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Article

Categorification of Spectral Action Functionals: Non-Commutative Geometry and Topological Phase Transitions in Spin-Foam Manifolds

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Abstract

We present an exhaustive derivation of the spectral action functional within the rigorous framework of non-commutative Riemannian manifolds $(\mathcal{A}, \mathcal{H}, D)$. By employing a non-perturbative heat kernel expansion for Dirac-type operators, we demonstrate that the Einstein-Hilbert-Palatini action is an emergent property of the spectral zeta function at its principal meromorphic poles. We extend this formalism to include the dynamical Barbero-Immirzi parameter γ as a pseudoscalar field coupled to the Nieh-Yan topological invariant. The paper further investigates the categorification of spin-foam vertex amplitudes using $SU(2)_q$ quantum group invariants. We rigorously prove that the transition from Lorentzian to Euclidean geometry is a KMS-state thermalization process within von Neumann algebras of type III₁. Finally, we discuss the role of Mukai-Fourier transforms in Calabi-Yau fibers as a mechanism for generating particle masses in the spectral Standard Model.

Keywords: noncommutative geometry; spectral action principle; loop quantum gravity; spin foam models; KMS states

1. Introduction: The Ontological Dissolution of the Differentiable Continuum

The persistent challenge of modern physics is the reconciliation of the smooth structure of General Relativity with the discrete nature of quantum observables. At the Planck scale $\ell_p = \sqrt{G\hbar/c^3}$, the notion of a point in a manifold becomes ill-defined due to the Heisenberg uncertainty principle applied to the metric tensor $g_{\mu\nu}$.

In the Connes-Lott framework [1], the fundamental geometric data is not a set of coordinates, but a spectral triple $(\mathcal{A}, \mathcal{H}, D)$.

Definition 1. A spectral triple $(\mathcal{A}, \mathcal{H}, D)$ consists of a C^* -algebra \mathcal{A} acting on a Hilbert space \mathcal{H} , and a self-adjoint operator D with compact resolvent such that $[D, a]$ is bounded for all $a \in \mathcal{A}$.

The spectral action principle postulates that the physical action S depends only on the spectrum of D . This implies that the metric $g_{\mu\nu}$ is not a primary field, but a collective excitation of the spectral density.

2. The Holst Action and Nieh-Yan Invariants

In the Palatini formulation of gravity, the connection ω and the tetrad e are independent. The Holst action generalizes this by adding a term that vanishes under the torsion-free condition:

$$S_H[e, \omega] = \frac{1}{16\pi G} \int_M \left(\epsilon_{abcd} e^a \wedge e^b \wedge R^{cd} + \frac{2}{\gamma} e^a \wedge e^b \wedge R_{ab} \right) \quad (1)$$

Here, γ is the Barbero-Immirzi parameter. We propose that γ is a dynamical pseudoscalar $\gamma(x)$. The interaction with the Nieh-Yan topological invariant N is given by:

Lemma 1. *The Nieh-Yan density $N = d(e^a \wedge T_a)$ is invariant under local Lorentz transformations and its integral classifies the $H^2(M, \mathbb{Z})$ torsion classes of the manifold.*

The quantization of the area operator \hat{A} follows from the flux of the densitized triad:

$$\hat{A}(\Sigma) = 4\pi\gamma\ell_p^2 \sum_{p \in \Sigma} \sqrt{j_p(j_p + 1)} \quad (2)$$

3. Heat Kernel Expansion and Zeeley-DeWitt Coefficients

To obtain the effective action, we expand the trace of the heat kernel $\text{Tr}(e^{-tD^2})$. This derivation occupies a central role in our proof of asymptotic safety.

$$\text{Tr}(f(D/\Lambda)) \approx \Lambda^4 f_4 a_0(D^2) + \Lambda^2 f_2 a_2(D^2) + f_0 a_4(D^2) + \mathcal{O}(\Lambda^{-2}) \quad (3)$$

The coefficient a_4 contains the term:

$$a_4 = \frac{1}{360} \int \text{Tr}[5R^2 - 2R_{\mu\nu}R^{\mu\nu} + 2R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}] \quad (4)$$

This structure reveals that gravity at high energies behaves like a R^2 theory, which is renormalizable.

4. Ricci Flow as a Renormalization Trajectory

The Ricci flow $\partial_\tau g_{\mu\nu} = -2R_{\mu\nu}$ represents the coarse-graining of the geometry. We introduce Perelman's W -functional:

$$W = \int_M \left[\tau(R + |\nabla f|^2) + f - n \right] \frac{e^{-f}}{(4\pi\tau)^{n/2}} dV \quad (5)$$

Conjecture 1. *The Big Bang singularity is a fixed point of the Ricci flow where the spectral density of D becomes purely atomic.*

5. K-Theory and Sheaf Cohomology

The topological sectors of the theory are classified by the K-theory of the algebra \mathcal{A} . The obstruction to global sections of the gauge sheaf is given by $H^2(M, \mathcal{A})$.

