \frac{d\mathbf{p}}{d\tau} = \mathbf{e}^\mu \cdot \mathbf{V} \quad (198)$$

We know that it may be seen as one-order tensor in practice, but it is really two order tensor.

As for velocity totally in contra variant space as $V_0^1 = \frac{dx^1}{dt}$, that would not be worked out by direct vector product.

In fact, it could be composed independently by dx^1 and dt .

$$V_0^1 = \frac{dx^1}{dt} = \frac{\mathbf{e}^1 \cdot d\mathbf{p}}{\mathbf{e}^0 \cdot d\mathbf{p}} = \frac{\mathbf{e}^1 dx^1}{\mathbf{e}^0 d\tau} = \mathbf{e}^1 \mathbf{e}_0 \frac{dx^1}{d\tau} \quad (199)$$

It is impossible to get $\mathbf{e}^1 \mathbf{e}_0$ from $\mathbf{e}^1 \cdot \mathbf{e}_0$ and the latter is zero. Notwithstanding, a velocity is a derivative on matter's trajectory rather than a direct derivative. That will be further discussed in next sections.

4.3. Derivation via Christoffel Symbols

Christoffel symbols were put forward to perform geometrical relationship that takes similar effects with that of bases. They are defined in the equations [3,13]

$$\frac{\partial \mathbf{e}^\mu}{\partial x^\lambda} = -\Gamma_{\lambda\nu}^\mu \mathbf{e}^\nu \quad \text{or} \quad \frac{\partial \mathbf{e}_\mu}{\partial x^\lambda} = \Gamma_{\lambda\mu}^\nu \mathbf{e}_\nu \quad (200)$$

Taking the first one for example, the purpose of the equation is to consider the derivatives to be a function of bases, so that the right hand item is really a kind of trivial types. In the summation items, \mathbf{e}^ν just act as direction indicators that would give out whole basic vectors of entire dimensions. And then, $\Gamma_{\lambda\nu}^\mu$ provide the coefficients of all directions. It is said, this definition has just provided an error-free frame for the functions of derivatives. It means there may be redundant designs for the coefficients.

Since there is probability of inequality of mixed derivatives of bases, we should define a specific sequence for subscripts of Christoffel symbols. For the traditional reasons, $\Gamma_{\lambda\nu}^\mu$ will be defined as the coefficient of a derivative of \mathbf{e}^μ that is derivated by x^λ , on a direction of \mathbf{e}^ν , that requires unexchangeable subscripts of $\Gamma_{\lambda\nu}^\mu$.

We know that a covariant differential is exactly a differential of a proper tensor

$$\frac{DA}{\partial x^\lambda} = \frac{\partial A}{\partial x^\lambda} \quad (201)$$

This highlightable concept is essentially carried out to perform general covariance.

The contra variant form also performs the same covariance as that

$$\frac{DA^\mu}{\partial x^\lambda} = \mathbf{e}^\mu \cdot \frac{DA}{\partial x^\lambda} \quad (202)$$

You might have found that the tensor component has been expressed by a total tensor is exactly partial expression. It is just of traditional operations. One can of course carry out whole form expression of \mathbf{A}^\dagger expressed by \mathbf{A} with base matrix $[\mathbf{e}^\dagger]$. But too more renovations in a performance will bring about more reading difficulties. So I prefer to present equations in traditional forms as far as possible.

A component of contra variant tensor transformed from covariant one

$$A^\mu = \mathbf{e}^\mu \cdot \mathbf{A} \quad (203)$$

Its derivatives is

$$\frac{\partial A^\mu}{\partial x^\nu} = \frac{\partial}{\partial x^\nu} (\mathbf{e}^\mu \cdot \mathbf{A}) = \frac{\partial \mathbf{e}^\mu}{\partial x^\nu} \cdot \mathbf{A} + \mathbf{e}^\mu \cdot \frac{\partial \mathbf{A}}{\partial x^\nu} = \frac{\partial \mathbf{e}^\mu}{\partial x^\nu} \cdot \mathbf{A} + \frac{DA^\mu}{\partial x^\nu} \quad (204)$$

so that

$$\frac{DA^\mu}{\partial x^\nu} = \frac{\partial A^\mu}{\partial x^\nu} - \frac{\partial \mathbf{e}^\mu}{\partial x^\nu} \cdot \mathbf{A} = \frac{\partial A^\mu}{\partial x^\nu} + \Gamma_{\nu\lambda}^\mu A^\lambda \quad (205)$$

It is easy to study those covariant derivatives for covariant tensors

$$\frac{DA_\mu}{dx^\nu} = \frac{\partial A_\mu}{\partial x^\nu} - \frac{\partial e_\mu}{\partial x^\nu} \cdot \mathbf{A} = \frac{\partial A_\mu}{\partial x^\nu} - \Gamma_{\nu\mu}^\lambda A_\lambda \quad (206)$$

We have seen that the methodologies of Christoffel symbols and the derivation directly from bases are actually equivalent treatments that present the covariant derivatives. That of course may be use to inspect the problems of equations of Christoffel symbols. Since we have known that part of Christoffel symbols with mixed subscripts do not equal in space time, it is necessary to do more discussions.

It is convenient to discuss the case that a space is defined by orthogonal bases. In practice, metrics are usually taken in to Christoffel connections analysis. For a series of bases $\mathbf{e}_0, \mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3$, there is the metric

$$g_{ij} = \mathbf{e}_i \cdot \mathbf{e}_j \quad (207)$$

For whole orthogonal coordinate spaces,

$$\begin{cases} g_{ij} = 0, i \neq j \\ g_{ij} = \mathbf{e}_i \cdot \mathbf{e}_i \neq 0, i = j \end{cases} \quad (208)$$

The nonvanishing derivatives

$$\frac{\partial g_{ij}}{\partial x^k} = \frac{\partial \mathbf{e}_i}{\partial x^k} \cdot \mathbf{e}_j + \frac{\partial \mathbf{e}_j}{\partial x^k} \cdot \mathbf{e}_i = 2 \frac{\partial \mathbf{e}_i}{\partial x^k} \cdot \mathbf{e}_j, i = j \quad (209)$$

Then the equation could have nonvanishing value, that

$$\frac{\partial g_{ij}}{\partial x^\lambda} = 2\Gamma_{\lambda i}^k \mathbf{e}_k \cdot \mathbf{e}_j = 2\Gamma_{\lambda i}^k g_{kj}, i = j \quad (210)$$

Then

$$\Gamma_{\lambda i}^k = \frac{1}{2} g^{kj} \frac{\partial g_{ij}}{\partial x^\lambda}, i = j \quad (211)$$

This equation could be called revised equation for Christoffel symbols in general relativity. In the equation, it also need $k = j$ for the $\Gamma_{\lambda i}^k$ to get nonvanishing.

For a result, this equation could be verified in a covariant derivative directly as that in Eq. (206) that

$$\Gamma_{\lambda i}^k A_k = \frac{1}{2} g^{kj} \frac{\partial g_{ij}}{\partial x^\lambda} A_k = \frac{1}{2} (\mathbf{e}^k \cdot \mathbf{e}^j) (2 \frac{\partial \mathbf{e}_i}{\partial x^\lambda} \cdot \mathbf{e}_j) A_k = A_k \mathbf{e}^k \cdot \frac{\partial \mathbf{e}_i}{\partial x^\lambda} = \frac{\partial \mathbf{e}_i}{\partial x^\lambda} \cdot \mathbf{A} \quad (212)$$

It should be pointed out that the Christoffel symbols are not necessary because that the issue only started from derivatives of bases, as consequences they surely might be taken the place by the operations of bases.

4.4. Derivatives on Matter's Trajectory

The calculation of time derivatives may cause to another mathematical abuse in classical theory. For a one source field, it could be seen as rest field in general. Thus, a time derivative of a field quantity should be zero. Case a matter moves in a space, it is the issue that the matter changes its position in a time interval and forms a motion trajectory. In this condition, to learn the acceleration is to study the position variation rather than field variation. It is one of the reasons that make errors in covariant derivative calculation in Eq. (43), in that the direct derivative has been used instead of trajectory derivative.

It is valuable to reclassify tensors to be field tensors and motion tensors, thus field tensors may vary with field while motion tensors should vary both with field and matter's motion. For example, the bases only depend on gravitational field, while velocity of a matter may vary due to positions changed. For example, case in one source field, a space derivative of a base may be nonvanishing, but a time derivative of a base must be zero, nevertheless, the base relating to a matter moving in space time would vary because the coordinates varied on trajectory.

A trajectory of motions of a matter should be a directed curve line in a space, from the start to the end. It is said that the trajectory must be single parametrical curve line. Theoretically, the parameter maybe natural i.e. the line it pass through, and as same it could be time that the motion experienced. What worthy of highlight is that these parameters are simultaneous. That is said a record of the parameter corresponds to a sole record of another. The parameter indicates the sequences.

It is no harm to discuss the trajectory vector λ as a curve line in contra variant space, in that the trajectory is a function of single variable. There is

$$\lambda(x^0, x^1, x^2, x^3) = \lambda(\lambda) = \lambda(t) \quad (213)$$

where λ is the length of trajectory.

It is said that the trajectory is of

$$x^0 = x^0(t), \quad x^1 = x^1(t), \quad x^2 = x^2(t), \quad x^3 = x^3(t) \quad (214)$$

A tensor variation ratio during a time interval on trajectory could be defined to be trajectory derivative that

$$\left(\frac{DA}{dx^\mu}\right)_{tr} = \frac{DA}{d\lambda} \frac{d\lambda}{dx^\mu} \quad (215)$$

For example

$$\left(\frac{DA}{dt}\right)_{tr} = \frac{DA}{d\lambda} \frac{d\lambda}{dt} \quad (216)$$

where, Einstein summation convention does not act on double λ because trajectory is just a single line. The differential length $d\lambda$ is the differential of matter's trajectory, so that $\frac{DA}{d\lambda} d\lambda$ is the covariant differential of tensor A between two neighbourhood positions on the trajectory. Thus, the so called trajectory derivative is really a kind of line derivative that is derivated by a parameter.

It should be noted that there is the substantial difference between trajectory derivatives and original derivatives. A differential on trajectory is the distance interval that a matter has past across, so that the velocity is a trajectory derivative

$$V_0^\dagger = \left(\frac{d\lambda}{dt}\right)_{tr} = \frac{d\lambda}{dt} = \left(\frac{dx^0(t)}{dt}, \frac{dx^1(t)}{dt}, \frac{dx^2(t)}{dt}, \frac{dx^3(t)}{dt}\right) \quad (217)$$

As mentioned above, bases are defined by direct derivatives such as

$$e^\mu = \frac{\partial r}{\partial x^\mu} = \left(\frac{\partial x^0}{\partial x^\mu}, \frac{\partial x^1}{\partial x^\mu}, \frac{\partial x^2}{\partial x^\mu}, \frac{\partial x^3}{\partial x^\mu}\right) \quad (218)$$

where, dx^μ is a proper coordinate differential, so that it is middle labeled.

For the velocity, the differential dx^μ is exactly defined on a trajectory of a matter, so that there is the probability that

$$V_0^\dagger \neq 0 \quad (219)$$

It is said that velocity tensor itself is literally trajectory derivatives.

Trajectory derivatives of general tensor should be labeled for discrepancy.

For example, the bases along a trajectory in rest field

$$\left(\frac{de^i}{dt}\right)_{tr} = \frac{de^i}{d\lambda} \frac{d\lambda}{dt} \neq 0 \quad (220)$$

while direct derivatives with $i \neq 0$

$$\frac{\partial e^i}{\partial t} = 0 \quad (221)$$

Eq. (216) could be calculated as

$$\left(\frac{DA}{dt}\right)_{tr} = \frac{D(A^i e_i)}{d\lambda} \frac{d\lambda}{dt} = e_i \frac{dA^i}{d\lambda} \frac{d\lambda}{dt} + A^i \frac{de_i}{d\lambda} \frac{d\lambda}{dt} \quad (222)$$

and $\frac{dA^i}{d\lambda} \frac{d\lambda}{dt}$ is also time derivative, and there is

$$\left(\frac{dA^i}{dt}\right)_{tr} = \frac{dA^i}{d\lambda} \frac{d\lambda}{dt} \quad (223)$$

where, Einstein summation convention does not act on double λ .

Thus, we have seen the difference between a trajectory derivative and an original derivative.

On the other hand, the value of matter's velocity

$$V_0^\lambda = \frac{d\lambda}{dt} \quad (224)$$

so that

$$\begin{aligned} \left(\frac{DA}{dt}\right)_{tr} &= \mathbf{e}_i \frac{dA^i}{d\lambda} \frac{d\lambda}{dt} + A^i \frac{d\mathbf{e}_i}{d\lambda} \frac{d\lambda}{dt} \\ &= \mathbf{e}_i \left(\frac{dA^i}{dt}\right)_{tr} + A^i \frac{d\mathbf{e}_i}{d\lambda} V_0^\lambda \\ &= \mathbf{e}_i \left(\frac{dA^i}{dt}\right)_{tr} + A^i \frac{\partial \mathbf{e}_i}{\partial x^\mu} \frac{\partial x^\mu}{\partial x^\lambda} V_0^\lambda \\ &= \mathbf{e}_i \left(\frac{dA^i}{dt}\right)_{tr} + A^i \frac{\partial \mathbf{e}_i}{\partial x^\mu} \frac{\partial x^\mu}{\partial x^\lambda} \cdot \mathbf{V}_0^\lambda \\ &= \mathbf{e}_i \left(\frac{dA^i}{dt}\right)_{tr} + A^i \frac{\partial \mathbf{e}_i}{\partial x^\mu} \alpha_\lambda^\mu \cdot \mathbf{V}_0^\lambda \\ &= \mathbf{e}_i \left(\frac{dA^i}{dt}\right)_{tr} + A^i \frac{\partial \mathbf{e}_i}{\partial x^\mu} V_0^\mu \quad (225) \end{aligned}$$

where, Einstein summation convention does not act on double λ , and $\alpha_\lambda^\mu = \frac{\partial x^\mu}{\partial x^\lambda}$ is the direction cosine on the direction μ of vector line λ , so that $V_0^\mu = \alpha_\lambda^\mu \cdot \mathbf{V}_0^\lambda$ is of component of \mathbf{V}_0^λ on that direction and \mathbf{V}_0^λ is vector form of V_0^λ so that it could also be written as \mathbf{V}_0^λ equivalently in this equation specially for matter's motion on the trajectory.

It could be expanded to be

$$\left(\frac{DA}{dt}\right)_{tr} = \mathbf{e}_i \left(\frac{dA^i}{dt}\right)_{tr} + A^i \left(\frac{\partial \mathbf{e}_i}{\partial x^0} V_0^0 + \frac{\partial \mathbf{e}_i}{\partial x^1} V_0^1 + \frac{\partial \mathbf{e}_i}{\partial x^2} V_0^2 + \frac{\partial \mathbf{e}_i}{\partial x^3} V_0^3 \right) \quad (226)$$

The expression in component forms could also be worked out as

$$\left(\frac{DA^\mu}{dt}\right)_{tr} = \left(\frac{dA^\mu}{dt}\right)_{tr} + A^i \left(\frac{\partial \mathbf{e}_i}{\partial x^0} \cdot \mathbf{e}^\mu V_0^0 + \frac{\partial \mathbf{e}_i}{\partial x^1} \cdot \mathbf{e}^\mu V_0^1 + \frac{\partial \mathbf{e}_i}{\partial x^2} \cdot \mathbf{e}^\mu V_0^2 + \frac{\partial \mathbf{e}_i}{\partial x^3} \cdot \mathbf{e}^\mu V_0^3 \right) \quad (227)$$

The expression in the way of Christoffel symbols as

$$\left(\frac{DA^\mu}{dt}\right)_{tr} = \left(\frac{dA^\mu}{dt}\right)_{tr} + \Gamma_{0i}^\mu A^i V_0^0 + \Gamma_{1i}^\mu A^i V_0^1 + \Gamma_{2i}^\mu A^i V_0^2 + \Gamma_{3i}^\mu A^i V_0^3 \quad (228)$$

We could find that the first component of 4-dimensional velocity is something special in that it is not true velocity. Generally, a velocity of massive matters in contra variant space is

$$\mathbf{V}_0^\lambda = (V_0^0, V_0^1, V_0^2, V_0^3) \quad (229)$$

And velocity composition expressed as

$$|\mathbf{V}_0^\lambda|^2 = -(V_0^0)^2 + (V_0^1)^2 + (V_0^2)^2 + (V_0^3)^2 \quad (230)$$

where $V_0^0 = \frac{cdt}{dt} = c$.

For light rays, the contra variant velocity will be composed to be nonvanishing

$$|\mathbf{c}_0^\lambda|^2 = -(c_0^0)^2 + (c_0^1)^2 + (c_0^2)^2 + (c_0^3)^2 \neq 0 \quad (231)$$

where $c_0^0 = \frac{cdt}{dt} = c$.

But their velocity composition in covariant space is zero

$$|c|^2 = -c^2 + (c1/0)^2 + (c2/0)^2 + (c3/0)^2 = 0 \quad (232)$$

in that the summation of the last 3 items is c^2 .

We know that invariant light speed is c . It is said that the composite velocity does not perform real velocity, so does the contra variant composite velocity. The velocity of V_0^0 or c_0^0 is something virtual quantity, and we have found more complexities in kinematics. One can argue that there are still some issues unsolved. That could be expected in next sections.

In rest fields, there is $\frac{\partial e_i}{\partial x^0} = 0$, so I prefer to suggest the real trajectory derivatives of rest field for discussion that

$$\left(\frac{DA}{dt}\right)_{tr} = e_i \left(\frac{dA^i}{dt}\right)_{tr} + A^i \left(\frac{\partial e_i}{\partial x^1} V_0^1 + \frac{\partial e_i}{\partial x^2} V_0^2 + \frac{\partial e_i}{\partial x^3} V_0^3\right) \quad (233)$$

and

$$\left(\frac{DA^\mu}{dt}\right)_{tr} = \left(\frac{dA^\mu}{dt}\right)_{tr} + \Gamma_{1i}^\mu A^i V_0^1 + \Gamma_{2i}^\mu A^i V_0^2 + \Gamma_{3i}^\mu A^i V_0^3 \quad (234)$$

It should be highlighted that the trajectory derivatives could also be defined in distance derivatives as

$$\left(\frac{DA^\mu}{d\lambda}\right)_{tr} = \frac{DA^\mu}{d\lambda} \quad (235)$$

which just performs a special appearance of trajectory derivatives.

For free falling trajectory, it is

$$\left(\frac{DA^\mu}{dr}\right)_{tr} = \frac{DA^\mu}{dr} \quad (236)$$

Trajectory derivative is derivative on a curve line in space time. Sometimes it is presented with time because that the time variable is used in the parametric equation. Christoffel symbols in the equations owe to derivatives of space differentials rather than time differentials. We will see that the concept of trajectory derivative help to describe frequency shift and acceleration, as well as to falsify the concept of geodesic line in section 8.2.

5. Theoretical Verifications on Gravitational Redshifts and Accelerations

Because of the inequality of mixed subscripts of Christoffel symbols, the classical Christoffel symbol equations could not be used any more in the theory of general relativity. The covariant derivatives in gravitational field should be considered in their correct forms.

5.1. On Gravitational Redshifts

Light rays travelling in gravitational field, are also the issue of matter's motions. Something special is that we would rather focus more on the frequency derivatives by distance, in that they perform more details of the concept of redshift.

Taking light propagation at vertical direction for example, a distance derivative of contra variant frequency is exactly trajectory derivative

$$\left(\frac{D\nu_0}{dr}\right)_{tr} = \frac{D\nu_0}{dr} = \frac{d\nu_0}{dr} - \Gamma_{10}^0 \nu_0 \quad (237)$$

It is sure to consider the tensor of frequency and its derivative to be vectors, but in traditions it is not of a rare necessity. It has been mentioned that Christoffel symbol Γ_{10}^0 were employed correctly with the form $\Gamma_{10}^0 = \frac{1}{2} g^{00} \frac{\partial g_{00}}{\partial x^1}$ as has been shown in Eq. (26). As a result, it will lead to a real answer

$$\Gamma_{10}^0 = \frac{1}{2} g^{00} \frac{\partial g_{00}}{\partial r} = \frac{1}{2} \left(1 - \frac{r^*}{r}\right)^{-1} \left(1 - \frac{r^*}{r}\right)' = \frac{r^*}{2r^2} \left(1 + \frac{r^*}{r-r^*}\right) \quad (238)$$

The covariant derivative will be calculated to be

$$\left(\frac{D\nu_0}{dr}\right)_{tr} = -\frac{r^*}{2r^2} \nu_{0(\infty)} - \frac{r^*}{2r^2} \left(1 + \frac{r^*}{r-r^*}\right) \left(1 + \frac{r^*}{2r}\right) \nu_{0(\infty)} \quad (239)$$

The approximate solution for weak field is that

$$\left(\frac{Dv_0}{\partial r}\right)_{tr} \approx 2 \frac{dv_0}{dr} \quad (240)$$

It is said that the solution Eq. (28) is confirmed again, and in weak field that reveals the covariant redshift is approximately double of the contra variant one.

Notwithstanding, one can also make another derivative as

$$\left(\frac{Dv_0}{dt}\right)_{tr} = \frac{dv_0}{dr} \frac{dr}{dt} - \Gamma_{10}^0 v_0 \frac{dr}{dt} \quad (241)$$

For light rays, $\frac{dr}{dt} = c_0^1$. It becomes

$$\left(\frac{Dv_0}{dt}\right)_{tr} = \left(\frac{dv_0}{dt}\right)_{tr} - c_0^1 \Gamma_{10}^0 v_0 \quad (242)$$

The first item of right hand side can also be transformed to be

$$\left(\frac{dv_0}{dt}\right)_{tr} = \frac{dv_0}{dr} \frac{dr}{dt} = c_0^1 \frac{dv_0}{dr} \quad (243)$$

so that

$$\left(\frac{Dv_0}{dt}\right)_{tr} = c_0^1 \left(\frac{dv_0}{dr} - \Gamma_{10}^0 v_0\right) = c_0^1 \left(\frac{Dv_0}{dr}\right)_{tr} \approx 2c_0^1 \frac{dv_0}{dr} \quad (244)$$

5.2. On Accelerations

One can take matter's freefalling for example to study the acceleration in gravitational fields. The entire contra variant acceleration is the derivative of contra variant velocity as that

$$a_{00}^1 = \left(\frac{Dv_0^1}{dt}\right)_{tr} = \frac{d^2 r}{dt^2} \quad (245)$$

And entire covariant acceleration is a pure covariant derivative

$$a_{1/00} = \left(\frac{DV_{1/0}}{d\tau}\right)_{tr} = \frac{d^2 \rho}{d\tau^2} \quad (246)$$

One may find that some tensors have been labeled with detailed middle index hence they may help to provide explicit expressions, in which / is employed to divide middle upper and middle lower indexes.