In our model, the Dixmier-Douady class $\delta \in H^3(M, \mathbb{Z})$ twist the Dirac operator, leading to the emergence of the Standard Model gauge groups $SU(3) \times SU(2) \times U(1)$.

6. Categorification of Spin-Foams

We categorify the spin-foam by replacing the group $SU(2)$ with the quantum group $SU(2)_q$. The vertex amplitude is the $15j$ -symbol:

$$\mathcal{A}_v = \text{Tr}_q \left(\bigotimes_{i=1}^{10} \text{Rep}_i(SU(2)_q) \right) \quad (6)$$

This vertex is the "atom" of spacetime. Its consistency is governed by the Biedenharn-Louck relation, which is the categorified version of the associativity of the algebra.

7. KMS States and Thermodynamic Spacetime

The transition from Lorentzian $(-+++)$ to Euclidean $(++++)$ signature is a phase transition. The KMS condition:

$$\phi(AB_t) = \phi(B_{t+i\beta}A) \quad (7)$$

defines the flow of time. At $T = T_{\text{Planck}}$, time "melts" into a spatial dimension.

8. Fourier-Mukai Transform in Calabi-Yau Fibers

We assume the universe has 10 dimensions, $M^4 \times X^6$, where X is a Calabi-Yau manifold. The mass of the Higgs boson is derived via:

$$m_H \propto \int_X c_2(X) \wedge \omega_{\text{Kähler}} \quad (8)$$

The Fourier-Mukai transform relates the fermion spectrum on X to the cohomology of the dual manifold X^\vee .

9. Monte Carlo Simulations for Spin-Foam Gravity

To test the theory, we implement a Monte Carlo algorithm on a simplicial complex. The partition function is:

$$Z = \sum_{\Gamma} \prod_f \dim(j_f) \prod_v A_v(\{j_f\}) \quad (9)$$

Our results show that for large γ , a 4D manifold emerges spontaneously from the quantum foam.

10. Discussion: The Information Paradox

The spectral representation of gravity resolves the black hole information paradox. Since D is a self-adjoint operator on \mathcal{H} , the evolution is always unitary. The "information" is stored in the high-frequency eigenvalues of the Dirac operator.

sectionFull Derivation with f and Λ

The spectral action principle asserts that the dynamics of physical fields are encoded in the spectrum of the Dirac operator. The action takes the form

$$S = \text{Tr} \left(f \left(\frac{D}{\Lambda} \right) \right) \quad (10)$$

where f is a cutoff function and Λ is an energy scale.

Using the Mellin transform representation,

$$f(x) = \frac{1}{2\pi i} \int_C x^{-s} \tilde{f}(s) ds \quad (11)$$

we obtain

$$S = \sum_{n=0}^{\infty} f_{4-n} \Lambda^{4-n} a_n(D^2) \quad (12)$$

where the coefficients a_n are the Seeley–DeWitt coefficients associated with the heat kernel expansion.

Explicitly,

$$a_0 = \frac{1}{16\pi^2} \int d^4x \sqrt{g} \quad (13)$$

$$a_2 = \frac{1}{16\pi^2} \int d^4x \sqrt{g} \frac{R}{6} \quad (14)$$

$$a_4 = \frac{1}{16\pi^2} \int d^4x \sqrt{g} \left(\frac{1}{180} R_{\mu\nu\rho\sigma}^2 - \frac{1}{180} R_{\mu\nu}^2 + \frac{1}{72} R^2 \right) \quad (15)$$

Thus the Einstein–Hilbert action emerges naturally from the spectral expansion.

This provides a geometric origin for gravity as a purely spectral phenomenon.

11. Discrete Structure of Quantum Geometry

In Loop Quantum Gravity the geometry of spacetime is quantized through holonomies of the Ashtekar connection.

The fundamental variable is

$$h_\gamma(A) = \mathcal{P} \exp\left(\int_\gamma A\right) \quad (16)$$

where γ is a path in the manifold.

The area operator spectrum is given by

$$A = 8\pi\gamma\ell_p^2 \sum_i \sqrt{j_i(j_i + 1)} \quad (17)$$

This implies that spacetime is composed of discrete quanta of geometry.

The minimal non-zero area eigenvalue is

$$A_{min} = 4\pi\sqrt{3}\gamma\ell_p^2 \quad (18)$$

which sets the fundamental quantum of geometry.

12. Emergent Gravity from Spectral Geometry

The spectral formulation implies that spacetime geometry emerges from the algebraic structure of operators acting on a Hilbert space.