With the relationship between covariant derivatives, it is drawn that

$$e^1 e_0 e_0 a_{1/00} = \left(\frac{DV_0^1}{dt}\right)_{tr} \quad (247)$$

To study the covariant derivatives in the way of Christoffel symbols, there is

$$\left(\frac{DV_0^1}{dt}\right)_{tr} = \left(\frac{dv_0^1}{dt}\right)_{tr} + \Gamma_{11}^1 \frac{dr}{dt} V_0^1 - \Gamma_{10}^0 \frac{dr}{dt} V_0^1 \quad (248)$$

With $\frac{dr}{dt} = V_0^1$, it becomes

$$\left(\frac{DV_0^1}{dt}\right)_{tr} = \left(\frac{dv_0^1}{dt}\right)_{tr} + \Gamma_{11}^1 (V_0^1)^2 - \Gamma_{10}^0 (V_0^1)^2 \quad (249)$$

With $\Gamma_{11}^1 = e^1 \frac{\partial e_1}{\partial r}$, $\Gamma_{10}^0 = e^0 \frac{\partial e_0}{\partial r}$, $e^0 \partial e_0 = -e_0 \partial e^0$ and $\left(\frac{dv_0^1}{dt}\right)_{tr} = a_{00}^1$ there is

$$\begin{aligned} \left(\frac{DV_0^1}{dt}\right)_{tr} &= a_{00}^1 + e^1 \frac{\partial e_1}{\partial r} (V_0^1)^2 - e^0 \frac{\partial e_0}{\partial r} (V_0^1)^2 \\ &= a_{00}^1 + e^1 \frac{\partial e_1}{\partial r} (V_0^1)^2 + e_0 \frac{\partial e^0}{\partial r} (V_0^1)^2 \\ &= a_{00}^1 + e^1 e_0 \frac{\partial}{\partial r} (e_1 e^0) (V_0^1)^2 \end{aligned} \quad (250)$$

With Eq. (247), there is

$$a1/00 = e^0 e^0 e_1 \left(\frac{DV_0^1}{dt} \right)_{tr} = e^0 e^0 e_1 a_{00}^1 + e^0 \frac{\partial}{\partial r} (e_1 e^0) (V_0^1)^2 \quad (251)$$

One can find that we have use Γ_{11}^1 and Γ_{10}^0 in the calculation of $a1/00$, rather than Γ_{00}^1 that has been used in Eq. (44). As we have discussed, the value of Γ_{00}^1 is really of zero.

As has been said that covariant derivatives could also be developed in a direct way without Christoffel symbols

$$\begin{aligned} a1/00 &= \left(\frac{DV1/0}{d\tau} \right)_{tr} = \left(\frac{d^2 \rho}{d\tau^2} \right)_{tr} = \left[\frac{d}{d\tau} \left(\frac{e_1 dr}{e_0 dt} \right) \right]_{tr} = e^0 \left[\frac{d}{dt} \left(\frac{e_1 dr}{e_0 dt} \right) \right]_{tr} \\ &= e^0 \frac{e_1}{e_0} \left[\frac{d}{dt} \left(\frac{dr}{dt} \right) \right]_{tr} + e^0 \frac{dr}{dt} \left[\frac{d}{dt} \left(\frac{e_1}{e_0} \right) \right]_{tr} \\ &= e^0 e^0 e_1 a_{00}^1 + e^0 V_0^1 \left[\frac{d}{dt} (e^0 e_1) \right]_{tr} \\ &= e^0 e^0 e_1 a_{00}^1 + e^0 V_0^1 \frac{\partial}{\partial r} (e^0 e_1) \frac{dr}{dt} \\ &= e^0 e^0 e_1 a_{00}^1 + e^0 \frac{\partial}{\partial r} (e^0 e_1) (V_0^1)^2 \quad (252) \end{aligned}$$

It could be found that this equation has been far different from the Eq. (44), because errors in Christoffel symbol equation have been eliminated and at the same time the concept of trajectory derivate help to calculate an acceleration in right way. These discussions have presented further verifications for the revised equation of Christoffel symbols of Eq. (211).

By the way, it is interesting to take some discussions on some trivial concepts such as a_{00}^0 and the covariant form $a0/00$ of massive matters. Since V_0^0 and $V0/0$ are the velocities of contra variant time and proper time but not the real velocity of light

$$V_0^0 = \frac{dx^0}{dt} = \frac{d(ct)}{dt} = c \quad (253)$$

and

$$V0/0 = \frac{e_0 dx^0}{d\tau} = \frac{e_0 d(ct)}{d\tau} = c \quad (254)$$

Then their derivatives are just the accelerations of time coordinates that

$$a_{00}^0 = \left(\frac{dV_0^0}{dt} \right)_{tr} = 0 \quad (255)$$

while

$$\left(\frac{dV_0^0}{dt} \right)_{tr} = \left(\frac{dV_0^0}{dt} \right)_{tr} + \Gamma_{10}^0 V_0^0 V_0^1 - \Gamma_{10}^0 V_0^0 V_0^1 = 0 \quad (256)$$

so that

$$a0/00 = 0 \quad (257)$$

As light propagation at a direction of a radius is concerned, we know that light speed c keeps invariant in covariant space, so that there is

$$\frac{Dc}{d\tau} = 0 \quad (258)$$

Case discussing the performance of contra variant light speed, with invariant distance, there is

$$ds^2 = -g_{00} c^2 dt^2 + g_{11} dr^2 = 0 \quad (259)$$

and then the light speed in contra variant space will be

$$c_0^1 = \frac{dr}{dt} = \sqrt{\frac{g_{00}}{g_{11}}} c = e_0 e^1 c \quad (260)$$

where, positive g_{00} is set instead of a minus g_{00} as has suggested previous.

Then the acceleration

$$a_{00}^1 = \left(\frac{dc_0^1}{dt}\right)_{tr} = \left[\frac{d}{dt}(e_0 e^1 c)\right]_{tr} = c \frac{\partial}{\partial r}(e_0 e^1) \frac{\partial r}{\partial t} = c c_0^1 \frac{\partial}{\partial r}(e_0 e^1) = c^2 e_0 e^1 \frac{\partial}{\partial r}(e_0 e^1) \quad (261)$$

With Eq. (250) the covariant derivative is

$$\begin{aligned} \left(\frac{Dc_0^1}{dt}\right)_{tr} &= \left(\frac{dc_0^1}{dt}\right)_{tr} - e_1 e^0 \frac{\partial}{\partial r}(e^1 e_0)(c_0^1)^2 \\ &= a_{00}^1 - e_1 e^0 \frac{\partial}{\partial r}(e^1 e_0)(c_0^1)^2 \\ &= c^2 e_0 e^1 \frac{\partial}{\partial r}(e_0 e^1) - c^2 e_0 e^1 \frac{\partial}{\partial r}(e^1 e_0) = 0 \quad (262) \end{aligned}$$

Of course, with Eq. (260) and Eq. (247), we could obtain the result only by a judgment that

$$\left(\frac{Dc_0^1}{dt}\right)_{tr} = e^1 e_0 e_0 \frac{Dc}{dt} = 0 \quad (263)$$

6. Experimental Verifications on Gravitational Redshifts and Accelerations

Every tensor involved with measurable quantities could have probabilities to be performed in practice with measured quantities to verify their theoretical expressions. In space time, space intervals and time intervals are all measurable quantities so that they surely could be employed to perform the space and time dependent tensors.

The methodology of the so called revisit gravitational redshift encourages me to sponsor a realistic analysis method to further verify the general covariance, which will present solutions all based on physical events of realities. Physical events always have substantial existence so that they can help to create irrefutable conclusions. We know that physical events may be record both in contra variant space and covariant space that might provide different values for physical quantities, but both of them actually represent the same physical realities.

6.1. On Measurable Experiments

Measurable quantities could be used to describe physical events, which may be coordinate independent or not. Coordinate independent quantities of course show invariance in physical events in different spaces, such as wave numbers, which could be record as images or texts at specific times and positions. However, coordinate dependent quantities measured in site maybe really dependent. For examples, distance measurements not only depend on in-site space intervals but also depend on the in-site rulers, so as well, time measurements also depend on both in-site time intervals and the in-site clocks. We imagine that the space rulers and clocks their selves maybe also vary. Logically, records of these quantities are recognizable even if they are in farthest distance to the bystanders.

Case a measurement equipment varies with time space, whether the measurement quantity measured is in contra variant space quantity or covariant space quantity? With general covariance, it has been believed that rulers will shrink when they go closer to the center source corresponding to the space interval to become shorter. And also, it has been expected that clocks will go variant corresponding to their dynamic conditions.

However, after those inspections in previous sections, we know that general covariance does not work in some circumstances. Energy and momentum of a matter may not keep covariant in covariant space, while light speed may keep covariant spectacularly. On another side, our discussions may have led to a theoretical inference that matters may experience relativistic emission when they go to a center source and then shrink because of the variation of energy structure. That could be called covariant deformations.

Once we measure space and time intervals at a position, maybe they are not committed to be contra variant quantities or covariant quantities, because our rulers may vary uncommitted. But we could do made measurements anyway. In another word, we could indeed measure something so that they will correspond to any others. Thus, it is not harmful to suppose one of the series of measurable

quantities could be measured in following discussions, for example, the contra variant distances or contra variant time intervals. And then they will be valid to be transformed from one to another. That will help us to do more analysis for comparisons and discussions.

6.2. Measurable Verifications for Gravitational Redshift

For the issue of redshift, we are going to sponsor the physical events of wave number counting. It is known that light frequency investigation should be accomplished by indirect techniques and sometimes it may come out with deviations. But it is supposed here that the wave number of the light is countable, or it is believed that light wave could be seen and record. This assumption actually may not do harm to our understanding to the realities, because it indeed will not change the realities and the events of wave counting in that the measurements themselves are also physical processes.

The event of wave number counting could be specified as the record of a number of waves to past a position in a time interval, and it could also be simplified to be one wave corresponding to a time lasting of the light ray propagating a wave length distance. On another side indirectly, one can get wave number by measuring wave length, based on the assumption of invariant light speed. But the apparent light speed might be variable so that the indirect method is not a good idea.

If there is a photon propagating from position 1 to position 2 in a one source field as shown in Figure 16, which correspond to coordinates $r_{(1)}$ and $r_{(2)}$,

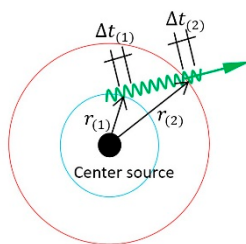


Figure 16. A photon traveling in center source field.

The wave number counting events should be carried out at the time that the photon passes the position 1 and position 2. In very short time $\Delta t_{(1)}$ and $\Delta t_{(2)}$, we will count the corresponding wave number $\Delta n_{(1)}$ and $\Delta n_{(2)}$.

Since the frequencies should be calculated as

$$v_{0(1)} = \frac{\Delta n_{(1)}}{\Delta t_{(1)}} \text{ and } v_{0(2)} = \frac{\Delta n_{(2)}}{\Delta t_{(2)}} \quad (264)$$

Redshift has been defined as

$$z_{\text{contra}} = \frac{v_{0(1)} - v_{0(2)}}{v_{0(2)}} \quad (265)$$

With the measurable records, redshift in contra variant space is

$$z_{\text{contra}} = \frac{\Delta n_{(1)} \Delta t_{(2)}}{\Delta n_{(2)} \Delta t_{(1)}} - 1 \quad (266)$$

In the event of wave counting, wave numbers are invariant in both contra variant space and covariant space, but the time intervals $\Delta t_{(1)}$ and $\Delta t_{(2)}$ will vary to be $\Delta \tau_{(1)}$ and $\Delta \tau_{(2)}$. In fact, every physical event keeps the only one event, whereas the different describing metrics lead to different results in the different spaces.

Naturally, gravitational redshift in covariant space is

$$z_{\text{revisit}} = \frac{v_{(1)} - v_{(2)}}{v_{(2)}} = \frac{\Delta n_{(1)} \Delta \tau_{(2)}}{\Delta n_{(2)} \Delta \tau_{(1)}} - 1 \quad (267)$$

where, this redshift symbol labeled with revisit is because it corresponds to that one named in classical equations.

As we know,

$$\Delta\tau = e_0\Delta t \quad (268)$$

It turns to be

$$z_{\text{revisit}} = \frac{e_{0(2)} \Delta n_{(1)} \Delta t_{(2)}}{e_{0(1)} \Delta n_{(2)} \Delta t_{(1)}} - 1 \quad (269)$$

As in a field of center source, the metric takes the forms of Schwarzschild solution, it is drawn that

$$z_{\text{revisit}} = \left(1 + \frac{\phi_2 - \phi_1}{c^2} + o\right) \frac{\Delta n_{(1)} \Delta t_{(2)}}{\Delta n_{(2)} \Delta t_{(1)}} - 1 \quad (270)$$

We know that the z_{contra} in Eq. (266) could have been measured in the physical event of wave number counting that of course equals to that in Eq. (9), so that

$$\frac{\Delta n_{(1)} \Delta t_{(2)}}{\Delta n_{(2)} \Delta t_{(1)}} = \left(1 + \frac{\phi_2 - \phi_1}{c^2} + o\right) \quad (271)$$

Thus, the covariant redshift in weak field is obtained

$$z_{\text{revisit}} = \left(1 + \frac{\phi_2 - \phi_1}{c^2} + o\right)^2 - 1 \approx 2 \frac{\Delta\phi}{c^2} \approx 2(z)_{\text{contra}} \quad (272)$$

It is said that, the revisit gravitational redshift is double of that of contra variant one.

As the equation of contra variant redshift is concerned, we know that it could be of course drawn by counting two wave numbers in two equivalent specified time intervals. For example, set $\Delta t_{(1)} = \Delta t_{(2)} = \Delta t$ which are measured at positions of $r_{(1)}$ and $r_{(2)}$, then $\Delta n_{(1)}$ and $\Delta n_{(2)}$ should represent the difference of frequency without time intervals. So that

$$z_{\text{contra}} = \frac{\Delta n_{(1)} \Delta t}{\Delta n_{(2)} \Delta t} - 1 = \frac{\Delta n_{(1)}}{\Delta n_{(2)}} - 1 \quad (273)$$

We know that $\Delta n_{(1)}$ and $\Delta n_{(2)}$ present the wave numbers with respect to $\Delta t_{(1)} = \Delta t$ and $\Delta t_{(2)} = \Delta t$.

As for revisit redshift, one will still get different covariant time intervals because the metrics go varied. Thus it is again doubled of the previous.

$$z_{\text{revisit}} = \frac{\Delta n_{(1)} e_{0(2)} \Delta t}{\Delta n_{(2)} e_{0(1)} \Delta t} - 1 \approx 2(z)_{\text{contra}} \quad (274)$$

It should be pointed out that in some experiments on gravitational redshift, only one timing clock was designed for time interval measurement. In this case, a wrong setting of proper time intervals may be taken into consideration, so that the experimental redshift may be presented as

$$z_{\text{experimental}} = \frac{\Delta n_{(1)} \Delta \tau_{(2)}}{\Delta n_{(2)} \Delta \tau_{(1)}} - 1 = \frac{\Delta n_{(2)} e_{0(x)} \Delta t_{(2)}}{\Delta n_{(1)} e_{0(x)} \Delta t_{(1)}} - 1 = (z)_{\text{contra}} \quad (275)$$

where, $e_{0(x)}$ is base component at clock position of $r_{(1)}$ or $r_{(2)}$ or any position other to them.

We can find out those completed experiments observations [14,15,16] on gravitational redshift will be easy to be verified to have only worked out the results of contra variant frequency shift.

Of course, one can calculate the real proper time intervals by time interval transformation between sole timing position and frequency shift positions. That will still help to work out revisit gravitational redshift as have discussed.

6.3. Measurable Verifications for Acceleration

6.3.1. Measurable Quantities and Measurable Acceleration

Firstly, I prefer to rise a controversy of a free falling on the Earth that if a matter freefalls from rest with velocity $V_{0(1)}^1 = 0$ as well as $V1/0_{(1)} = 0$ by nature, we do know that it will move quite

faster with velocity $V_{0(2)}^1$ after traveling a distance and a time interval because of gravity. In traditional theory, we know that the proper velocity $V1/0_{(2)} = e_0 e^1 V_{0(2)}^1$. Considering the weak field effect, there is $e_0 e^1 \approx 1$. Hence comes the controversy that the covariant acceleration must be great than zero because the matter has started from rest to a quite apparent motion. That is really contradicted with the principle of general covariance with a requirement of zero covariant acceleration. Nevertheless, considering that $V_{0(2)}^1$ and $V1/0_{(2)}$ are still non-relativistic velocities, it is easy to estimated that the accelerations are also approximately equal that $a1/00 \approx a_{00}^1$. The following works of so called realistic verifications in this section are exactly to be sponsored to solve these controversies thoroughly.

One cannot count on an investigation only by measuring a distance and a time interval of a freefalling to get a contra variant acceleration and then transforming to proper one to discover the difference. It is because that a distance a matter flies across in a time interval, does not interpret accurate velocity variance but gives a mean velocity.

Then we know that it is difficult to measure acceleration only at a single position, so that I would rather sponsor an investigation based on two-position measurements. A freefalling test with initial velocity is going to be put forward, in which a matter freely falls to source center from a position $r_{(1)}$ to a position $r_{(2)}$ shown as Figure 17. Once the velocities at the two positions are measured, the average values could be estimated with the velocities difference and the interval distance. Considering the condition on the surface of the Earth, a freefalling with a rarely big velocity and a rarely small travel would be performed. For example, a velocity of more than 10000 m/s, could be seen as a constant accelerated motion even in covariant space, in that a covariant derivative is expected to be linear with velocity.

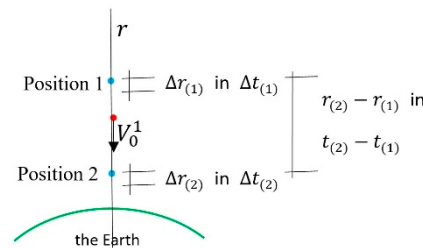


Figure 17. Free falling measurement.

It is supposed that contra variant distances and time intervals are measurable. Based on the measurements, velocity at position 1 could be written as

$$V_{0(1)}^1 = \frac{\Delta r_{(1)}}{\Delta t_{(1)}} \quad (276)$$

where the $\Delta r_{(1)}$ and $\Delta t_{(1)}$ are measured distance and time intervals when the matter goes by the position $r_{(1)}$, so that they are both tensors of contra variant space. For reasons of convenience their bracket sub indexes here are only employed to represent positions.

As well as that at position 2

$$V_{0(2)}^1 = \frac{\Delta r_{(2)}}{\Delta t_{(2)}} \quad (277)$$

Because the velocity at position 2 is the result of acceleration, it could be written in integral form

$$\begin{aligned} V_{0(2)}^1 &= V_{0(1)}^1 + \int_{t_{(1)}}^{t_{(2)}} a_{00}^1 dt \\ &= V_{0(1)}^1 + \overline{a_{00}^1} [t_{(2)} - t_{(1)}] \end{aligned} \quad (278)$$

where, the mean acceleration $\overline{a_{00}^1}$ is the integral point value, and $t_{(2)} - t_{(1)}$ is time interval for the matter traveling from $r_{(1)}$ to $r_{(2)}$.

So that the mean acceleration is

$$\overline{a_{00}^1} = \frac{V_{0(2)}^1 - V_{0(1)}^1}{t_{(2)} - t_{(1)}} \quad (279)$$

On the other hand, with covariant bases there are the relationships of contra variant quantities and proper ones

$$\Delta\rho_{(1)} = e_{1(1)}\Delta r_{(1)}, \quad \Delta\tau_{(1)} = e_{0(1)}\Delta t_{(1)}, \quad \Delta\rho_{(2)} = e_{1(2)}\Delta r_{(2)}, \quad \Delta\tau_{(2)} = e_{0(2)}\Delta t_{(2)} \quad (280)$$

One can use the mean metric to calculate the proper time intervals from position 1 to position 2

$$\tau_{(2)} - \tau_{(1)} = \int_{t_{(1)}}^{t_{(2)}} e_0 dt = \bar{e}_0(t_{(2)} - t_{(1)}) \quad (281)$$

where the \bar{e}_0 is the value at integral point, and it is suggested to be evaluated approximately as following in weak field

$$\bar{e}_0 \approx 0.5(e_{0(1)} + e_{0(2)}) \quad (282)$$

The proper velocities

$$V1/0_{(1)} = \frac{\Delta\rho_{(1)}}{\Delta\tau_{(1)}} \quad (283)$$

$$V1/0_{(2)} = \frac{\Delta\rho_{(2)}}{\Delta\tau_{(2)}} \quad (284)$$

And the integral relationship in covariant space that

$$\begin{aligned} V1/0_{(2)} &= V1/0_{(1)} + \int_{\tau_{(1)}}^{\tau_{(2)}} a1/00 d\tau \\ &= V1/0_{(1)} + \overline{a1/00}(\tau_{(2)} - \tau_{(1)}) \quad (285) \end{aligned}$$

And also, we get the mean covariant acceleration with Lagrangian mean value theorem of integration that

$$\overline{a1/00} = \frac{V1/0_{(2)} - V1/0_{(1)}}{\tau_{(2)} - \tau_{(1)}} \quad (286)$$

So that the mean covariant derivative

$$\overline{a1/00} = \frac{\frac{e_{1(2)}\Delta r_{(2)}}{e_{0(2)}\Delta t_{(2)}} - \frac{e_{1(1)}\Delta r_{(1)}}{e_{0(1)}\Delta t_{(1)}}}{\bar{e}_0(t_{(2)} - t_{(1)})} = \frac{\frac{e_{1(2)}V_{0(2)}^1}{e_{0(2)}} - \frac{e_{1(1)}V_{0(1)}^1}{e_{0(1)}}}{\bar{e}_0(t_{(2)} - t_{(1)})} \quad (287)$$

It is of course the measuring forms of an acceleration of a freefalling. And then it could be compared to that of contra variant one.

We would like substitute the equation of contra variant velocity 2 of Eq. (278) into this equation. That is

$$\begin{aligned} \overline{a1/00} &= \bar{e}^0 \frac{\frac{e_{1(2)}}{e_{0(2)}} [(t_{(2)} - t_{(1)})\overline{a_{00}^1} + V_{0(1)}^1] - \frac{e_{1(1)}}{e_{0(1)}} V_{0(1)}^1}{t_{(2)} - t_{(1)}} \\ &= \bar{e}^0 \frac{e_{1(2)}}{e_{0(2)}} \overline{a_{00}^1} + \bar{e}^0 \frac{\frac{e_{1(2)}}{e_{0(2)}} - \frac{e_{1(1)}}{e_{0(1)}}}{t_{(2)} - t_{(1)}} V_{0(1)}^1 \\ &= \bar{e}^0 e_{1(2)} e_{0(2)}^0 \overline{a_{00}^1} + \bar{e}^0 \frac{e_{1(2)}e_{0(2)}^0 - e_{1(1)}e_{0(1)}^0}{t_{(2)} - t_{(1)}} V_{0(1)}^1 \quad (288) \end{aligned}$$

It could be transformed to be

$$\begin{aligned} \overline{a1/00} &= \bar{e}^0 e_{1(2)} e_{0(2)}^0 \overline{a_{00}^1} + \bar{e}^0 \frac{e_{1(2)}e_{0(2)}^0 - e_{1(1)}e_{0(1)}^0}{t_{(2)} - t_{(1)}} V_{0(1)}^1 \frac{r_{(2)} - r_{(1)}}{r_{(2)} - r_{(1)}} \\ &= \bar{e}^0 e_{1(2)} e_{0(2)}^0 \overline{a_{00}^1} + \bar{e}^0 \frac{e_{1(2)}e_{0(2)}^0 - e_{1(1)}e_{0(1)}^0}{r_{(2)} - r_{(1)}} V_{0(1)}^1 \overline{V_0^1} \quad (289) \end{aligned}$$

Here we have got the transformed form of covariant acceleration of freefalling.

Nevertheless, with Lagrangian differential mean value theorem, we can write down the differential form as

$$\overline{a1/00} = \overline{e^0} e_{1(2)} e^0_{(2)} \overline{a^1_{00}} + \overline{e^0} \frac{\partial}{\partial r} (e_1 e^0) V^1_{0(1)} \overline{V^1_0} \quad (290)$$

Or the form of reverse bases

$$\overline{a1/00} = \overline{e^0} e_{1(2)} e^0_{(2)} \overline{a^1_{00}} - \overline{e^0} e_{1(2)} e_{1(1)} e^0_{(2)} e^0_{(1)} \frac{\partial}{\partial r} (e^1 e_0) V^1_{0(1)} \overline{V^1_0} \quad (291)$$

Thus by the way, another kind of proof of differential analysis of the Eq. (252) and Eq. (249) has been completed, in the way of measurable experiment.

6.3.2. Examples

Some terrestrial experiments are going to be put forward, that matters with initial velocity freefall in vacuum circumstance with in 1000m height to the ground. Both at the start point position 1 and end point position 2, the matter’s velocities will be measured. And of course, the space and time intervals between position 1 and 2 that depend on the so called geodesic line will be measured together so that to calculate the mean accelerations.

Some basic data of the Earth have already been tested certainly, so that we can take the standard value for our experiments, such as the total mass of the Earth $M = 5.97237 \times 10^{24} kg$, and the position on the ground could be assigned to have a radial coordinate $R_{\oplus} = 6.371393 \times 10^6 m$. We could also take the gravitational constant $G = 6.67259 \times 10^{-11} Nm^2/kg^2$, with the light speed $c = 299792458 m/s$ thus the gravitational radius will be calculated as

$$r^* = \frac{2GM}{c^2} = 8.8680827 \times 10^{-3} m \quad (292)$$

With Newtonian equation and Schwarzschild’s solution, some positional data could be list in following Table 1.

Table 1. Base components , derivatives and gravity at experimental positions.

r	e_1	e_0	GM/r^2	$\partial(e_1 e^0)/\partial r$
6372.393	1.00000000069582045	0.99999999930417955	9.81377376	2.183859184×10 ⁻¹⁶
6371.493	1.00000000069591874	0.99999999930408126	9.81654643	2.184476187×10 ⁻¹⁶
6371.393	1.00000000069592966	0.99999999930407034	9.81685457	2.184544758×10 ⁻¹⁶

So far as we have discussed, the accelerations a^1_{00} and $a1/00$ are really geological quantities, and now it is necessary to make an extending study. We know that all kinds of interactions could be seen as momentum exchanges between matters, as that

$$dP = d(mv) = vdm + m dv \quad (293)$$

For the convenience, some quantities discussed in this section will not be marked with tensor index anymore.

In conditions of low velocity motions, the theory of special relativity indicates small mass variations, thus

$$P \approx m dv \quad (294)$$

For the cases of high velocity motions, one should take a total analysis. Now the total acceleration could be defined

$$\Lambda = \frac{d(mv)}{mdt} = \frac{vdm}{mdt} + \frac{dv}{dt} \quad (295)$$

It is said, the total acceleration includes mass variant acceleration and velocity variant acceleration, and the latter also could be called geometrical acceleration.

With a momentum variation, kinetic energy will vary a difference

$$dE_k = v dP = v^2 dm + m v dv \quad (296)$$

At the same time, the mass energy equation of differential form is

$$dE_k = c^2 dm \quad (297)$$

Thus, there will be

$$c^2 dm = v^2 dm + m v dv \quad (298)$$

To be divided by time differential, there is geometrical acceleration

$$a_{00}^1 = \frac{dv}{dt} = \frac{c^2 - v^2}{mv} \frac{dm}{dt} \quad (299)$$

Now one can define a coefficient of geometrical acceleration

$$\eta = \frac{\frac{dv}{dt}}{\Lambda} = \frac{\frac{c^2 - v^2}{mv} \frac{dm}{dt}}{\frac{v dm}{mdt} + \frac{c^2 - v^2}{mv} \frac{dm}{dt}} = 1 - \frac{v^2}{c^2} \quad (300)$$

In one source field, single acting gravitational geometrical acceleration is

$$a_{00}^1 = \eta \frac{GM}{r^2} = \eta g \quad (301)$$

where g is the total acceleration of gravity.

We will see that geometrical acceleration declines as velocity goes up to a relativistic level, and it goes to zero as velocity closely catches up to light speed.

If a matter is accelerated from rest, the total energy includes rest part and kinetic part

$$mc^2 = m_0 c^2 + \xi m v^2 \quad (302)$$

where m is relativistic mass and m_0 is rest mass.

We know that in special relativity there is

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (303)$$

Back to Eq. (302), there is

$$\xi = (1 - \sqrt{1 - \frac{v^2}{c^2}}) \frac{c^2}{v^2} \quad (304)$$

Then the total kinetic energy

$$E_k = \xi m v^2 = (1 - \sqrt{1 - \frac{v^2}{c^2}}) \frac{c^2}{v^2} m v^2 \quad (305)$$

Case v/c is a small value, ξ would be close to 0.5, so that

$$E_k \approx \frac{1}{2} m v^2 \quad (306)$$

On the occasion of freefalling the variation of kinetic energy

$$\xi_2 m_2 v^2 - \xi_1 m_1 v_1^2 = \int_{r(1)}^{r(2)} m g ds = \bar{m} \bar{g} s \quad (307)$$

Again with Eq. (298) the energy difference in an experiment

$$\Delta m c^2 \approx \Delta m v^2 + m_1 v \Delta v \quad (308)$$

Thus, there is

$$\Delta m = m_1 \frac{v \Delta v}{c^2 - v^2} \quad (309)$$

Considering $\eta_1 \approx \eta_2$ in experiments, the difference of kinetic energy could be written as

$$\begin{aligned} \Delta E_k &= \xi_2 m_2 v^2 - \xi_1 m_1 v_1^2 = \xi_2 (m_1 + \Delta m) (v_1 + \Delta v)^2 - \xi_1 m_1 v_1^2 \\ &\approx \xi_1 (m_1 + \Delta m) (v_1^2 + 2 v_1 \Delta v) - \xi_1 m_1 v_1^2 \\ &= \xi_1 (\Delta m v_1^2 + 2 m_1 v_1 \Delta v) \end{aligned}$$

$$\begin{aligned} &= \xi_1 m_1 \left(\frac{v_1 \Delta v}{c^2 - v_1^2} v_1^2 + 2 v_1 \Delta v \right) \\ &\approx m_1 g s \quad (310) \end{aligned}$$

so that

$$\Delta v \approx \frac{gs}{\xi_1 v_1 \left(\frac{v_1^2}{c^2 - v_1^2} + 2 \right)} \quad (311)$$

Thus,

$$t_{(2)} - t_{(1)} = \Delta t = \Delta v / \overline{a_{00}^1} \quad (312)$$

Unfortunately, this solution cannot come up with a higher accuracy than that

$$\Delta t = \frac{s}{v_1 + 0.5 \Delta v} \quad (313)$$

After then, we are going to sponsor series of freefalling experiments. Contra variant accelerations and covariant accelerations for every position are easy to calculate. While measurable covariant acceleration $\overline{a1/00}$ could be obtained with measured distances and time intervals via Eq. (287). But it is convenient to calculate with Eq. (289), in that the latter is just a transformation of the previous. And in this equation, space intervals would be gained with Newtonian equations and time intervals $t_{(2)} - t_{(1)}$ with Eq. (313) for convenience. One may argue that the measured quantities might come from calculation. That doesn't matter, because the equation has been verified for hundreds of years, therefore it is sure that the calculated quantities have equal value with that by measuring. And then the mean covariant acceleration will be taken to compare with the contra variant acceleration a_{00}^1 calculated with Eq. (301) and covariant acceleration $a1/00$ calculated with Eq. (249) or Eq. (252). Calculation results have been listed in Tables 2 and 3.

Table 2. Terrestrial freefalling experiments from position 6371.493 to 6371.393.

Analysis method	Theoretical analysis				Measurable analysis	
Positions	r(1)=6371.493		r(2)=6371.393		From r(1) to r(2)	
Initial velocity	Contra variant acceleration	Covariant acceleration	Contra variant acceleration	Covariant acceleration	measured time intervals	Covariant acceleration
$V_{0(1)}^1$ (m/s)	$a_{00(1)}^1$ (m/s ²)	$a1/00_{(1)}$ (m/s ²)	$a_{00(2)}^1$ (m/s ²)	$a1/00_{(2)}$ (m/s ²)	$t_{(2)} - t_{(1)}$ (s)	$\overline{a1/00}$ (m/s ²)
0	9.8165464	9.8165464	9.8168546	9.8168546	4.5137	9.8165
100000	9.8165453	9.8165453	9.8168535	9.8168535	1.0×10 ⁻³	9.8165
10000000	9.8056240	9.8058425	9.8059318	9.8061503	1.0×10 ⁻⁵	9.8275
100000000	8.7243083	10.908784	8.7245822	10.909127	1.0×10 ⁻⁶	10.908
200000000	5.4475940	14.185499	5.4477651	14.185944	5.0×10 ⁻⁷	14.184
250000000	2.9900583	16.643034	2.9901522	16.643557	4.0×10 ⁻⁷	16.640
299792458	0	19.633093	0	19.633709	3.3356×10 ⁻⁷	19.629

Table 3. Terrestrial freefalling experiments from position 6372.393 to 6371.393.

Analysis method	Theoretical analysis				Measurable analysis	
Positions	$r_{(1)}=6372.393$		$r_{(2)}=6371.393$		From $r_{(1)}$ to $r_{(2)}$	
Initial velocity	Contra variant acceleration	Covariant acceleration	Contra variant acceleration	Covariant acceleration	measured time intervals	Covariant acceleration
$V^1_{0(1)}$ (m/s)	$a^1_{00(1)}$ (m/s ²)	$a1/00_{(1)}$ (m/s ²)	$a^1_{00(2)}$ (m/s ²)	$a1/00_{(2)}$ (m/s ²)	Δt (s)	$\overline{a1/00}$ (m/s ²)
0	9.8137738	9.8137738	9.8168546	9.8168546	14.276	9.8138
100000	9.8137730	9.8137751	9.8168535	9.8168535	1.0×10^{-2}	9.8138
10000000	9.8028548	9.8246934	9.8059318	9.8061503	1.0×10^{-4}	9.8247
100000000	8.7218446	10.905704	8.7245822	10.909127	1.0×10^{-5}	10.906
200000000	5.4460559	14.181493	5.4477651	14.185944	5.0×10^{-6}	14.182
250000000	2.9892144	16.383343	2.9901522	16.643557	4.0×10^{-6}	16.639
299792458	0	19.627547	0	19.633709	3.3356×10^{-6}	19.629

7. Conclusions and Inferences and Their Applications

7.1. Conclusions

Previous discussions will lead to two conclusions for matters’ motions in gravitational fields:

1) For light: Light speed keeps general covariance, but light frequency keeps conservation in contra variant space.

2) For massive matters: Massive matter’s velocity does not perform general covariance.

These two conclusions have been drawn based on three items, which are light speed invariance, gravitational redshift measurement and acceleration measurement. Among these items, light speed invariance is a theoretical setting. This setting is come from special relativity and observational verifications. Gravitational redshifts and accelerations could be measured in realistic events that guarantees the conclusions in a very high reliability.

It is only the general covariance of light speed that has been observed. That relates to the inacceleratability of light rays even in contra variant space, which is one of the performances of light speed invariance.

7.2. Inferences

These conclusions are really different from classical theory of general relativity and they will then lead to natural inferences. I prefer to focus on the inferences on kinematics and relativistic release:

1) For kinematics: General covariance goes break by a large range. During the motions in gravitational field, all matters, including light rays, will keep energy and momentum conservation in contra variant space rather than that in covariant space. Only for light rays they may keep velocity invariant in covariant space, but their energy and momentum will still keep conservation in contra variant space. Energy momentum conservation is the conservation under the condition of gravitational potential conversions. It is said that there is only one exception in realities, the light speed invariance, which will lead to the validity of light ray propagation Lagrangian. While for massive matters, Lagrangian goes invalid. In any positions in gravity fields, massive matters always have opportunities to be accelerated up to and keep velocities close to absolute-light-speed.

2) For relativistic release: Since apparent light speed may vary in gravitational field, that will bring changes to interaction ratio in particles of massive matters to influence fine structures. For electromagnet force there will be of variation of momentum exchange. It is also reasonable to predict that the speed of gluons relating to the strong interactions is general covariant like that of photons. On the other side, these interactions keep energy momentum conservations at the same time. Therefore, case massive matters inflow enough intervals in gravitational fields, they might get an excited state and release, which could be called relativistic release, just as excited electrons might do. The difference is that relativistic releases may experience thoroughly exciting in whole intrinsic structures, including exciting of electrons. Matters may also experience covariant deformation after relativistic release because of equivalent state.

7.3. Applications

Detailed discussions on some applications will be sponsored consequently that will greatly support the conclusions and inferences.

1) On kinematics: General covariance and conservation principle are the two key handles to rectify the classical equations, especially the principle of mass energy conservation. It would be seen that those efforts to employ the geodesic equation or covariant derivatives to build kinetic equations have already gone failed, in that covariant derivatives may be actually nonvanishing.

2) On relativistic release: The concept of equivalent state would be carried out to estimate the energy exceeding for inflow matters so that to discuss energy release, which will then lead to relativistic redshift of emission and absorption. Equivalent state also relates to relativistic deformation that might perform another kind of covariance.

8. Kinematics and Dynamics

8.1. The Most Important

The second Newtonian law interprets the mechanism of accelerative motions of massive matters so that to form the dynamics. Case in the conditions that matters have relativistic velocities, forces acting on matters will cause not only the variations of velocities but also the variations of matter's mass. It should be pointed out that a force really is of a statistic quantity rather than an essential physical quantity. In fact, a force is just a performance of exchange of momentum, as well as mass energy. Thus, that physics could be called the relativistic dynamics.

But for light propagations, the second Newtonian law will not take effects anymore. Even in the case that a force is vertical to a light ray, we will see that the second Newtonian law remains invalid. That is the reason we suggest the concept of kinematics that others to the concept of dynamics. If we persistently employ the concept of dynamics, it should be a new one.

No matter the space time been determined by what kind of metrics and labeled by what kind of coordinates, it is just a methodology for descriptions for physical events. None of them would have priorities. Physics is on earth depends on its nature rather on spaces. The most important is the conservation principles in realities.

It is easy to imagine that geodesic line could be employed for the solution of kinematic trajectories of matters, because general relativity expects conservations in curve space. But we will find out that geodesic equation or covariant derivatives have not really taken effect in the solving of the kinematics in the past century. We know the reason is that covariant derivatives may be nonvanishing so that those imposed settings of vanishing covariant derivatives might cause discrepancies to realities.

Most of methodologies for kinematic trajectories published were based on the so called Lagrangian. Besides these conditions, contra variant angular momentum conservation has been used in all of those solutions. One can imagine that this condition is apparently contradicted with general covariance. In fact, it is always the greatest reason for me to persist in this issue with more efforts.

Finally, the Lorentz covariance is also a kind of constraint condition, since it has been involved in the settings of Minkowski space and pseudo Riemannian space.

8.2. Discussions on Geodesic Equation

The geodesic line equations presented by Weinberg [17] with metrics given by

$$ds^2 = B(r)dt^2 - A(r)dr^2 - r^2d\theta^2 - r^2\sin^2\theta d\varphi^2 \quad (314)$$

is

$$\frac{d^2x^\mu}{d\tau^2} + \Gamma_{\nu\lambda}^\mu \frac{dx^\nu}{d\tau} \frac{dx^\lambda}{d\tau} = 0 \quad (315)$$

That has been calculated to perform as components as [17]

$$0 = \frac{d^2r}{d\tau^2} + \frac{A'(r)}{2A(r)} \left(\frac{dr}{d\tau}\right)^2 - \frac{r}{A(r)} \left(\frac{d\theta}{d\tau}\right)^2 - r \frac{\sin^2\theta}{A(r)} \left(\frac{d\varphi}{d\tau}\right)^2 + \frac{B'(r)}{2A(r)} \left(\frac{dt}{d\tau}\right)^2 \quad (316)$$

$$0 = \frac{d^2\theta}{d\tau^2} + \frac{2}{r} \frac{d\theta}{d\tau} \frac{dr}{d\tau} - \sin\theta \left(\frac{d\varphi}{d\tau}\right)^2 \quad (317)$$

$$0 = \frac{d^2\varphi}{d\tau^2} + \frac{2}{r} \frac{d\varphi}{d\tau} \frac{dr}{d\tau} + 2\cot\theta \frac{d\varphi}{d\tau} \frac{d\theta}{d\tau} \quad (318)$$

$$0 = \frac{d^2t}{d\tau^2} + \frac{B'(r)}{B(r)} \frac{dt}{d\tau} \frac{dr}{d\tau} \quad (319)$$

And then, with $\theta = \pi/2$, the so called kinematic equation, were finally drawn as

$$A(r) \left(\frac{dr}{d\tau}\right)^2 + \frac{L^2}{r^2} \left(\frac{dr}{d\tau}\right)^2 - \frac{1}{B(r)} = -E \quad (320)$$

where, $L = r^2 \frac{d\varphi}{d\tau}$ and E are set constants.

It seems that the kinematic equation has been created by covariant derivatives. But it should be pointed out that Eq. (315) to Eq. (320) have gone wrong. Because we have discussed that the covariant derivatives could be nonvanishing in some occasions, so that there must be something wrong involved.

We are going to sponsor investigations in two ways for a comparison. Firstly, a transformation from the tangent space defined as $(dt, dr, d\theta, d\varphi)$ to that defined as $(cd\tau, d\rho, rd\theta, r\sin\theta d\varphi)$ will be put into considerations.

The Eq. (228) could be employed for the solution, so that the derivative components could be performed with and then

$$\begin{aligned} \left(\frac{DV^1}{d\tau}\right)_{tr} &= \left(\frac{D}{d\tau} \frac{dr}{d\tau}\right)_{tr} = \left(\frac{dV^1}{d\tau}\right)_{tr} + \Gamma_{0i}^1 V^i V^0 + \Gamma_{1i}^1 V^i V^1 + \Gamma_{2i}^1 V^i V^2 + \Gamma_{3i}^1 V^i V^3 \\ &= \left(\frac{dV^1}{d\tau}\right)_{tr} + \Gamma_{01}^1 V^1 V^0 + \Gamma_{11}^1 V^1 V^1 + \Gamma_{21}^1 V^1 V^2 + \Gamma_{31}^1 V^1 V^3 \\ &= \left(\frac{dV^1}{d\tau}\right)_{tr} + \Gamma_{11}^1 V^1 V^1 \\ &= \frac{d^2r}{d\tau^2} + \frac{A'(r)}{2A(r)} \left(\frac{dr}{d\tau}\right)^2 \quad (321) \end{aligned}$$

$$\begin{aligned} \left(\frac{DV^2}{d\tau}\right)_{tr} &= \left(\frac{D}{d\tau} \frac{d\theta}{d\tau}\right)_{tr} = \left(\frac{dV^2}{d\tau}\right)_{tr} + \Gamma_{0i}^2 V^i V^0 + \Gamma_{1i}^2 V^i V^1 + \Gamma_{2i}^2 V^i V^2 + \Gamma_{3i}^2 V^i V^3 \\ &= \left(\frac{dV^2}{d\tau}\right)_{tr} + \Gamma_{02}^2 V^2 V^0 + \Gamma_{12}^2 V^2 V^1 + \Gamma_{22}^2 V^2 V^2 + \Gamma_{32}^2 V^2 V^3 \\ &= \left(\frac{dV^2}{d\tau}\right)_{tr} + \Gamma_{12}^2 V^2 V^1 \\ &= \frac{d^2\theta}{d\tau^2} + \frac{1}{r} \frac{d\theta}{d\tau} \frac{dr}{d\tau} \quad (322) \end{aligned}$$

$$\begin{aligned}
\left(\frac{DV^3}{\partial\tau}\right)_{tr} &= \left(\frac{D}{d\tau} \frac{d\varphi}{d\tau}\right)_{tr} = \left(\frac{dV^3}{d\tau}\right)_{tr} + \Gamma_{0i}^3 V^i V^0 + \Gamma_{1i}^3 V^i V^1 + \Gamma_{2i}^3 V^i V^2 + \Gamma_{3i}^3 V^i V^3 \\
&= \left(\frac{dV^3}{d\tau}\right)_{tr} + \Gamma_{03}^3 V^3 V^0 + \Gamma_{13}^3 V^3 V^1 + \Gamma_{23}^3 V^3 V^2 + \Gamma_{33}^3 V^3 V^3 \\
&= \left(\frac{dV^2}{d\tau}\right)_{tr} + \Gamma_{13}^3 V^3 V^1 + \Gamma_{23}^3 V^3 V^2 \\
&= \frac{d^2\varphi}{d\tau^2} + \frac{1}{r} \frac{d\varphi}{d\tau} \frac{dr}{d\tau} + \cot\theta \frac{d\varphi}{d\tau} \frac{d\theta}{d\tau} \quad (323)
\end{aligned}$$

$$\begin{aligned}
\left(\frac{DV^0}{\partial\tau}\right)_{tr} &= \left(\frac{D}{d\tau} \frac{dt}{d\tau}\right)_{tr} = \left(\frac{dV^0}{d\tau}\right)_{tr} + \Gamma_{0i}^0 V^i V^0 + \Gamma_{1i}^0 V^i V^1 + \Gamma_{2i}^0 V^i V^2 + \Gamma_{3i}^0 V^i V^3 \\
&= \left(\frac{dV^0}{d\tau}\right)_{tr} + \Gamma_{00}^0 V^0 V^0 + \Gamma_{10}^0 V^0 V^1 + \Gamma_{20}^0 V^0 V^2 + \Gamma_{30}^0 V^0 V^3 \\
&= \left(\frac{dV^2}{d\tau}\right)_{tr} + \Gamma_{10}^0 V^0 V^1 \\
&= \frac{d^2t}{d\tau^2} + \frac{B'(r)}{2B(r)} \frac{dt}{d\tau} \frac{dr}{d\tau} \quad (324)
\end{aligned}$$

where, $V^\mu = \frac{dx^\mu}{d\tau}$ is number μ component of contra variant velocity.

We have seen that some errors in the Eq. (316) to Eq. (320) have been rectified. In fact, it is easy to find out calculation errors. If any $i \neq j$ the Christoffel symbol of $\Gamma_{\mu i}^j = 0$, in that the bases we discussed are orthogonal.