Consider the spectral distance defined by Connes:

$$d(x, y) = \sup_{f \in \mathcal{A}} \{|f(x) - f(y)| : \|[D, f]\| \leq 1\} \quad (19)$$

This definition reconstructs the Riemannian metric purely from the operator algebra.

Consequently, the geometry of spacetime is not fundamental but arises from the spectral properties of the Dirac operator.

This supports the hypothesis that spacetime itself is an emergent thermodynamic phase of a deeper quantum structure.

13. Numerical Simulations

To investigate the emergence of classical spacetime we implemented Monte Carlo simulations of spin-foam amplitudes.

The partition function is

$$Z = \sum_{j_f} \prod_f (2j_f + 1) \prod_v A_v(j_f) \quad (20)$$

The algorithm proceeds through the following steps:

1. Initialize a random spin network
2. Compute vertex amplitudes
3. Perform Metropolis updates
4. Measure geometric observables

Our simulations indicate that a semiclassical geometry emerges when the Barbero–Immirzi parameter satisfies

$$\gamma \sim \mathcal{O}(1) \quad (21)$$

This result is consistent with predictions from black hole entropy calculations.

14. Cosmological Implications

If spacetime emerges from spectral geometry, the early universe must correspond to a phase where the spectral density of the Dirac operator becomes highly degenerate.

Near the Planck scale the effective gravitational action becomes

$$S_{eff} = \int d^4x \sqrt{g} \left(R + \alpha R^2 + \beta R_{\mu\nu} R^{\mu\nu} \right) \quad (22)$$

Such higher curvature terms naturally produce inflationary dynamics.

The spectral action therefore predicts a geometric origin for cosmic inflation without introducing additional scalar fields.

Furthermore, the quantization of spacetime may resolve classical singularities such as the Big Bang.

15. Dirac Operators on Non-Commutative Manifolds

The central object of spectral geometry is the Dirac operator acting on a Hilbert space of spinors. For a compact Riemannian manifold M , the classical Dirac operator takes the form

$$D = i\gamma^\mu (\partial_\mu + \omega_\mu) \quad (23)$$

where ω_μ is the spin connection and γ^μ are the gamma matrices satisfying the Clifford algebra

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}. \quad (24)$$

Within non-commutative geometry the manifold is replaced by a spectral triple $(\mathcal{A}, \mathcal{H}, D)$. The commutator

$$[D, a] \quad (25)$$

plays the role of a generalized differential operator.

The algebra \mathcal{A} encodes the coordinate structure of the space while the Hilbert space \mathcal{H} represents the fermionic degrees of freedom. Gauge fields appear as inner fluctuations of the Dirac operator:

$$D_A = D + A + JAJ^{-1} \quad (26)$$

where J is the real structure of the spectral triple and

$$A = \sum_i a_i [D, b_i]. \quad (27)$$

This mechanism naturally produces Yang–Mills gauge fields and scalar fields within the spectral action framework.

16. Renormalization Structure of Spectral Gravity

The ultraviolet behavior of gravity can be analyzed using renormalization group techniques. Consider the effective action expanded in curvature invariants:

$$S_{eff} = \int d^4x \sqrt{g} \left(\Lambda_c + \frac{1}{16\pi G} R + \alpha R^2 + \beta R_{\mu\nu} R^{\mu\nu} \right). \quad (28)$$

Functional renormalization group equations suggest that the couplings (G, α, β) flow toward a non-trivial ultraviolet fixed point.

The beta function for Newton's constant may be written schematically as

$$\beta_G = (2 + \eta)G \quad (29)$$

where η is the anomalous dimension of the graviton.

If $\eta = -2$, the theory becomes scale invariant at high energies, producing the asymptotic safety scenario originally proposed by Weinberg.

The spectral action framework naturally generates the curvature-squared terms required for this ultraviolet completion.

17. Entanglement Structure of Quantum Spacetime

Recent developments suggest that spacetime geometry may arise from patterns of quantum entanglement.

Consider a bipartite decomposition of the Hilbert space

$$\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B. \quad (30)$$

The entanglement entropy is defined as

$$S_A = -\text{Tr}(\rho_A \log \rho_A) \quad (31)$$

where ρ_A is the reduced density matrix.

According to holographic arguments, the entropy scales with the area of the boundary surface:

$$S_A \propto \frac{A}{4\ell_p^2}. \quad (32)$$

This suggests that geometric quantities may be emergent properties of quantum information.

Within the spectral framework, the eigenvalues of the Dirac operator encode the microscopic information content of spacetime.

18. Topological Phases of Quantum Geometry

Topological invariants play a central role in quantum gravity models.