Secondly, we are going to study another condition that the tangent spaces from $(cdt, dr, rd\theta, r\sin\theta d\varphi)$ to $(cd\tau, d\rho, rd\theta, r\sin\theta d\varphi)$. The invariant distance could be written as

$$ds^2 = -B(r)(cdt)^2 + A(r)dr^2 + (rd\theta)^2 + (r\sin\theta d\varphi)^2 \quad (325)$$

One may argue that this transformation has overcome Riemannian manifold definition because the contra variant space is not a R^4 . But on earth in mathematics, that doesn't matter because we know that the space could map to a R^4 at all. The derivation regulars are still available. Then the derivatives should be performed with Eq. (228) that

$$\begin{aligned}
\left(\frac{DV^1}{\partial\tau}\right)_{tr} &= \left(\frac{D}{d\tau} \frac{dr}{d\tau}\right)_{tr} = \left(\frac{dV^1}{d\tau}\right)_{tr} + \Gamma_{0i}^1 V^i V^0 + \Gamma_{1i}^1 V^i V^1 + \Gamma_{2i}^1 V^i V^2 + \Gamma_{3i}^1 V^i V^3 \\
&= \left(\frac{dV^1}{d\tau}\right)_{tr} + \Gamma_{01}^1 V^1 V^0 + \Gamma_{11}^1 V^1 V^1 + \Gamma_{21}^1 V^1 V^2 + \Gamma_{31}^1 V^1 V^3 \\
&= \left(\frac{dV^1}{d\tau}\right)_{tr} + \Gamma_{11}^1 V^1 V^1 \\
&= \frac{d^2r}{d\tau^2} + \frac{A'(r)}{2A(r)} \left(\frac{dr}{d\tau}\right)^2 \quad (326)
\end{aligned}$$

$$\begin{aligned}
\left(\frac{DV^2}{\partial\tau}\right)_{tr} &= \left(\frac{D}{d\tau} \frac{rd\theta}{d\tau}\right)_{tr} = \left(\frac{dV^2}{d\tau}\right)_{tr} + \Gamma_{0i}^2 V^i V^0 + \Gamma_{1i}^2 V^i V^1 + \Gamma_{2i}^2 V^i V^2 + \Gamma_{3i}^2 V^i V^3 \\
&= \left(\frac{dV^2}{d\tau}\right)_{tr} + \Gamma_{02}^2 V^2 V^0 + \Gamma_{12}^2 V^2 V^1 + \Gamma_{22}^2 V^2 V^2 + \Gamma_{32}^2 V^2 V^3 \\
&= \left(\frac{dV^2}{d\tau}\right)_{tr} + 0 \\
&= r \frac{d^2\theta}{d\tau^2} + \frac{d\theta}{d\tau} \frac{dr}{d\tau} \quad (327)
\end{aligned}$$

$$\left(\frac{DV^3}{\partial\tau}\right)_{tr} = \left(\frac{D}{d\tau} \frac{r\sin\theta d\varphi}{d\tau}\right)_{tr} = \left(\frac{dV^3}{d\tau}\right)_{tr} + \Gamma_{0i}^3 V^i V^0 + \Gamma_{1i}^3 V^i V^1 + \Gamma_{2i}^3 V^i V^2 + \Gamma_{3i}^3 V^i V^3$$

$$\begin{aligned}
&= \left(\frac{dV^3}{d\tau}\right)_{tr} + \Gamma_{03}^3 V^3 V^0 + \Gamma_{13}^3 V^3 V^1 + \Gamma_{23}^3 V^3 V^2 + \Gamma_{33}^3 V^3 V^3 \\
&= \left(\frac{dV^2}{d\tau}\right)_{tr} + 0 \\
&= r \sin \theta \frac{d^2 \varphi}{d\tau^2} + \sin \theta \frac{d\varphi}{d\tau} \frac{dr}{d\tau} + r \cos \theta \frac{d\varphi}{d\tau} \frac{d\theta}{d\tau} \quad (328) \\
\left(\frac{DV^0}{\partial\tau}\right)_{tr} &= \left(\frac{D}{d\tau} \frac{cdt}{d\tau}\right)_{tr} = \left(\frac{dV^0}{d\tau}\right)_{tr} + \Gamma_{0i}^0 V^i V^0 + \Gamma_{1i}^0 V^i V^1 + \Gamma_{2i}^0 V^i V^2 + \Gamma_{3i}^0 V^i V^3 \\
&= \left(\frac{dV^2}{d\tau}\right)_{tr} + \Gamma_{10}^0 V^0 V^1 \\
&= c \frac{d^2 t}{d\tau^2} + \frac{B'(r)}{2B(r)} \frac{cdt}{d\tau} \frac{dr}{d\tau} \quad (329)
\end{aligned}$$

In comparisons of the last two calculations, we could find subtle nuance in that they are settled by different x^μ and V^μ . But they have really given the equivalent results, in that both of them could be transformed to uniform covariant derivatives $\frac{DV}{d\tau}$. Because the latter calculation is very easy to be done, it is also a kind of verification to the previous. And moreover, the most important, the comparison calculations have verified the conclusions on inequality of mixed subscript Christoffel symbols, because the last result is easily worked out and approved to be right, and then one could find that the simplified expression of the second step could be used to verified the solution of the first step. That will finally indicate the errors in classical theory, as well as that in Weinberg's calculations on geodesic equations.

We know many efforts [11,12,17] have been made to attempt to prove the conservation principles after the equation of geodesic equations that it is expected

$$r^2 \frac{d\varphi}{d\tau} = \text{const.} \quad (330)$$

come from the Eq. (318), and

$$\left(1 - \frac{r^*}{r}\right) \frac{dt}{d\tau} = \text{const.} \quad (331)$$

come from the Eq. (319).

It is easy to find that all of the works involve with errors. In comparison on the results of $\left(\frac{DV^2}{\partial\tau}\right)_{tr}$ and $\left(\frac{DV^3}{\partial\tau}\right)_{tr}$ in previous two kinds of strategies, we will find that the two derivatives do nothing with gravity influence, and they are just come from transformation of spherical coordinates so that any doctrines after that to form angular momentum conservation principle would be lack of supports. We will make further verifications in next sections that these two equations are all false. In fact, the Eq. (330) is not a correct form of angular momentum, and the Eq. (331) does nothing with energy. We will see that, motion trajectories cannot be calculated based on covariant derivatives, in that they are really not the geodesic lines.

8.3. Classical Equations of Light Ray Deflection

It is indicated in some books that Lagrangian relates to Euler-Lagrangian equation and geodesic equation [3,4]. It is trivial to continue the discussions on whatever of the origins. I will say that the Lagrangian equation for light rays is absolutely correct, because we will see that it is just the expression of composition of light speed components in covariant space. It is the reason that the Lagrangian is employed for the equation of matter's trajectories in most publications. In fact, velocity composition equation could be easily employed to solve the trajectory of Newtonian problems. But it should be pointed out that the Lagrangian equation is not proper for massive matters, which will be presented in following discussions.

Incomprehensibly, classical solving processes for Lagrangian equations seem like to do nothing with geodesic line equation and covariant differentials. On the other side, we could find that those solved trajectories all involved with contra variant angular momentum conservation, which indicates that the solved trajectories may be not real geodesic lines.

To take the problem of the light rays passing across the Sun for granted, as shown in Figure 18.

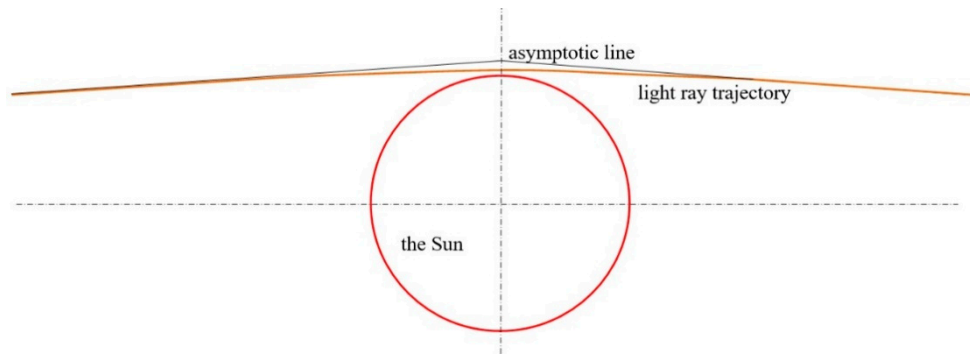


Figure 18. Light rays passing across the Sun.

It is expected in classical theories that for the motion of light rays, the Lagrangian is zero

$$\mathcal{L} = -(1 - \frac{r^*}{r})c^2\dot{t}^2 + (1 - \frac{r^*}{r})^{-1}\dot{r}^2 + r^2\dot{\phi}^2 = 0 \quad (332)$$

The following two items were always set to be constant in most publications [3,4] as

$$\frac{1}{2} \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = r^2 \dot{\phi} = L \quad (333)$$

$$\frac{1}{2} \frac{\partial \mathcal{L}}{\partial \dot{t}} = (1 - \frac{r^*}{r})c^2\dot{t} = c^2E \quad (334)$$

NB because of the multiplier r^2 used, L is really an equivalent contra variant angular momentum rather than a covariant one, so that this setting seems like to insert the contra variant conservation in. It is said that the setting of L does not coincide with general covariance. Even so, we will find out that it is exactly wrong setting so that the setting of constant E should be required to give not bad results.

Then the Lagrangian would be treated to be

$$c^2E^2 - \dot{r}^2 - (1 - \frac{r^*}{r})\frac{L^2}{r^2} = 0 \quad (335)$$

Setting $r' = \frac{dr}{d\phi} = \frac{\dot{r}}{\dot{\phi}}$, $u = \frac{1}{r}$ and then $r' = \frac{dr}{d\phi} = -\frac{u'}{u^2}$, there is

$$\frac{c^2E^2}{L^2} - u'^2 - (1 - r^*u)u^2 = 0 \quad (336)$$

Derivated by ϕ , it is

$$u'' + u = \frac{3}{2}r^*u^2 \quad (337)$$

At a farthest position the right hand item goes to zero, then the equation will have a solution $u = \sin\phi/b$, where b is a distance from solar center to the asymptotic line. Replacing the very small right hand item with the solution, there is

$$u'' + u = \frac{3r^*}{2b^2}\sin^2\phi \quad (338)$$

There is a particular solution

$$u_1 = \frac{3r^*}{4b^2}(1 + \frac{1}{3}\cos 2\phi) \quad (339)$$

Then the approximate solution of Eq. (337) is obtained

$$u = \frac{\sin\varphi}{b} + \frac{3r^*}{4b^2} \left(1 + \frac{1}{3}\cos 2\varphi\right) \quad (340)$$

For an infinite large r , the u is infinitesimal one, there is

$$\varphi_\infty = -\frac{r^*}{b} \quad (341)$$

Observational deflection case $b \approx R_\odot$ is

$$\Delta = 2|\varphi_\infty| = 1.75'' \quad (342)$$

where R_\odot is the radius of the Sun.

At the position $\varphi = \frac{\pi}{2}$ and $r = R$, there is

$$\frac{1}{R} = \frac{1}{b} + \frac{3r^*}{4b^2} \left(1 - \frac{1}{3}\right) \quad (343)$$

or

$$b \approx R + \frac{r^*}{2} \quad (344)$$

That will bring about contradictions for the setting of Eq. (333) that at peak point there is

$$L_R = r^2\dot{\varphi} = Rc \quad (345)$$

while at a very far point, there is

$$L_b = r^2\dot{\varphi} = bc = \left(R + \frac{r^*}{2}\right)c \neq L_R \quad (346)$$

It is said that the equations have been solved incorrectly. Might as well, we could find out an inevitable solution after the inappropriate setting of Eq. (333) in next section.

8.4. Revisit Equations of Light Ray Deflection

8.4.1. Errors Hidden in Classical Equations

It is obvious that a wrong setting has been made in Eq. (335), because in gravitational field, $(1 - \frac{r^*}{r})c^2\dot{t}$ really varies with r . In fact, the Eq. (332) could be solved directly as following.

Considering $\dot{t} = \frac{dt}{d\tau} = (1 - \frac{r^*}{r})^{-1/2}$, the Lagrangian really is

$$-c^2 + (1 - \frac{r^*}{r})^{-1}\dot{r}^2 + r^2\dot{\varphi}^2 = 0 \quad (347)$$

or

$$-c^2 + \dot{r}^2 + r^2\dot{\varphi}^2 = 0 \quad (348)$$

It is really the composition of light speed in covariant space.

Transform the equation as

$$(1 - \frac{r^*}{r})c^2 - \dot{r}^2 - (1 - \frac{r^*}{r})r^2\dot{\varphi}^2 = 0 \quad (349)$$

If setting peak point radius as R and the angular momentum at that point $L = Rc = r^2\dot{\varphi} = \text{const}$, with $u = \frac{1}{r}$, it becomes

$$(1 - r^*u)c^2 - u'^2R^2 - (1 - r^*u)u^2R^2 = 0 \quad (350)$$

To be derivated by φ , it is

$$u'' + u = \frac{3}{2}r^*u^2 - \frac{r^*}{2R^2} \quad (351)$$

With $\frac{3r^*}{2b^2}\sin^2\varphi$ instead of $\frac{3}{2}r^*u^2$, it becomes

$$u'' + u = \frac{3r^*}{2b^2}\sin^2\varphi - \frac{r^*}{2R^2} \quad (352)$$

Because $R \approx b$, it could be solved as

$$u = \frac{\sin\varphi}{b} + \frac{r^*}{4b^2} (1 + \cos 2\varphi) \quad (353)$$

One will obtain the solution of the equation as

$$\varphi_\infty = -\frac{r^*}{2b} \quad (354)$$

At the peak point as $\varphi = \frac{\pi}{2}$, we will get the constant b that

$$\frac{1}{R} = \frac{1}{b} + \frac{r^*}{4b^2} (1 - 1) \quad (355)$$

so that

$$b = R \quad (356)$$

Mathematically, this equality indicates that the solved line is not a natural line, in that if the curved line and its asymptotic line have same coordinates at start point and end point, there must be an inflection point on the curved line.

And the setting of Eq. (333) has been well kept that

$$L_R = r^2 \dot{\varphi} = Rc \quad (357)$$

and

$$L_b = r^2 \dot{\varphi} = bc = Rc = L_R \quad (358)$$

This is really the inevitable solution for classical equations but it is not a true result for realities. We have seen that the classical equations to have been solved to an answer Eq. (340) accurately up to the observation results is just caused by the wrong settings of energy momentum conservations of Eq. (333) and Eq. (334). It is said the classical equations do involve with problems while the wrong settings do.

8.4.2. Momentum, Energy and Angular Momentum Conservation

We have drawn the conclusion that light momentum keeps conservation in contra variant space rather than covariant space, and then of course, so does the light mass energy. In fact, apparent light speed or so called contra variant light speed may varies in contra variant space, but light momentum and energy will not be affected by apparent speeds. In fact, neither geodesic equation nor the derivations of Lagrangian could help proving the Eq.(333) and Eq.(334), in that Eq.(333) and Eq.(334) are substantially not correct.

Considering a light ray goes a vertical distance on the Earth, one could gain the mass variation as

$$m_r = \frac{h\nu_0(\infty)}{c^2} \left(1 + \frac{r^*}{2r}\right) = m_\infty \left(1 + \frac{r^*}{2r}\right) \quad (359)$$

It could be called simplified equation of mass in gravitational field.

Light momentum could be expressed as

$$P = \frac{h\nu_0}{c} = m_r c \quad (360)$$

or the momentum square

$$P^2 = m_r^2 c^2 \quad (361)$$

Case in contra variant space, apparent light speed varies with position so that that speed cannot be used in expressions of light momentum directly. The invariant light speed $c = \text{const.}$ could be seen as absolute light speed. Eq. (360) performs full variation with gravity by m_r , which is the performance of momentum conservation. In fact, light momentum depends on frequency, just as m_r does. If we ask more for a deep reason, that should be mass energy equation.

Case in covariant space, if light momentum will also be expressed by frequency, that will vary with bases additionally.

We know the Lagrangian

$$c^2 = (1 - \frac{r^*}{r})^{-1} \dot{r}^2 + r^2 \dot{\varphi}^2 \quad (362)$$

With Lagrangian substituted in conservative momentum square, it turns to be

$$P^2 = m_r^2 [(1 - \frac{r^*}{r})^{-1} \dot{r}^2 + r^2 \dot{\varphi}^2] \quad (363)$$

or

$$P^2 = m_r^2 [(1 - \frac{r^*}{r})^{-2} (\frac{dr}{dt})^2 + (1 - \frac{r^*}{r})^{-1} r^2 (\frac{d\varphi}{dt})^2] \quad (364)$$

As we have discussed, light speed cannot be directly composed in contra variant space but can be done in covariant space. This is a reason the Lagrangian is employed in conservative momentum square.

In one source fields, the momentum vector could be discomposed to be components of centripetal and tangent

$$\mathbf{P} = \mathbf{P}_c + \mathbf{P}_t \quad (365)$$

or

$$P^2 = P_c^2 + P_t^2 \quad (366)$$

Obviously, the tangent momentum relates to tangent velocity and centripetal momentum relates to centripetal velocity. It could also be inferred that the tangent component varies with the corresponding velocity, and so does the centripetal one.

So that there must be

$$P_c^2 = m_r^2 (1 - \frac{r^*}{r})^{-1} \dot{r}^2 = m_r^2 (1 - \frac{r^*}{r})^{-2} (\frac{dr}{dt})^2 \quad (367)$$

and

$$P_t^2 = m_r^2 r^2 \dot{\varphi}^2 = m_r^2 (1 - \frac{r^*}{r})^{-1} r^2 (\frac{d\varphi}{dt})^2 \quad (368)$$

In one source field, the angular momentum conservation could be expressed as

$$L^2 = r^2 P_t^2 = m_r^2 r^4 \dot{\varphi}^2 = m_r^2 (1 - \frac{r^*}{r})^{-1} r^4 (\frac{d\varphi}{dt})^2 = \text{const.} \quad (369)$$

It should be highlighted that we are talking about the moment conservation in contra variant space. It is amazing that the angular momentum should be expressed in the form of $r^2 m_r \dot{\varphi}$ rather than the form of $r^2 m_r \frac{d\varphi}{dt}$, or we have seen that light momentum could be only directly composed in covariant form. The real reason is that the invariant light speed c is just employed for the expressions by invariance.

The issues of light momentum have always been one of the controversies in physics for more than a hundred years [18]. The main problem is the difficulty of assessing the light momentum in transparent materials between Minkowski's equation [19] and Abraham's equation [20]. To one's surprise, we would have made the conclusion different from both of them, after the discussions in previous sections, because of Lorentz covariance.

Nevertheless, angular momentum brings about new surprises on it. We will find that the surprises not only rise up from the expressions, but also hide behind the kinematics of light propagations in gravitational fields. These efforts might bring about tiny contributions for the attempt to answer the question of Einstein about 'What are light quanta?' [21] I appreciate what Leonhardt has said that light continues to surprise [22].

8.4.3. Revisit Equations for Light Ray Trajectory

We have recognized that it is conservation principle that really controls the solutions. In fact, light rays in gravitational field may experience mass energy variation.

As has mentioned previous, the light mass at the peak point is

$$m_R = \frac{h\nu_{0(\infty)}}{c^2} \left(1 + \frac{r^*}{2R}\right) \quad (370)$$

And it varies at position r

$$m_r = \frac{h\nu_{0(\infty)}}{c^2} \left(1 + \frac{r^*}{2r}\right) \quad (371)$$

These two equations involve the energy conservation in contra variant space rather than that in covariant space.

For a light ray passing by a one source field, there is the angular momentum conservation as

$$L = r^2 m_r \dot{\phi} = R m_R c = \text{const.} \quad (372)$$

It should be highlighted again that the contra variant light momentum has been expressed by c and $\dot{\phi}$ which are of covariant space quantities rather than c_0^μ and $\frac{d\phi}{dt}$ of contra variant ones. In fact, it could be proved that c_0^μ and $\frac{d\phi}{dt}$ cannot be taken to form momentum conservation, if one takes efforts to have a try. That is because momentum variation is depend on m_r that perform the effect of gravity, or in another words, the gravity input energy into the m_r . The invariant light speed c employed reveals that light momentum variation really depends on frequency rather than real velocity, just as that light propagates in transparent materials. That perhaps is really surprise.

Considering the Eq. (370) and Eq. (371), there is

$$\dot{\phi} = \frac{R m_R c}{m_r r^2} = \frac{R c}{r^2} \left(\frac{1 + \frac{r^*}{2R}}{1 + \frac{r^*}{2r}}\right) = \frac{\tilde{L}}{r^2 (1 + \frac{r^*}{2r})} \quad (373)$$

where $\tilde{L} = \left(R + \frac{r^*}{2}\right) c$. In weak field, the item $1/(1 + \frac{r^*}{2r}) \approx (1 - \frac{r^*}{2r})$, thus

$$\dot{\phi} = \frac{1}{r^2} \left(1 - \frac{r^*}{2r}\right) \tilde{L} \quad (374)$$

I prefer to present the Lagrangian again

$$-c^2 + \left(1 - \frac{r^*}{r}\right)^{-1} \dot{r}^2 + r^2 \dot{\phi}^2 = 0 \quad (375)$$

Define $r' = \frac{dr}{d\phi}$, so that $\dot{r} = \frac{dr}{d\phi} \frac{d\phi}{dt} = r' \dot{\phi}$. There is

$$-c^2 + \left(1 - \frac{r^*}{r}\right)^{-1} r'^2 \dot{\phi}^2 + r^2 \dot{\phi}^2 = 0 \quad (376)$$

Insert the Eq. (374) into it, so that

$$-c^2 + \left(1 - \frac{r^*}{r}\right)^{-1} r'^2 \left[\frac{1}{r^2} \left(1 - \frac{r^*}{2r}\right) \tilde{L}\right]^2 + r^2 \left[\frac{1}{r^2} \left(1 - \frac{r^*}{2r}\right) \tilde{L}\right]^2 = 0 \quad (377)$$

With $\left(1 - \frac{r^*}{2r}\right)^2 \approx 1 - \frac{r^*}{r}$, it is

$$-c^2 + r'^2 \frac{\tilde{L}^2}{r^4} + \left(1 - \frac{r^*}{r}\right) \frac{\tilde{L}^2}{r^2} = 0 \quad (378)$$

With $u = \frac{1}{r'}$ and $r' = \frac{dr}{d\phi} = -\frac{u}{u^2}$ it turns to be

$$-\frac{c^2}{\tilde{L}^2} + u'^2 + (1 - ur^*) u^2 = 0 \quad (379)$$

We have seen the similar form of Eq. (336), but this equation comes from the settings of the real conservation principles.

Differentiation results in

$$u'' + u = \frac{3}{2} r^* u^2 \quad (380)$$

Case r is very big value, the equation could be simplified as

$$u'' + u = 0 \quad (381)$$

It could be solved to be

$$u = \frac{\sin\varphi}{b} \quad (382)$$

It is a horizontal line with a perpendicular distance b to the center of the Sun.

The right item of Eq. (380) could be replaced with the simple solution, because the deviation is also very small. So that the equation could be reformed to be

$$u'' + u = \frac{3r^*}{2b^2} \sin^2\varphi \quad (383)$$

Once again, we could obtain the solution

$$u = \frac{\sin\varphi}{b} + \frac{3r^*}{4b^2} \left(1 + \frac{1}{3} \cos 2\varphi\right) \quad (384)$$

And then the deflection angle

$$\varphi_\infty = -\frac{r^*}{b} \quad (385)$$

Case $\varphi = \frac{\pi}{2}$, there is

$$\frac{1}{R} = \frac{1}{b} + \frac{r^*}{2b^2} \quad (386)$$

Because $b \gg r^*$ it becomes

$$R \approx b - \frac{r^*}{2} \quad (387)$$

or

$$b = R + \frac{r^*}{2} \quad (388)$$

To verify the momentum conservation that

$$L_R = m_R R c \quad (389)$$

At a position $r \gg R$

$$L_b = m_b b c = \frac{1+\frac{r^*}{2R}}{1+\frac{r^*}{2R}} m_R b c \approx \frac{1}{R+\frac{r^*}{2}} R m_R b c = m_R R c = L_R \quad (390)$$

It seems that we have got the same results as that of classical equations. But the truth is that the solution is the results after the conclusions of momentum and mass conservations which completely others to that of classical theory and at the same time the assumptions of Eq. (333) and Eq. (334) are thoroughly given up.

The most important is that the real kinematics of light propagation has been discovered.

8.5. More Discussions

The trajectories of light ray in gravitational fields have more details behind the previous solution. Further analyses may help to discover more realities.

8.5.1. Detailed Discussions on the Coordinates of the Light Ray Trajectory

Further discussions are going to be sponsored to make more detailed analysis for understanding of some items, positions and their calculations.

Item 1 is to recognize various kinds of lines. The most important line is the real light ray that is emitted from a farthest star to the observers on the Earth. This line should be curved as it goes closely to the sun. Prolonging the straight parts of the light ray, one will gain the crossed straight lines, the asymptotic lines of light trajectory. They will be parallel to the two radial coordinate vectors r_∞ left and r_∞ right. There is another important line is the straight line from farthest star to the observer, that will present the real star direction from observer to the star.

Item 2 is to recognize those angles. φ_∞ is the second coordinate of farthest point on the light ray. Because the Earth is far enough to the Sun, the elevation angle of observer view line could be

seen as φ_∞ . And because the star is very very far from the sun, the angle between straight line to the star and horizontal line could be also seen as φ_∞ . Thus the total deflect angle of Δ_{observe} is approximate $2\varphi_\infty$. They could be shown in Figure 19.

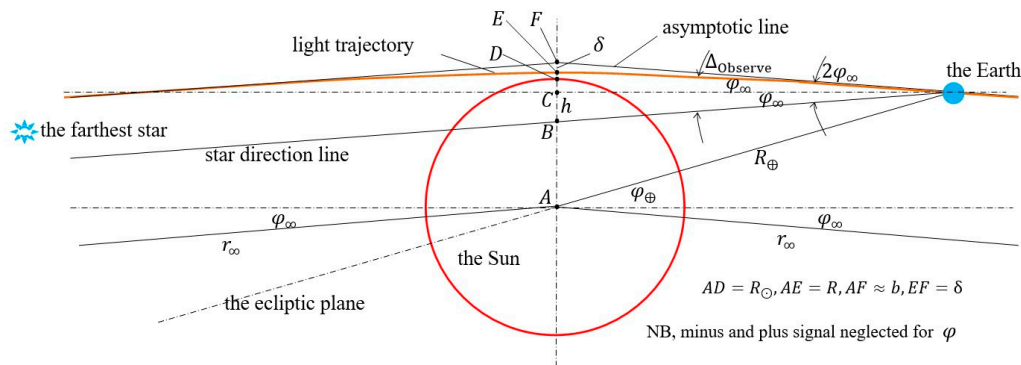


Figure 19. Detailed relationships for angles of light rays and view lines and coordinates.

For the coordinate of the Earth, $u = 1/A_u$, where A_u is astronomical unit, the simple solution could be used

$$\frac{1}{A_u} \approx \frac{\varphi_\oplus}{b} \quad (391)$$

where, φ_\oplus is second coordinate of the Earth.

Thus, there is

$$\varphi_\oplus \approx \frac{b}{A_u} \quad (392)$$

Because we know that $b \approx R_\odot$, where R_\odot is solar radius, so that

$$\varphi_\oplus \approx 0.004654478 \text{ rad} \quad (393)$$

The height of the horizon line to light ray

$$CE = R - A_u \sin \varphi_\oplus \quad (394)$$

It is difficult to get a not bad accurate solution. However, we can turn to discuss the value of $h = BC$ instead. For the approximate of $\varphi_\infty \approx \frac{1}{2} \Delta_{\text{observer}}$, it could be gain

$$h \approx A_u \sin \varphi_\infty \approx 635 \text{ km} \quad (395)$$

in the upper height $h = CF$, there is a very small difference between real ray trajectory and asymptotic line, the $\delta = EF$.

For the condition that light rays run from farthest position to the position they pass by the Sun, there is the equation after angular momentum conservation

$$\left(1 + \frac{r^*}{2r_\infty}\right)b = \left(1 + \frac{r^*}{2R}\right)R \quad (396)$$

NB, b is perpendicular length to asymptotic line, which is the moment distance for farthest positions, and only in approximate cases, it could be seen as $R + \delta$. R is peak point radius, and it need not be determined to be R_\odot in these discussions theoretically.

Then it is easy to obtain

$$b = R + \frac{r^*}{2} \quad (397)$$

Peak difference between R and b is

$$\delta = \frac{r^*}{2} \approx 1.5 \text{ km} \quad (398)$$

It is more difficult to investigate such a fine distance in practice that not only because of the observational accuracy but also due to the coordinating of the peak point of light ray. The

development of very-long-baseline interferometry have the capability of measuring angular separations and changes in angles as small as 10^{-4} seconds of arc [23]. That shows probabilities for the quite good accuracy for fine angle measuring. This issue perhaps cannot be solved easily.

We can study the right branch of asymptotic line with an overlaid φ_{∞} as shown in Figure 20.

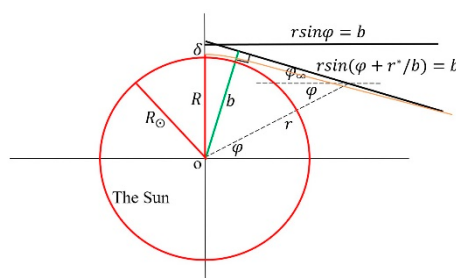


Figure 20. Asymptotic line with an overlaid φ_{∞} .

There is the asymptotic line equation as

$$r \sin(\varphi + |\varphi_{\infty}|) = b \quad (399)$$

or

$$r \sin\left(\varphi + \frac{r^*}{b}\right) = b \quad (400)$$

Case $\varphi = \frac{\pi}{2}$, we will see that the top point is close to the peak point

$$r = b / \sin\left(\frac{\pi}{2} + \frac{r^*}{b}\right) \approx b \quad (401)$$

Case φ is not very close to $\frac{\pi}{2}$, considering $\frac{r^*}{b}$ is very small, the approximate form of sine function is

$$\sin\left(\varphi + \frac{r^*}{b}\right) \approx \sin\varphi + \frac{r^*}{b} \cos\varphi \quad (402)$$

Thus, the deformed equation of asymptotic line could be written as

$$r = \frac{b}{\sin\varphi + r^* \cos\varphi / b} \quad (403)$$

It should be pointed out that the equation $u = \sin\varphi/b$ or $r \sin\varphi = b$ is a horizontal line with a perpendicular distance b to the Sun center.

For the case that a light passes closely by the edge of the sun, there might be some influences from solar corona. It is a good idea to left a distance from the edge, for example, the position of $1.5R_{\odot}$ or even further. But this idea is only for theoretical discussions. Hitherto, we could only observe light deflections in the conditions that light rays pass by the sun edge just fine, because otherwise, peak point could not be coordinated at all.

8.5.2. The Invalidity of Newtonian Second Law in Light Propagation

It is the most interesting that perhaps there is the probability to carry out new numerical method to calculate the light trajectory in gravitational field, which could be called ballistic trajectory method. Considering that the gravity component parallel to the motion trajectory will not bring about changes to further motion, we could only consider the calculation on motion variation due to vertical

component of gravity so that to determine a differential coordinate on the trajectory. Thus, in the way incremental, the trajectory could be solved at last.

Firstly, it is proficient to take the motion of close-to-light-speed massive particle into discussions, as shown in Figure 21.

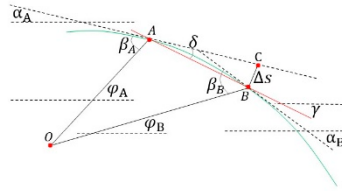


Figure 21. Invalidity of Newtonian second law for light deviation.

It is believed that for a massive particle in weak field, the vertical deviation could be solved by Newtonian second law

$$\Delta s = \frac{1}{2} g_v \Delta t^2 \quad (404)$$

The coordinates at position A are known to be (r_A, φ_A) or (x_A, y_A) . As well, the trajectory will have a direction with angle α_A . Then, the distance in a time interval Δt the particle travelling is

$$AB = c \Delta t \quad (405)$$

The deviation angle of AB to AC is

$$\delta \approx \Delta s / AB \quad (406)$$

and

$$\gamma = \alpha_A + \delta \quad (407)$$

Thus, the coordinates at position B could be calculated as

$$x_B = x_A + AB \cos \gamma \quad (408)$$

$$y_B = y_A + AB \sin \gamma \quad (409)$$

or

$$r_B = \sqrt{x_B^2 + y_B^2} \quad (410)$$

$$\varphi_B = \arccos \frac{x_B}{r_B} \quad (411)$$

It should be pointed out that the angle γ is not the direction angle α_B . The α_B could be calculated by angular momentum conservation.

$$\left(1 + \frac{r^*}{2R}\right)R = \left(1 + \frac{r^*}{2r_B}\right)r_B \sin \beta_B \quad (412)$$

where, $\beta_B = \alpha_B + \varphi_B$.

Now back to the fly of light ray from position A to position B . It might be imagined to calculate the deviation Δs also by Newtonian second law. That sounds naturally to see the photons as light speed particles with dynamic mass so that to deviate in the same way as massive particles do. Unfortunately, it is impossible in that that kind of calculation will lead to the result very close to that of massive particles. It seems like that the Newtonian second law sounds invalid in light propagation even at the vertical direction.

But there is still an opportunity to explore the ballistic trajectory method for it. That is to perform by trial method. We can imagine that after point A there will be a wave front after Huygens' postulation, then, some points on it may be selected for further considerations. With angular momentum conservation and Lagrangian, one can calculate the velocity and angle α_B as shown in Figure 22. Thus, the deviations of the points could be estimated to help for further trial. I have made more efforts to try but not done, although I still believe the probabilities.

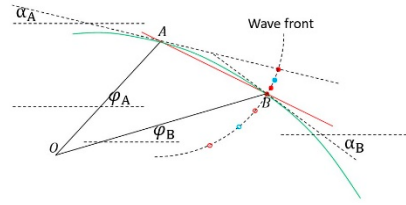


Figure 22. Probability of ballistic trajectory method for light deviation.

8.5.3. A Wrong Treatment for Light Propagations

If the covariant angular momentum conversation would be insisted for light propagations as

$$L = Rm_Rc = r^2m_r \frac{d\varphi}{dt} \quad (413)$$

It is seemingly that $\frac{d\varphi}{dt}$ presents the contra variant angular velocity of photons which is expected to correspond to contra variant angular momentum. And moreover, it could be argued to take $Rm_Rc_0^\mu$ as expressions of contra variant angular momentum, that will exactly lead to conservation break as gravitational redshift is concerned.

Now for the equation Eq. (413), there is

$$\frac{d\varphi}{dt} = \frac{Rm_Rc}{m_r r^2} = \frac{Rc}{r^2} \left(\frac{1+\frac{r^*}{2r}}{1+\frac{r^*}{2r}} \right) = \frac{\tilde{L}}{r^2(1+\frac{r^*}{2r})} \quad (414)$$

where $\tilde{L} = \left(R + \frac{r^*}{2}\right)c$. In weak field, the item $1/(1 + \frac{r^*}{2r}) \approx (1 - \frac{r^*}{2r})$, and $(1 - \frac{r^*}{2r})^2 \approx 1 - \frac{r^*}{r}$, thus

$$\frac{d\varphi}{dt} = \frac{(1-\frac{r^*}{2r})\tilde{L}}{r^2} \quad (415)$$

Because the Lagrangian with coordinate time is

$$-c^2 + (1 - \frac{r^*}{r})^{-2} (\frac{dr}{dt})^2 + (1 - \frac{r^*}{r})^{-1} r^2 (\frac{d\varphi}{dt})^2 = 0 \quad (416)$$

It could be transformed to be

$$c^2 - (1 - \frac{r^*}{r})^{-1} r'^2 \frac{\tilde{L}^2}{r^4} - \frac{\tilde{L}^2}{r^2} = 0 \quad (417)$$

or

$$\left(1 - \frac{r^*}{r}\right) c^2 - r'^2 \frac{\tilde{L}^2}{r^4} - (1 - \frac{r^*}{r}) \frac{\tilde{L}^2}{r^2} = 0 \quad (418)$$

Consider $\tilde{L} \approx bc$ again

$$(1 - r^*u) - u'^2 b^2 - (1 - r^*u)u^2 b^2 = 0 \quad (419)$$

To be derivated to be

$$u'' + u = \frac{3}{2} r^* u^2 - \frac{r^*}{2b^2} \quad (420)$$

It could be solved as

$$u = \frac{\sin\varphi}{b} + \frac{r^*}{4b^2} + \frac{r^*}{4b^2} \cos 2\varphi \quad (421)$$

One will gain obtain that

$$\varphi_\infty = -\frac{r^*}{2b} \quad (422)$$

One can verify the peak point on the case of $\varphi = \pi/2$, and $r = R$, and then will gain

$$b = R \quad (423)$$

That has broken that principle of angular momentum conservation. Of course, it is a wrong answer for light ray propagations, because of a wrong setting.

8.5.4. Equations of Close-to-Light-Speed Massive Particles

We have known that massive particles will not run with general covariance. They run with Newtonian laws. The velocity composition for close-to-light-speed massive particles really is

$$-c^2 + \left(\frac{dr}{dt}\right)^2 + r^2 \left(\frac{d\varphi}{dt}\right)^2 = 0 \quad (424)$$

The conservation of angular momentum

$$L = R m_R c = r^2 m_r \frac{d\varphi}{dt} \quad (425)$$

One will find that $\frac{d\varphi}{dt}$ is used here is because the velocity composition equation Eq. (424), which is rare different from the Lagrangian for light. The constant c is not the speed of light while it really is the approximate speed of massive matter. So that the c in Eq. (424) is just a velocity composition of close-to-light-speed motion. The Eq. (424) and Eq. (425) perform the differences of conservation between light fly and the motion of massive matter. We will see these differences in the equations of low velocity motions of massive matters in next sections.

The angular velocity

$$\frac{d\varphi}{dt} = \frac{R m_R c}{m_r r^2} = \frac{R c}{r^2} \left(\frac{1 + \frac{r^*}{2R}}{1 + \frac{r^*}{2r}} \right) = \frac{\tilde{L}}{r^2 (1 + \frac{r^*}{2r})} \quad (426)$$

where $\tilde{L} = \left(R + \frac{r^*}{2}\right)c$. In weak field, the item $1/(1 + \frac{r^*}{2r}) \approx (1 - \frac{r^*}{2r})$, and $(1 - \frac{r^*}{2r})^2 \approx 1 - \frac{r^*}{r}$, thus

$$\frac{d\varphi}{dt} = \frac{1}{r^2} \left(1 - \frac{r^*}{2r}\right) \tilde{L} \quad (427)$$

Together with $\frac{dr}{dt} = \frac{dr}{d\varphi} \frac{d\varphi}{dt} = r' \frac{d\varphi}{dt}$, the Eq. (424) turns to be

$$c^2 - \left(1 - \frac{r^*}{r}\right) r'^2 \frac{\tilde{L}^2}{r^4} - \left(1 - \frac{r^*}{r}\right) \frac{\tilde{L}^2}{r^2} = 0 \quad (428)$$

or

$$\left(1 + \frac{r^*}{r}\right) c^2 - r'^2 \frac{\tilde{L}^2}{r^4} - \frac{\tilde{L}^2}{r^2} = 0 \quad (429)$$

With $u = \frac{1}{r'}$ and $r' = \frac{dr}{d\varphi} = -\frac{u'}{u^2}$, and considering $\tilde{L} \approx bc$, there is

$$(1 + r^* u) - u'^2 b^2 - u^2 b^2 = 0 \quad (430)$$

To be differentiated as

$$u'' + u = \frac{r^*}{2b^2} \quad (431)$$

It could be solved as

$$u = \frac{\sin\varphi}{b} + \frac{r^*}{2b^2} \quad (432)$$

Obviously, there is

$$\varphi_\infty = -\frac{r^*}{2b} \quad (433)$$

It is easy to calculate in Eq. (432) that at the peak point, $\varphi = \pi/2$, and $r = R$, so that

$$b = R + \frac{r^*}{2} \quad (434)$$

Or with angular momentum conservation there is

$$\left(1 + \frac{r^*}{2r_\infty}\right) b = \left(1 + \frac{r^*}{2R}\right) R \quad (435)$$

Again, we obtain

$$b = R + \frac{r^*}{2} \quad (436)$$

The asymptotic line

$$r = \frac{b}{\sin\varphi + r^* \cos\varphi / b/2} \quad (437)$$

It is said that the trajectories of light rays and close-to-light-speed particles present different φ_∞ but the same peak difference δ . The same peak difference is just the result of momentum conservation.

In fact, these studies have help to create the dynamics of close-to-light-speed particles in gravitational field. By the way, one can compare this solution with that in section 8.4.1 that they have both given the same result of φ_∞ , but they gave different value of b .

8.5.5. Comparisons of Numerical Solutions and Algebraic Solutions

To make comparisons with numerical and algebraic solutions is not only a kind of further verification but also a further support to the conclusion of energy moment conservation. As has discussed previous, ballistic trajectory method could be developed to calculate the trajectories of close-to-light massive particles. It could be applied by dividing trajectory to finite segments and calculating coordinates with Eq. (404) to Eq. (412) step by step. The invalidity of Newtonian second law for light propagation is actually another kind of support to the inferences of kinematics.

Notwithstanding, difference method for all differential equations could also be employed for more comparisons. Difference method has shown great advantages in scientific calculations and many excellent schemes have been developed to serve for more complex requirements.

For the equation of light rays

$$u'' + u = \frac{3}{2} r^* u^2 \quad (438)$$

The simple central difference scheme could be suggested that

$$u'' = \frac{u(\varphi + \Delta\varphi) - 2u(\varphi) + u(\varphi - \Delta\varphi)}{(\Delta\varphi)^2} \quad (439)$$

Thus, the difference equation could be built as

$$\frac{u(\varphi + \Delta\varphi) - 2u(\varphi) + u(\varphi - \Delta\varphi)}{(\Delta\varphi)^2} + u(\varphi) = \frac{3}{2} r^* u^2(\varphi) \quad (440)$$

Pre-exercises show that if step intervals defined by $\Delta\varphi < 0.1^\circ$, the calculations would get quite good accuracy.

As for the equation of close-to-light massive particles, the differential equation

$$u'' + u = \frac{r^*}{2b^2} \quad (441)$$

The difference equation could also be built simply as

$$\frac{u(\varphi + \Delta\varphi) - 2u(\varphi) + u(\varphi - \Delta\varphi)}{(\Delta\varphi)^2} + u(\varphi) = \frac{r^*}{2b^2} \quad (442)$$

Thus, we would sponsor a comparison with the analytical solutions, difference method solutions and the ballistic trajectory method solutions mentioned previously, together with the asymptotic lines and horizontal line for references.

Numerical methods of difference and ballistic were carried in office computer and others were completed by the calculator of my mobile phone, in that the phone calculator could provide 8 floating point precisions more than that of the Fortran software in the computer.

It is shown in Table 3 that difference method and ballistic method still perform not bad precisions in the results of upper parts of the trajectories. Of course, both of the two methods are of step by step arithmetic so as to accumulate quite amount of deviations at the rear parts of the trajectories, especially for that of ballistic method. Anyway, the two numerical methods are reliable, that greatly supports the inferences of kinematics.

It easy to develop more optimal schemes for difference method and the way of optimal secant seeking instead of the tangent seeking for ballistic method or any other effective technologies to improve the precisions of numerical analyses. That will confirm the conclusions and inferences that have been drawn previous.

Table 3. Comparison of the analytical solutions, difference method solutions and the ballistic trajectory method solutions.

coord (degr	coordinate radius (m) of			coordinate radius (m) of solution of close-				coordin horizon
		differenc	asympto	analytica	differenc	ballistic	asympto	
90	69550000	69550000	69550147	69550000	69550000	69550000	69550147	6955014
80	70622914	70622916	70623018	70622918	70622921	70622922	70623044	7062307
70	74013534	74013536	74013606	74013554	74013557	74013559	74013664	7401372
60	80309346	80309348	80309396	80309396	80309399	80309402	80309494	8030959
50	90790914	90790915	90790946	90791018	90791021	90791025	90791108	9079126
40	10820025	10820025	10820027	10820046	10820046	10820047	10820054	1082008
30	13909926	13909925	13909927	13909970	13909970	13909971	13909978	1391002
20	20334865	20334864	20334865	20334976	20334976	20334978	20334984	2033510
10	40051369	40051363	40051369	40051844	40051843	40051847	40051851	4005233
0.26	15312377	15312266	15312377	15319538	15319514	15319539	15319538	1532670
0.01	38902784	38895571	38902784	39370353	39368732	39370335	39370353	3984929
0	16378489	16251576	16378489	32756978	32645005	32761649	32756978	∞

8.6. Time Delay of Radar Echoes

For a light ray passing by the Sun, the velocity varies with the positions and directions on the trajectory. There should be a difference between the real time interval and that calculated via invariant light speed instead of apparent light speed. Shapiro proposed new tests of time delay of radar signals which transmitted from the Earth to pass by the edge of the Sun to another planet or satellite and then reflected back to the Earth [24,25]. The observations on time delay of radar echoes would forcefully support the theory of general relativity as well as that of light ray deflect.

However, the solutions of time delay must have involved in the problems with light trajectory, that the solution process has inherited the errors in classical equations of light trajectory, so that it is necessary to make a detailed discussion to rectify.

8.6.1. Classical Solution

In classical procedure, with the assumptions of $\dot{t}(1 - \frac{r^*}{r}) = E$ and $r^2\dot{\phi} = L$, the Lagrangian could be transformed to be

$$\dot{r}^2 = c^2E^2 - (1 - \frac{r^*}{r})\frac{L^2}{r^2} = 0 \tag{443}$$

And also with

$$\dot{r} = \frac{dr}{dt}\dot{t} = \frac{dr}{dt}E(1 - \frac{r^*}{r})^{-1} \tag{444}$$

The Lagrangian becomes

$$(1 - \frac{r^*}{r})^{-3} (\frac{dr}{dt})^2 = c^2 (1 - \frac{r^*}{r})^{-1} - \frac{1}{r^2} \frac{L^2}{E^2} \quad (445)$$

At the peak point, $\frac{dr}{dt} = 0$ so that there is

$$\frac{L^2}{E^2} = c^2 R^2 (1 - \frac{r^*}{R})^{-1} \quad (446)$$

Where, R is the coordinate of the peak point.