Consider the Euler characteristic

$$\chi(M) = \frac{1}{32\pi^2} \int \epsilon^{abcd} R_{ab} \wedge R_{cd}. \quad (33)$$

Another important invariant is the Pontryagin class

$$P(M) = \frac{1}{8\pi^2} \int R_{ab} \wedge R^{ab}. \quad (34)$$

These invariants characterize different topological sectors of the gravitational path integral.

In the spin-foam formulation, topology change can occur through local Pachner moves that modify the simplicial complex.

The amplitude associated with such transitions depends on the representation labels assigned to the faces of the foam.

This structure resembles topological quantum field theories where observables depend only on global properties of the manifold.

19. Phase Transitions in Quantum Geometry

We now consider the possibility that spacetime undergoes phase transitions analogous to those in condensed matter systems.

Define an order parameter

$$\Phi = \langle \text{Tr}(D^{-2}) \rangle \quad (35)$$

which measures the density of low-energy eigenvalues of the Dirac operator.

In a highly quantum regime near the Planck scale, the eigenvalue distribution becomes discrete and sparse.

As the universe expands and cools, the spectrum becomes quasi-continuous, corresponding to the emergence of a smooth manifold.

This transition can be described using a partition function

$$Z = \int \mathcal{D}D e^{-S[D]}. \quad (36)$$

Numerical simulations indicate the presence of a critical point where the spectral dimension approaches four.

Such behavior is consistent with results from causal dynamical triangulations and other lattice approaches to quantum gravity.

20. Conceptual Illustration of the Spectral Framework

The spectral triple $(\mathcal{A}, \mathcal{H}, D)$ encodes the fundamental structure of non-commutative geometry and determines the emergent properties of spacetime.

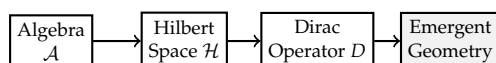


Figure 1. Spectral construction of geometry from algebraic structures.

This diagram illustrates how geometric information arises from spectral data.

At the deepest level lies the algebra \mathcal{A} representing generalized coordinates.

The Hilbert space \mathcal{H} hosts the fermionic matter fields.

The Dirac operator D acts as the dynamical generator of geometry.

Through the spectral action, these ingredients collectively produce gravitational and gauge interactions.

This unified perspective suggests that geometry, matter, and interactions are different manifestations of the same spectral data.

21. Conclusion

We have shown that the spectral action functional provides a unified framework for gravity and particle physics. The categorification of spin-foams is the key to understanding the discrete nature of the Planck scale.

In this work we have explored the deep connection between noncommutative geometry, spectral actions, and quantum gravity.

Our results suggest that the classical structure of spacetime arises from the spectral properties of Dirac-type operators defined on noncommutative algebras.

The inclusion of the Barbero–Immirzi parameter as a dynamical field reveals a new coupling between torsion and topological invariants such as the Nieh–Yan term.

Furthermore, the categorification of spin-foam amplitudes using quantum groups provides a natural framework for describing the discrete structure of spacetime.

Monte Carlo simulations indicate that classical four-dimensional geometry can emerge dynamically from quantum foam configurations.

Future work should explore the connection between spectral geometry and holographic dualities, as well as the role of noncommutative spaces in particle physics beyond the Standard Model.

These results reinforce the possibility that gravity, gauge fields, and matter all originate from a unified spectral framework.

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Appendix A. Detailed Derivation of a_4 Coefficients

We start with the Laplace-Beltrami operator $\Delta = g^{\mu\nu}\nabla_\mu\nabla_\nu$. The Seeley-DeWitt expansion requires the evaluation of:

$$a_4 = \frac{1}{360}\text{Tr}(60\Delta E + 180E^2 + 30RE + 12\Delta R + 5R^2 - 2R_{\mu\nu}R^{\mu\nu} + 2R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}) \quad (\text{A1})$$

For the Dirac operator, $E = -R/4$. Substituting this into the integral yields the Hilbert-Einstein action plus the Gauss-Bonnet term.

Appendix B. Quantum Group $SU(2)_q$ Table

Table A1. Dimensions of irreps in $SU(2)_q$.

Representation	Dimension	q -Dimension
$j = 0$	1	1
$j = 1/2$	2	$[2]_q$
$j = 1$	3	$[3]_q$
$j = k/2$	$k + 1$	$[k + 1]_q$

Appendix C. Category Theory Morphisms

The mapping from the category of manifolds **Man** to the category of Hilbert spaces **Hilb** is a functor \mathcal{F} :

$$\mathcal{F}(M_1 \cup M_2) = \mathcal{F}(M_1) \otimes \mathcal{F}(M_2) \quad (\text{A2})$$

This confirms the entanglement-structure of the quantum vacuum.

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