Taking it back into the Eq. (445), there is

$$(1 - \frac{r^*}{r})^{-3} (\frac{dr}{dt})^2 - c^2 (1 - \frac{r^*}{r})^{-1} + c^2 \frac{R^2}{r^2} (1 - \frac{r^*}{R})^{-1} = 0 \quad (447)$$

or

$$(\frac{dr}{dt})^2 = c^2 (1 - \frac{r^*}{r})^2 - c^2 \frac{R^2}{r^2} (1 - \frac{r^*}{r})^3 (1 - \frac{r^*}{R})^{-1} \quad (448)$$

The differential relationship could be integrated that

$$t|_R^r = \frac{1}{c} \int_R^r (1 - \frac{r^*}{r})^{-1} [1 - (1 - \frac{r^*}{r})(1 - \frac{r^*}{R})^{-1} \frac{R^2}{r^2}]^{-1/2} dr \quad (449)$$

Because $\frac{r^*}{r}$ and $\frac{r^*}{R}$ are very small, it could be written as

$$t|_R^r = \frac{1}{c} \int_R^r (1 - \frac{r^*}{r})^{-1} [1 - (1 - \frac{r^*}{r} + \frac{r^*}{R}) \frac{R^2}{r^2}]^{-1/2} dr \quad (450)$$

It has an approximate form [4] as

$$\begin{aligned} t|_R^r &\approx \frac{1}{c} \int_R^r (1 - \frac{R^2}{r^2})^{-1/2} [1 + \frac{r^*}{r} + \frac{r^* R}{2r(r+R)}] dr \\ &= \frac{1}{c} [\sqrt{r^2 - R^2} + r^* \ln \left(\frac{r + \sqrt{r^2 - R^2}}{R} \right) + \frac{r^*}{2} \sqrt{\frac{r-R}{r+R}}] \quad (451) \end{aligned}$$

8.6.2. Errors in

We have known that the assumption in classical equations is incorrect. In fact, the equation could be solved without the assumption. Such as the Lagrangian

$$-c^2 + (1 - \frac{r^*}{r})^{-1} \dot{r}^2 + r^2 \dot{\phi}^2 = 0 \quad (452)$$

With $L = Rc = r^2 \dot{\phi} = \text{const.}$ it is

$$-c^2 + (1 - \frac{r^*}{r})^{-1} \dot{r}^2 + \frac{c^2 R^2}{r^2} = 0 \quad (453)$$

And we know that

$$\dot{r} = \frac{dr}{dt} \dot{t} = \frac{dr}{dt} \frac{dt}{d\tau} = \frac{dr}{dt} (1 - \frac{r^*}{r})^{-1/2} \quad (454)$$

so that

$$(\frac{dr}{dt})^2 = c^2 (1 - \frac{r^*}{r})^2 - c^2 (1 - \frac{r^*}{r})^2 \frac{R^2}{r^2} \quad (455)$$

The differential relationship could be integrated that

$$t|_R^r = \frac{1}{c} \int_R^r (1 - \frac{r^*}{r})^{-1} (1 - \frac{R^2}{r^2})^{-1/2} dr \quad (456)$$

Approximate solution is

$$t|_R^r = \frac{1}{c} [\sqrt{r^2 - R^2} + r^* \ln \left(\frac{r + \sqrt{r^2 - R^2}}{R} \right)] \quad (457)$$

One can find that this treatment has just brought about a little deviation from the previous. That is because the trajectory is only a little different from that one at deflect angle.

8.6.3. Revisit Solution for Radar Echoes

The revisit equation of trajectory may help to get new performances of the issue.

For the Lagrangian

$$-c^2 + (1 - \frac{r^*}{r})^{-1} \dot{r}^2 + r^2 \dot{\phi}^2 = 0 \quad (458)$$

In previous section, we have got the angular velocity expression based on energy momentum conservation as discussed in previous sections that $Rm_{RC} = r^2 m_r \dot{\phi}$ and $\dot{\phi} = \frac{Rm_{RC}}{m_r r^2} = \frac{Rc}{r^2} \left(\frac{1 + \frac{r^*}{2R}}{1 + \frac{r^*}{2r}} \right) = \frac{\tilde{L}}{r^2 (1 + \frac{r^*}{2r})}$, where

$$\tilde{L} = \left(R + \frac{r^*}{2} \right) c \quad (459)$$

In weak field, the item $1/(1 + \frac{r^*}{2r}) \approx (1 - \frac{r^*}{2r})$, thus

$$\dot{\phi} = \frac{1}{r^2} (1 - \frac{r^*}{2r}) \tilde{L} \quad (460)$$

Then the Lagrangian becomes

$$-c^2 + (1 - \frac{r^*}{r})^{-1} \dot{r}^2 + \frac{1}{r^2} (1 - \frac{r^*}{2r}) \tilde{L}^2 = 0 \quad (461)$$

A transformation could be made as

$$-c^2 + (1 - \frac{r^*}{r})^{-2} (\frac{dr}{dt})^2 + (1 - \frac{r^*}{r}) \frac{\tilde{L}^2}{r^2} = 0 \quad (462)$$

With Eq. (459) it becomes

$$-c^2 + (1 - \frac{r^*}{r})^{-2} (\frac{dr}{dt})^2 + (1 - \frac{r^*}{r}) (1 + \frac{r^*}{R}) \frac{R^2 c^2}{r^2} = 0 \quad (463)$$

so that there is

$$(\frac{dr}{dt})^2 = c^2 [(1 - \frac{r^*}{r})^2 - (1 - \frac{r^*}{r})^3 (1 + \frac{r^*}{R}) \frac{R^2}{r^2}] \quad (464)$$

or

$$(dt)^2 = \frac{1}{c^2} [(1 - \frac{r^*}{r})^2 - (1 - \frac{r^*}{r})^3 (1 + \frac{r^*}{R}) \frac{R^2}{r^2}]^{-1} (dr)^2 \quad (465)$$

or

$$dt = \frac{1}{c} [(1 - \frac{r^*}{r})^2 - (1 - \frac{r^*}{r})^3 (1 + \frac{r^*}{R}) \frac{R^2}{r^2}]^{-1/2} dr \quad (466)$$

Integrate above differential to be

$$t|_R^r = \frac{1}{c} \int_R^r (1 - \frac{r^*}{r})^{-1} (1 - \frac{R^2}{r^2} + \frac{r^* R^2}{r^2} - \frac{r^* R^2}{R^2 r^2})^{-1/2} dr \quad (467)$$

or

$$t|_R^r = \frac{1}{c} \int_R^r (1 + \frac{r^*}{r}) [1 - \frac{R^2}{r^2} + (\frac{r^*}{r} - \frac{r^*}{R}) \frac{R^2}{r^2}]^{-1/2} dr \quad (468)$$

By building a function

$$f(x) = x^{-1/2} \quad (469)$$

where, $x = 1 - \frac{R^2}{r^2}$. The derivative

$$f'(x) = -\frac{1}{2} x^{-3/2} \quad (470)$$

Considering $\frac{r^*}{r}$ is very small, the rear part of the integral could be simplified as

$$[1 - \frac{R^2}{r^2} + (\frac{r^*}{r} - \frac{r^*}{R}) \frac{R^2}{r^2}]^{-1/2} \approx (1 - \frac{R^2}{r^2})^{-\frac{1}{2}} - \frac{1}{2} (1 - \frac{R^2}{r^2})^{-3/2} (\frac{r^* R^2}{r^3} - \frac{r^* R}{r^2}) \quad (471)$$

Then the integration could be gain as

$$t|_R^r \approx \frac{1}{c} \int_R^r [(1 - \frac{R^2}{r^2})^{-1/2} (1 + \frac{r^*}{r}) - \frac{1}{2} (1 - \frac{R^2}{r^2})^{-3/2} (\frac{r^* R^2}{r^3} - \frac{r^* R}{r^2})] dr \quad (472)$$

where in the last item, a $(1 + \frac{r^*}{r})$ has been simplified to be 1.

The first part of the integration could be calculated to be

$$\text{part1} = \frac{1}{c} \int_R^r (1 - \frac{R^2}{r^2})^{-1/2} (1 + \frac{r^*}{r}) dr = \frac{1}{c} [\sqrt{r^2 - R^2} + r^* \ln(\frac{r + \sqrt{r^2 - R^2}}{R})] \quad (473)$$

and the last part is done as

$$\text{part2} = -\frac{1}{c} \int_R^r \frac{1}{2} (1 - \frac{R^2}{r^2})^{-3/2} (\frac{r^* R^2}{r^3} - \frac{r^* R}{r^2}) dr = \frac{r r^*}{2\sqrt{r^2 - R^2}} - \frac{r^* R}{2\sqrt{r^2 - R^2}} \quad (474)$$

so that the total integration is

$$t|_R^r \approx \frac{1}{c} [\sqrt{r^2 - R^2} + r^* \ln(\frac{r + \sqrt{r^2 - R^2}}{R}) + \frac{r^*}{2} \sqrt{\frac{r-R}{r+R}}] \quad (475)$$

8.6.4. Solution for Close-to-Light-Speed Particles

Now let's discuss another kind of no delay time spend, the time spend of close-to-light-speed particles. The velocity composition of close-to-light-speed particles could be expressed as

$$-c^2 + (\frac{dr}{dt})^2 + r^2 (\frac{d\varphi}{dt})^2 = 0 \quad (476)$$

With angular momentum conservation, there is

$$-c^2 + (\frac{dr}{dt})^2 + (1 - \frac{r^*}{r}) \frac{\bar{L}^2}{r^2} = 0 \quad (477)$$

or

$$-c^2 + (\frac{dr}{dt})^2 + (1 - \frac{r^*}{r})(1 + \frac{r^*}{R}) \frac{R^2 c^2}{r^2} = 0 \quad (478)$$

so that the integration could be built

$$\begin{aligned} t|_R^r &= \frac{1}{c} \int_R^r [1 - (1 - \frac{r^*}{r})(1 + \frac{r^*}{R}) \frac{R^2}{r^2}]^{-1/2} dr \\ &= \frac{1}{c} \int_R^r [1 - \frac{R^2}{r^2} - (\frac{r^*}{r} - \frac{r^*}{R}) \frac{R^2}{r^2}]^{-1/2} dr \\ &\approx \frac{1}{c} \int_R^r [(1 - \frac{R^2}{r^2})^{-1/2} - \frac{1}{2} (1 - \frac{R^2}{r^2})^{-3/2} (\frac{r^* R^2}{r^3} - \frac{r^* R}{r^2})] dr \\ &= \frac{1}{c} [\sqrt{r^2 - R^2} + \frac{r^*}{2} \sqrt{\frac{r-R}{r+R}}] \quad (479) \end{aligned}$$

We have found that the revisit solution of time spend is the same with classical. The reality is that the classical treatment has got the same trajectory by additional settings, so that it is undoubtful to gain a same time spend with.

8.6.5. Equations of Time Delay

In fact, the problem of time delay of radar echoes is just a kind of performance of light deflection. The results in these discussions are also just extensions of that in the problem of light deflection. The classical equation for time delay is to defined the difference between time interval of light rays and motions of absolute light speed, and the length of the trajectory has been coarsely set to be

$$l = \sqrt{r^2 - R^2} \quad (480)$$

So that the classical equation of half branch time delay could be calculated as

$$\Delta t = \frac{1}{c} \left[r^* \ln \left(\frac{r + \sqrt{r^2 - R^2}}{R} \right) + \frac{r^*}{2} \sqrt{\frac{r-R}{r+R}} \right] \quad (481)$$

But the half branch time delay of light rays, with respect to the time interval of close-to-light-speed particles Eq. (479), will be a little different

$$\Delta t_p \approx \frac{r^*}{c} \ln \left(\frac{r + \sqrt{r^2 - R^2}}{R} \right) \quad (482)$$

One may argue that the trajectory of close-to-light-speed particles will be different to light rays. In fact the real length of the light trajectory also does not equal to $\sqrt{r^2 - R^2}$. One can easily take measures to work out the real length of that trajectory. I prefer to give a more accurate value than Eq. (480) that could be estimated by geometrical relationship while $r \gg R \approx b$ that

$$l_{\text{accurate}} = \sqrt{r^2 - R^2} + R \sin(-\varphi_\infty) = \sqrt{r^2 - R^2} + r^* \quad (483)$$

where, $\varphi_\infty = -r^*/b$ is deflect angle of light ray.

So that a real half branch time delay of light rays with respect to an absolute motion on the very trajectory is

$$\Delta t_{\text{real}} = \frac{1}{c} \left[r^* \ln \left(\frac{r + \sqrt{r^2 - R^2}}{R} \right) + \frac{r^*}{2} \sqrt{\frac{r-R}{r+R}} - r^* \right] \approx \frac{1}{c} \left[r^* \ln \left(\frac{r + \sqrt{r^2 - R^2}}{R} \right) - \frac{r^*}{2} \right] \quad (484)$$

We know that the time delay has been verified to be very high accurate value with respect to the classical equation, that is because the trajectory length has always been set to be $\sqrt{r^2 - R^2}$, which is indeed not accurate length of trajectory line.

If consider the close-to-light-speed particles, the real length of its trajectory line could be estimated to be

$$l_p = \sqrt{r^2 - R^2} + R \sin(-\varphi_\infty) = \sqrt{r^2 - R^2} + \frac{r^*}{2} \quad (485)$$

where, $\varphi_\infty = -r^*/(2b)$ is deflect angle of particles.

With Eq. (479), the half branch time delay of close-to-light-speed particles is

$$\Delta t_p = 0 \quad (486)$$

This result of close-to-light-speed particles is really a spectacular inevitability of kinematics rather than occasionality.

Something different from trajectory investigation, the test of time delay of light and close-to-light-speed particle propagation might allow a quite big separation to the Sun edge thereby to provide not bad accuracy. Furthermore, it could be expected to sponsor an experiment to emit light rays and massive particles on a straight line from a point not very close to the Sun at the same time for time delay comparisons.

8.7. Equations for Un-Close-to-Light-Speed Massive Particles

Let us investigate a un-close-to-light-speed motion for massive particles from specific position R to another position r . Considering a variable mass m_r , the gravity

$$F = -\frac{GMm_r}{r^2} = -m_r c^2 \frac{r^*}{2r^2} \quad (487)$$

Dynamic energy converted from gravitational potential by a motion from R to r is

$$E_{\text{conv}} = -\int_R^r m_p c^2 \frac{r^*}{2\rho^2} d\rho \quad (488)$$

and the total energy

$$E = m_r c^2 = m_R c^2 + E_{\text{conv}} \quad (489)$$

The function of total energy could be written as

$$m_r c^2 = m_R c^2 - \int_R^r m_p c^2 \frac{r^*}{2\rho^2} d\rho \quad (490)$$

Setting energy at position R as a known quantity, it could be differentiated that

$$m_r' = -m_r \frac{r^*}{2r^2} \quad (491)$$

It could be integrated to be

$$\ln m_r - \ln m_R = \frac{r^*}{2r} - \frac{r^*}{2R} \quad (492)$$

Then we gain the expression of variable mass that

$$m_r = m_R e^{\left(\frac{r^*}{2r} - \frac{r^*}{2R}\right)} = m_\infty e^{\frac{r^*}{2r}} \quad (493)$$

where m_∞ is the mass at a farthest position. This equation could be called general mass equation in gravitational field.

And we would like to carry out the expanded expression of gravity equation that

$$F = -m_r c^2 \frac{r^*}{2r^2} = -m_R c^2 \frac{r^*}{2r^2} e^{\left(\frac{r^*}{2r} - \frac{r^*}{2R}\right)} = -m_\infty c^2 \frac{r^*}{2r^2} e^{\frac{r^*}{2r}} \quad (494)$$

One could find that the relativistic mass may not keep mass conservation any more that may surprise us, but it really has been revealed in realities. This is another kind of comprehensive physics which I will not make further discussions in this section. Mass of matter does matter [26]. The truths stay in realities.

Dynamic energy will be converted as

$$E_{conv} = m_r c^2 - m_R c^2 = m_R c^2 [e^{\left(\frac{r^*}{2r} - \frac{r^*}{2R}\right)} - 1] \quad (495)$$

It should be pointed out that the expression of exponential equations of mass variation of course could be used for light ray propagations and close-to-light-speed massive particles for higher accuracies case in strong fields.

For massive matters, with special relativity, the equation of dynamic energy at a position r is

$$E_k = \xi m_r V_r^2 \quad (496)$$

where, $\xi = (1 - \sqrt{1 - \frac{V_r^2}{c^2}}) \frac{c^2}{V_r^2}$, see Eq. (304).

In gravitational field the dynamic energy varies with position that

$$E_k = \xi_r m_r V_r^2 = \xi_R m_R V_R^2 + E_{conv} = \xi_R m_R V_R^2 + m_R c^2 [e^{\left(\frac{r^*}{2r} - \frac{r^*}{2R}\right)} - 1] \quad (497)$$

Conversion energy E_{conv} could be positive case potential release or negative case potential withdrawn. However, dynamic energy E_k always great than zero, so that the variable r will be limited in some specific cases that depends on initial conditions.

NB, for the convenience of expressions, the tensors of velocities, frequencies or the components maybe not written in tensor format anymore case they may not bring about confusions for understandings, for examples, V_r and V_R refer to the velocity at position r and R .

It should be further discussed here that we have seen dynamic mass energy may come from the release of gravitational potential. That will then perform as inertial mass and gravitational mass. If in two source system, they move closer or farther will cause mass increase or lose in that we incline to realize that potential does not act as mass. We are not sure that in this condition mass conservation is available or not. I am inclining to say no. In this section, this controversy does not really matter. This discussion just presents the issue for more focus.

We now prefer to study the conditions of irrelativistic velocities so that $\xi_r \approx \xi_R \approx 0.5$, and after Eq. (493) and Eq. (497), there is the velocity square

$$\begin{aligned} V_r^2 &= e^{\left(\frac{r^*}{2R} - \frac{r^*}{2r}\right)} V_R^2 + 2e^{\left(\frac{r^*}{2R} - \frac{r^*}{2r}\right)} [e^{\left(\frac{r^*}{2r} - \frac{r^*}{2R}\right)} - 1] c^2 \\ &= e^{\left(\frac{r^*}{2R} - \frac{r^*}{2r}\right)} V_R^2 + 2[1 - e^{\left(\frac{r^*}{2R} - \frac{r^*}{2r}\right)}] c^2 \quad (498) \end{aligned}$$

As we have discussed, Lagrangian could no longer be employed for analysis on the motion of massive particles. In some publications, the so called Lagrangian for massive particles is classically defined to be $-c^2$ as [3,10]

$$\mathcal{L} = -(1 - \frac{r^*}{r})c^2\dot{t}^2 + (1 - \frac{r^*}{r})^{-1}\dot{r}^2 + r^2\dot{\varphi}^2 = -c^2 \quad (499)$$

In fact, it could be easily deformed to be

$$\mathcal{L} = -c^2 + (1 - \frac{r^*}{r})^{-1}\dot{r}^2 + r^2\dot{\varphi}^2 = -c^2 \quad (500)$$

That will lead to obvious errors that

$$(1 - \frac{r^*}{r})^{-1}\dot{r}^2 + r^2\dot{\varphi}^2 = 0 \quad (501)$$

That means a zero velocity composition. It is impossible to interpret realities with. We know that in classical equations, the assumption of $\dot{t}(1 - \frac{r^*}{r}) = E$ has still been employed to solve the problem. One might be involved with more errors to solve one more error.

After conclusions and inferences previous, for a massive particle in one source gravitational field, the velocity composition could be expressed as

$$-V_r^2 + (\frac{dr}{dt})^2 + r^2(\frac{d\varphi}{dt})^2 = 0 \quad (502)$$

This equation could also be performed for that of close-to-light-speed massive particles as $V_r \rightarrow c$, as Eq. (424) and Eq. (476), where these discussions have not been released because of consideration of reducing arguments.

With angular momentum conservation, there is

$$r^2 m_r \frac{d\varphi}{dt} = R m_R V_R = \text{const.} \quad (503)$$

Of course, the angular momentum conservation for lower velocity motion will be the same as that of close-to-light-speed massive particles. But they are really different from that of light rays.

So that

$$\frac{d\varphi}{dt} = \frac{R}{r^2} \frac{m_R}{m_r} V_R = \frac{R}{r^2} V_R e^{(\frac{r^*}{2R} - \frac{r^*}{2r})} \quad (504)$$

In which V_R and m_R are velocity and mass at peak point, where $r = R$.

Together with the Eq. (498) and $\frac{dr}{dt} = \frac{dr}{d\varphi} \frac{d\varphi}{dt} = r' \frac{d\varphi}{dt}$, the Eq. (502) becomes

$$-e^{(\frac{r^*}{2R} - \frac{r^*}{2r})} V_R^2 - 2[1 - e^{(\frac{r^*}{2R} - \frac{r^*}{2r})}]c^2 + (r')^2 \frac{1}{r^4} e^{(\frac{r^*}{R} - \frac{r^*}{r})} R^2 V_R^2 + \frac{1}{r^2} e^{(\frac{r^*}{R} - \frac{r^*}{r})} R^2 V_R^2 = 0 \quad (505)$$

or

$$-e^{(\frac{r^*}{2r} - \frac{r^*}{2R})} V_R^2 - 2[e^{(\frac{r^*}{r} - \frac{r^*}{R})} - e^{(\frac{r^*}{2r} - \frac{r^*}{2R})}]c^2 + (r')^2 \frac{1}{r^4} R^2 V_R^2 + \frac{1}{r^2} R^2 V_R^2 = 0 \quad (506)$$

Setting $u = \frac{1}{r}$, it is

$$-e^{(\frac{r^*}{2}u - \frac{r^*}{2R})} V_R^2 - 2[e^{(ur^* - \frac{r^*}{R})} - e^{(\frac{r^*}{2}u - \frac{r^*}{2R})}]c^2 + (u')^2 R^2 V_R^2 + u^2 R^2 V_R^2 = 0 \quad (507)$$

To be derivated by φ , it becomes

$$-\frac{r^*}{4} e^{(\frac{r^*}{2}u - \frac{r^*}{2R})} V_R^2 - [r^* e^{(ur^* - \frac{r^*}{R})} - \frac{r^*}{2} e^{(\frac{r^*}{2}u - \frac{r^*}{2R})}]c^2 + u'' R^2 V_R^2 + u R^2 V_R^2 = 0 \quad (508)$$

To take the one-order approximation for those exponent functions, the equation becomes

$$-\frac{r^*}{4} (1 + \frac{r^*}{2}u - \frac{r^*}{2R}) V_R^2 - [r^* (1 + ur^* - \frac{r^*}{R}) - \frac{r^*}{2} (1 + \frac{r^*}{2}u - \frac{r^*}{2R})]c^2 + u'' R^2 V_R^2 + u R^2 V_R^2 = 0 \quad (509)$$

One can also get this equation in simplified ways, but that will also experience comprehensive complexities while the way above seems more physically explicit. The exponent function forms used here is also to present more primary expressions of mass energy. In the conditions of strong fields,

the exponent forms may take effects. One can see that the kinematics of light or close-to-light particles could also be expressed in more primary ways.

In weak fields, those very small items in Eq. (509) will be neglected. That becomes

$$-\frac{3}{4}r^{*2}c^2u - \frac{1}{2}r^*c^2 + u''R^2V_R^2 + uR^2V_R^2 = 0 \quad (510)$$

or

$$u'' + (1 - \frac{3}{4}\frac{r^{*2}c^2}{R^2V_R^2})u = \frac{1}{2}\frac{r^*c^2}{R^2V_R^2} \quad (511)$$

Setting

$$\omega^2 = 1 - \frac{3}{4}\frac{r^{*2}c^2}{R^2V_R^2} \quad (512)$$

For planets in solar system, the last item in Eq. (512) is far less than 1.0, so that

$$\omega \approx 1 - \frac{3}{8}\frac{r^{*2}c^2}{R^2V_R^2} \quad (513)$$

For planet trajectory, the equation could be solved as [27]

$$u = A[1 - \varepsilon \sin(\omega\varphi)] \quad (514)$$

With this solution substituted back into Eq. (511)

$$A = \frac{1}{2}\frac{r^*c^2}{R^2V_R^2}\omega^{-2} \quad (515)$$

One-order approximation of Eq. (507) is

$$-(1 + \frac{r^*}{2}u - \frac{r^*}{2R})V_R^2 - 2[(1 + ur^* - \frac{r^*}{R}) - (1 + \frac{r^*}{2}u - \frac{r^*}{2R})]c^2 + (u')^2R^2V_R^2 + u^2R^2V_R^2 = 0 \quad (516)$$

Neglecting very small items, it is

$$-V_R^2 + \frac{r^*}{R}c^2 - r^*uc^2 + (u')^2R^2V_R^2 + u^2R^2V_R^2 = 0 \quad (517)$$

With Eq. (514), it becomes

$$-V_R^2 + \frac{r^*}{R}c^2 - r^*Ac^2 + r^*A\varepsilon \sin(\omega\varphi)c^2 + A^2\omega^2\varepsilon^2[\cos(\omega\varphi)]^2R^2V_R^2 + (A - A\varepsilon \sin(\omega\varphi))^2R^2V_R^2 = 0 \quad (518)$$

Case $\varphi = 0$ there is

$$-V_R^2 + \frac{r^*}{R}c^2 - r^*Ac^2 + A^2\omega^2\varepsilon^2R^2V_R^2 + A^2R^2V_R^2 = 0 \quad (519)$$

So that the eccentricity is

$$\varepsilon \approx [\frac{(1-A^2R^2)V_R^2 - (1-AR)\frac{r^*}{R}c^2}{A^2\omega^2R^2V_R^2}]^{-2} \quad (520)$$

Some publications have given the following expression for eccentricity

$$L^2 = a(1 - \varepsilon^2)GM \quad (521)$$

where, L has been set the angular momentum.

But this is not a real solution for eccentricity, because that the Eq. (514) shows $a = A^{-1}(1 - \varepsilon^2)^{-1}$. We will then prefer to present the perihelion precession of

$$\Delta\varphi \approx \frac{3}{8}\frac{r^{*2}c^2}{R^2V_R^2}\varphi \quad (522)$$

or the precession per revolution

$$\Delta\varphi_{2\pi} \approx \frac{3\pi}{4}\frac{r^{*2}c^2}{R^2V_R^2} \quad (523)$$

The most surprising is that this solution of perihelion precession is really irrelativistic. It is a half of the value of the classical solution. It is practicable to carry out more experiments of motions closely around the Sun to verified the conclusions of Eq. (522).

One can focus on more sophisticated conditions for massive matters traveling in gravitational fields, based on the Eq. (508) or the Eq. (511), especially for those motions in strong fields and with high velocities.

Perhaps this result is the only one in which we have focused on that is in contradictions with the observations that have declared perihelion precessions of planets in solar system [17]. But some observations on PSR J0737-3039A/B [28,29] have shown quite big deviations from classical predictions, which had been expected to have perihelion precession together with geodetic precession. Nevertheless, the experiment of Gravity Probe B [30] shows that geodetic item has got accurate results while frame-dragging item hasn't, in which it could be recognized that the geodetic item is really the effect of special relativity. Observations on S2 [31] and PSR 1913+16 [32] also cannot provide positive supports for the predictions of perihelion precession of classical theory.

9. Relativistic Release and Relativistic Frequency Shift

It is predicted that in one source field, the inflow of matters may cause relativistic release. Sole matters moving to the source center for a separation is exactly a kind of inflow, of course. But I want to point out that the inflows of accretions of active galactic nuclei are really normal performances in celestial evolutions. As the matter of facts, they have not been well investigated, so that it is necessary to cast a few concentrations on.

9.1. Dynamics of Active Galactic Nuclei Accretions

9.1.1. Galactic Nuclei Accretions and Planet Rings

The discovery of active galactic nuclei is one of the greatest advances in the proceedings of the astronomy in the 1960s [33,34]. Active galactic nuclei are of the most spectacular extragalactic objects not only because of their extraordinary giant radiations [35] but also due to the complicated spectrum and high red shifts. The emission spectrums of active galactic nuclei are really non-thermal continuum [36]. Those issued models for active galactic nuclei in the past decades, such as the so called standard model [37] or unified model [38], have not really interpreted the mechanism of accretion inflow and energy release. This is the reason that I want to sponsor a brief discussion on the dynamics of active galactic nuclei accretions.

Active galactic nuclei accretion disks exactly experience resemblant kinematics with planet ring systems in that both of them are of surrounding dusts revolving the center source, loosely inside the Roche limit [39]. I am going to sponsor a study firstly with respect to the planet ring systems so that to renovate those recognitions upon the evolution and fueling mechanism relating to active galactic nuclei accretions.

Planet rings are not static structures [40,41]. Their evolutions involve with numerous internal and external processes in which Keplerian shear acts the key role. It causes rings to spread [42] in all process. As we have seen, Saturn rings usually have that striking refined structures.

In the condition that ring particles are small enough, the electromagnet interactions will then overcome the inertials so that to control the motion within the rings. Such particles or together with gases getting to form colloid, will as a whole to perform fluid behaviors so as to bear pressure. Under the driving effects from outer boundaries, a fluid ring may go cubically bolder because pressure increasing helps particles to move deviate vertically to the ring plane, such as the Halo ring of Jupiter [43]. Spectrum observations predict that the particles in Halo ring are less than submicron. Optical observations also show that Halo ring seems blue and gray with respect to that the main rings show red color [43]. It is because that the fine particles scatter more blue rays, just like our sky does, while coarse particles scatter more red.

Firstly, I want to make a discussion based on a fluid ring in the condition that the center velocity equals to Keplerian velocity.

9.1.2. The Structure of Fluid Rings

For Keplerian motion of matters around center object, there is

$$mr\omega_k^2 = \frac{GMm}{r^2} \quad (524)$$

where, m is mass of a particle, G is gravitational constant, M is the mass of center source, r is distance to the center and $\omega_k = \frac{v_k}{r}$ is angular velocity while v_k is tangent velocity.

The tangent velocity and angular velocity will increase as radius getting smaller that

$$v_k = \frac{(GM)^{0.5}}{r^{0.5}} \text{ and } \omega_k = \frac{(GM)^{0.5}}{r^{1.5}} \quad (525)$$

Keplerian kinetic energy could be expressed as

$$E_k = \frac{1}{2}mv_k^2 = \frac{1}{2}m \frac{GM}{r} \quad (526)$$

Keplerian motion means constant angular momentum for a particle at specific position.

$$L_k = rmv_k = mr^{0.5}(GM)^{0.5} \quad (527)$$

Namely, there is Keplerian shear at the specific position, and the shearing rate can be derived to be

$$\frac{dv_k}{dr} = -\frac{(GM)^{0.5}}{2r^{1.5}} \text{ and } \frac{d\omega_k}{dr} = -\frac{3(GM)^{0.5}}{2r^{2.5}} \quad (528)$$

We will start from a simple condition that a fluid ring has a Keplerian velocity at its center position. For a differential element at a position of radius r in the ring, there could be unbalanced force between gravity and inertial force on the plane of rings. The reasons of unbalanced force may come from low velocity matters falling from outer rings, the driving force, or the resistance of inner parts which have higher velocities and be restrained by outer parts. These interactions in a ring may lead to bold ring as Figure 23 shows, that really others to the flat rings as that of Saturn.

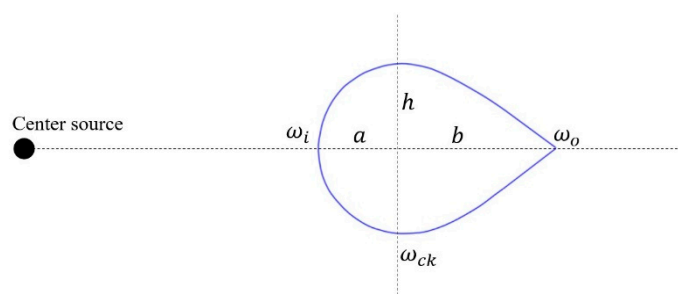


Figure 23. Fluid ring in one source field.

This unbalanced force will be rebalanced by pressure gradient, which will take the part to form the fluid pressure. For example, in the outer part of a fluid ring, if the velocity is lower to form driving force, the pressure could be calculated as

$$p = \int_{r_o}^r (\rho \frac{GM}{x^2} - \rho x \omega^2) dx \quad (529)$$

where, ρ is density of fluid, r_o is radius respective to the outer edge of fluid ring, and the angular velocity ω is less than Keplerian angular velocity ω_k of the very point and greater than the center one so that to give driving force. If a linear function is defined as

$$\omega^2 = [\omega_k^2 - \frac{r-r_c}{r_o-r_c} \Delta\omega_o^2] \quad (530)$$

where, r_c refers to radius of pressure center and $\Delta\omega_o^2$ represents the angular velocity square difference with respect to ω_k^2 for the position of outer edge point. It interprets active falling down of the outer part because of losing of velocity.

Concern that at every position there is the relationship

$$\rho \frac{GM}{r^2} = \rho r \omega_k^2 \quad (531)$$

The Eq. (529) becomes

$$p = \int_{r_o}^r \rho x \frac{x-r_c}{r_o-r_c} \Delta \omega_o^2 dx \quad (532)$$

We know that the pressure of inner edge and outer edge equal to zero. Concern $r_o - r_c \ll r_c$ There is an approximate solution of center pressure

$$p_c \approx \frac{1}{2} \rho r_c (r_o - r_c) \Delta \omega_o^2 = \frac{1}{2} \rho r_c b \Delta \omega_o^2 \quad (533)$$

where, $b = r_o - r_c$ represents outer radius of fluid ring.

On another side, the resistant force determined by the higher velocity of inner parts, with velocity as

$$\omega^2 = [\omega_k^2 + \frac{r-r_i}{r_c-r_i} \Delta \omega_i^2] \quad (534)$$

In which ω^2 is greater than ω_k^2 so that to form a passive resistant force.

There will be the center pressure

$$p_c \approx \frac{1}{2} \rho r_c a \Delta \omega_i^2 \quad (535)$$

where, $a = r_c - r_i$ represents inner radius of fluid ring.

On the direction vertical to the plane, it is easily to calculate that

$$p_c \approx \int_0^h \rho r \omega_{ck}^2 \frac{y}{r} dy = \frac{\rho h^2}{2} \omega_{ck}^2 \quad (536)$$

where, h represents vertical height from top or the bottom to the center of the ring, y is the distance to pressure center and ω_{ck} is Keplerian angular velocity at pressure center.

Thus, the relationships between semi axes are

$$a \approx \frac{b \Delta \omega_i^2}{\Delta \omega_o^2} \quad (537)$$

$$a \approx \frac{h^2 \omega_{ck}^2}{r_c \Delta \omega_i^2} \quad (538)$$

$$b \approx \frac{h^2 \omega_{ck}^2}{r_c \Delta \omega_o^2} \quad (539)$$

The outer side of the ring may be called driving side, depending on a random falling of outer particles, that forms an ambiguous boundary. The inner side of the ring comes from passive driving, may be called driven side, so that to form a sharp edge as seen in Halo ring [43], just like water surface in an accelerated container. Of course, fluid rings also can be flattened and split to be refined structures so that performs more complicated than Halo ring.

I just want to point out that there is a special condition that

$$\omega_i = \omega_o = \omega_c = \omega_{ck} \quad (540)$$

Thus, the whole ring will have a unified velocity so that there will be no Keplerian shear in the ring. And then

$$\Delta \omega_o^2 = \frac{GM}{r_c^3} - \frac{GM}{r_o^3} \quad (541)$$

as well as

$$\Delta \omega_i^2 = \frac{GM}{r_i^3} - \frac{GM}{r_c^3} \quad (542)$$

and

$$a \approx b \quad (543)$$

That will lead to an equilibrium state that the ring stops evolving because of Keplerian shear vanishing. But this state may need strict conditions and may be destroyed by outer influences in a long term. Halo ring might have reached an approximately equilibrium state which could be investigated by detailed observations in the future.

Secondly, I want to discuss the split of a ring under the condition of developed Keplerian shear. If a fluid ring is well Keplerian sheared, it could be predicted that the ring will go thinner. We know that there are Keplerian shear at every position in the ring. But we can make an analysis on the angular momentum conversion in the middle point which could then at last split the ring to two parts. For convenience, it could be defined as two rings originally, the ring i and the ring $i + 1$ as shown in Figure 24. They will have their own Keplerian velocity points at r_i and r_{i+1} . In a time interval, the two rings will complete quite amount of angular momentum conversion because of the Keplerian shear, which is just like viscous friction acts in the interface that makes the conversion.

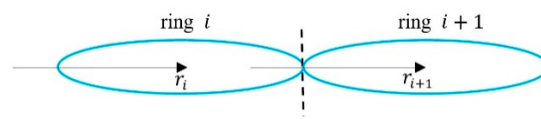


Figure 24. Angular momentum conversion between neighboring rings.

Setting the mass of rings $m_i = m_{i+1} = m$, the relative angular momentum difference between two neighboring rings could be estimated as

$$rm\Delta v = -m \frac{(GM)^{0.5}}{2r^{0.5}} \Delta r \quad (544)$$

where $\Delta r = r_{i+1} - r_i$ and $r \approx 0.5(r_{i+1} + r_i)$

The exchange of angular momentum may be expressed as

$$\Delta L = -\beta m \frac{(GM)^{0.5}}{2r^{0.5}} \Delta r \quad (545)$$

where, momentum exchange ratio $0 \leq \beta \leq 1$.

It is said that after a Keplerian shear, the outer one should get increase of angular momentum and the inner one should get that decline. For the reasons of kinetic mechanism, both the inner ring and outer ring will be driven apart from original position. If inner ring goes a difference δr , the difference of Keplerian angular momentum could be calculated as

$$\delta L_k = m \frac{(GM)^{0.5}}{2r^{0.5}} \delta r \quad (546)$$

Let δL_k equals to ΔL , it is obtained that

$$\delta r = \beta \Delta r \quad (547)$$

For the inner ring, it goes to shrink by falling a (minus) δr separation so that to form a new state. And for the outer ring, it expands by going up to a (positive) δr separation.

We have seen that this is not a whole dynamic analysis because that the pressure in the ring has also taken the effect of angular momentum conversion, but that does no harm for the conclusion that Keplerian shear, i.e. the viscous shear in the ring, leads to angular momentum conversion and rings split. One can make detailed study with Navier Stokes equations.

It is said that a process of rings spreading around a source center is an irreversible process with entropy production.

9.1.3. Shearing Dissipation in Fluid Rings

This issue is carried out to estimate the shearing dissipation so that to discuss the evolution by Keplerian shear. Of course, the evolution would involve together with momentum conversion and

rings split. But in the aspect of theoretical methodology, it is better to be investigated in two steps, the angular momentum conversion by Keplerian shear, and the energy momentum conservation during rings split.

Angular momentum conversion must be done across the interface between the two rings, so that the first step analysis is reasonable. I prefer to focus on the states before and after momentum conversion of two neighboring rings. Before conversion, the neighboring rings have been separated by $\Delta r = r_{i+1} - r_i$, ($r_{i+1} > r_i$), with velocities

$$v_i = \frac{(GM)^{0.5}}{r_i^{0.5}} \text{ and } v_{i+1} = \frac{(GM)^{0.5}}{r_{i+1}^{0.5}} \quad (548)$$

The difference between v_i and v_{i+1} is

$$\Delta v = -\frac{(GM)^{0.5}}{2r^{1.5}} \Delta r \quad (549)$$

where Δv is defined as minus quantity while Δr as plus. Namely,

$$v_{i+1} = v_i + \Delta v \quad (550)$$

Considering ring mass equals to each other, two conditions would be taken into study.

As angular momentum conversion happens, there are

$$v_{i,after} = v_i + \beta \Delta v \text{ and } v_{i+1,after} = v_i + \Delta v - \beta \Delta v \quad (551)$$

As Δv has been defined minus, the inner ring is really slowed down and the outer ring is accelerated. Because of $0 \leq \beta \leq 1.0$, kinetic energy may not keep conservation. With original kinetic energy of

$$E_{before} = \frac{1}{2} m v_i^2 + \frac{1}{2} m (v_i + \Delta v)^2 \quad (552)$$

and the kinetic energy after conversion of

$$E_{after} = \frac{1}{2} m (v_i + \beta \Delta v)^2 + \frac{1}{2} m (v_i + \Delta v - \beta \Delta v)^2 \quad (553)$$

Thus, there is the dissipation energy without considering potential variations

$$Dissipation_1 = E_{before} - E_{after} = \beta m \Delta v^2 - \beta^2 m \Delta v^2 \quad (554)$$

Considering the Eq. (549), it turns to be

$$Dissipation_1 = \beta(1 - \beta) m \frac{GM}{4r^3} \Delta r^2 \quad (555)$$

Then we will go to the second step. After the first step of angular momentum conversion, the outer part of inner ring loses momentum so that provides driving force to drive the inner ring to inflow and the inner part of outer ring will be accelerated to higher momentum so that to compress the outer ring to outflow. This evolution will perform a little sophisticated in that the inflow ring should be accelerated by gravity and the outflow ring should be slow down by gravity. We know that the inflow of a ring will cause half potential release surplus with respect to the Keplerian energy requirement and the outflow will experience potential withdraw. And also, we know that the gravity is perpendicular to the rings. These issues could be solved in fluid flow anyway, for example the evolution of vortexes. But we are just going to check the energy conservation of final state.

After the inner ring inflows a separation δr and the outer ring outflows a separation δr , they come to a new state of Keplerian balance. Taking the inner ring for granted, before Keplerian shear, its total energy involves with Keplerian kinetic energy and potential

$$E_{i,before} = \frac{1}{2} m \frac{GM}{r_i} + 0 \quad (556)$$

After Keplerian shear and inflowing a separation δr , total energy becomes

$$E_{i,after} = \frac{1}{2} m \frac{GM}{r_i - \delta r} - m \frac{GM}{r_i^2} \delta r \quad (557)$$

and for outer ring, the energy before

$$E_{i+1,before} = \frac{1}{2}m \frac{GM}{r_{i+1}} + 0 \quad (558)$$

The energy of final state

$$E_{i+1,after} = \frac{1}{2}m \frac{GM}{r_{i+1}+\delta r} + m \frac{GM}{r_{i+1}^2} \delta r \quad (559)$$

Energy difference could be estimated by

$$\begin{aligned} & E_{i,before} - E_{i,after} + E_{i+1,before} - E_{i+1,after} \\ &= \left(\frac{1}{2}m \frac{GM}{r_i} - \frac{1}{2}m \frac{GM}{r_i - \delta r} \right) + \left(\frac{1}{2}m \frac{GM}{r_{i+1}} - \frac{1}{2}m \frac{GM}{r_{i+1} + \delta r} \right) + \left(\frac{GM}{r_i^2} \delta r - \frac{GM}{r_{i+1}^2} \delta r \right) \quad (560) \end{aligned}$$

The first two items could be transformed as

$$\begin{aligned} & \left(\frac{1}{2}m \frac{GM}{r_i} - \frac{1}{2}m \frac{GM}{r_i - \delta r} \right) \\ &= \frac{1}{2}m \frac{GM}{r_i} - \frac{1}{2}mGM \left(\frac{1}{r_i} + \frac{1}{r_i^2} \delta r + \frac{1}{r_i^3} \delta r^2 \right) \\ &= GMm \left(-\frac{1}{2r_i^2} \delta r - \frac{1}{2r_i^3} \delta r^2 \right) \quad (561) \end{aligned}$$

The middle two items could be transformed as

$$\begin{aligned} & \left(\frac{1}{2}m \frac{GM}{r_{i+1}} - \frac{1}{2}m \frac{GM}{r_{i+1} + \delta r} \right) \\ &= \frac{1}{2}m \frac{GM}{r_{i+1}} - \frac{1}{2}mGM \left(\frac{1}{r_{i+1}} - \frac{1}{r_{i+1}^2} \delta r - \frac{1}{r_{i+1}^3} \delta r^2 \right) \\ &= GMm \left(\frac{1}{2r_{i+1}^2} \delta r + \frac{1}{2r_{i+1}^3} \delta r^2 \right) \\ &= GMm \left[\frac{1}{2(r_i + \Delta r)^2} \delta r + \frac{1}{2r_{i+1}^3} \delta r^2 \right] \\ &= GMm \left[\frac{1}{2r_i^2} \left(1 - \frac{2\Delta r}{r_i} \right) \delta r + \frac{1}{2r_{i+1}^3} \delta r^2 \right] \quad (562) \end{aligned}$$

so that the previous four items are

$$-GMm \frac{\Delta r \delta r}{r_i^3} = -GMm \beta \frac{\Delta r^2}{r_i^3} \quad (563)$$

The last two items are

$$\begin{aligned} & \left(\frac{GMm}{r_i^2} \delta r - \frac{GMm}{r_{i+1}^2} \delta r \right) \\ &= \frac{GMm}{r_i^2} \delta r - \frac{GMm}{(r_i + \Delta r)^2} \delta r \\ &= \frac{GMm}{r_i^2} \delta r - \frac{GMm}{r_i^2} \delta r \left(1 - \frac{2\Delta r}{r_i} \right) \\ &= 2 \frac{GMm}{r_i^3} \Delta r \delta r \\ &= 2GMm \beta \frac{\Delta r^2}{r_i^3} \quad (564) \end{aligned}$$

The energy difference is total energy dissipation

$$\begin{aligned} \text{Dissipation} &= E_{i,before} - E_{i,after} + E_{i+1,before} - E_{i+1,after} \\ &= \beta GMm \frac{\Delta r^2}{r_i^3} \quad (565) \end{aligned}$$

One can calculate the gravity potential release of a single ring inflow, although it does not sound reasonable for a comparison with dissipations because there is another ring outflow. Anyway, that is

$$dW = m \frac{GM}{r^2} \delta r = \beta m \frac{GM}{r^2} \Delta r \quad (566)$$

Thus, we get the ratio of dissipation and potential release

$$\frac{\text{Dissipation}}{dW} = \frac{\Delta r}{r} \quad (567)$$

Or the ratio of dissipation-1 and potential release

$$\frac{\text{Dissipation}_1}{dW} = \frac{(1-\beta)\Delta r}{4r} \quad (568)$$

where $\beta \neq 0$.

Rings evolution dynamics indicates that the inner half of an accretion inclines to inflow and outer half outflows by and large. Additionally, the process of vortex evolution would dissipate, even though the two dissipations are not independent.

$$\text{Dissipation}_2 = \text{Dissipation} - \text{Dissipation}_1 = \frac{1}{4}\beta(3 + \beta)m \frac{GM}{r^3} \Delta r^2 \quad (569)$$

It shows that the dissipation in vortex evolution might be greater than that in momentum exchange between rings.

We know that inter-rings separation Δr would be far less than the radius r of relative rings. It is said that in the condition of Keplerian shear, the dissipative energy is a higher rank infinitesimal quantity of that of potential release of inflow, no matter that inner ring potential release will support the outer ring potential withdraw. Classical quasars could have luminosities of 10^{45} erg s^{-1} to 10^{46} erg s^{-1} [44]. It is unimaginable that those releases completely come from gravitational potential energy.

9.1.4. On Particle Rings

For solid particle rings, it is difficult to make a simple analysis on angular momentum conversions and on ring's evolutions. In fact, any collisions between two particles might cause great changes for their trajectories. It could be imagined that for numerous particles in neighboring rings, assuming very small velocity difference between every Keplerian collision particles, we can predict quasi fluid behavior for Keplerian shear acting on neighboring rings, that in a whole, Keplerian shear help to make angular momentum conversion and rings split. One can also write down those equations of energy momentum just as that of fluid rings. But all in all, particle rings cannot form pressure so that they will only spread as a very thin disk with perfect and refined structures in the equator plane, unlike the fluid rings, which will form a bold 3-dimensional torus structures. I believe that numerical method could be expected for quasi fluid behavior simulation for Keplerian shear of particle rings, as the effects of the rings of Saturn have shown us.

9.2. Relativistic Release

Any efforts to conceiving the mechanism of relativistic release would encounter great difficulties, because of two reasons that relativistic release has not been well recognized and we still know less about the intrinsic structures of fundamental particles that would be seriously related to relativistic emissions. But on the other hand, tremendous of observations have shown incredible probabilities for intrinsic emissions and correlated pseudo redshifts, as well as intrinsic redshift of absorptions, which are so intensely corresponding to the inference of relativistic release that I cannot help to make a try.

Despite the complexities of relativistic release for a massive matter inflow into the center source, I suggest a mathematic analysis based on an assumption of continuum release. It is still not a true expression of realities, but that may help for more understandings on that topic.

The apparent light speed c_0^1 could be employed to describe the concept of equivalent state in gravitational fields. It may be a state after relativistic release, although this state should be not really stable because there should be energy deficiency in the other two dimensions.

One can discuss the momentum exchange efficiency under the conditions of different apparent light speeds so that to talk about the variation of equivalent state. Another important clue that indicates equivalent state is the fine structure constant. We know that fine structure constant determines the dimensions of condensed matters, of course that may determine the dimensions of intrinsic structures. We can imagine that the variation of apparent light speed could bring about the variation of fine structure constant so that to bring about the variation of the energy of equivalent state.

I don't think all of energy of a matter would take part in the relativistic release, for example, the most fundamental particles may be expected to be made of quasi photons, in common recognitions, they might not be split anymore. Thus, we would rather define the concept of releasable mass \tilde{m} . I cannot give an estimation whether the maximum releasable mass should be up to $0.9m_0$ or not, but it could be believed to be quite amount, where m_0 is total rest mass of original equivalent state, which corresponding to the state that the matters do not experience relativistic release.

Now for a center source field, equivalent state of matters is proposed to be

$$\tilde{E} = \tilde{m}(c_0^1)^2 \quad (570)$$

Matters located at different positions in the gravitational field perform different equivalent states. As a matter goes a separation dr along the radial direction, the expression of exceeding energy is

$$d\tilde{E} = 2\tilde{m}c_0^1(c_0^1)'dr \quad (571)$$

As a matter goes from the farthest point to the position of radius r , the relativistic released energy could be integrated to be

$$\tilde{E}_{released} = \int_r^\infty d\tilde{E} = \tilde{m}c^2 - \tilde{m}(c_0^1)^2 = \tilde{m}c^2[1 - (1 - \frac{r^*}{r})^2] \quad (572)$$

Thereafter, we could also define a corresponding concept of the relativistic residual energy that, at a position, matters keep an amount of energy for subsequent releases,

$$\tilde{E}_{residual} = \tilde{m}c^2(1 - \frac{r^*}{r})^2 \quad (573)$$

It seems that the residual energy may go zero as matters reaches r^* . It is just a solution of theoretical model. In practice, relativistic release will be intermittent, and matters perhaps to experience flattenizations as they close to r^* so that to bring about instabilities. As results, residual energy may not go zero at the end in practice. This effort is just to manage to perform kind of probabilities of relativistic release.

It can be calculated that most part of energy is kept before $10r^*$

$$\tilde{E}_{residual}(10r^*) = 0.81\tilde{m}c^2 \quad (574)$$

Now it is naturally to define the release rate by derivative of release energy along the radial direction to the center source. Released energy could be the integration of release distribution function from a position farthest to the position r or the minus inversely

$$\tilde{E}_{released} = \int_\infty^r e dr = \int_r^\infty -e dr \quad (575)$$

On the other hand, the residual energy could also be defined as an integration of release distribution from position r^* to position r or the minus form

$$\tilde{E}_{residual} = \int_{r^*}^r e dr = \int_r^{r^*} -e dr \quad (576)$$

Thus, the release distribution could be derivated to be

$$e = 2\tilde{m}c^2(1 - \frac{r^*}{r}) \frac{r^*}{r^2} \quad (577)$$

It is easy to calculate that the peak value locates at the position

$$r_{\max} = \frac{3}{2} r^* \quad (578)$$

with the maximum release distribution

$$e_{\max} = \frac{8}{27} \frac{\tilde{m} c^2}{r^*} \quad (579)$$

One can calculate the release intensity of accretion inflow. Setting equivalent mass for every single ring in an accretion, the luminosity could be derived as

$$l = 2\tilde{\rho}^* v c^2 \left(1 - \frac{r^*}{r}\right) \frac{r^{*2}}{r^3} \quad (580)$$

where, $\tilde{\rho}^*$ is matter's assuming density at the position r^* and v is inflow velocity.

Suppose the accretion spreading sufficiently, it can be calculated that the most luminous area is at the position

$$r_{l\max} = \frac{4}{3} r^* \quad (581)$$

As well as maximum luminosity

$$l_{\max} = \frac{27}{128} \frac{m\tilde{\rho}^* v c^2}{r^*} \quad (582)$$

The Eq, (580) represents intensity of release at a specific position in an accretion disk. It is revealed that the peak luminosity should not happen at the edge of horizon event, but at a little outer position, just as that was shown in the event-horizon-scale images of M87 taken by the Event Horizon Telescope Collaboration with wavelength of 1.3mm [45].

9.3. Relativistic Emission Lines and Relativistic Redshift

Eyeing on the stimulated release of electrons in atoms, we could carry out a concept of exceeding ratio for relativistic release of an energy structure.

$$\sigma = \frac{\tilde{E}_{\text{excited}} - \tilde{E}_{\text{ground}}}{\tilde{E}_{\text{ground}}} \quad (583)$$

where $\tilde{E}_{\text{excited}}$ and $\tilde{E}_{\text{ground}}$ are energy of excited state and ground state of specific structure. Exceeding ratios for relativistic release may be more different from that of stimulated release of electrons, probably, they might be quite small or big. We know less about that.

In one source field, exceeding ratio would be variable with position as the so called equivalent state has shown.

$$\sigma = \sigma_{\infty} \left(1 - \frac{r^*}{r}\right)^2 \quad (584)$$

Thus, frequencies of specific emission rays could be written as

$$\nu_r = \nu_{\infty} \left(1 - \frac{r^*}{r}\right)^2 \quad (585)$$

Case $r \gg r^*$ it could be simplified to be

$$\nu_r = \nu_{\infty} \left(1 - \frac{2r^*}{r}\right) \quad (586)$$

These emissions of course include but not limit to electron transition emissions that we are more familiar to that could be easily certificated comparing to stimulated release on the Earth. This kind of frequency will be quite different with which we have known well in weak field, in that it looks like redshifted after photons reach the Earth. But it is just a pseudo redshift because the emissions have not experienced real redshift at the emission time. The fact is that it is only verified to be redshifted by comparing to the spectrum characteristics of a matter in weak fields. In this case, it is just redshift seemingly. It could be called relativistic redshift.

In practice, emission lines will experience gravitational redshift to arrive a farthest position. The Eq. (493) could also be employed to interpret light mass energy variation in gravitational field, so that the detectable frequency could be expressed as

$$\nu_{r \rightarrow d} = \nu_{\infty} (1 - \frac{r^*}{r})^2 e^{-\frac{r^*}{2r}} \tag{587}$$

Case $r \gg r^*$ it could be simplified to be

$$\begin{aligned} \nu_{r \rightarrow d} &\approx \nu_{\infty} (1 - \frac{r^*}{r})^2 (1 + \frac{r^*}{2r})^{-1} \\ &= \nu_{\infty} (1 - \frac{r^*}{r})^{2.5} \approx \nu_{\infty} (1 - 2.5 \frac{r^*}{r}) \tag{588} \end{aligned}$$

Propose a definition of relativistic frequency shift based on frequency

$$z_{rel} = \frac{\nu_{\infty} - \nu_r}{\nu_r} = (1 - \frac{r^*}{r})^{-2} - 1 \tag{589}$$

Case $r \gg r^*$ it could be simplified to be

$$z_{rel} = \frac{\nu_{\infty} - \nu_r}{\nu_r} = \frac{2r^*}{r} \tag{590}$$

It is an intrinsic frequency shift rather than cosmological redshift. Note that ν_r keep invariant during propagation process if neglect the effect of gravitational redshift. ν_{∞} is the comparative frequency of equivalent radiation in no gravity condition.

Detectable redshift should be calculated as

$$z_{r \rightarrow d} = \frac{\nu_{\infty} - \nu_{r \rightarrow d}}{\nu_{r \rightarrow d}} = (1 - \frac{r^*}{r})^{-2} e^{\frac{r^*}{2r}} - 1 \tag{591}$$

Case $r \gg r^*$ it could be simplified to be

$$z_{r \rightarrow d} = (1 - \frac{r^*}{r})^{-2.5} - 1 \approx \frac{2.5r^*}{r} \tag{592}$$

Relativistic frequency shift and detectable frequency shift could be calculated for comparison in Table 4.

Table 4. Relativistic and detectable frequency shifts at different positions.

Emission positions	Relativistic frequency shifts	Detectable frequency shifts
$1.3r^*$	17.778	26.585
$1.4r^*$	11.250	16.508
$1.5r^*$	8.000	11.561
$2.0r^*$	3.000	4.136
$3.0r^*$	1.250	1.658
$10.0r^*$	0.235	0.298
$20.0r^*$	0.108	0.136

The highest redshift we have observed on quasars is up to 16.4 [46,47]. It could be expected that more higher redshift will be seen to be up to more than 20.0 in recent future. Perhaps higher redshifts have already been observed but not certificated.

9.4. Broad Lines and Narrow Lines

In the emissions of a quite amount of continuous inflows, the exciting width that the inflow involved sustains at radial direction may cause specific continuous emission distribution. As a result, that could be certificated to be a broadened line in spectrum diagram. For a massive accretion around

a galactic nucleus, the Kepler shear may be employed to interpret the mechanism of inflow of a ring. If there is an inflow of a ring at position r with exciting width of Δr , the emission frequency may distribute from ν_r to $\nu_{r+\Delta r}$. Considering the effect of gravitational redshift, one can get a line width expressed by frequency after Eq. (587) that

$$w_\nu = \nu_{r+\Delta r \rightarrow d} - \nu_{r \rightarrow d} = [(1 - \frac{r^*}{r+\Delta r})^2 e^{-\frac{r^*}{2(r+\Delta r)}} - (1 - \frac{r^*}{r})^2 e^{-\frac{r^*}{2r}}] \nu_\infty \tag{593}$$

Case $r \gg r^*$ it could be simplified to be

$$w_\nu \approx (1 - \frac{2.5r^*}{r+\Delta r} - 1 + \frac{2.5r^*}{r}) \nu_\infty = (\frac{1}{r} - \frac{1}{r+\Delta r}) 2.5r^* \nu_\infty \approx \frac{2.5\Delta r r^*}{r^2} \nu_\infty \tag{594}$$

or by wave length

$$\begin{aligned} w_\lambda &= \lambda_r - \lambda_{r+\Delta r} = c/\nu_{r \rightarrow d} - c/\nu_{r+\Delta r \rightarrow d} \\ &= [(1 - \frac{r^*}{r})^{-2} e^{\frac{r^*}{2r}} - (1 - \frac{r^*}{r+\Delta r})^{-2} e^{\frac{r^*}{2(r+\Delta r)}}] \lambda_\infty \end{aligned} \tag{595}$$

Case $r \gg r^*$ it could be simplified to be

$$w_\lambda \approx (1 + \frac{2.5r^*}{r} - 1 - \frac{2.5r^*}{r+\Delta r}) \lambda_\infty \approx \frac{2.5\Delta r r^*}{r^2} \lambda_\infty \tag{596}$$

which could also be amount to a so called Doppler velocity by frequency as

$$v_{D\nu} = \frac{w_\nu}{\nu_r} c \tag{597}$$

or Doppler velocity by wave length as

$$v_{D\lambda} = \frac{w_\lambda}{\lambda_r} c \tag{598}$$

where, c is light speed.

Case $r \gg r^*$ it could be simplified to be

$$v_{D\nu} \approx v_{D\lambda} \tag{599}$$

These equations show that in the region near to r^* , relativistic emissions should have broader line widths.

Calculations on line widths of various conditions will be presented in Table 5.

Table 5. Emission line widths at different positions with specific exciting widths.

Emission positions	Exciting widths	w_ν/ν_∞	$v_{D\nu}$ km/s	w_λ/λ_∞	$v_{D\lambda}$ km/s
$1.5r^*$	$0.1r^*$	0.02327	87681	2.8408	67850
$1.5r^*$	$0.05r^*$	0.01158	43634	1.5949	38093
$3.0r^*$	$0.1r^*$	0.01433	11425	0.09752	11006
$3.0r^*$	$0.05r^*$	0.007239	5772	0.05018	5663
$10.0r^*$	$0.1r^*$	0.002079	809	0.003492	807
$10.0r^*$	$0.05r^*$	0.001044	406	0.001756	406
$20.0r^*$	$0.1r^*$	0.0005706	194	0.0007360	194
$20.0r^*$	$0.05r^*$	0.0002860	97	0.0003690	97
$100.0r^*$	$0.1r^*$	0.00002455	7.6	0.00002582	7.6
$100.0r^*$	$0.05r^*$	0.00001228	3.8	0.00001291	3.8

It is said that line widths depend on inflow exciting width, but in totally analysis, they highly relate to the inflow positions. The regions close to r^* may perform broader line than farther regions. Observations show that the sizes of broad line regions are estimated within 0.01 or 0.1pc and line

widths are about 1000 kms^{-1} to 10000 kms^{-1} , while the sizes of narrow line regions are about 0.1 pc to 1 kpc [48,49,50]. In fact, those certificated narrow lines incline to have been certificated as lower redshift lines. For radio emissions, some of them were observed to have compact cores and extended components [48]. The compact cores could be interpreted that equivalent state in inner areas of accretions would have lower equivalent energy to release lower frequency emissions, while giant extended components of emission pictures might be the jet outflow who emit radio lines.

Relativistic release must perform in sophisticated conditions. The most probable condition is that releases in one of the rings could overwhelming all of others, especially that of the most inner ring. Thus, in most cases, releases of an active galactic nucleus might be certificated by the emissions of the sole ring. One can find that most of narrow lines have been certificated lower redshift, such as NGC 4151 with $z=0.0033$ [51] and MCG-5-23-16 with $z=0.00849$ [52] have narrow Fe K α lines. But on the other hand, their spectrums both involve with broad lines [53], that indicate controversies in emission line certifications.

The variability, line asymmetry shift, wave length shift and so much as line broadening of active nuclei emissions [54] indicate more sophisticated dynamic conditions than we imagine for the evolution processes of accretions.

9.5. Relativistic Absorption

Most of active galactic nuclei have giant accretion bodies spreading around their equatorial planes. Case a pulse of light rays go through some parts of an accretion body, that light ray may experience different absorptions due to their passing positions. If a pulse of light crosses several separate rings or blocks of an accretion, it must experience multiple absorptions. Absorption frequencies depend on equivalent state of the matters that have experienced relativistic release, so that the absorption frequencies could be expressed as

$$\nu_r = \nu_\infty \left(1 - \frac{r^*}{r}\right)^2 \quad (600)$$

and the detectable frequency is

$$\nu_{r \rightarrow d} = \nu_\infty \left(1 - \frac{r^*}{r}\right)^2 e^{-\frac{r^*}{2r}} \quad (601)$$

where, ν_r is the frequency of absorption line in that the light rays cross through a ring located at a position of r , and ν_∞ is the absorption frequency case that structure would absorb in no-gravity fields. One can of course calculate the widths of absorption lines just like that of emission lines.

$$w_\nu = \left[\left(1 - \frac{r^*}{r+\Delta r}\right)^2 e^{-\frac{r^*}{2(r+\Delta r)}} - \left(1 - \frac{r^*}{r}\right)^2 e^{-\frac{r^*}{2r}}\right] \nu_\infty \quad (602)$$

or

$$w_\lambda = \left[\left(1 - \frac{r^*}{r}\right)^{-2} e^{\frac{r^*}{2r}} - \left(1 - \frac{r^*}{r+\Delta r}\right)^{-2} e^{\frac{r^*}{2(r+\Delta r)}}\right] \lambda_\infty \quad (603)$$

where, Δr is absorption width of a ring that the light ray passes across.

or expressed by Doppler velocity in frequency as

$$v_{D\nu} = \frac{w_\nu}{\nu_r} c \quad (604)$$

or Doppler velocity in wave length as

$$v_{D\lambda} = \frac{w_\lambda}{\lambda_r} c \quad (605)$$

But more different from emissions, absorption spectrums must have more narrow lines and lower redshift than emissions, because the absorption rings are mostly at outer regions with respect to the inner shining emission source, otherwise they would be difficult to be detected.

As we have discussed above, the inflow of inner rings of accretion may act as key role of entire emissions. So, we can image that brighter inner rings emit light rays, and on our view line, they go

through some surrounding rings to be absorbed. It is said that, absorbing dusts might be the matters located at surrounding outer positions rather than those so called insert bodies far away from the emission areas. This image may lead to the so called relative blueshift, with respect to that of emission lines. Of course, there might be still seldom absorptions happening in inner areas so that we will also observe relative redshift absorption lines occasionally. Given a separation Δr between emission position and absorption position more exterior, the emission frequency could be calculated as

$$\nu_{em \rightarrow d} = \nu_{\infty} \left(1 - \frac{r^*}{r_{em}}\right)^2 e^{-\frac{r^*}{2r_{em}}} \quad (606)$$

with redshift of the line

$$z_{em \rightarrow d} = \frac{\nu_{\infty} - \nu_{em \rightarrow d}}{\nu_{em \rightarrow d}} \quad (607)$$

and absorption frequency varies because a separation Δr between absorption and emission positions

$$\nu_{ab \rightarrow d} = \nu_{\infty} \left(1 - \frac{r^*}{r_{ab}}\right)^2 e^{-\frac{r^*}{2(r_{ab} + \Delta r)}} \quad (608)$$

so that the redshift of absorption line is

$$z_{ab \rightarrow d} = \frac{\nu_{\infty} - \nu_{ab \rightarrow d}}{\nu_{ab \rightarrow d}} \quad (609)$$

And then, there is a blueshift of an absorption line

$$\Delta z = z_{ab \rightarrow d} - z_{em \rightarrow d} = \left(1 - \frac{r^*}{r_{ab}}\right)^{-2} e^{\frac{r^*}{2(r_{ab} + \Delta r)}} - \left(1 - \frac{r^*}{r_{em}}\right)^{-2} e^{\frac{r^*}{2r_{em}}} \quad (610)$$

Given $r_{em} = 2.5r^*$ and $\Delta r = 0.5r^*$, it can be calculated that $z_{em} = 2.39$ and $z_{ab} = 1.66$. Here, the absorption lines may show a blueshift of 0.73 with respect to the emission lines.

We have found that there were also absorption broad line region and absorption narrow line region [55] in single system in which those absorption lines were all corresponding to that of emission lines.

Case a continuum spectrum of light rays passes across multiple rings and arrive at an observatory on the Earth, one can then get a set of multiple frequency shift of absorptions. It is because absorptions only depend on position in accretions. This does interpret the multiple redshift absorption lines in tremendous of observations [48,49,54] on those active galactic nuclei. Especially for that of so called Lyman-alpha forest [56], a series of regular Lyman-alpha absorption lines, with descending order of redshifts, queue up amazingly at the left side of the great main emission line.

There is a special condition for a matter out flowing from center source, for example the center jet outflow. When a matter at ground state moves a separation toward outer direction, its structural energy should experience energy deficiency in that the energy of equivalent state gets increasing. We then see that the matter is in the state lack of energy, and generally, the state should be kept until surrounding environment happens to give amount of specific light sources, so as the result, the matter gets absorptions. It is said that, this kind of absorptions would be far different from normal relativistic absorptions as that have been discussed above, so that they could be called the super relativistic absorption. As for the mentioned absorption energy, it depends on the position it getting absorbing as well as the position it has ever been in ground state. Note that this kind of absorptions really distinguish from the general relativistic absorptions in that the super relativistic absorption energy depends on the difference between the final emission position and specific absorption position, while the general relativistic absorption is due to property of structure state. It could be calculated as

$$\Delta \nu = \nu_{ab} - \nu_{gr} \quad (611)$$

where, $\nu_{ab} = \left(1 - \frac{r^*}{r_{ab}}\right)^2 e^{-\frac{r^*}{2r_{ab}}} \nu_{\infty}$, $\nu_{gr} = \left(1 - \frac{r^*}{r_{gr}}\right)^2 e^{-\frac{r^*}{2r_{gr}}} \nu_{\infty}$, r_{ab} is the radius at absorption position, and r_{gr} is the radius of the position of equivalent ground state before outflow. This kind of absorption might need greater absorption energy than that of general relativistic absorption, and it

does nothing with original absorption energy. Therefore, it should be called the super relativistic absorption.

To the observer at a farthest position, the absorption line will also perform special frequency shift that

$$z_{super} = \frac{v_{\infty} - \Delta v}{\Delta v} = \frac{1}{\frac{-r^*}{(1-\frac{r^*}{r_{ab}})^2 e^{2r_{ab}}} - \frac{-r^*}{(1-\frac{r^*}{r_{gr}})^2 e^{2r_{gr}}}} - 1 \quad (612)$$

Case the jet flow particles of an active galactic nucleus are assigned in a broad area, it could cause a broader absorption pit for a crossing light ray. One can calculate the average redshift and the width of absorption pit.

Additionally speaking on the topic of jet flow, massive particles may be easily accelerated up and keep approaching absolute light speed so as to escape from inner event horizon. On the other hand, apparent light speed may be at lower level in the nearby r^* area so that to form pseudo super luminal motions. A high pseudo super luminal motion may cause Cherenkov radiation, which could generally show polarized propagation. That might also be the reasons for some high energy light rays. One can imagine that the events of γ -ray burst detection delay in observations on SN1987A [57] and GW170817 [58] might be interpreted by pseudo super luminal propagations of neutrinos and gravitational waves.

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